

Road to Minimal WIMPs

Particle Physics Seminar

UC Davis

Oct 14, 2024

Rodolfo Capdevilla

Fermilab

• Minimal WIMPs:

*Minimal Dark Matter
Electroweak Interacting Dark Matter
Minimal WIMPs*

Cirelli, Fornengo, Strumia, Nucl. Phys. B 753 (2006) 178-194
Cirelli, Strumia, New J. Phys. 11 (2009) 105005
Hisano, Ishiwata, Nagata, Takesako, JHEP 07 (2011) 005
Low, Wang, JHEP 08 (2014) 161
DelNobile, Nardecchia, Panci, JCAP 04 (2016) 048
Baumgart et al., JHEP 01 (2019) 036

$$\chi_2 \equiv \chi_{\tilde{H}} = \begin{pmatrix} \chi_{\tilde{H}}^+ \\ \chi_{\tilde{H}}^0 \\ \chi_{\tilde{H}}^- \end{pmatrix}, \quad \chi_3 \equiv \chi_{\tilde{W}} = \begin{pmatrix} \chi_{\tilde{W}}^+ \\ \chi_{\tilde{W}}^0 \\ \chi_{\tilde{W}}^- \end{pmatrix} \dots \quad \chi_5 \dots \quad \chi_9 \dots$$

*Doublets
(Higgsino-like)*

*Triplets
(Wino-like)*

Fiveplets...

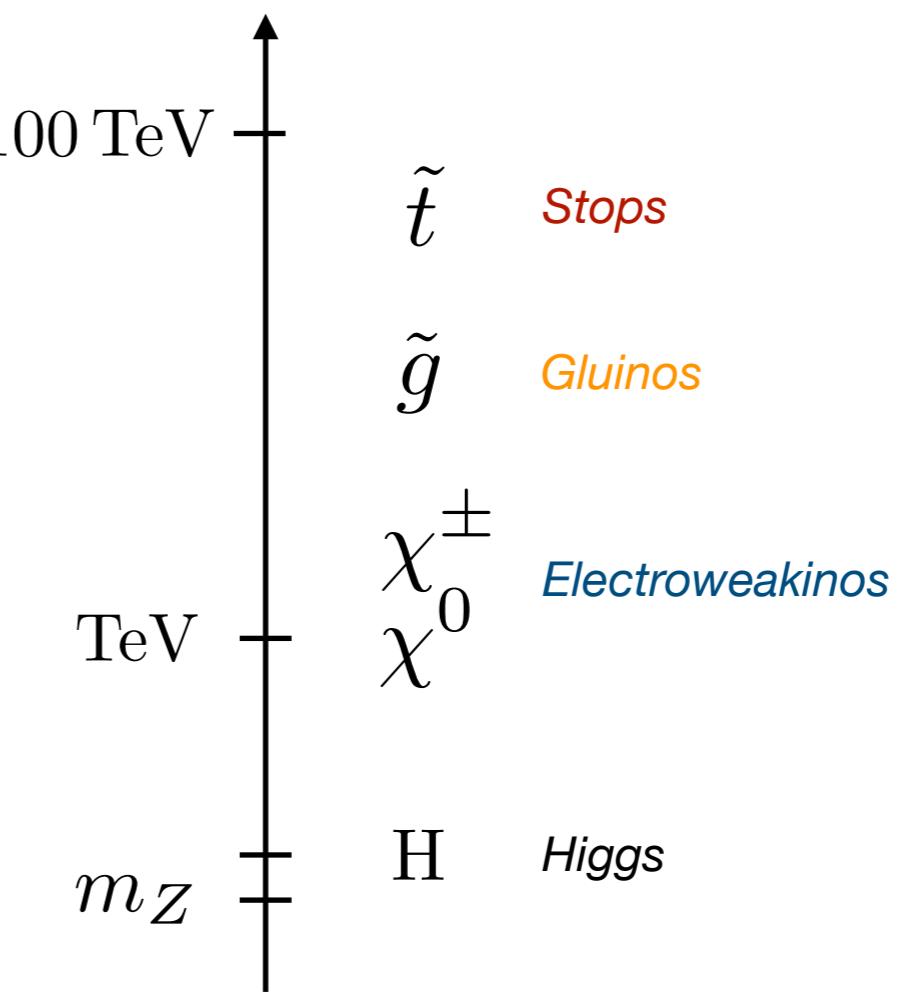
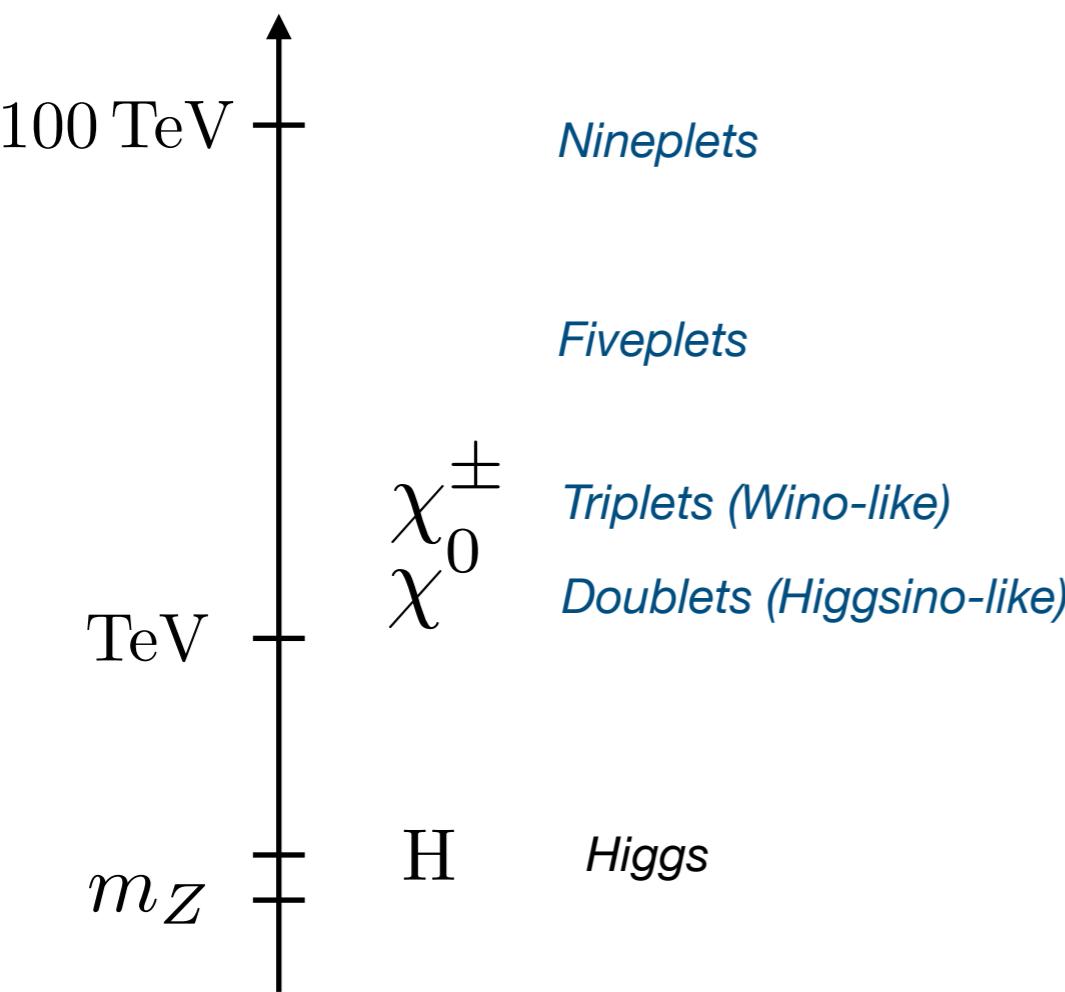
Nineplets...

- **Minimal WIMPs:**

Minimal Dark Matter
Electroweak Interacting Dark Matter
Minimal WIMPs

Simplified Models

(Split)Supersymmetry



Outline

1. Introduction

2. Minimal WIMP Searches

- Direct Detection
- Indirect Detection
- LHC

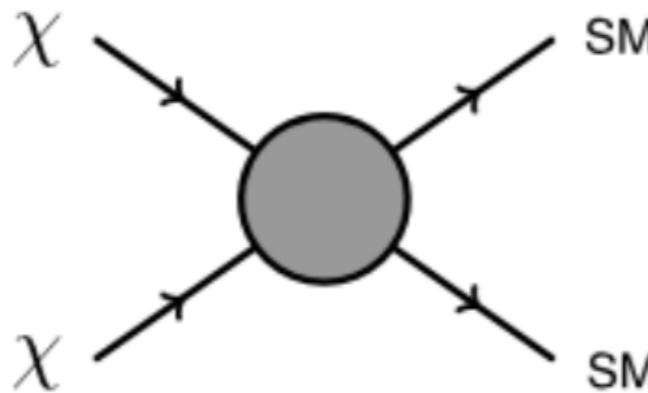
3. Future p/e Colliders

4. Muon Collider

- Disappearing Tracks
- Soft Tracks

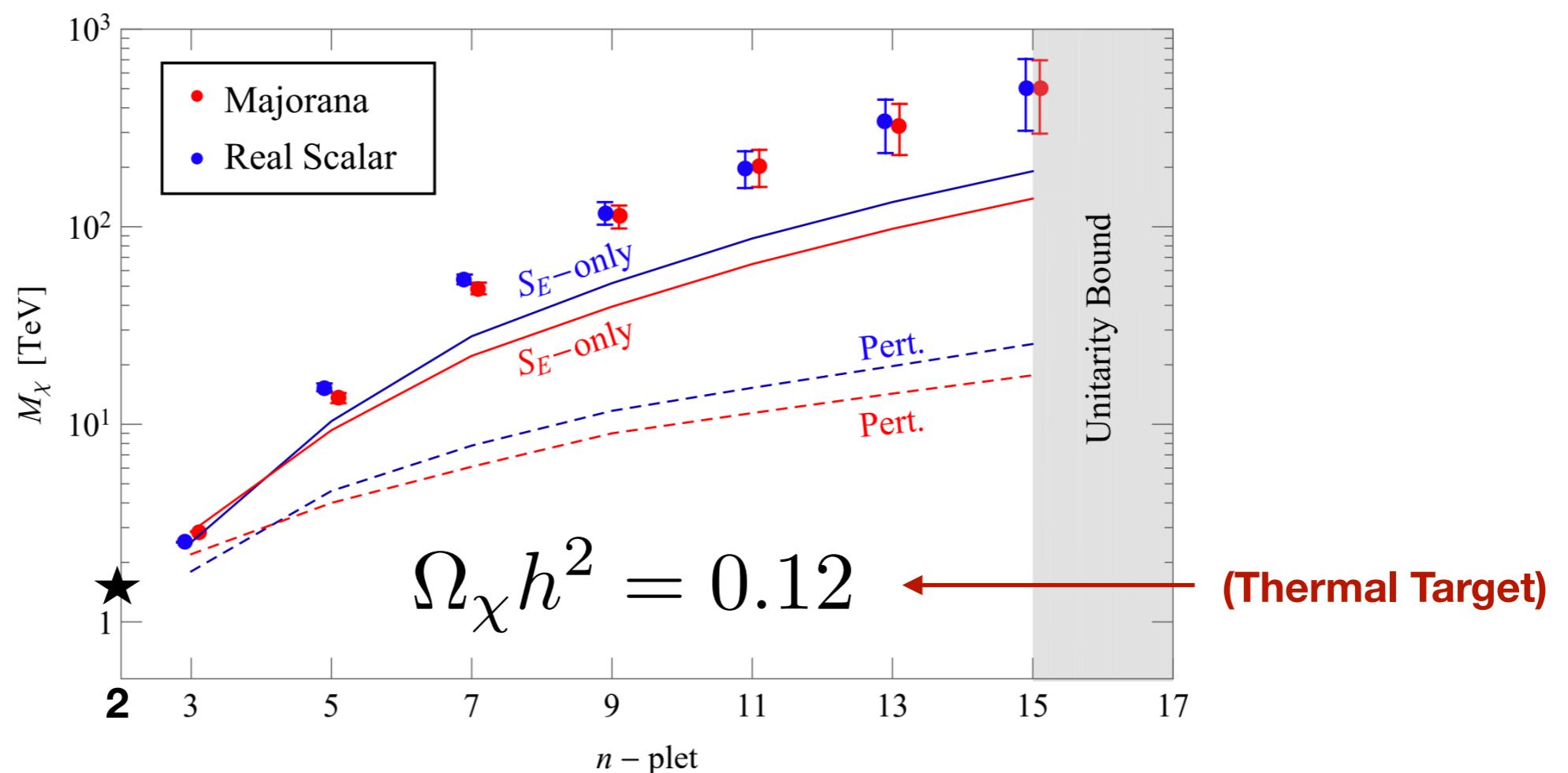
Summary

- **Mass:**



$$\langle\sigma v\rangle \sim \frac{g^4 n^3 + 8g^2 g_Y^2 Y^2 n}{128\pi M_\chi^2} \quad \begin{matrix} \textbf{(Majorana)} \\ \textbf{Large n} \end{matrix}$$

Cirelli, Fornengo, Strumia,
Nucl. Phys. B 753 (2006) 178-194



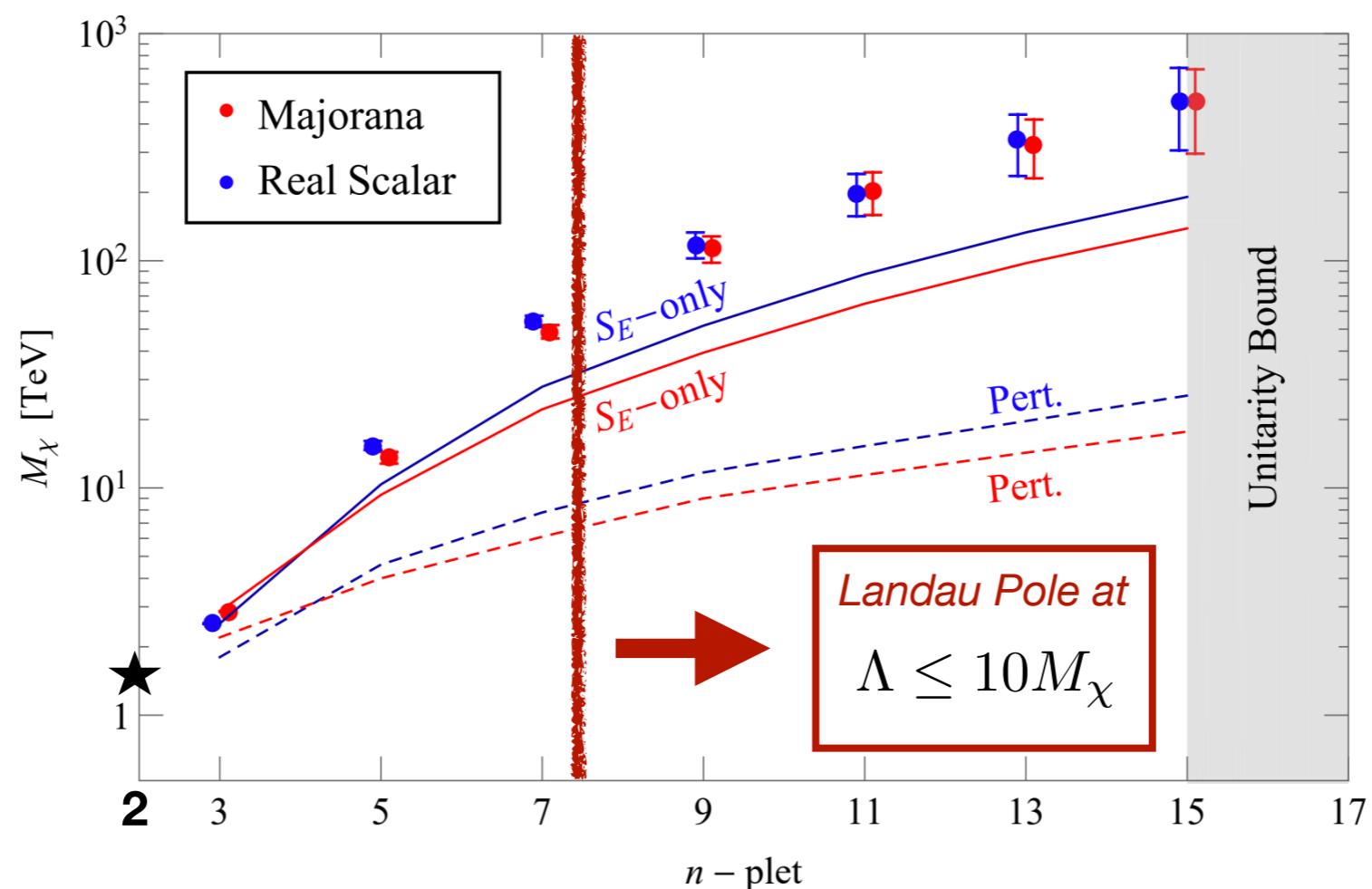
Bottaro et al., Eur. Phys. J. C 82 (2022) 1, 31

- **Mass:**

Very large multiplets become less motivated

A good target can be:

- *EW multiplets: 2, 3, ..., 7*
- *Masses: 1-50 TeV*



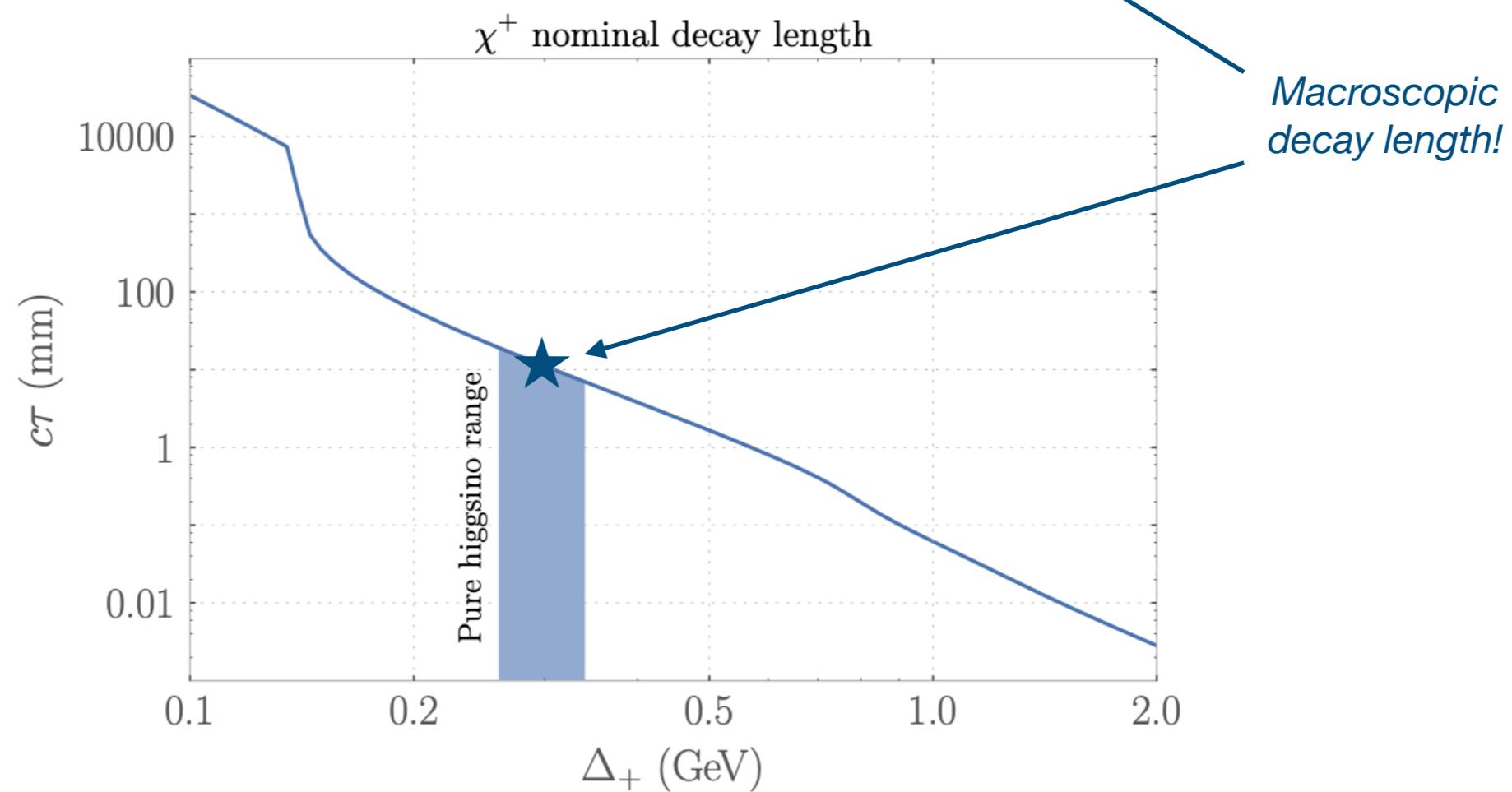
- Lifetime

For a given multiplet

$$\chi_{\tilde{H}} = \begin{pmatrix} \chi_{\tilde{H}}^+ \\ \chi_0^0 \\ \chi_{\tilde{H}}^- \end{pmatrix}$$

*Loop effects generate
chargino-neutralino splitting*

$$\Delta M = M_{\chi^\pm} - M_{\chi^0} > 0$$



R. Mahbubani, P. Schwaller, J. Zurita, JHEP 06 (2017) 119

- Thermal Targets

$$\chi_2 \equiv \chi_{\tilde{H}} = \begin{pmatrix} \chi_{\tilde{H}}^+ \\ \chi_{\tilde{H}}^0 \end{pmatrix},$$

*Doublets
(Higgsino-like)*

$$\chi_3 \equiv \chi_{\tilde{W}} = \begin{pmatrix} \chi_{\tilde{W}}^+ \\ \chi_{\tilde{W}}^0 \\ \chi_{\tilde{W}}^- \end{pmatrix} \dots \quad \chi_5 \dots \quad \chi_9 \dots$$

*Triplets
(Wino-like)*

Fiveplets... Nineplets...

