

Early Universe

Universe expands with scale factor $a(t)$

Einstein:

$$H^2 = \left(\frac{\dot{a}(t)}{a(t)} \right)^2 = \frac{8\pi}{3} G \rho$$

hot:

$$\begin{aligned} k_B T &\gg mc^2 \\ pc &\gg mc^2 \end{aligned}$$

neglect μ

Early Universe

$$\rho_b = \frac{d_s}{2\pi^2} \int_0^\infty \frac{(\hbar k c) k^2 dk}{e^{\hbar k c / k_B T} - 1}$$

$$\rho_f = \frac{d_s}{2\pi^2} \int_0^\infty \frac{(\hbar k c) k^2 dk}{e^{\hbar k c / k_B T} + 1}$$

$$\int_0^\infty dx \frac{x^{\nu-1}}{e^{ax}-1} = a^{-\nu} \Gamma(\nu) \zeta(\nu)$$

$$\int_0^\infty dx \frac{x^{\nu-1}}{e^{ax}+1} = (1 - 2^{1-\nu}) a^{-\nu} \Gamma(\nu) \zeta(\nu)$$

$$\rho_b = \frac{d_s \pi^2}{30} T^4$$

$$\rho_f = \frac{7}{8} \frac{d_s \pi^2}{30} T^4$$

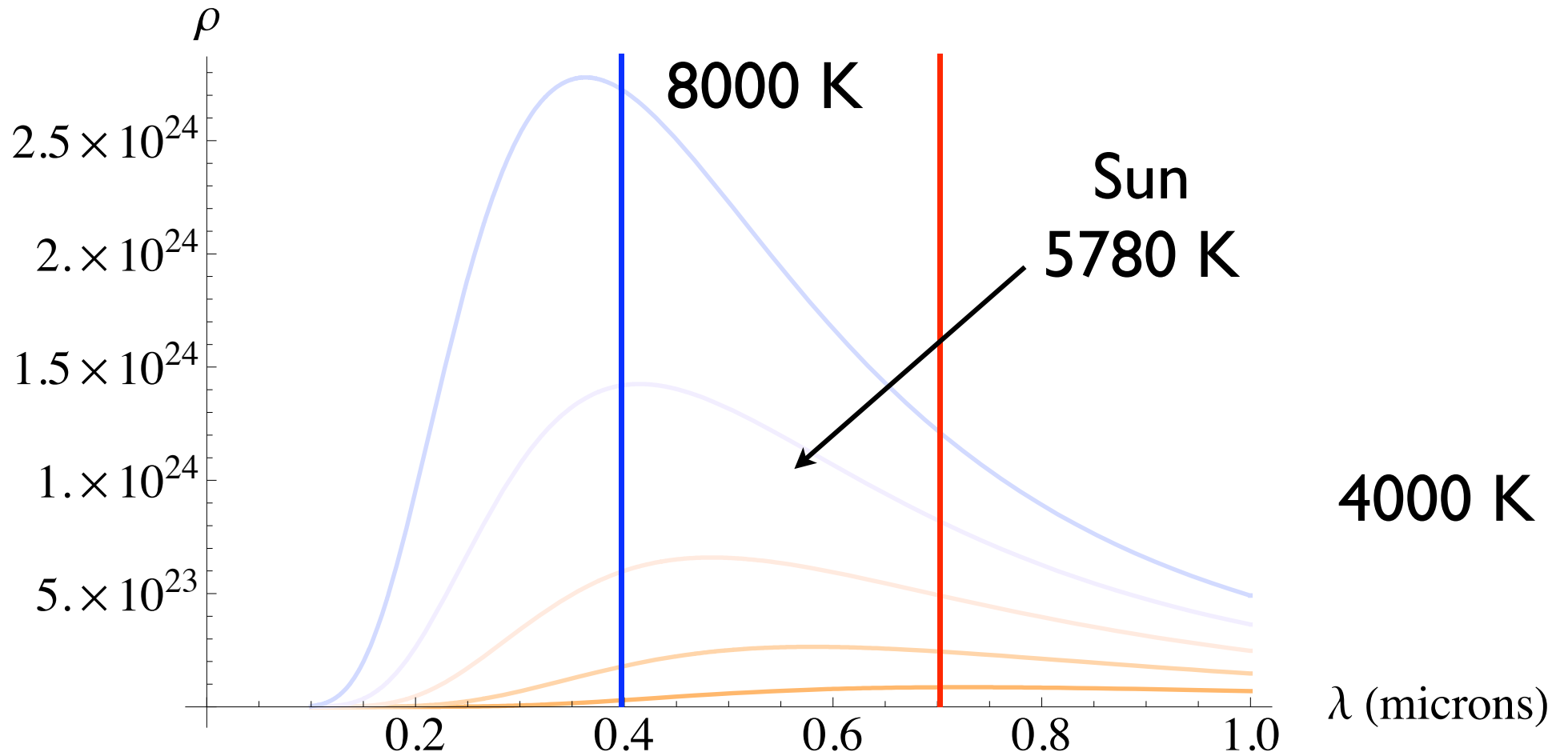
energy density \Rightarrow expansion rate