Mass: 1.1 TeV 2.7 TeV 14 TeV

Lifetime: 0.02 ns 0.2 ns 0.2 ns

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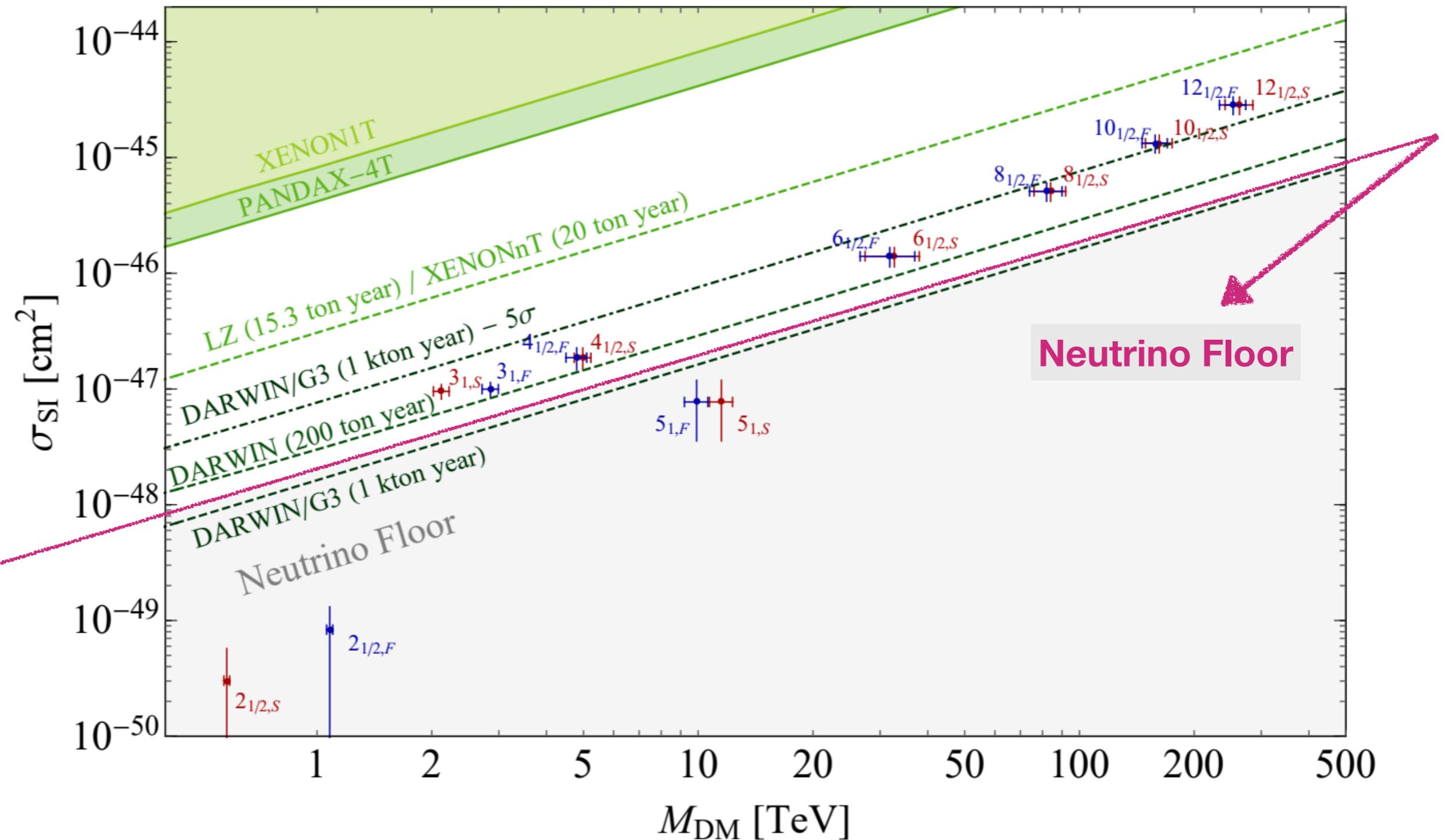
4. Muon Collider

- Disappearing Tracks
- Soft Tracks

Summary

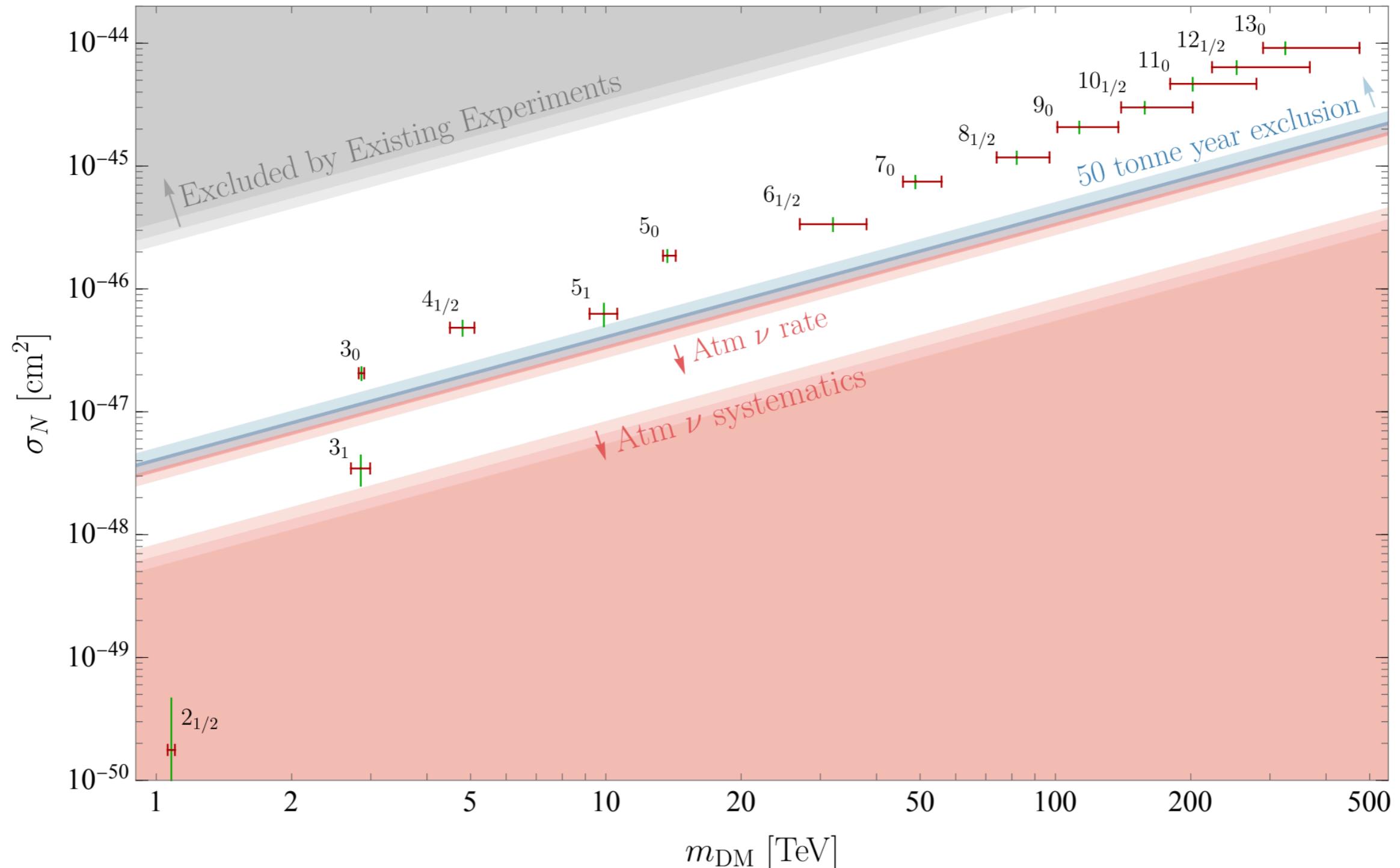
- Direct Detection (DD)

Bottaro, Buttazzo, Costa, Franceschini, Panci, Redigolo, Vittorio,
Eur. Phys. J. C 82 (2022) 11, 992



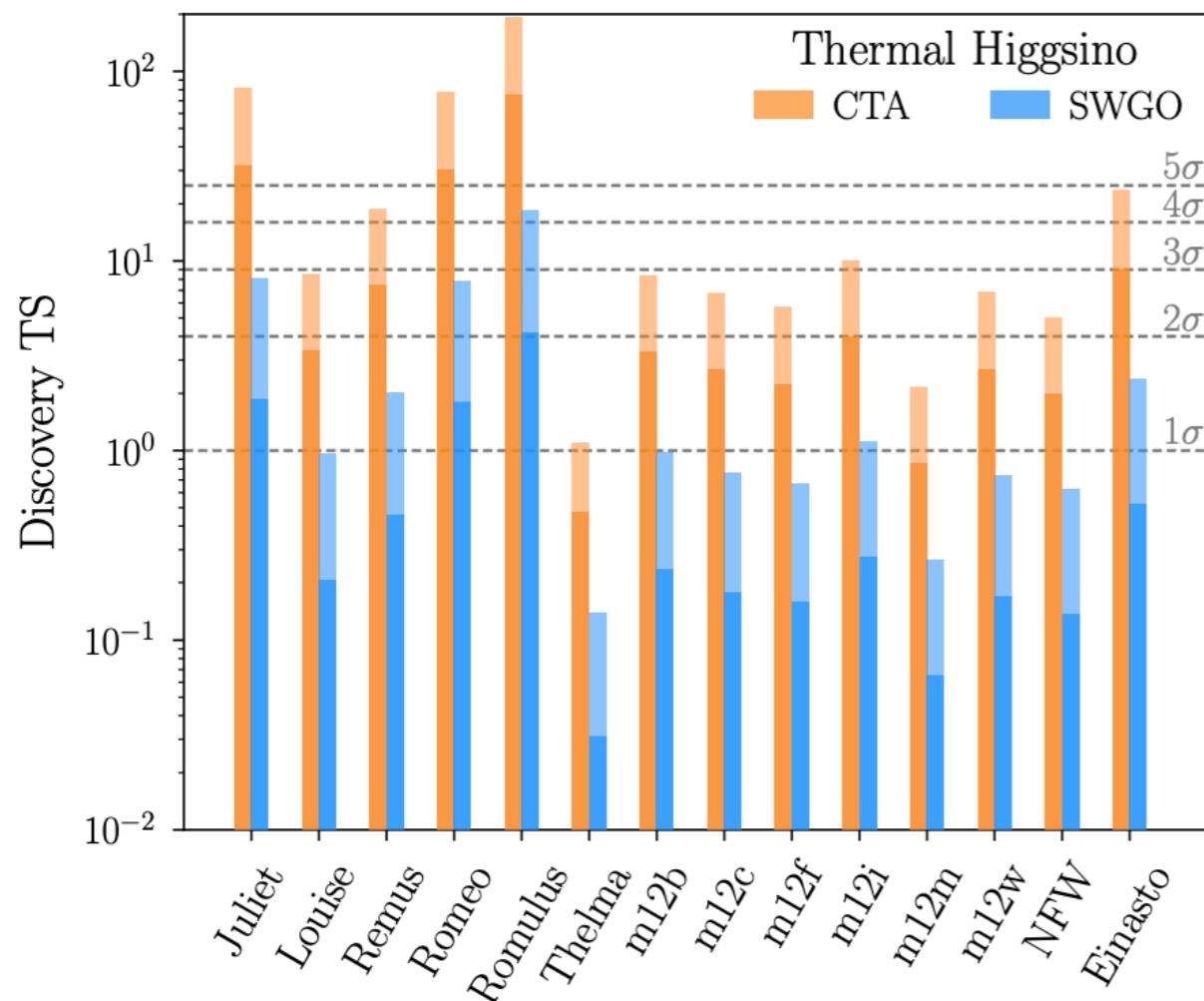
- **Direct Detection (DD)**

I. Bloch, S. Bottaro, D. Redigolo, L. Vittorio, ArXiv:2410.02723



- Indirect Detection (ID)

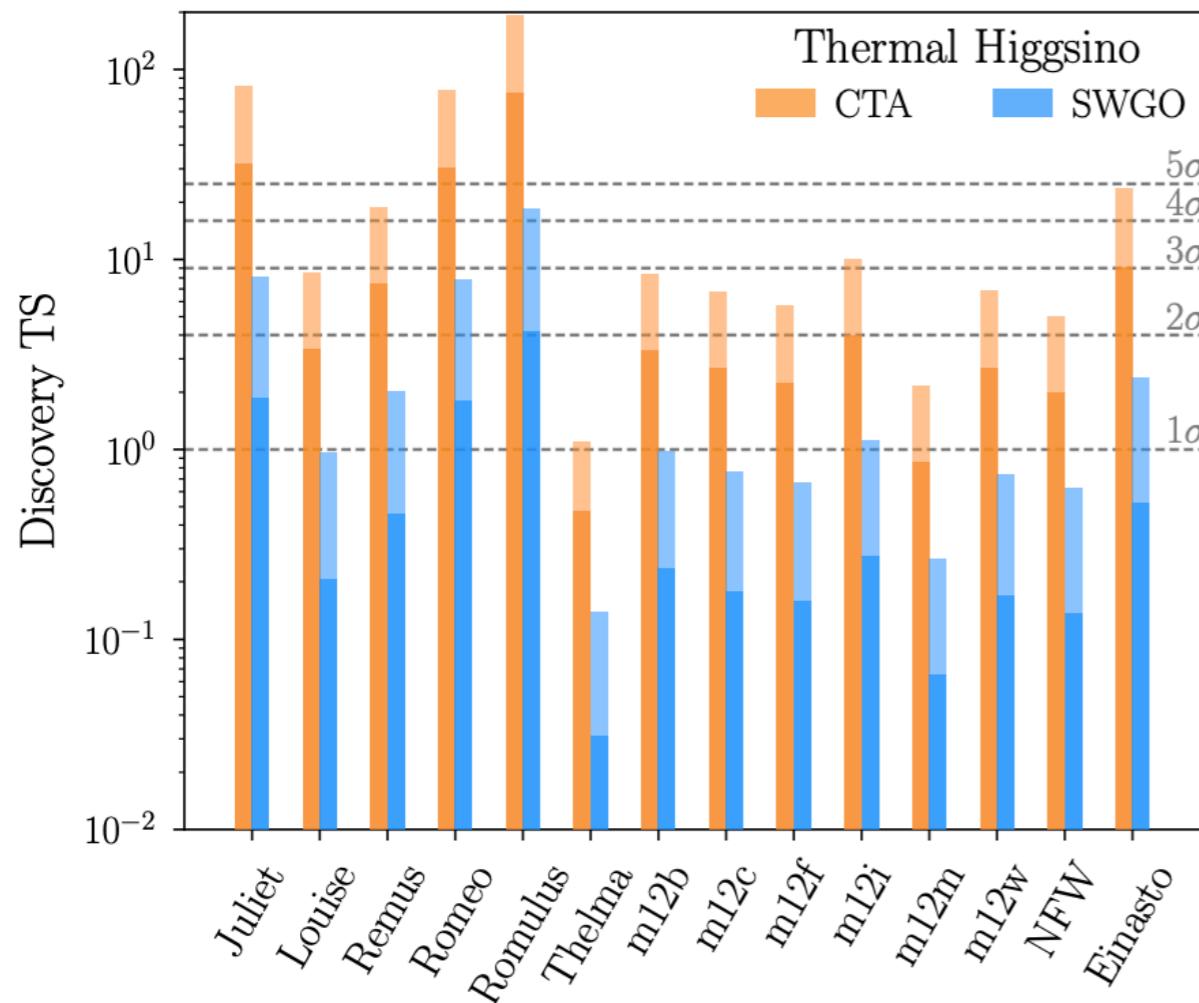
Doublet



Nicholas L. Rodd, Benjamin R. Safdi, Weishuang Linda Xu,
Phys. Rev. D 110 (2024) 4, 043003

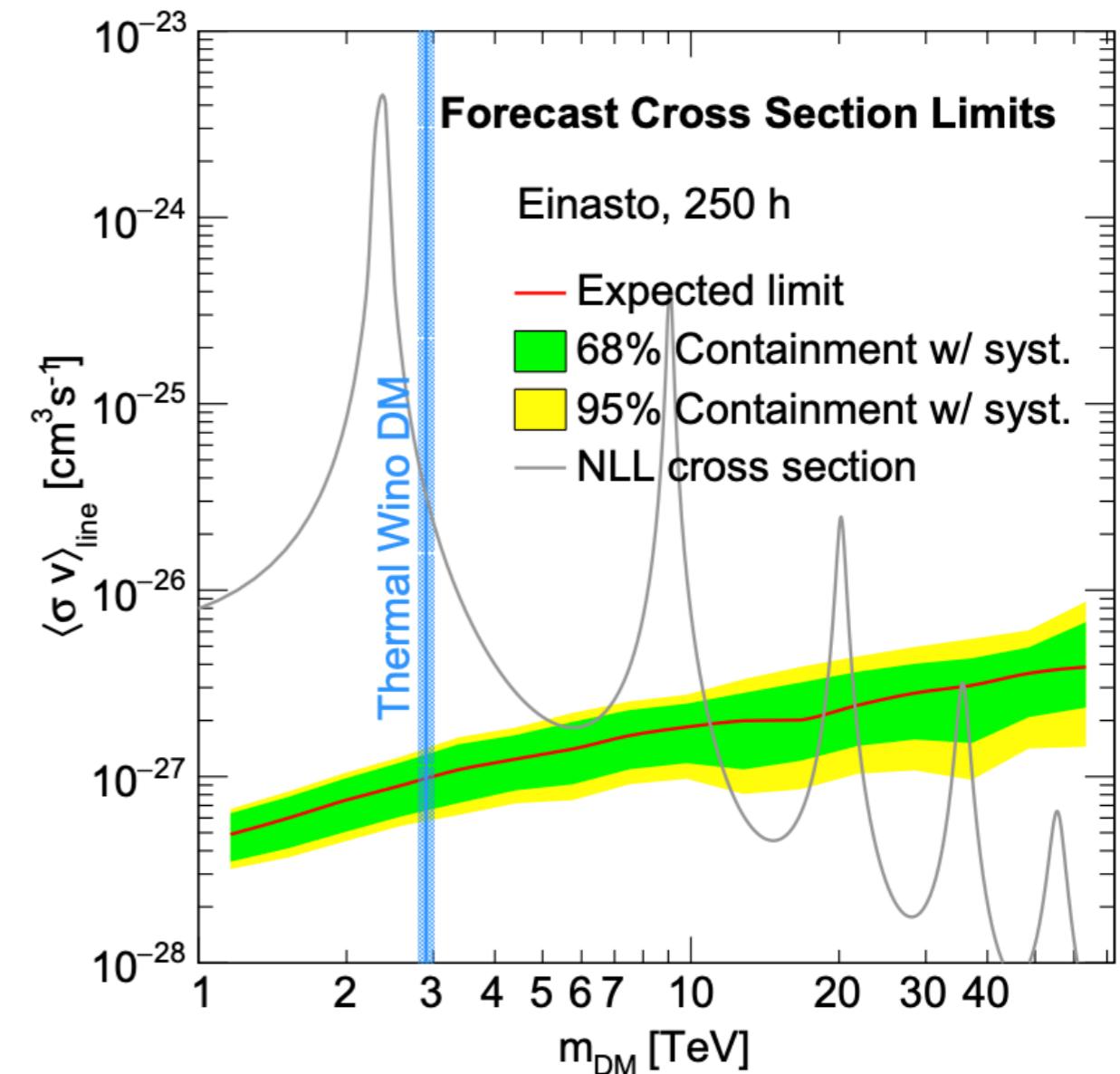
- Indirect Detection (ID)

Doublet



Nicholas L. Rodd, Benjamin R. Safdi, Weishuang Linda Xu,
Phys. Rev. D 110 (2024) 4, 043003

Triplet



Rinchiuso et al., Phys. Rev. D 98 (2018) 12, 123014
Cohen, Lisanti, Pierce, Slatyer, JCAP 10 (2013) 061

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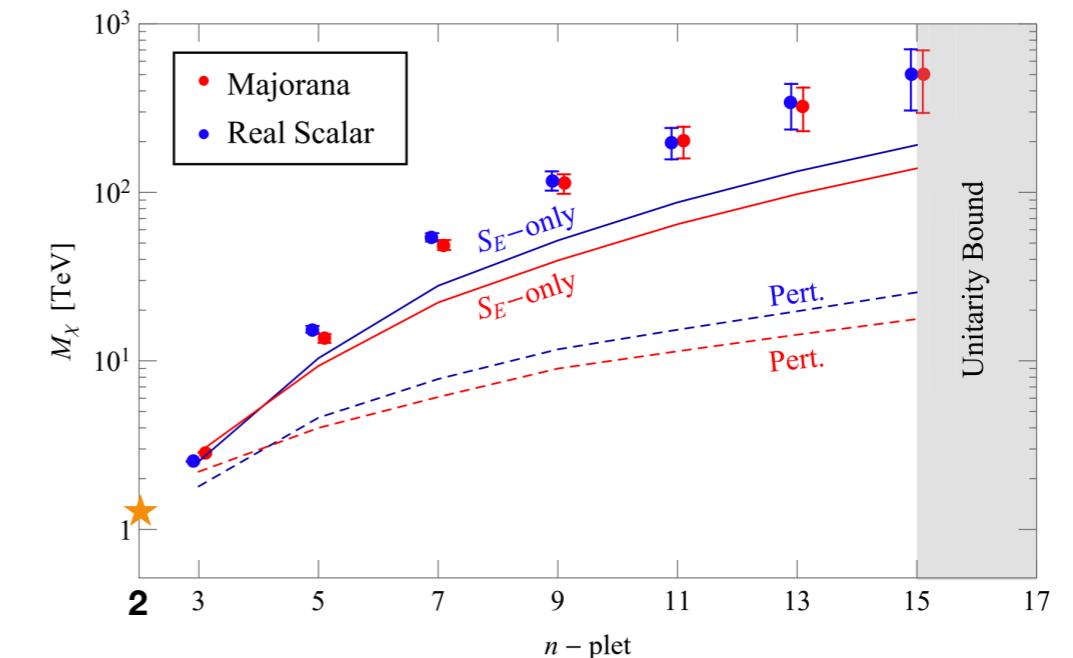
- Towards the Thermal Targets:

- Electroweak states
- 1-100 TeV masses
- 100% of Dark Matter

A task for future colliders!

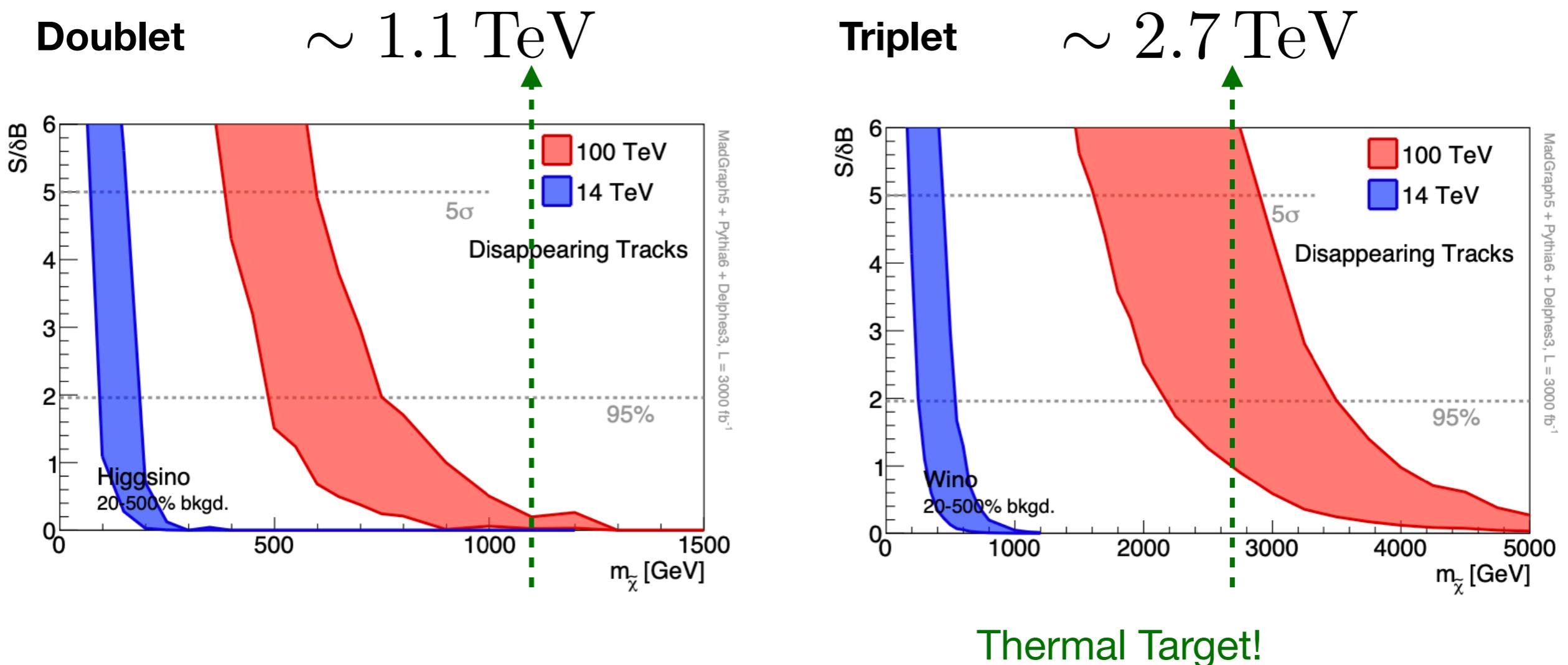


CLIC
3 TeV
e+e-

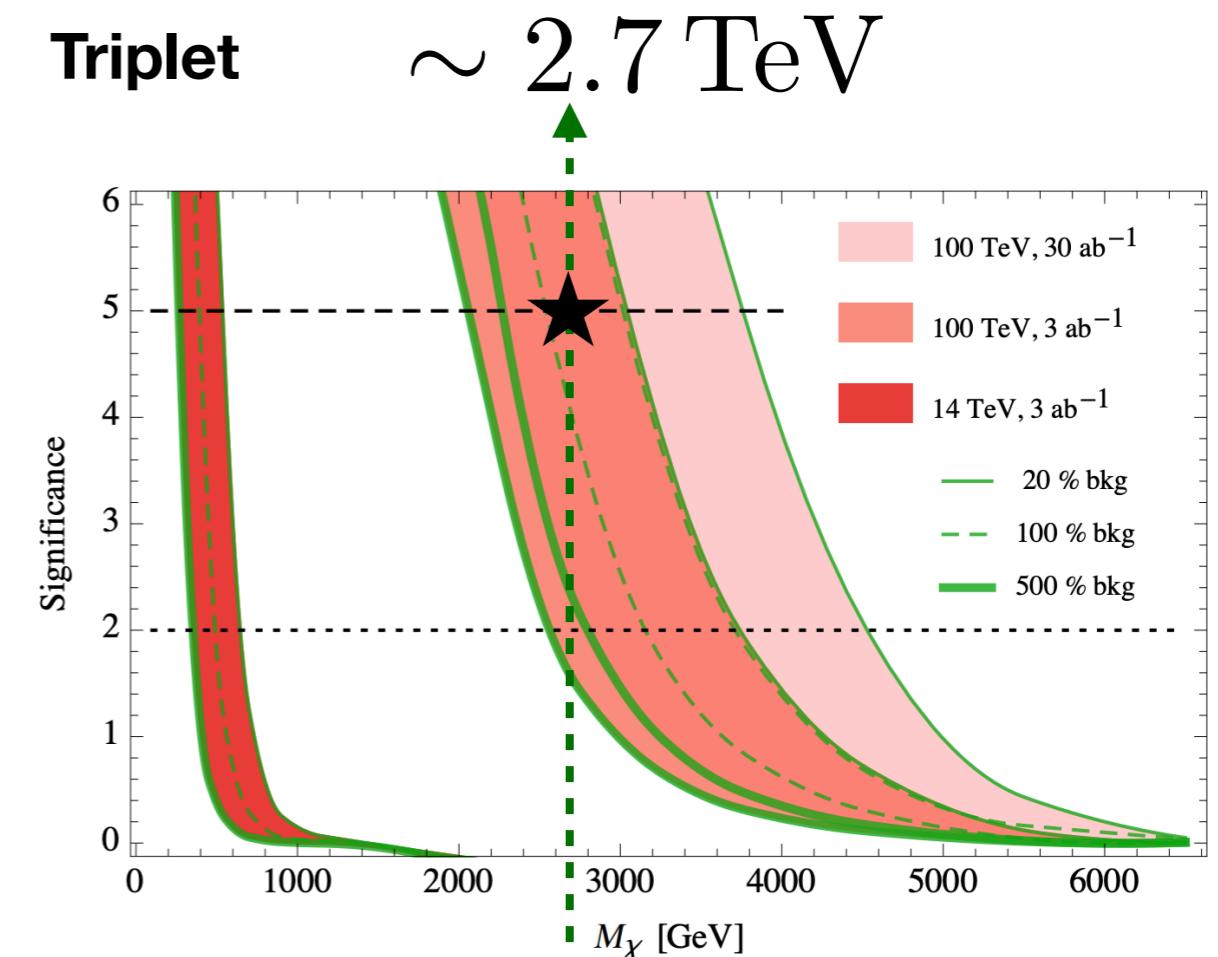
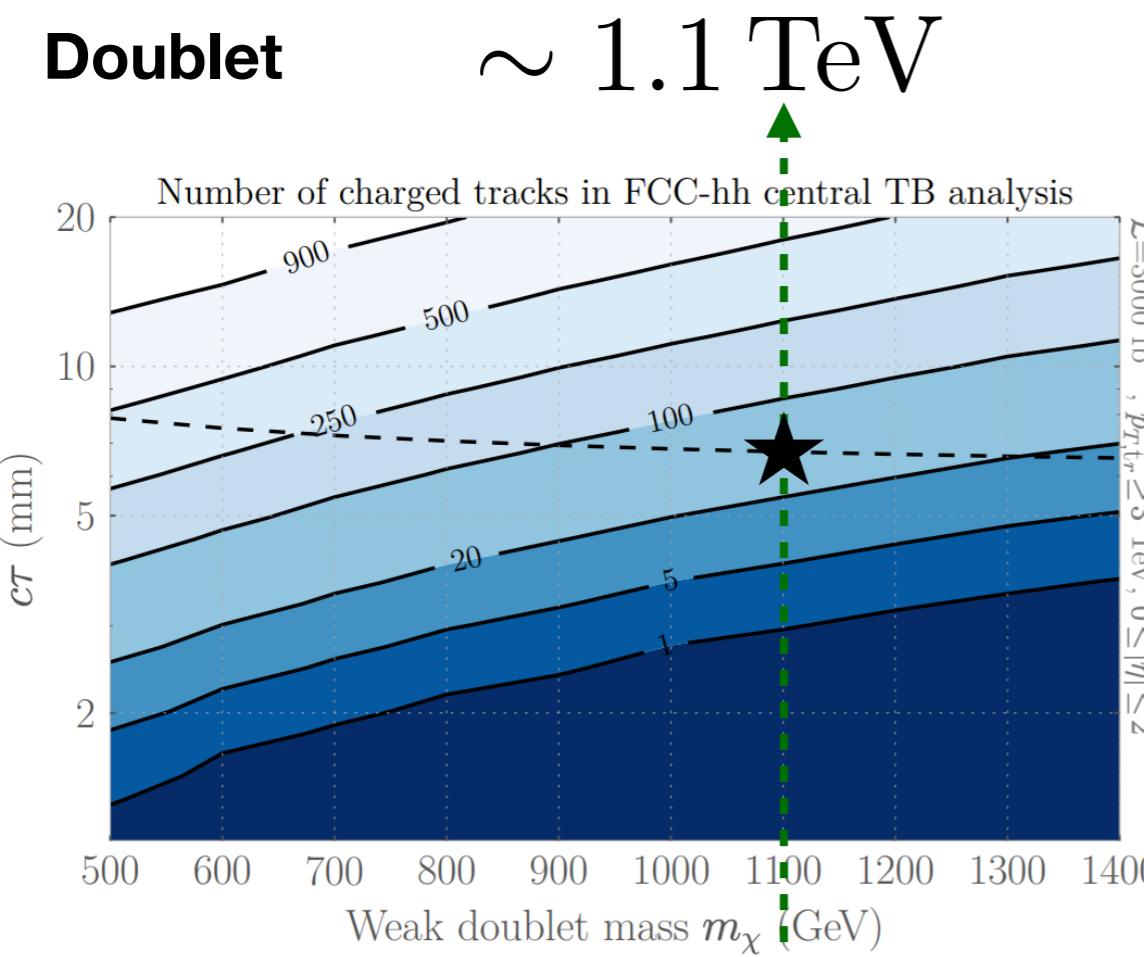


- pp (100 TeV):

M. Low, L. Wang, JHEP 08 (2014) 161



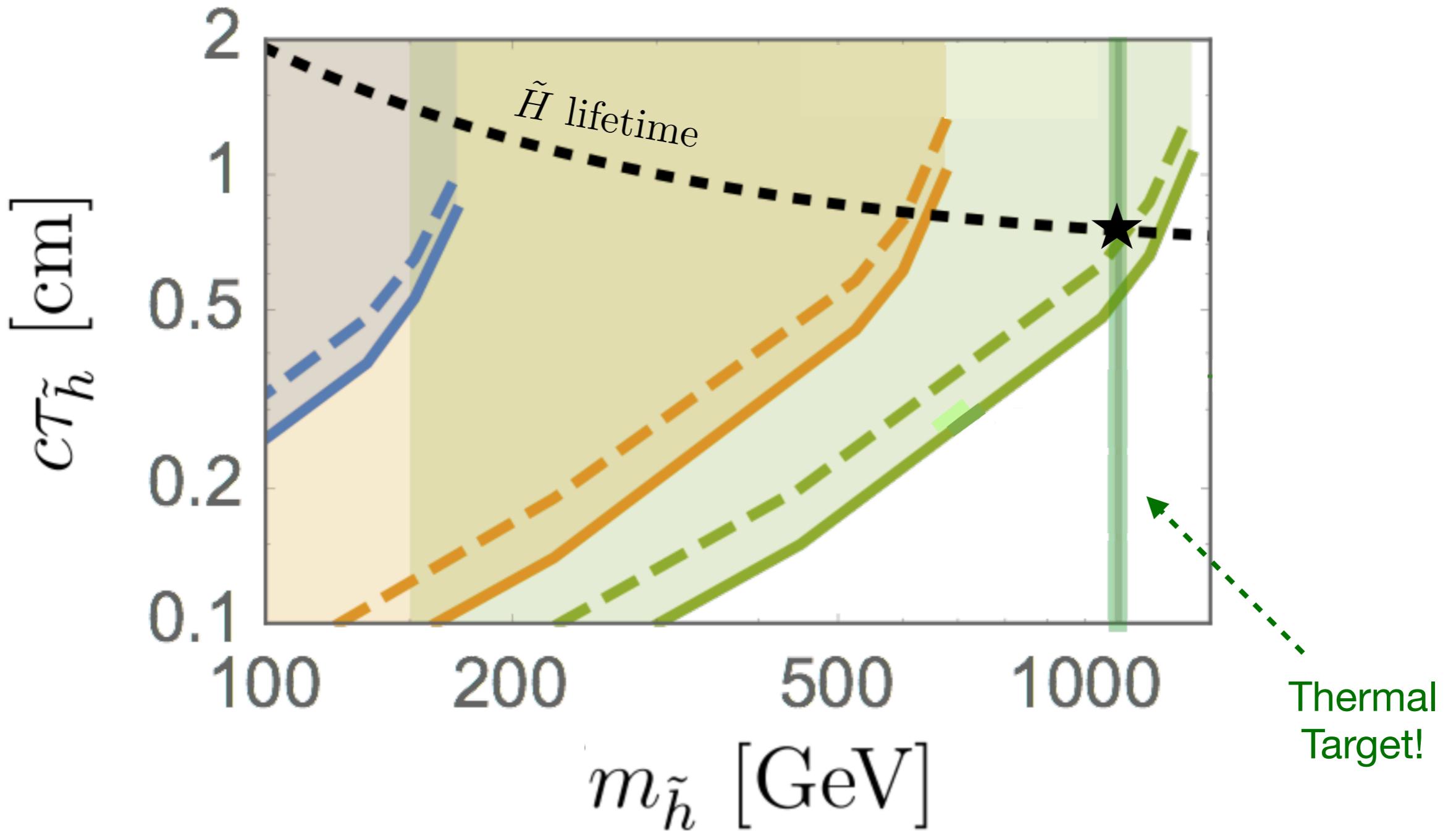
- pp (100 TeV):



R. Mahbubani, P. Schwaller, J. Zurita, JHEP 06 (2017) 119

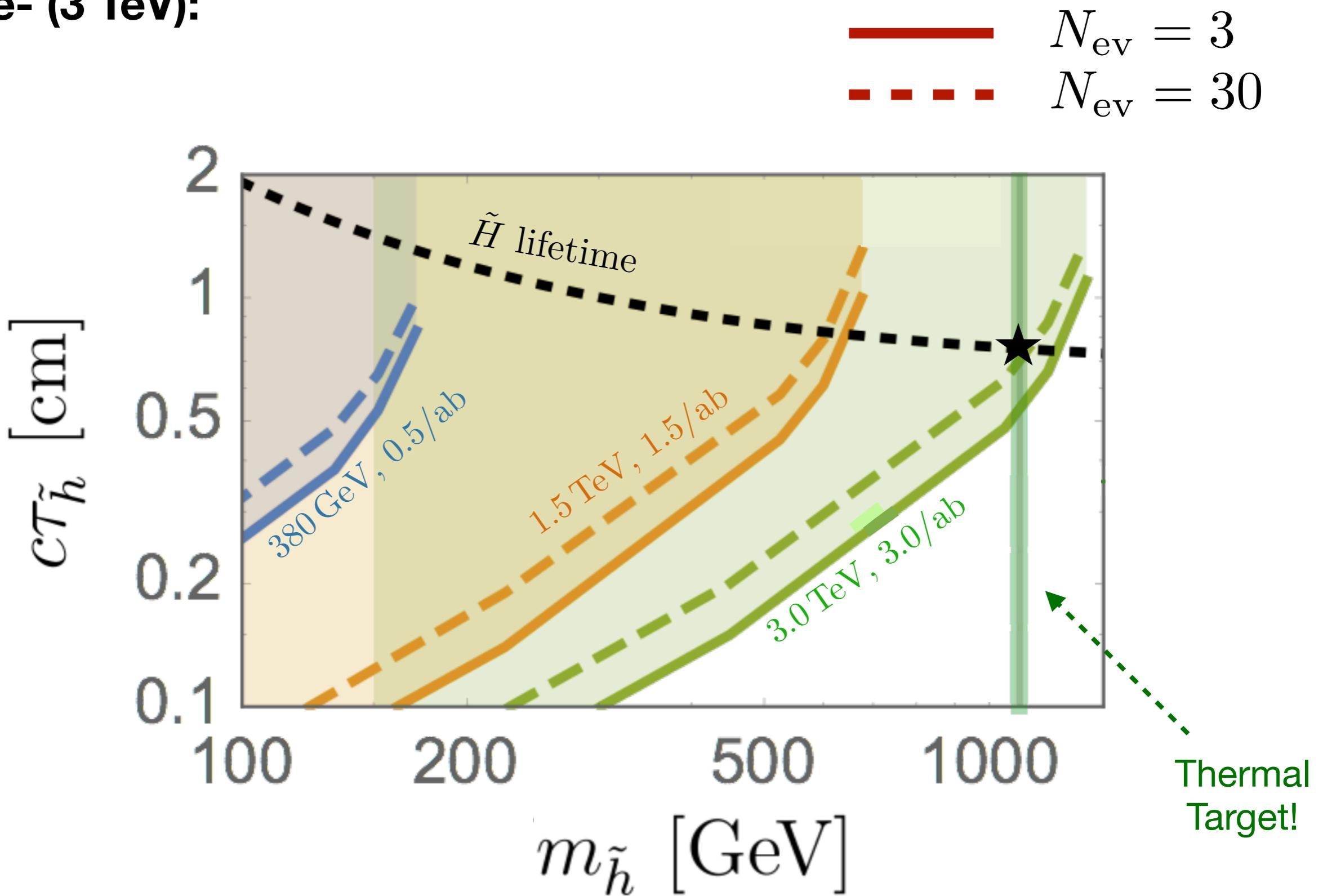
M. Cirelli, F. Sala, M. Taoso, JHEP 10 (2014) 033

- e+e- (3 TeV):



CLIC Collaboration, CERN Yellow Rep. Monogr. 3 (2018)

- e+e- (3 TeV):