\Rightarrow relative abundance of He, D, Li

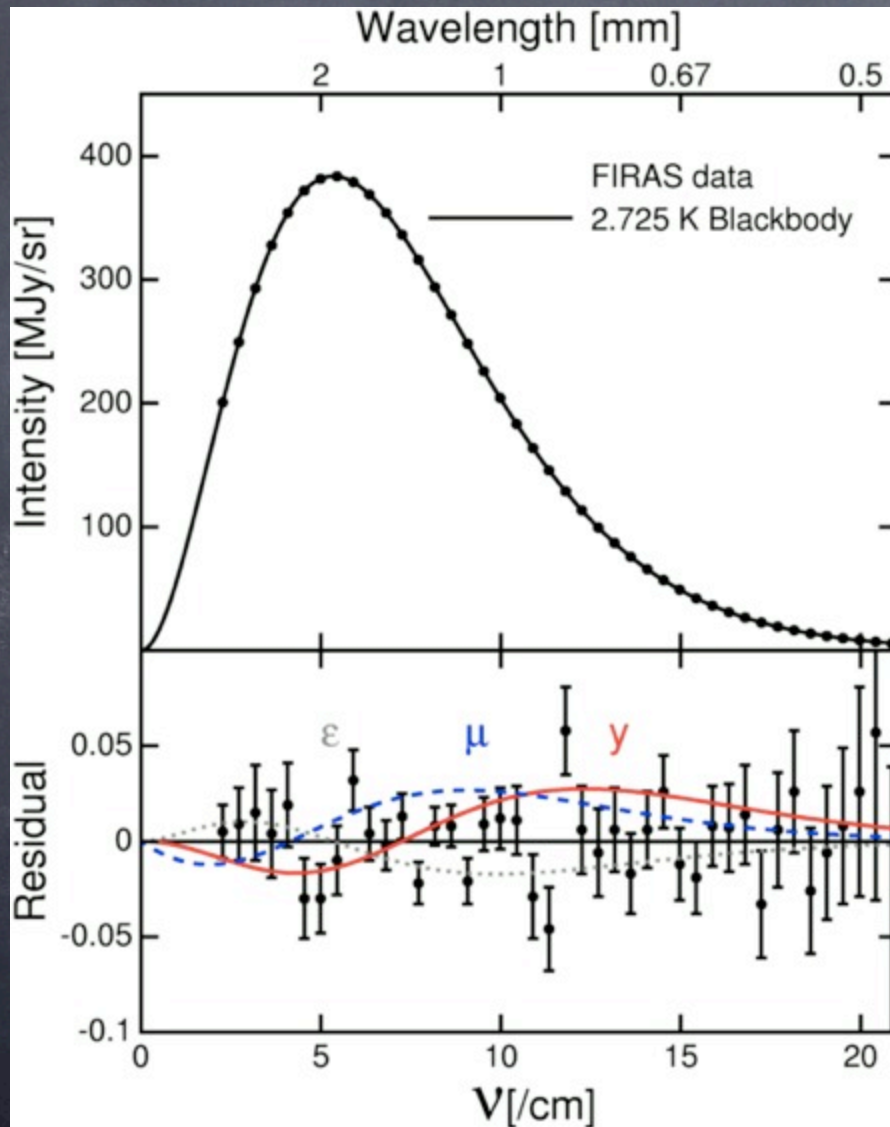
requires $N_\nu = 3 \pm 0.1$

LEP found $N_\nu = 2.994 \pm 0.012$

Planck Distribution

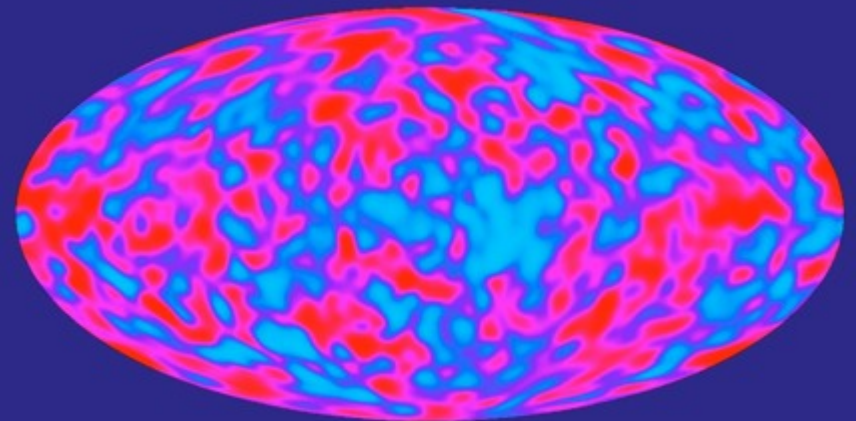


Cosmic Microwave Background



FIRAS
experiment
on
COBE

DMR's Two Year CMB Anisotropy Result



2006 Mather and Smoot



Bose-Einstein Condensation

$$n_b(\epsilon) = \frac{1}{e^{(\epsilon-\mu)/k_B T} - 1} > 0$$

$$e^{(\epsilon-\mu)/k_B T} > 1$$