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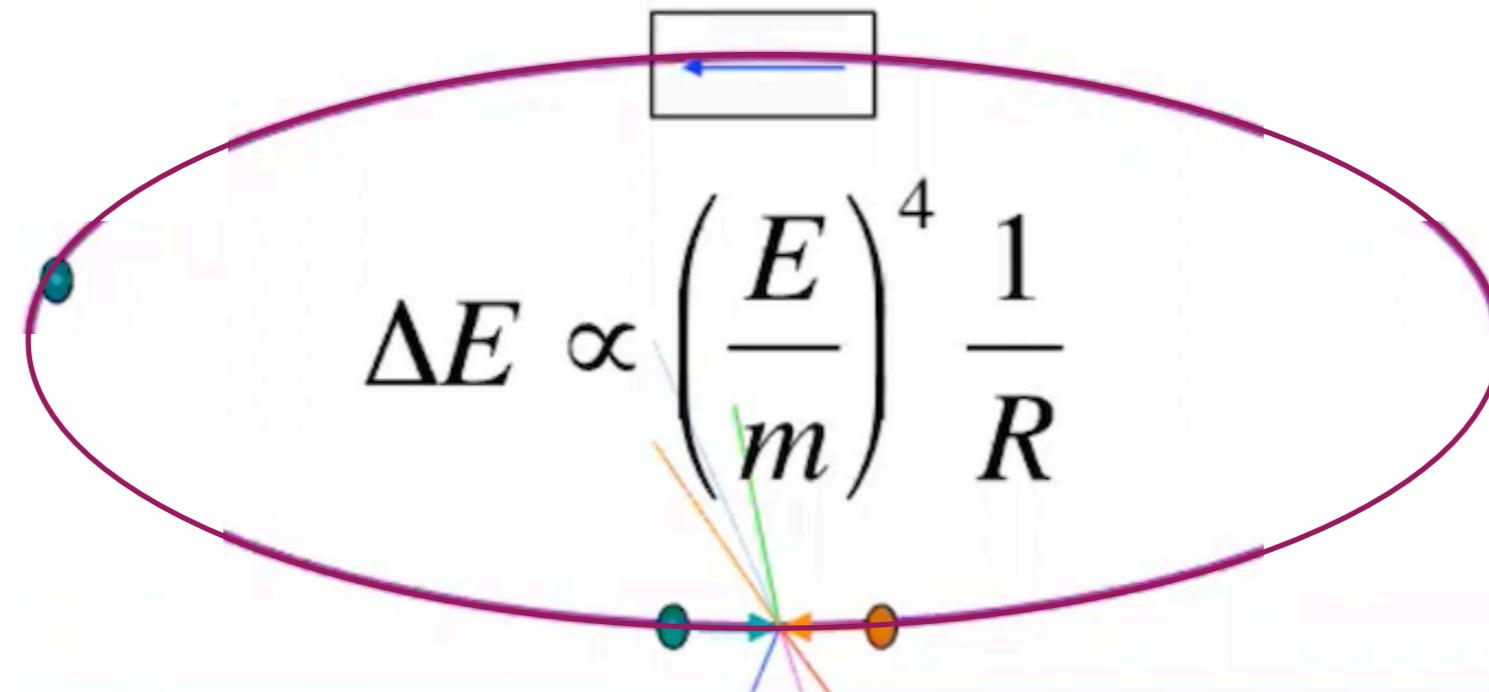
- Disappearing Tracks
- Soft Tracks

Summary

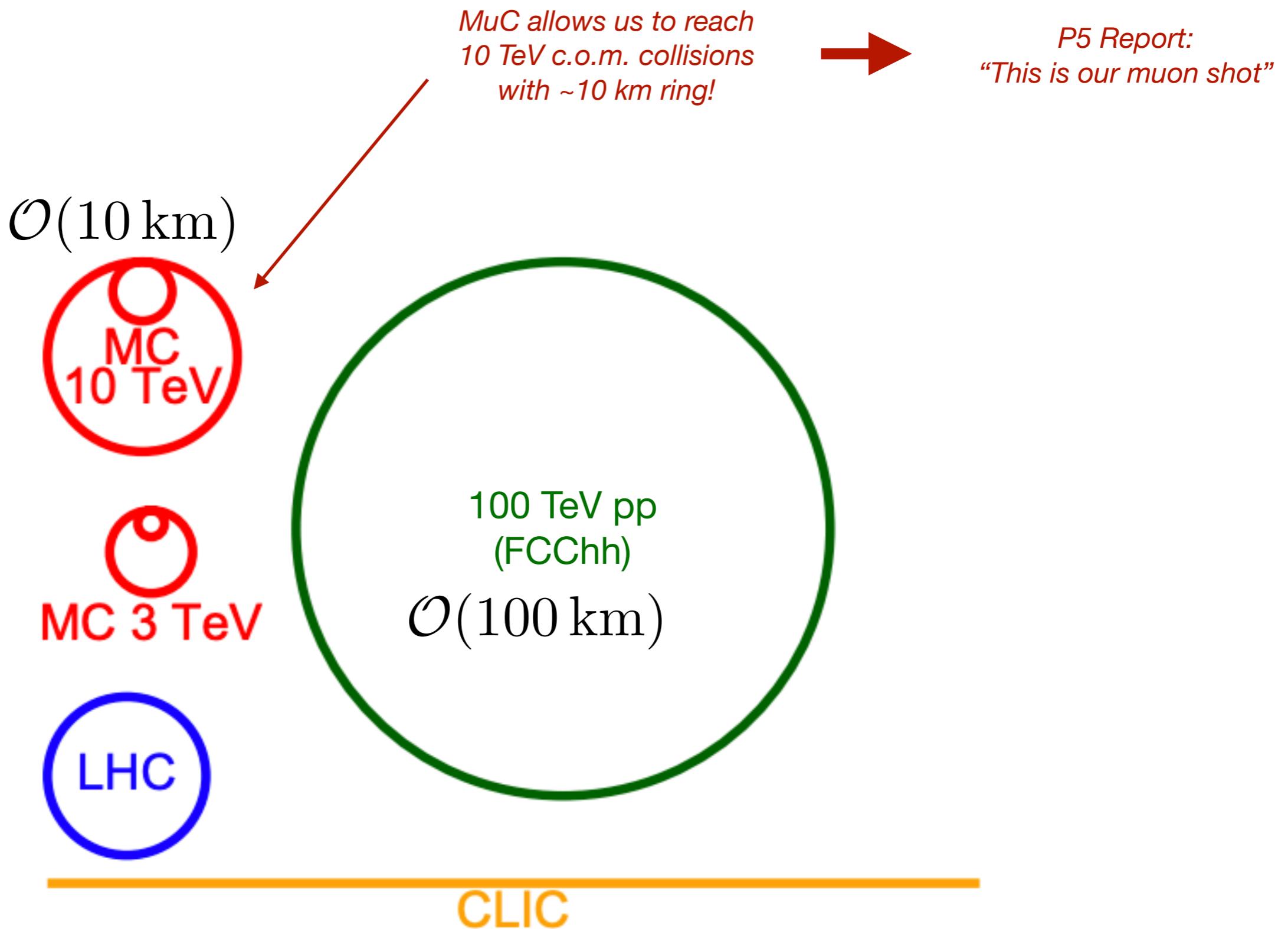
- **Muon Colliders:**

Advantages of circular lepton colliders:

- (i) Multi-pass! (recycle beam; lumi, power)
- (ii) All CoM energy is available
- (iii) Clean (from huge hadronic backgrounds!)

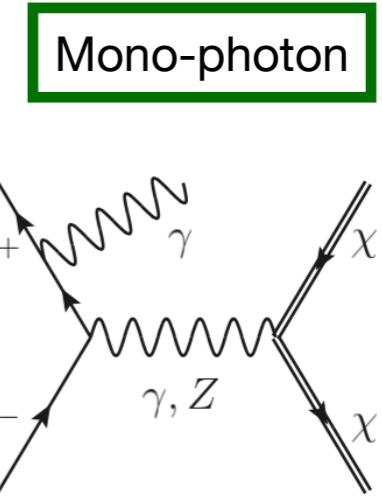
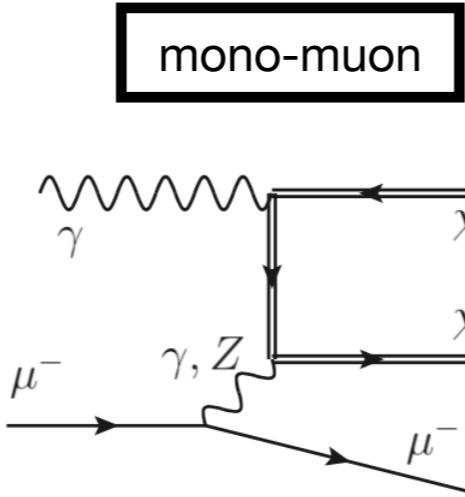
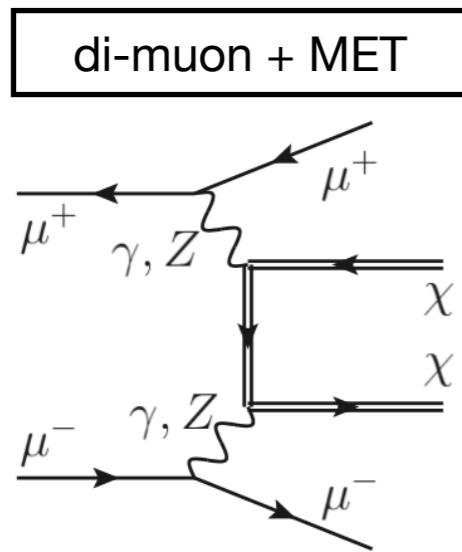


- Muon Colliders:



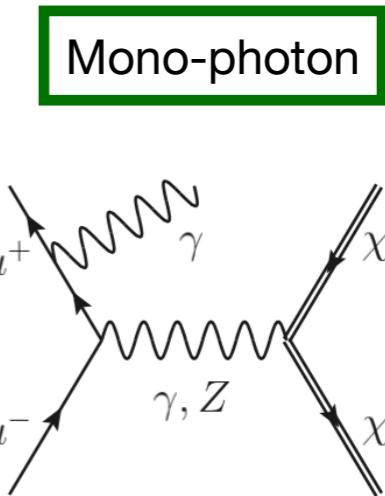
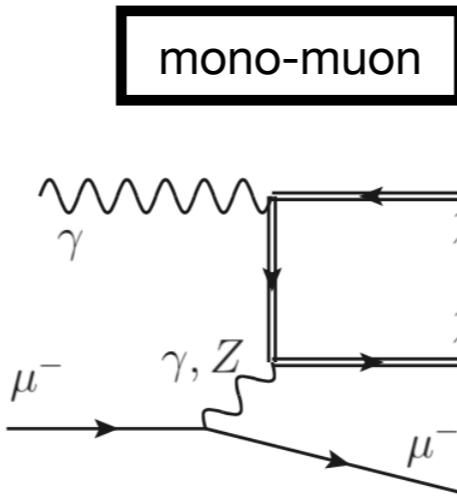
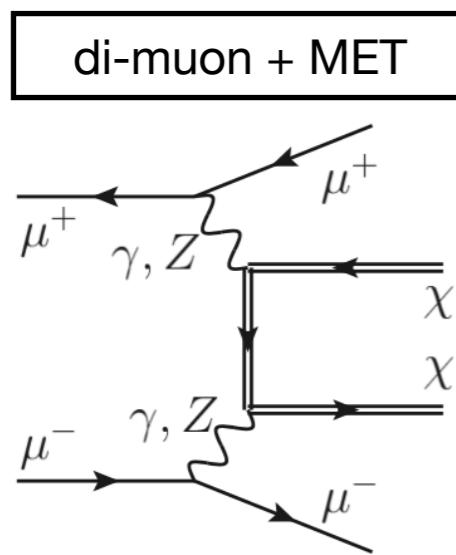
- **Muon Colliders:**

T. Han, Z. Liu, L. Wang, X. Wang,
Phys. Rev. D 103 (2021) 7, 075004



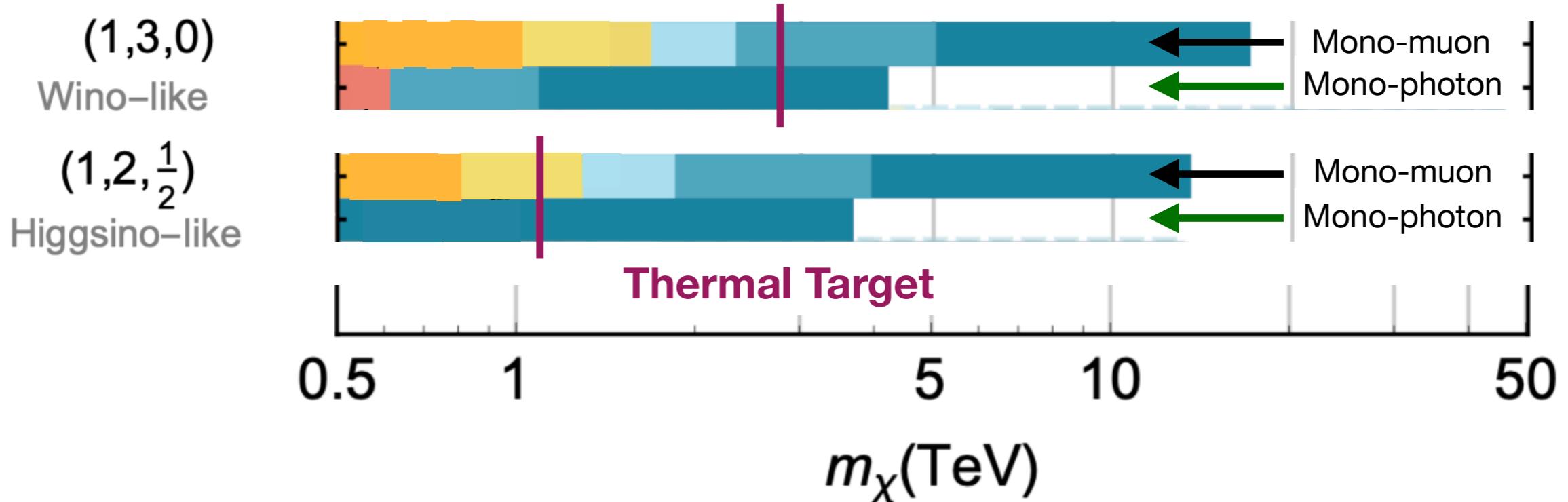
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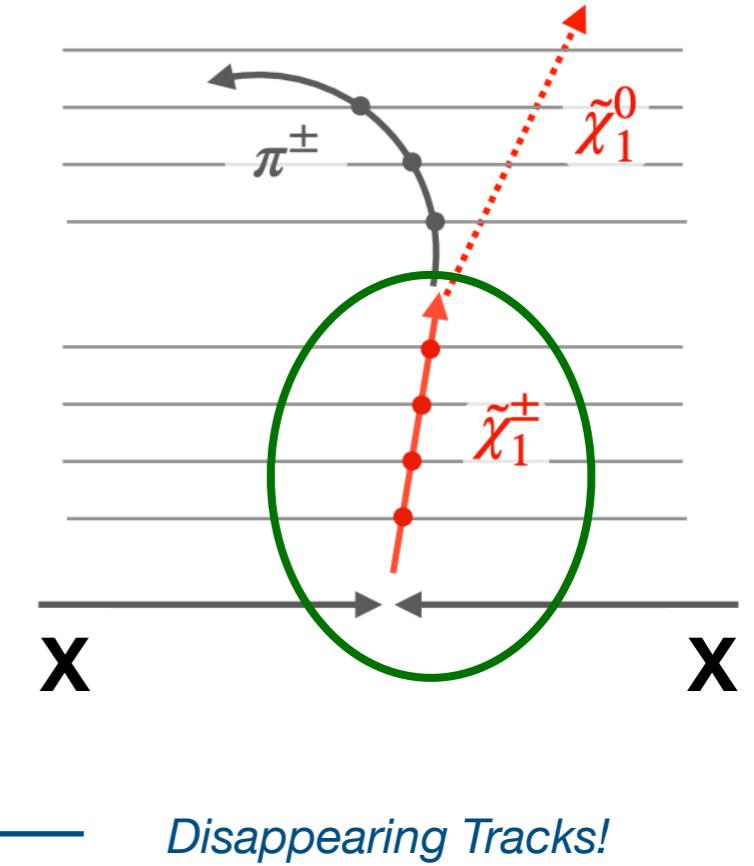
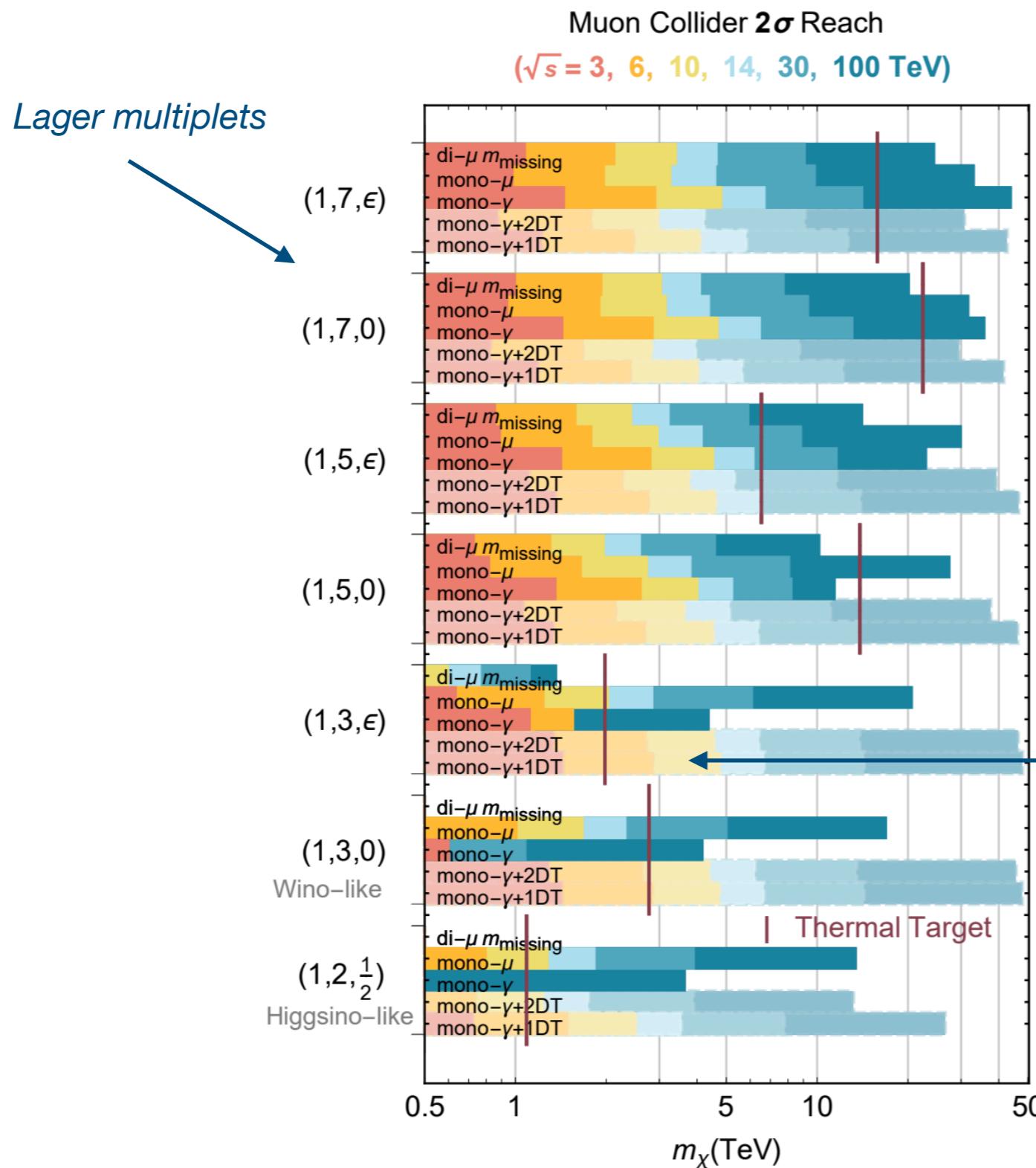
Muon Collider 2σ Reach

($\sqrt{s} = 3, 6, 10, 14, 30, 100 \text{ TeV}$)

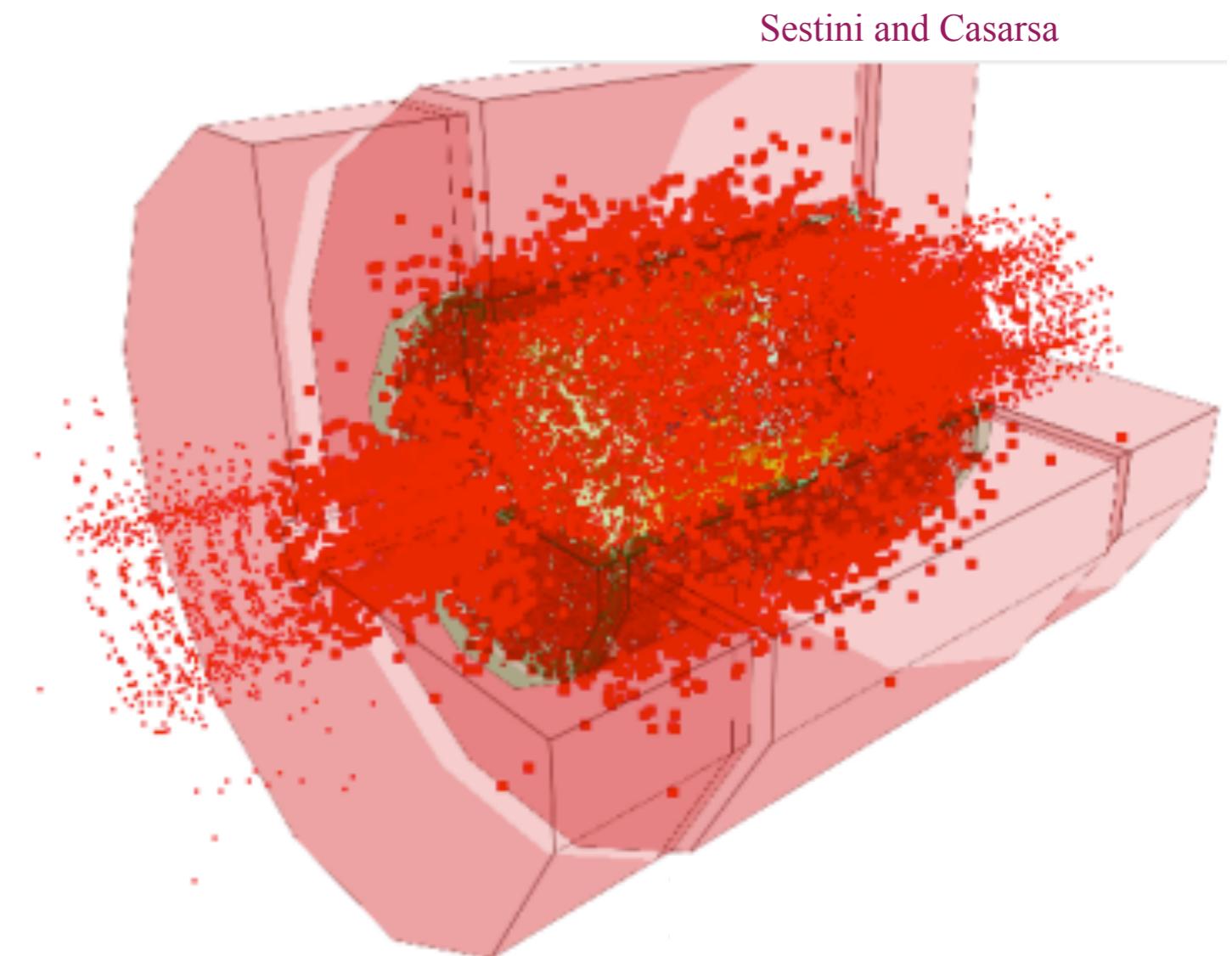
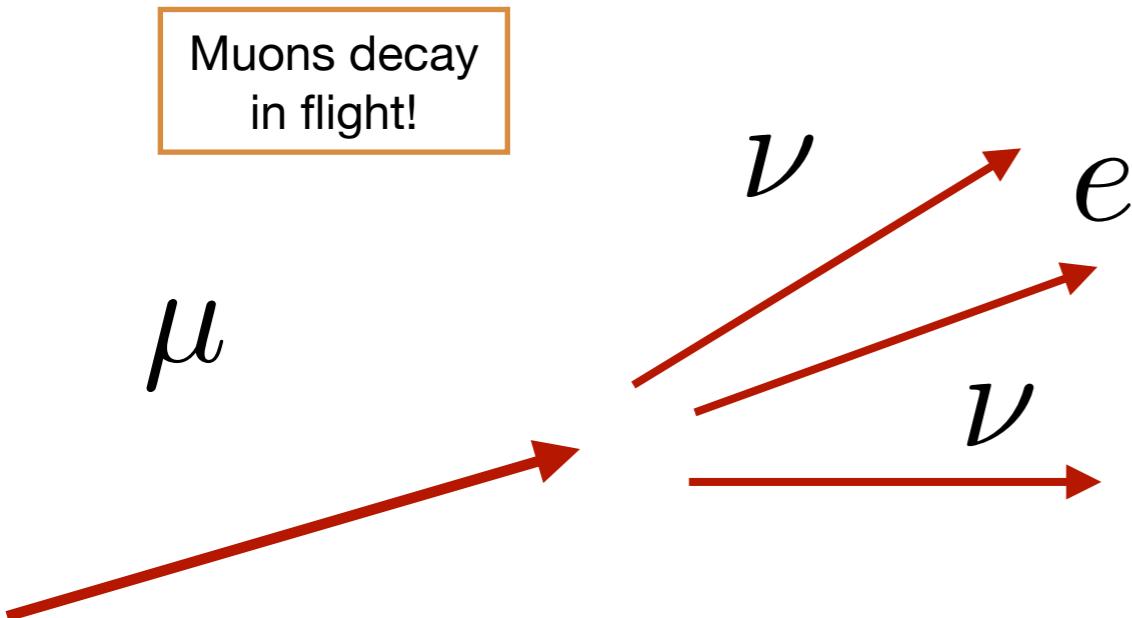


• Muon Colliders:

T. Han, Z. Liu, L. Wang, X. Wang,
Phys. Rev. D 103 (2021) 7, 075004

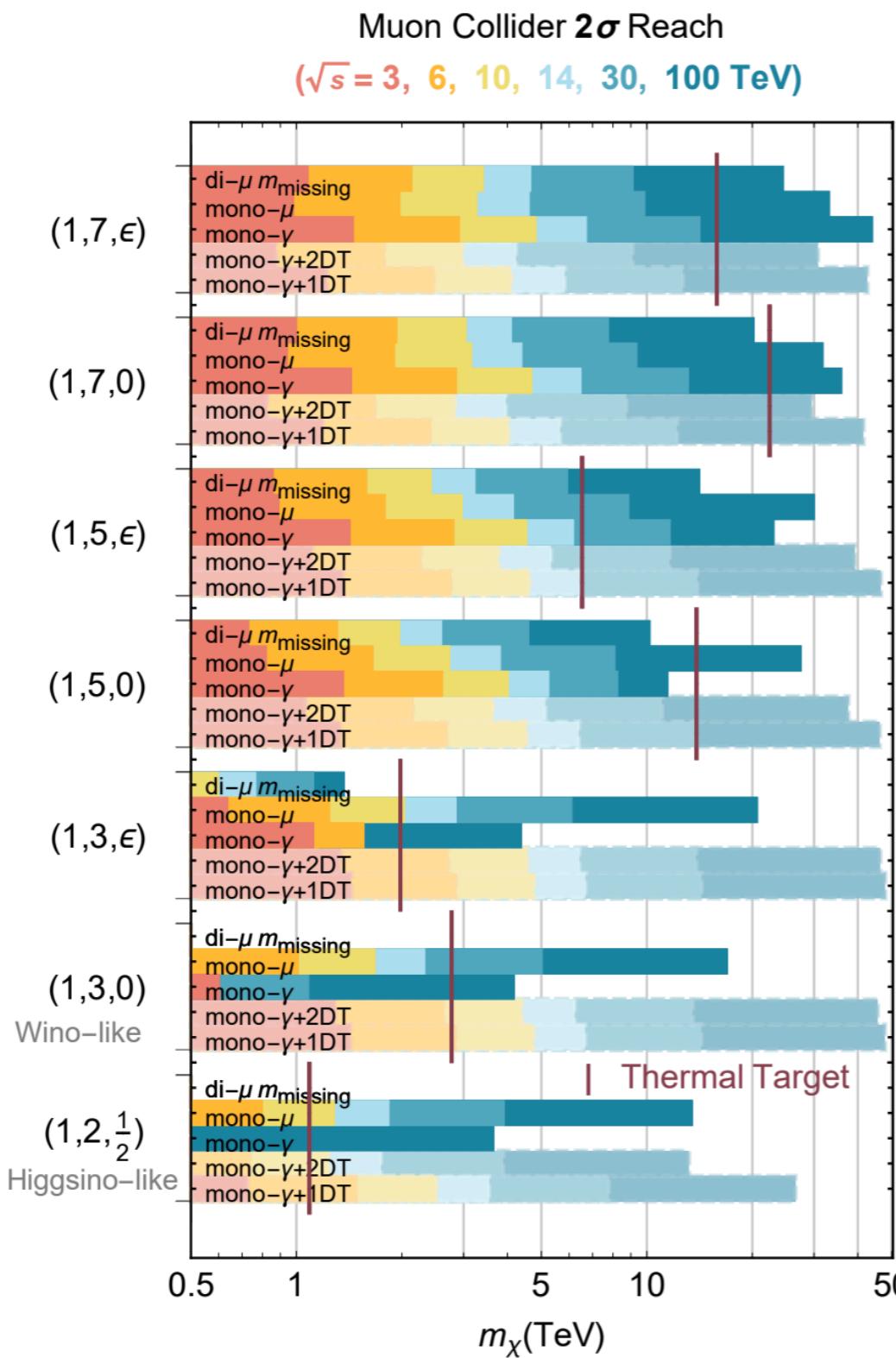


- Muon Colliders:



• Muon Colliders:

T. Han, Z. Liu, L. Wang, X. Wang,
Phys. Rev. D 103 (2021) 7, 075004



How about full detector simulation?

How about the BIB?

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- **Disappearing Tracks**
- Soft Tracks

Summary

- Full Detector Simulation + BIB:

DESY 21-019 IFIC/21-03

Hunting wino and higgsino dark matter at the muon collider with disappearing tracks

Rodolfo Capdevilla,^{a,b} Federico Meloni,^c Rosa Simoniello,^d Jose Zurita^e

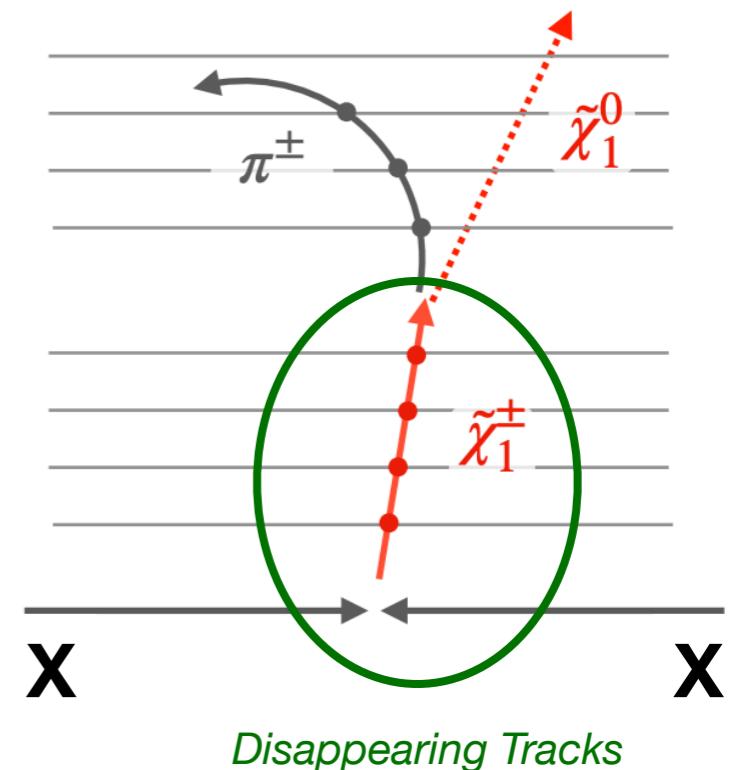
^aDepartment of Physics, University of Toronto, Canada