$$(\epsilon - \mu) > 0$$

for any allowed ϵ

$$\epsilon > \mu(T)$$

for free particles

$$\epsilon = \frac{\hbar^2 k^2}{2m} \geq 0$$

$$\mu(T) < 0$$

Bose-Einstein Condensation

fixed N , V and vary T

$$\frac{N}{V} = \frac{1}{2\pi^2} \int_0^\infty \frac{k^2 dk}{e^{(\hbar^2 k^2 / (2m) - \mu) / k_B T} - 1}$$

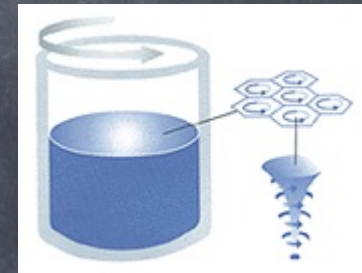
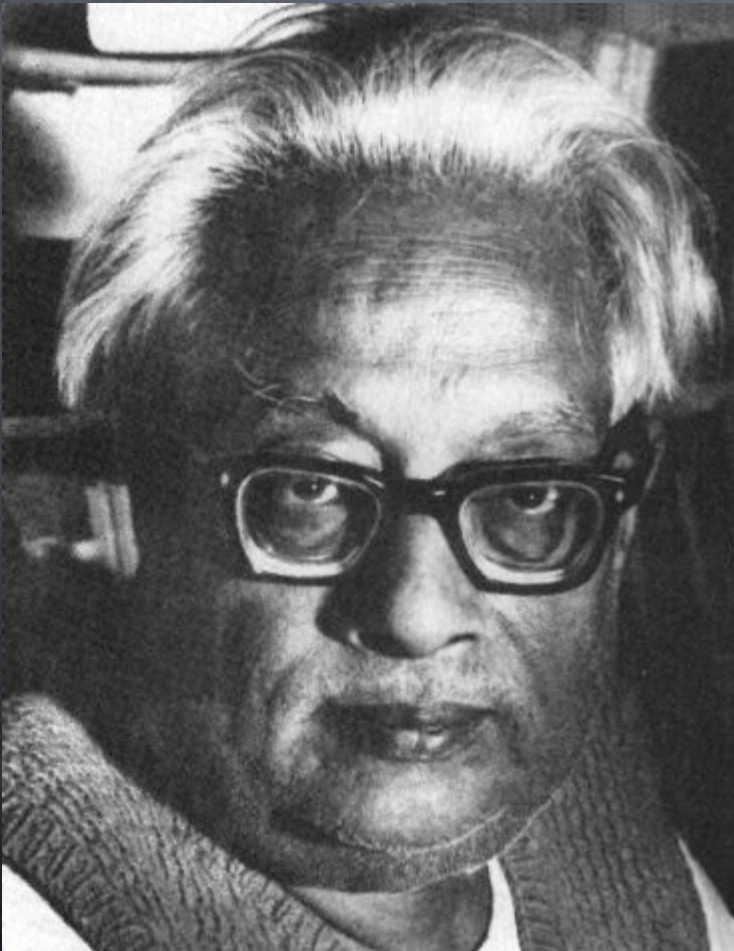
$$T \downarrow \text{ requires } \left(\frac{\hbar^2 k^2}{2m} - \mu \right) \downarrow$$