^bPerimeter Institute for Theoretical Physics, Waterloo, Ontario, Canada

^cDeutsches Elektronen-Synchrotron DESY, Hamburg, Germany

^dCERN, Geneva, Switzerland

^eInstituto de Física Corpuscular, CSIC-Universitat de València, Valencia, Spain



RC, Federico Meloni, Rosa Simoniello, Jose Zurita, JHEP 06 (2021) 133



Federico
Meloni, DESY

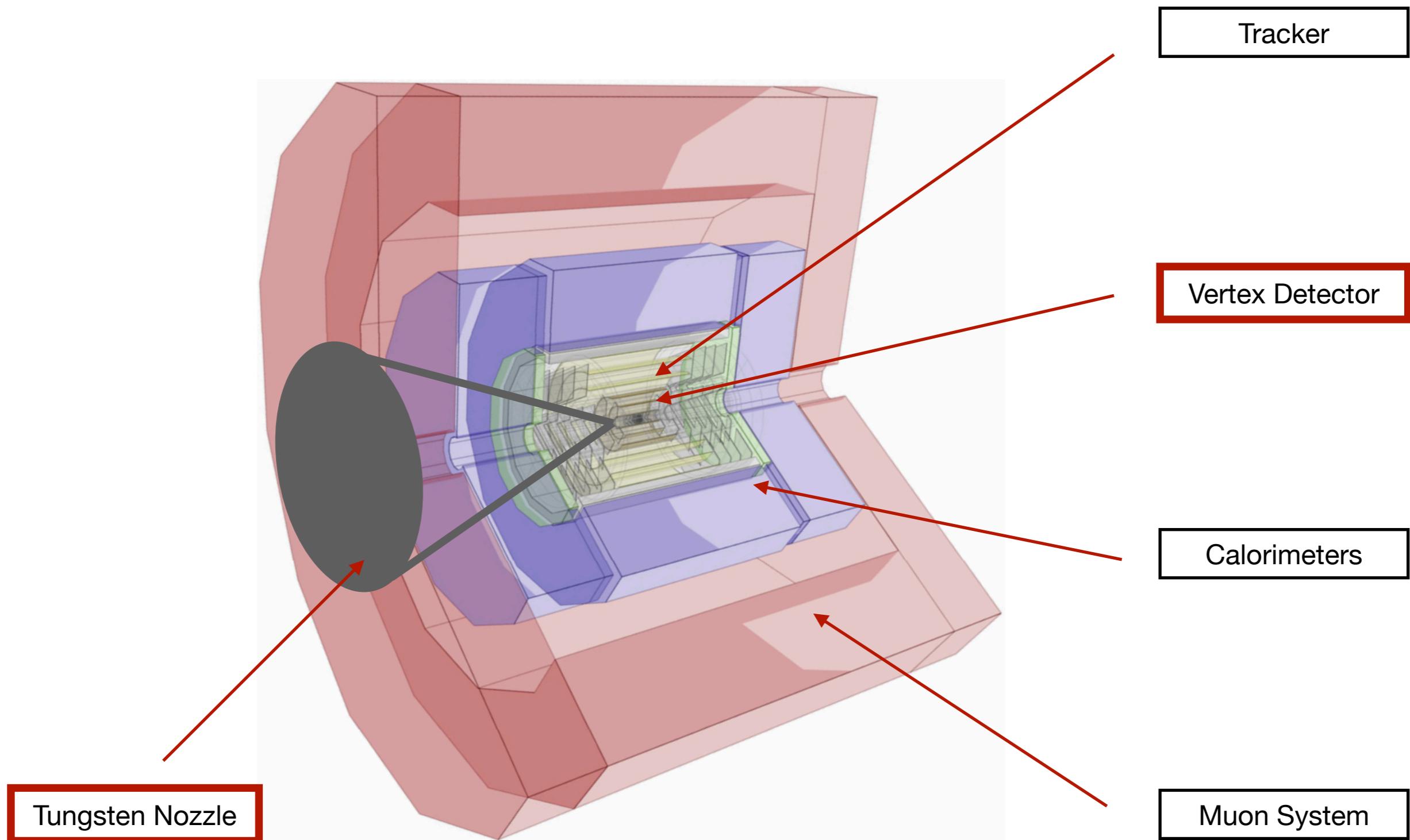


Rosa Simoniello,
CERN

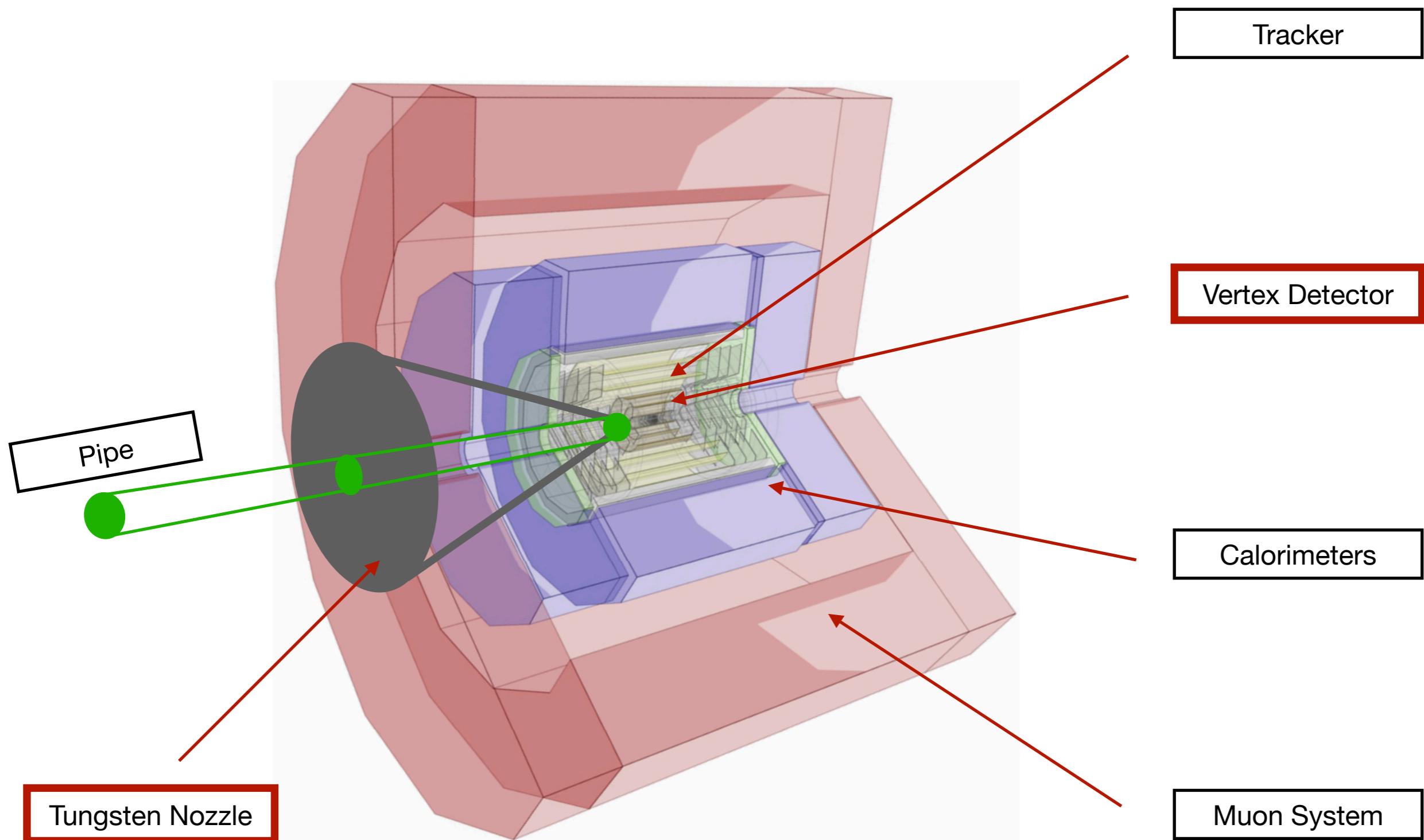


Jose Zurita,
U. Valencia

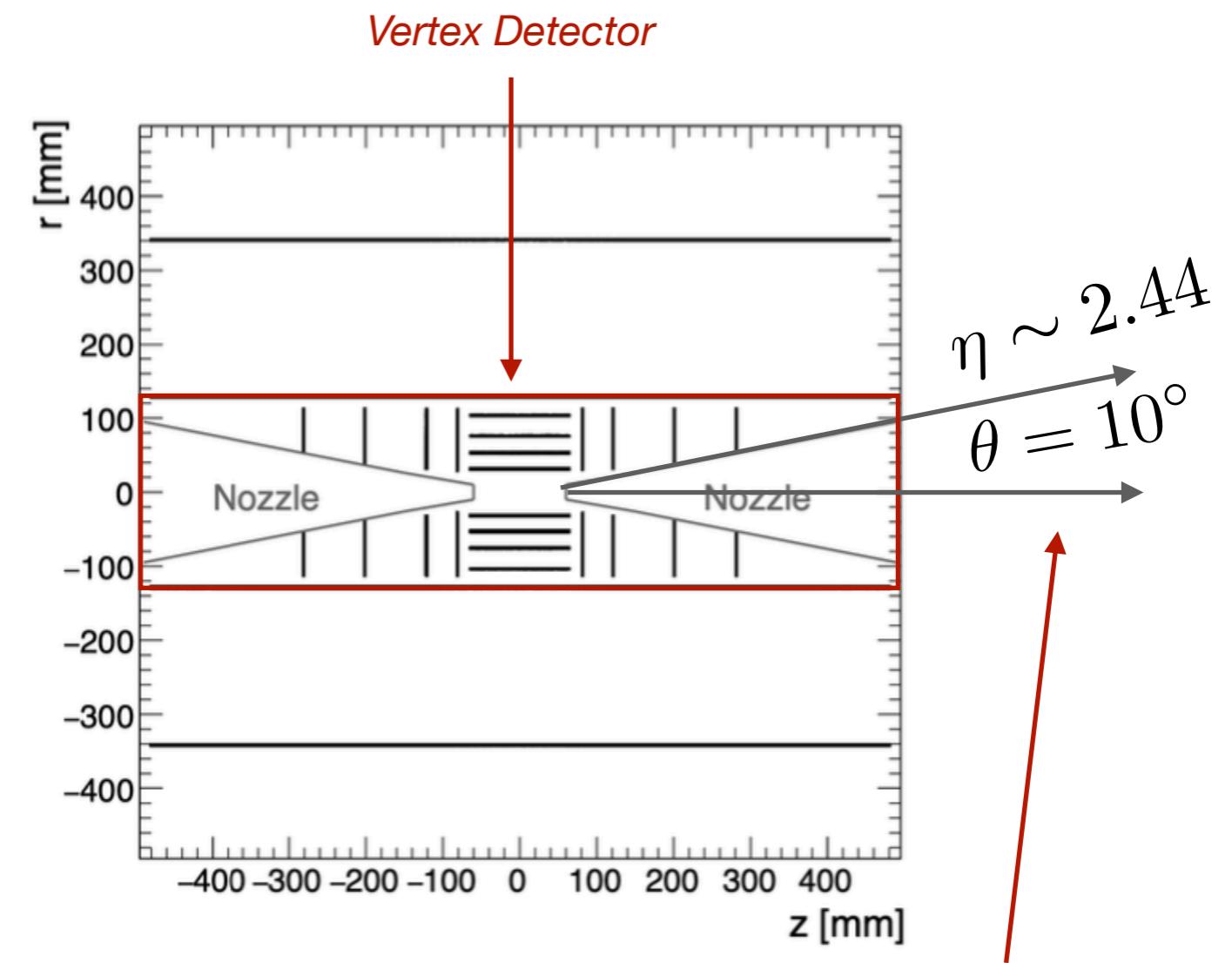
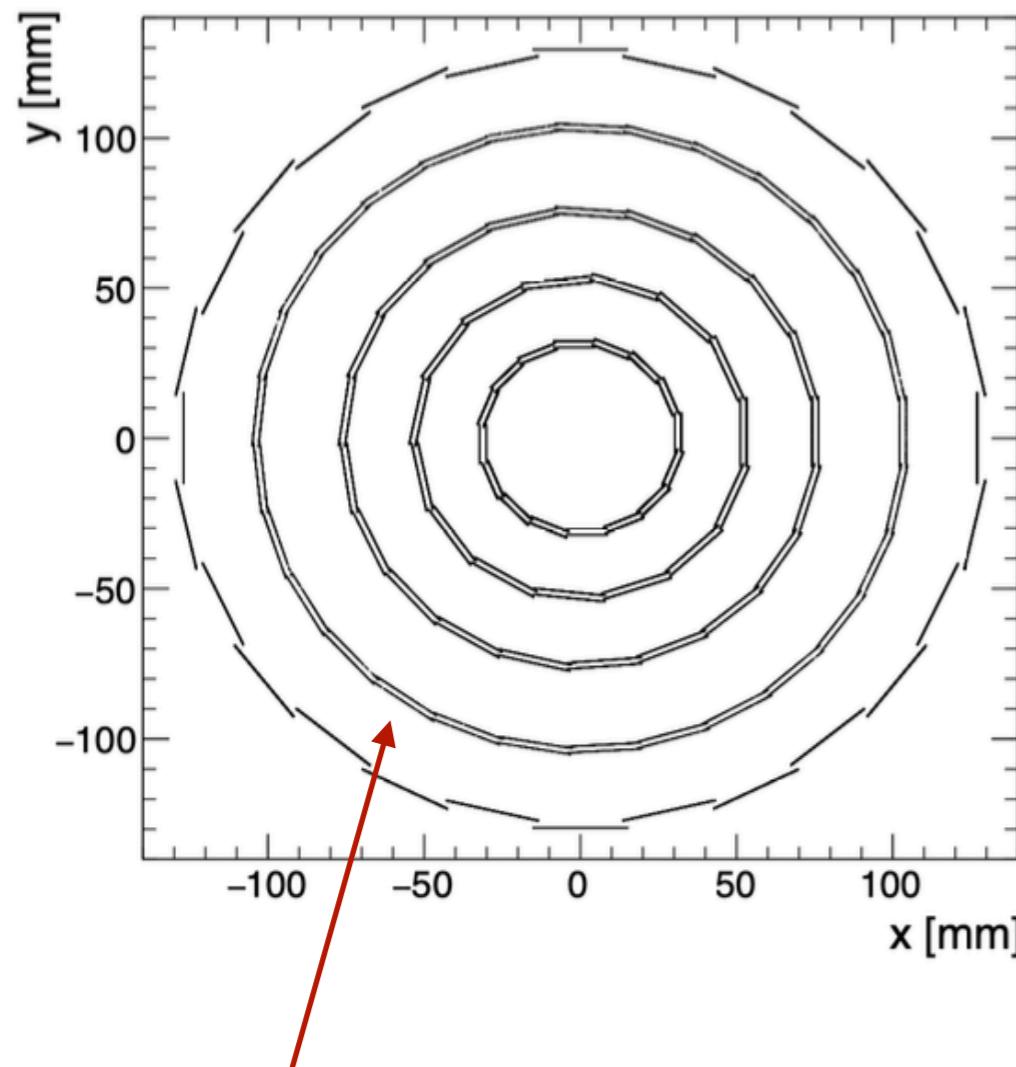
- Design Adapted from CLIC:



- Design Adapted from CLIC:



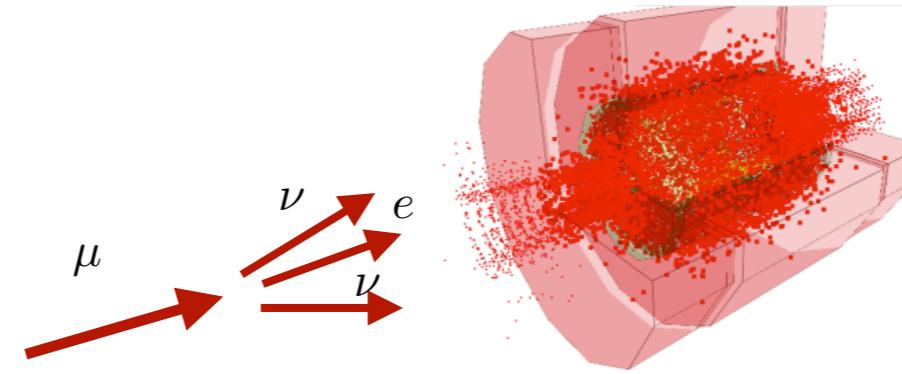
- Focus on the Vertex Detector:



Double layer
tracker

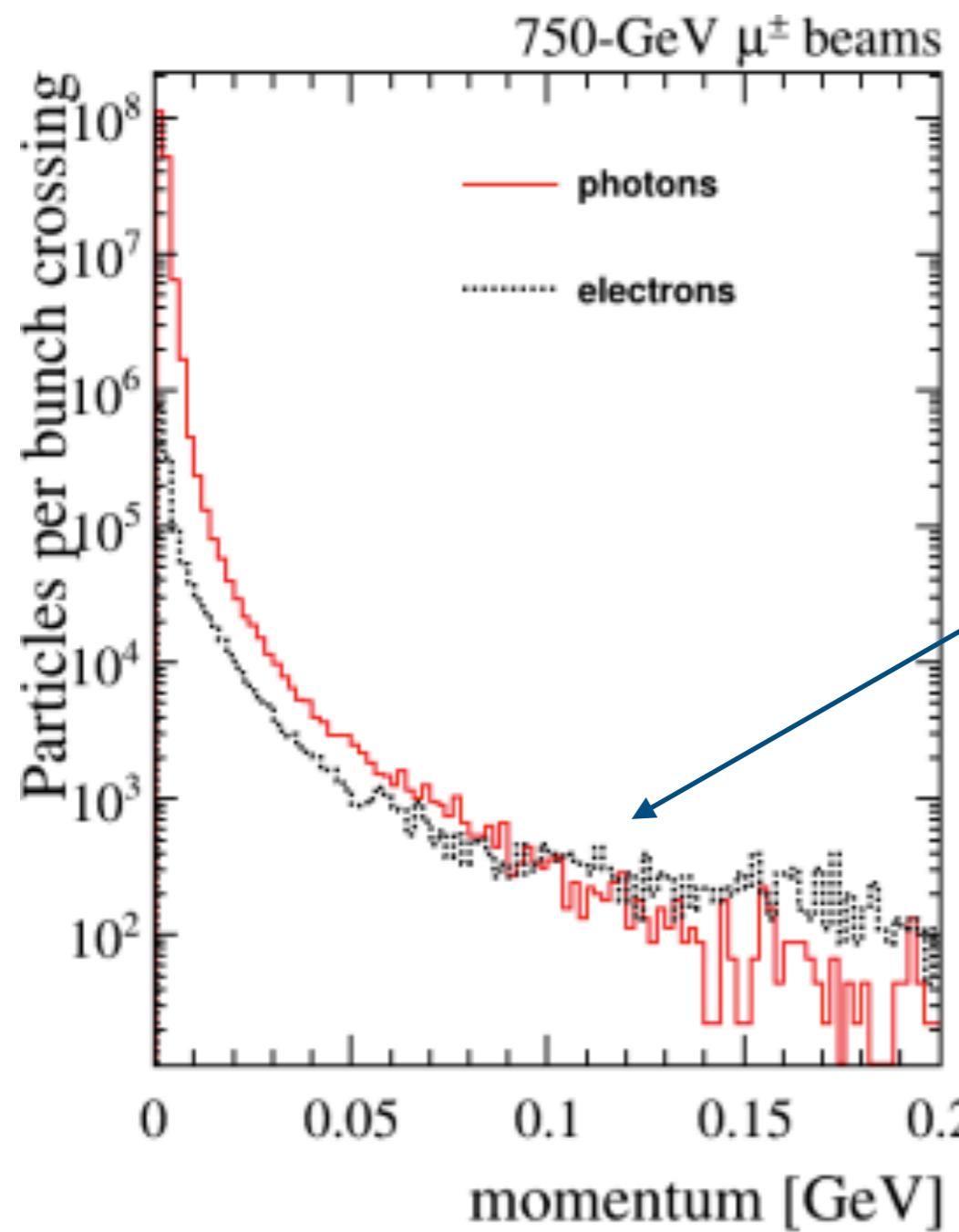
The nozzle covers 10
degrees of the angular
acceptance

- **BIB:**



1. Soft
2. Arrives late
3. Mostly forward

N. Bartosik et al., 2020 JINST 15 P05001



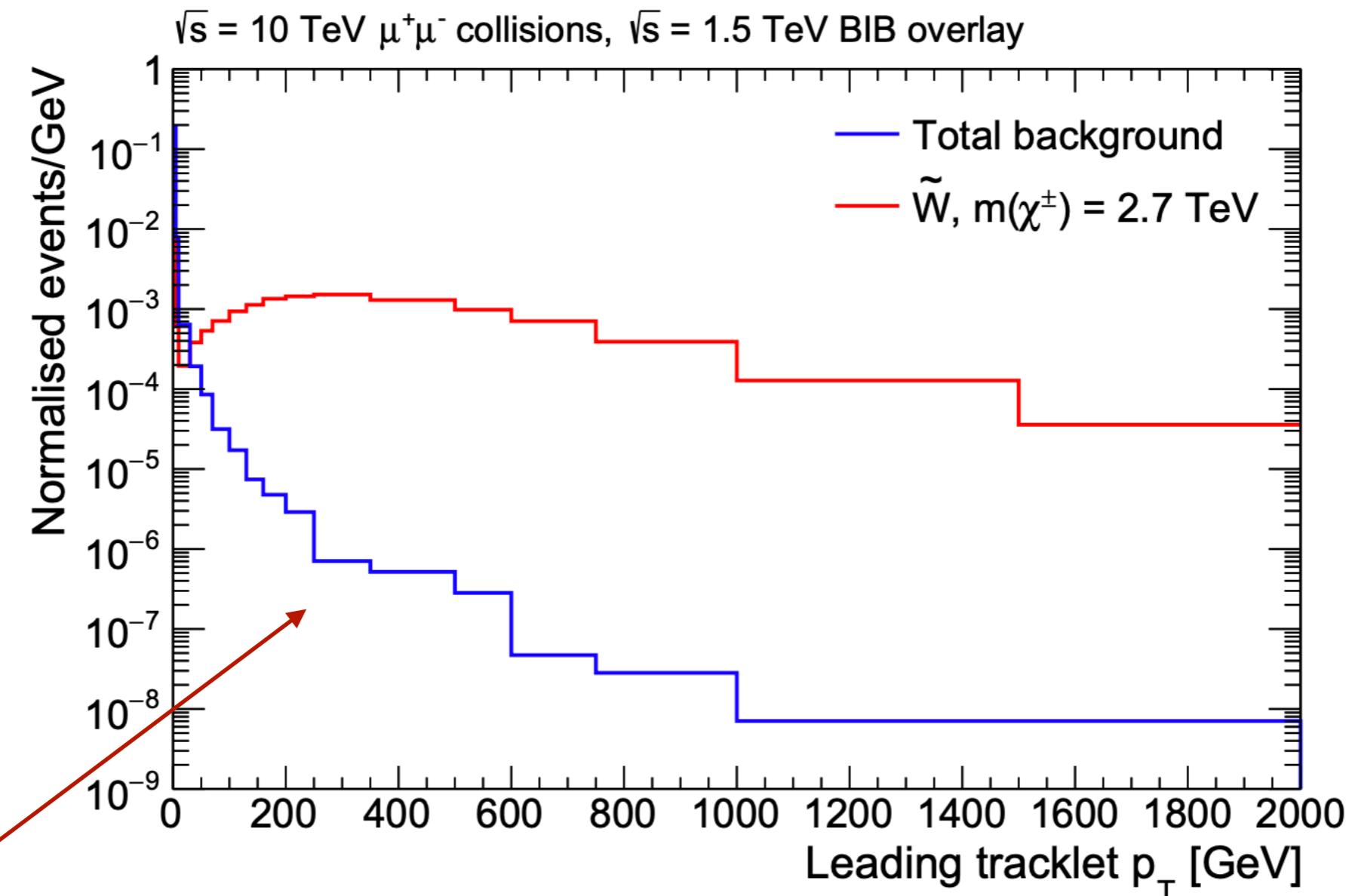
BIB is mainly composed of photons and electrons with a rapidly falling momentum distribution

- Removing the BIB:

1. Soft

2. Arrives late

3. Mostly forward



BIB tracklets tend to have low p_T

- **BIB:**

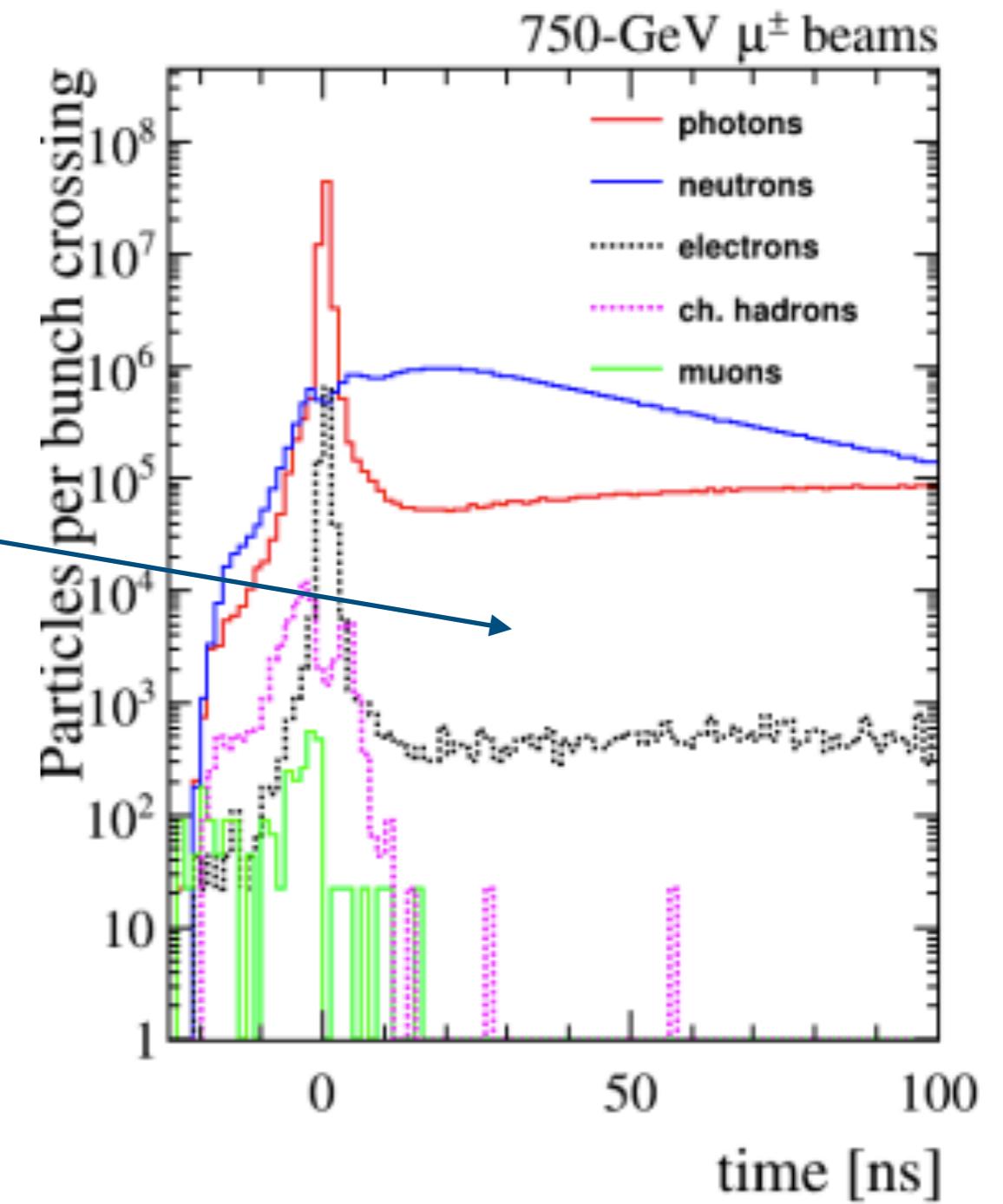
1. Soft

2. Arrives late

3. Mostly forward

N. Bartosik et al., 2020 JINST 15 P05001

The BIB spreads over a wide range of time

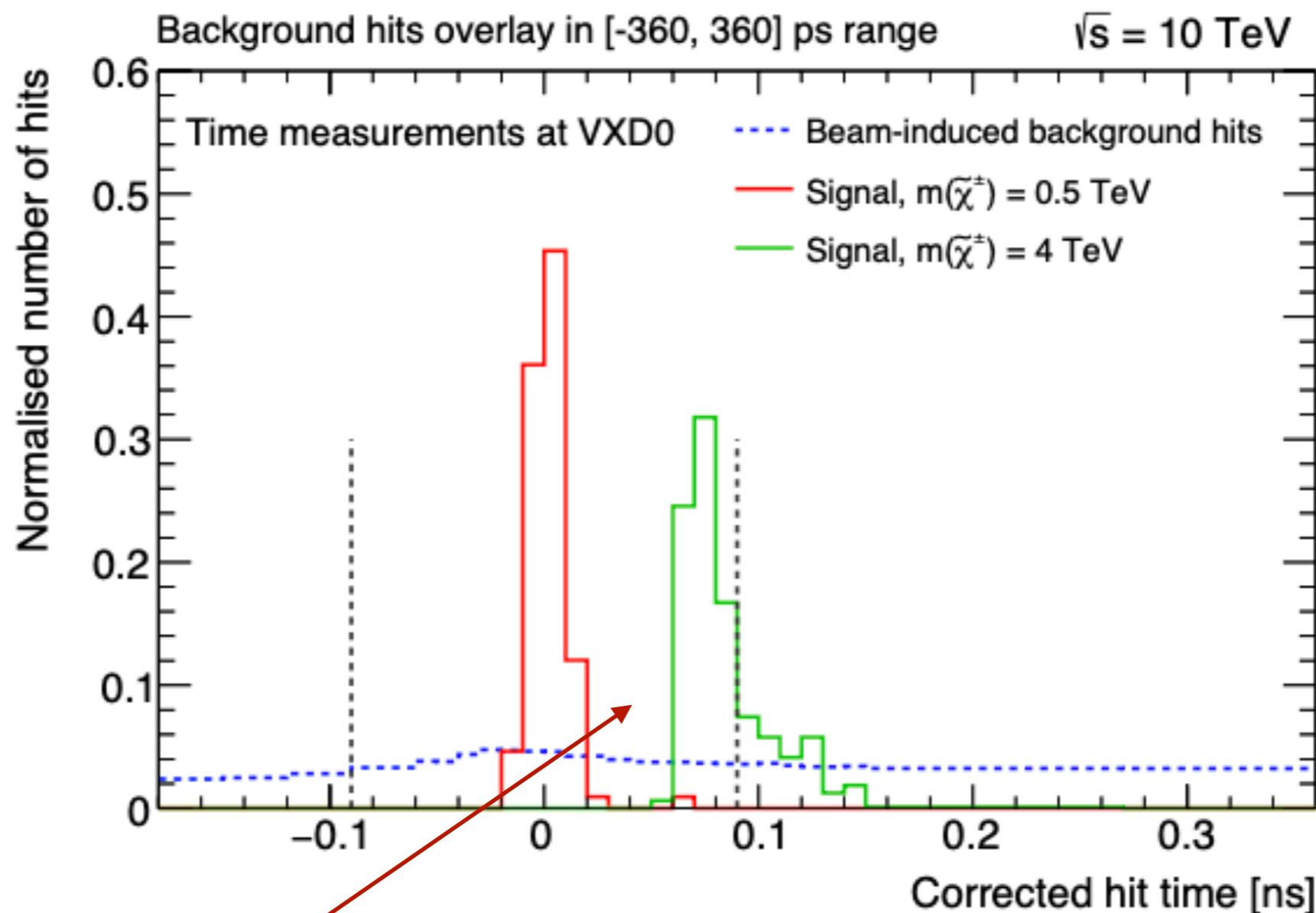


- **Removing the BIB:**

1. Soft

2. Arrives late

3. Mostly forward



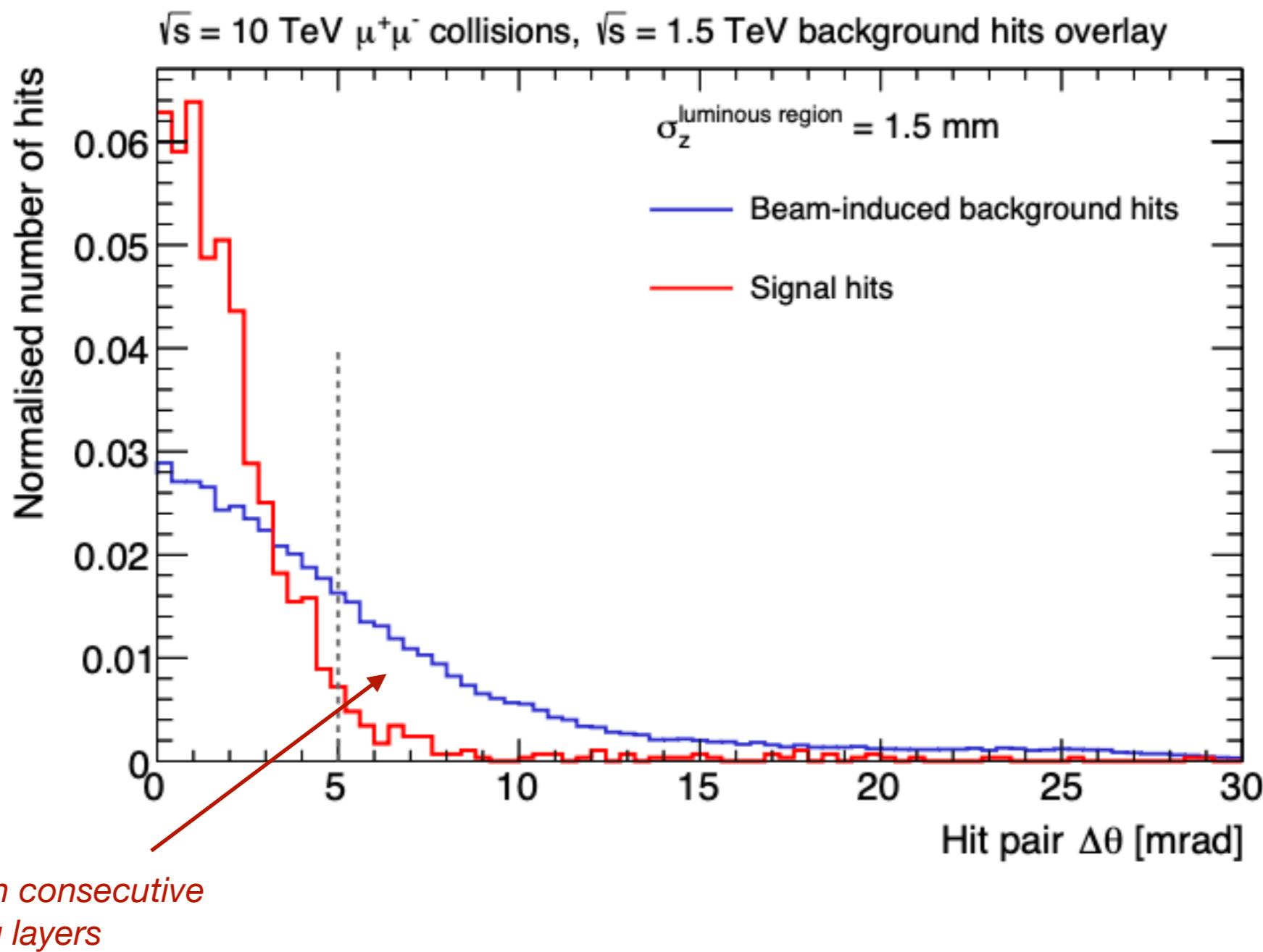
For a given mass we choose the appropriate time window!

- Removing the BIB:

1. Soft

2. Arrives late

3. Mostly forward



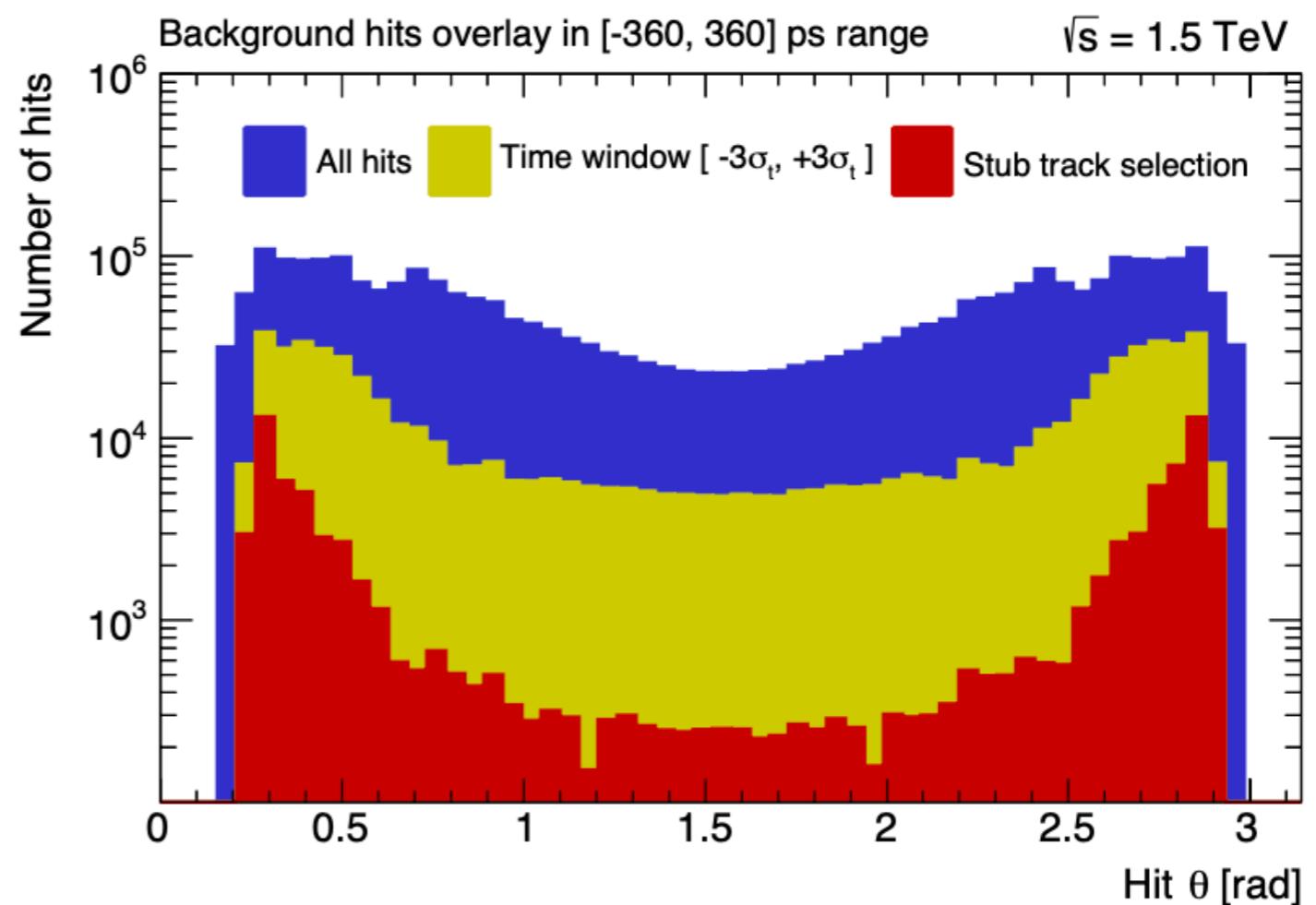
- **Removing the BIB:**

	SR_{1t}^γ	SR_{2t}^γ
Total background	187.8 ± 0.6	0.16 ± 0.05
\tilde{W} , 2.7 TeV, $\tau = 0.2$ ns	313 ± 5	168 ± 2
\tilde{H} , 1.1 TeV, $\tau = 0.02$ ns	53.0 ± 0.7	3.92 ± 0.05

*Three orders of magnitude
background suppression*



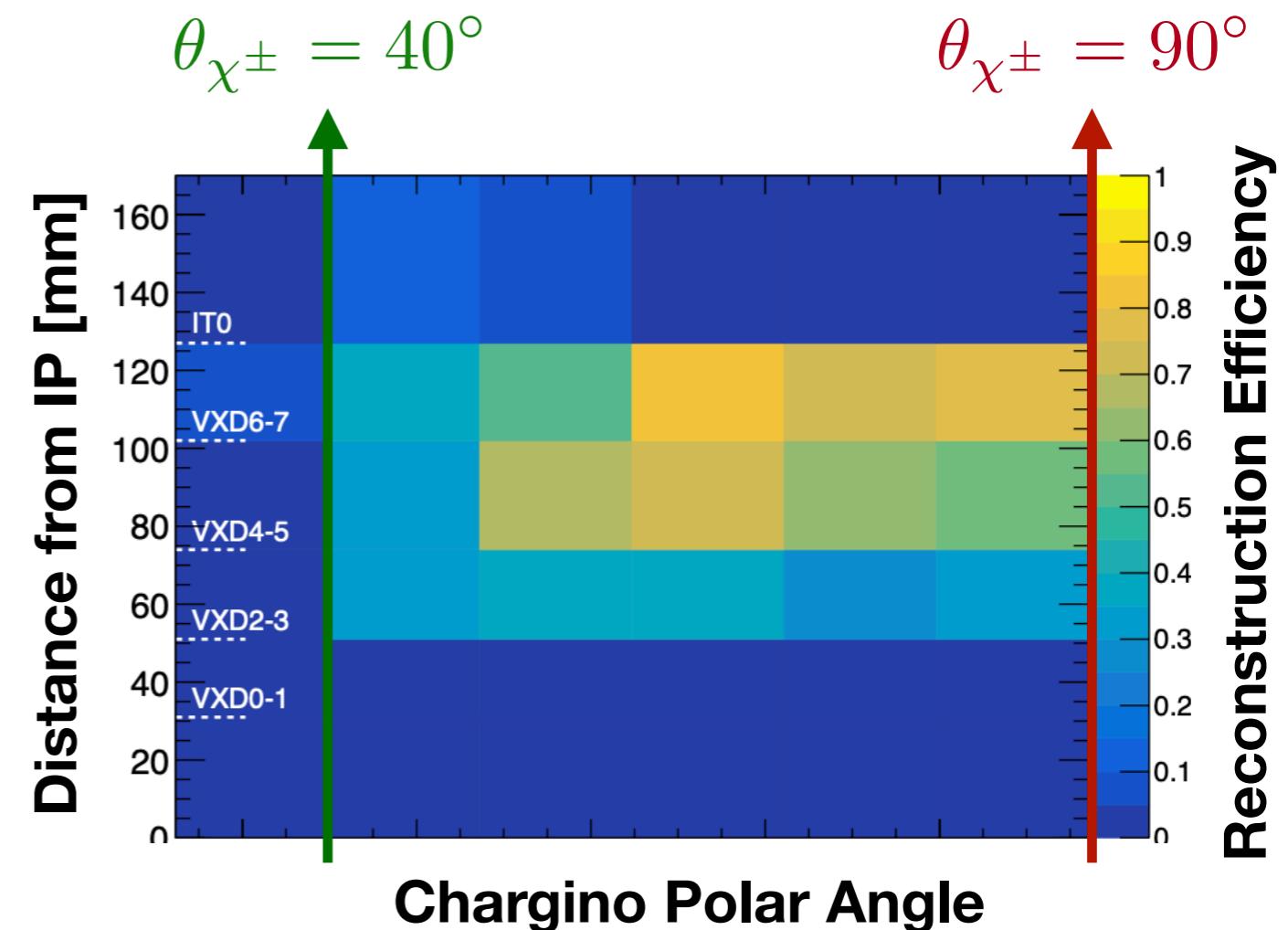
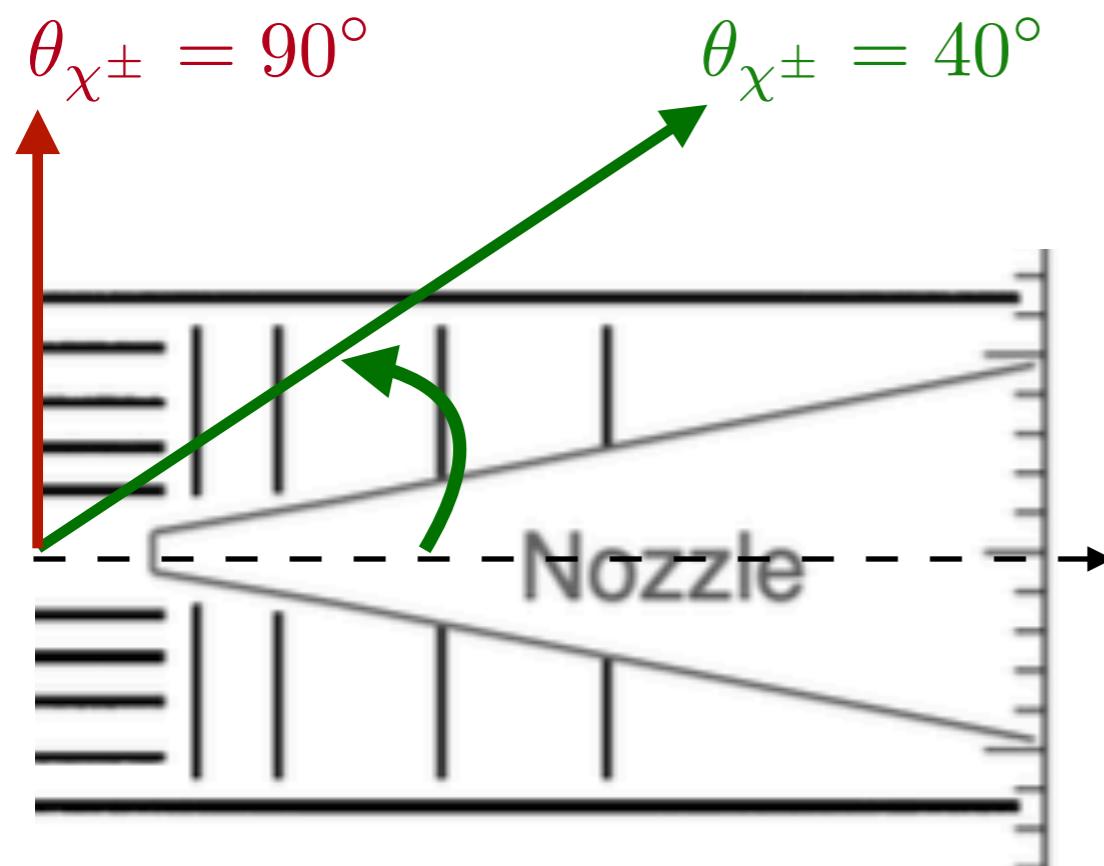
1. Soft
2. Arrives late
3. Mostly forward



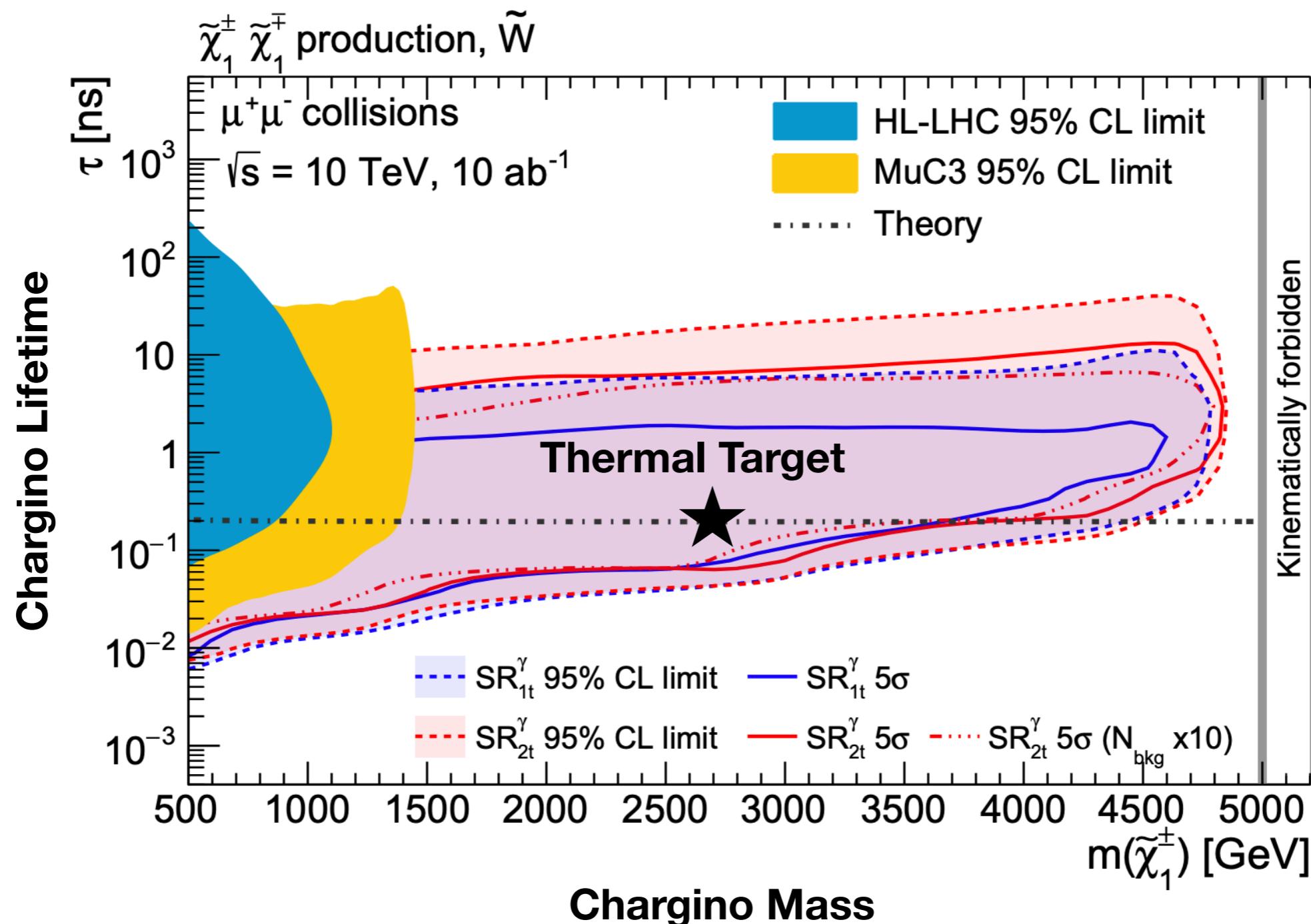
Requirement / Region	SR_{1t}^γ	SR_{2t}^γ
Veto	leptons and jets	
Leading tracklet p_T [GeV]	> 300	> 20
Leading tracklet θ [rad]	$[2/9\pi, 7/9\pi]$	
Subleading tracklet p_T [GeV]	-	> 10
Tracklet pair Δz [mm]	-	< 0.1
Photon energy [GeV]	> 25	> 25