$$T \downarrow \text{ requires } \mu \uparrow$$

$$\text{But } \mu(T) < 0$$

$$\text{at } T_c, \mu(T_c) = 0$$

Bose-Einstein Condensation



Bose-Einstein Condensation

$$\frac{N}{V} = \frac{1}{2\pi^2} \int_0^\infty \frac{k^2 dk}{e^{(\hbar^2 k^2 / (2m)) / k_B T} - 1}$$

$$k = \frac{\sqrt{2mk_B T_c} \sqrt{x}}{\hbar}$$
$$dk = \frac{\sqrt{mk_B T_c}}{\sqrt{2} \hbar} \frac{dx}{\sqrt{x}}$$

$$\begin{aligned} \frac{N}{V} &= \frac{1}{4\pi^2} \left(\frac{2mk_B T_c}{\hbar^2} \right)^{3/2} \int_0^\infty \frac{\sqrt{x} dx}{e^x - 1} \\ &= \frac{1}{4\pi^2} \left(\frac{2mk_B T_c}{\hbar^2} \right)^{3/2} \Gamma\left(\frac{3}{2}\right) \zeta\left(\frac{3}{2}\right) \\ &= \frac{1}{8} \left(\frac{2mk_B T_c}{\pi \hbar^2} \right)^{3/2} 2.612 \end{aligned}$$

$$T_c = \frac{2\pi \hbar^2}{mk_B} \left(\frac{N}{2.612V} \right)^{2/3}$$

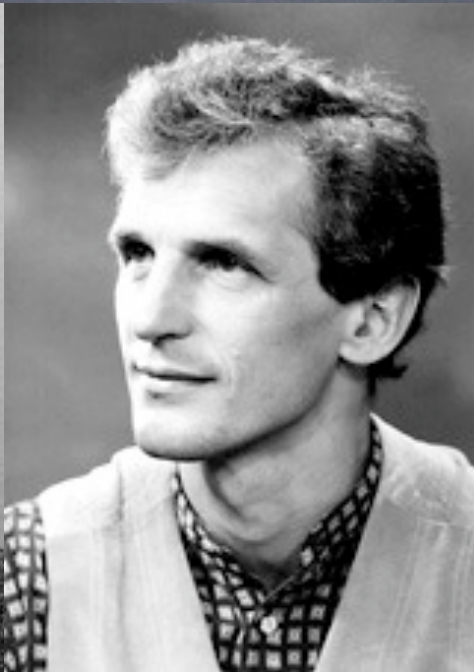
for ${}^4\text{He}$:

$$\frac{N}{V} = 0.15 \text{ g/cm}^3 \frac{1}{4.167 \times 10^{-27} \text{ kg/atom}} = 2.2 \times 10^8 \text{ 1/m}^3$$

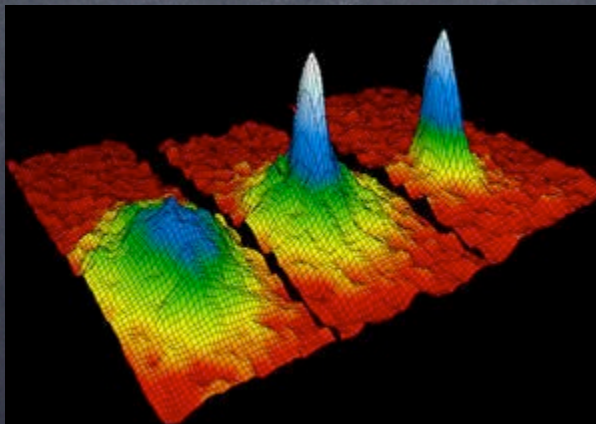
$$T_c = 3.1 \text{ K}$$

$$T_{c,exp} = 2.17 \text{ K}$$

2001 Cornell, Ketterle, Wieman



velocity
distribution



dilute gas
of rubidium
atoms