- **Efficiency:**

Good signal efficiency ~30-80%

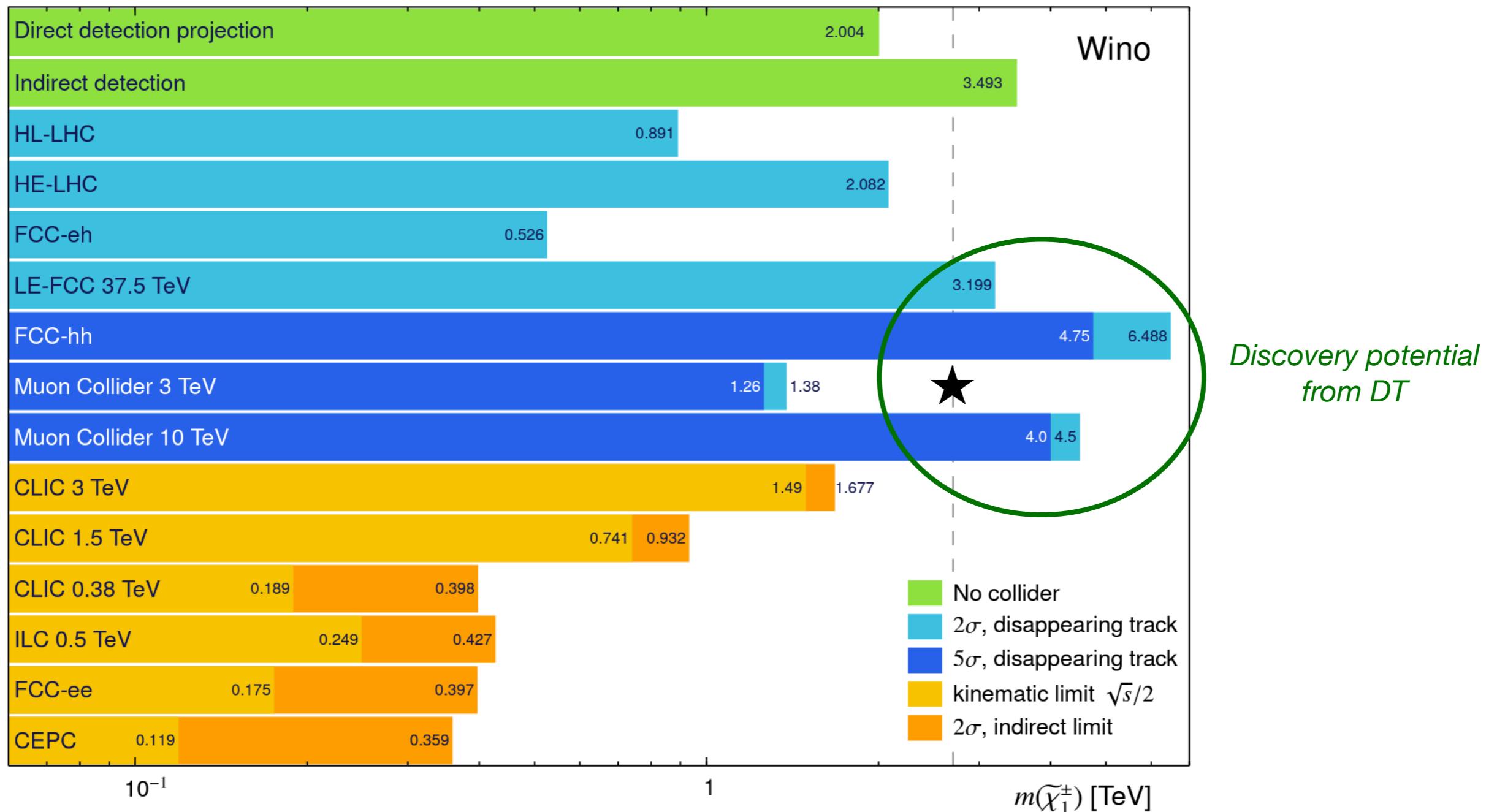


- DT Reach: Wino

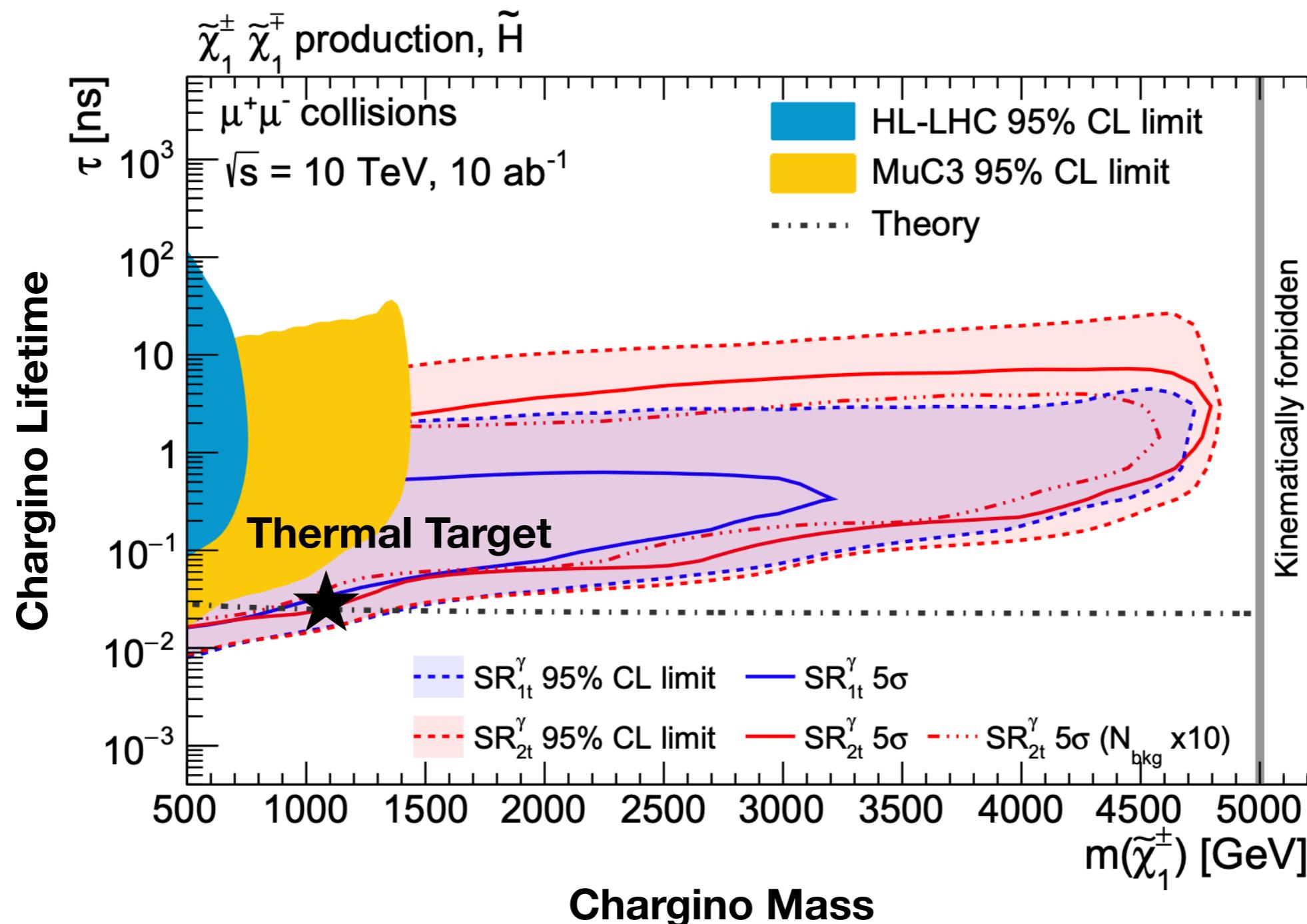


- DT Reach: Wino

RC, Meloni, Simoniello, Zurita, JHEP 06 (2021) 133

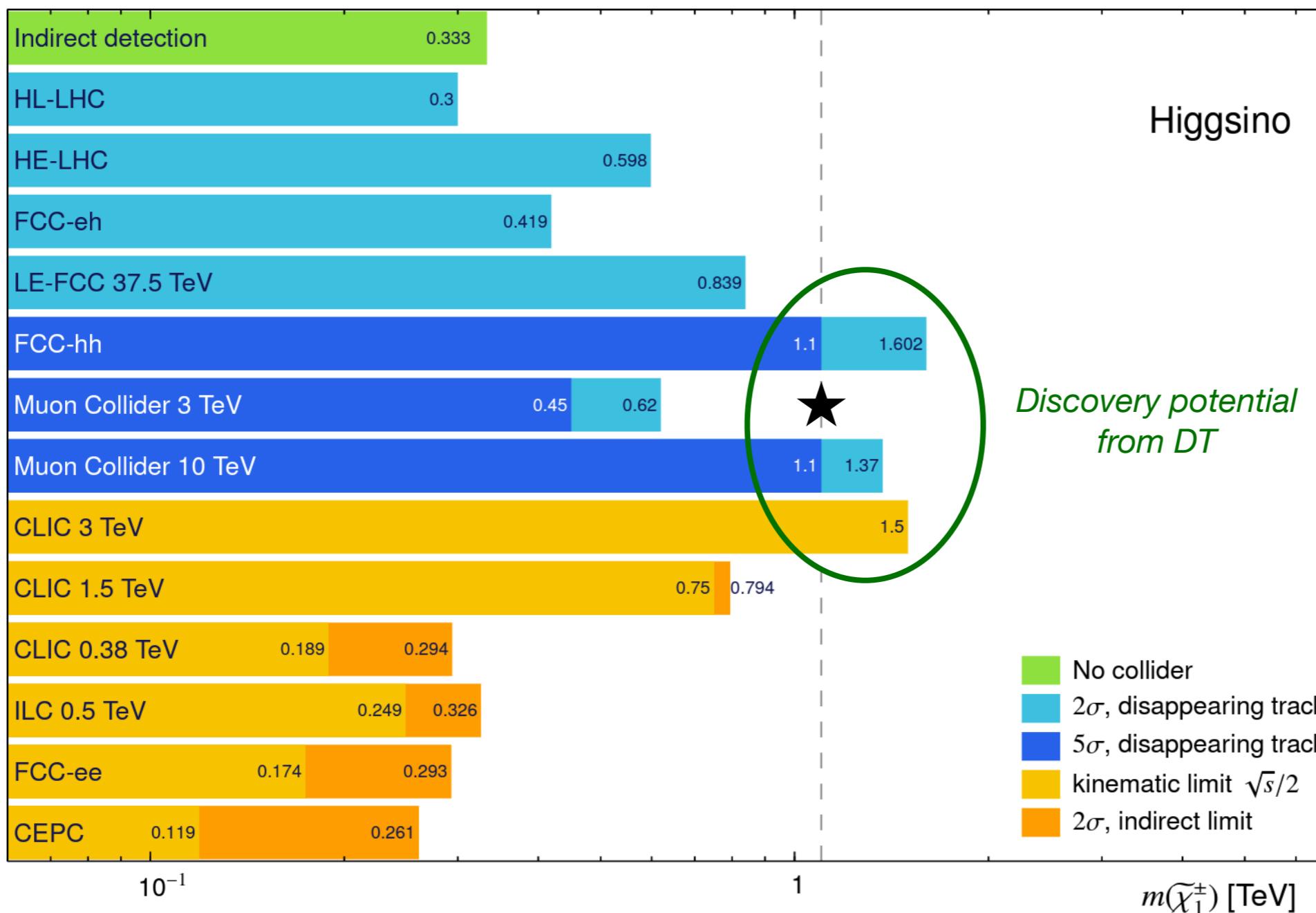


- DT Reach: Higgsino



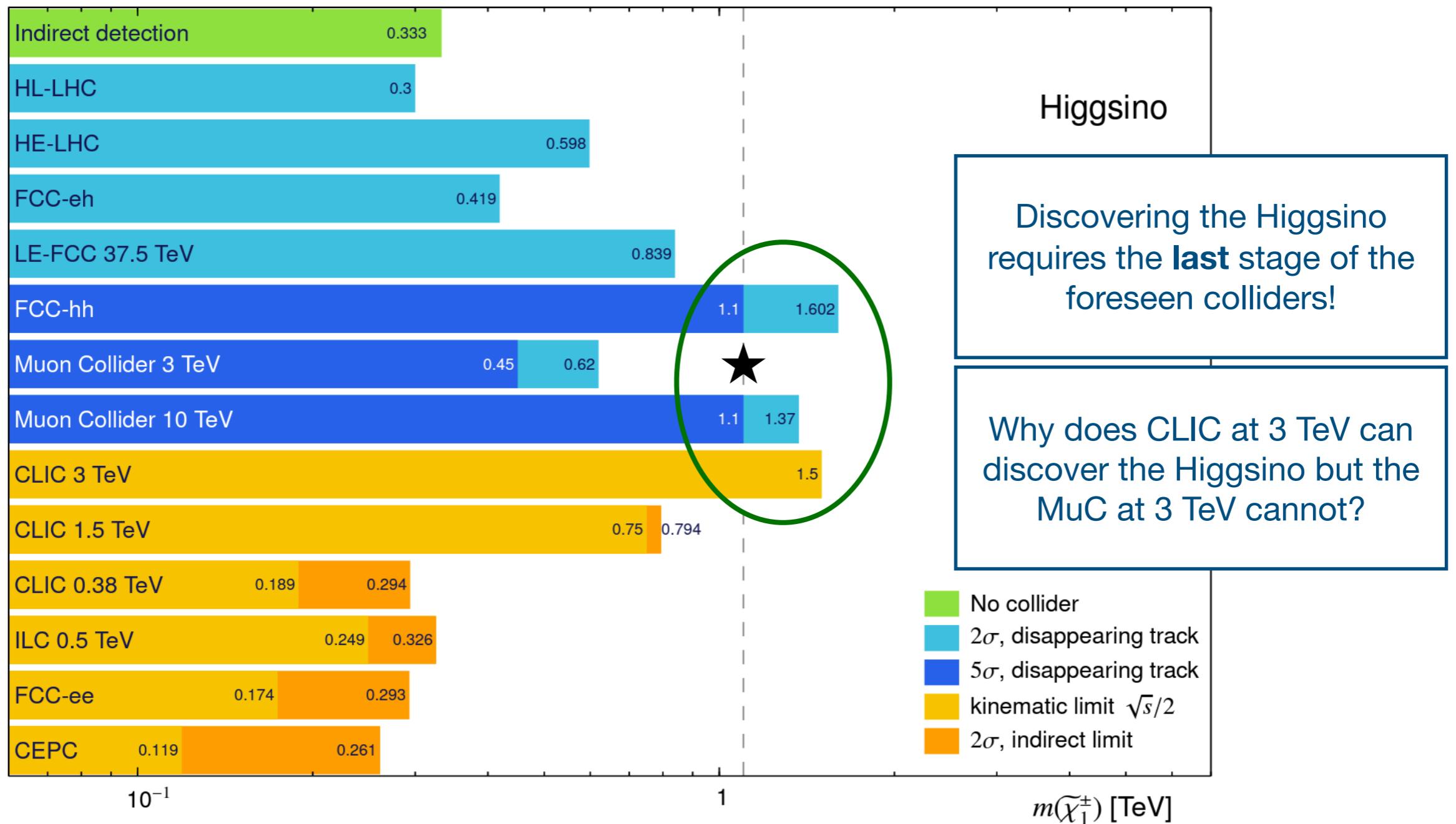
- DT Reach: Higgsino

RC, Meloni, Simoniello, Zurita, JHEP 06 (2021) 133



- DT Reach: Higgsino (can we do better?)

RC, Meloni, Simoniello, Zurita, JHEP 06 (2021) 133



- DT Reach: Higgsino (can we do better?)

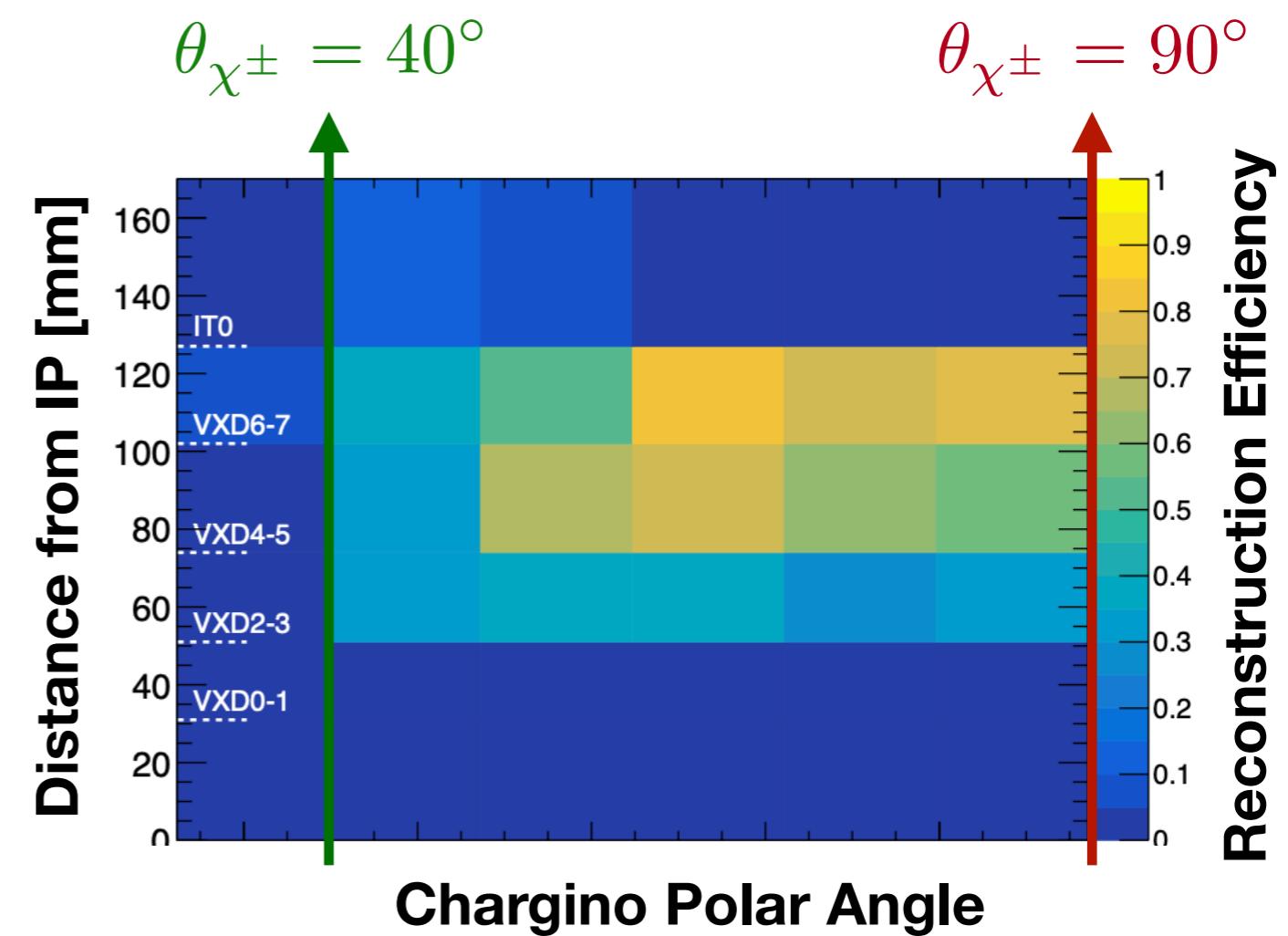
Why does CLIC at 3 TeV can discover the Higgsino but the MuC at 3 TeV cannot?

*Doubles
(Higgsino-like)*

$$\chi_2 \equiv \chi_{\tilde{H}} = \begin{pmatrix} \chi_{\tilde{H}}^+ \\ \chi_{\tilde{H}}^0 \end{pmatrix}$$

Mass = 1.1 TeV

Lifetime = 0.02 ns



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- Disappearing Tracks
- **Soft Tracks**

Summary

- A New Search Strategy: Soft Tracks!

FERMILAB-PUB-23-0832-T

DESY-24-069

APS/123-QED

Discovering Electroweak Interacting Dark Matter at Muon Colliders using Soft Tracks

Rodolfo Capdevilla,¹ Federico Meloni,² and Jose Zurita³

¹*Particle Theory Department, Fermi National Accelerator Laboratory , Batavia, IL 60510, USA**

²*Deutsches Elektronen-Synchrotron DESY, Notkestr. 85, 22607 Hamburg, Germany[†]*

³*Instituto de Física Corpuscular, CSIC-Universitat de València, Valencia, Spain[‡]*

(Dated: May 16, 2024)

Focus on:

1. Higgsino Thermal Target
2. Muon Collider 3 TeV

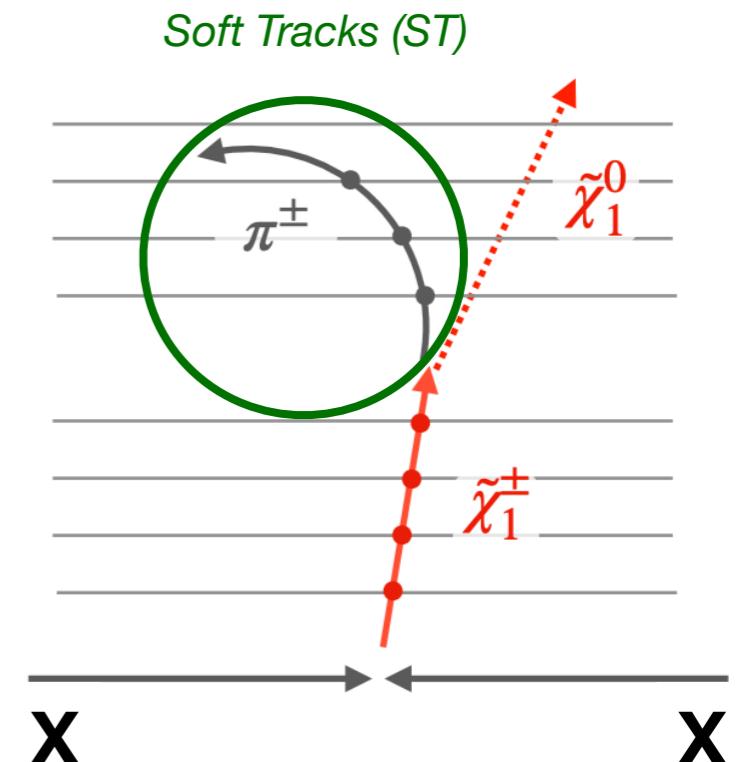


Federico
Meloni, DESY



José Zurita,
U. Valencia

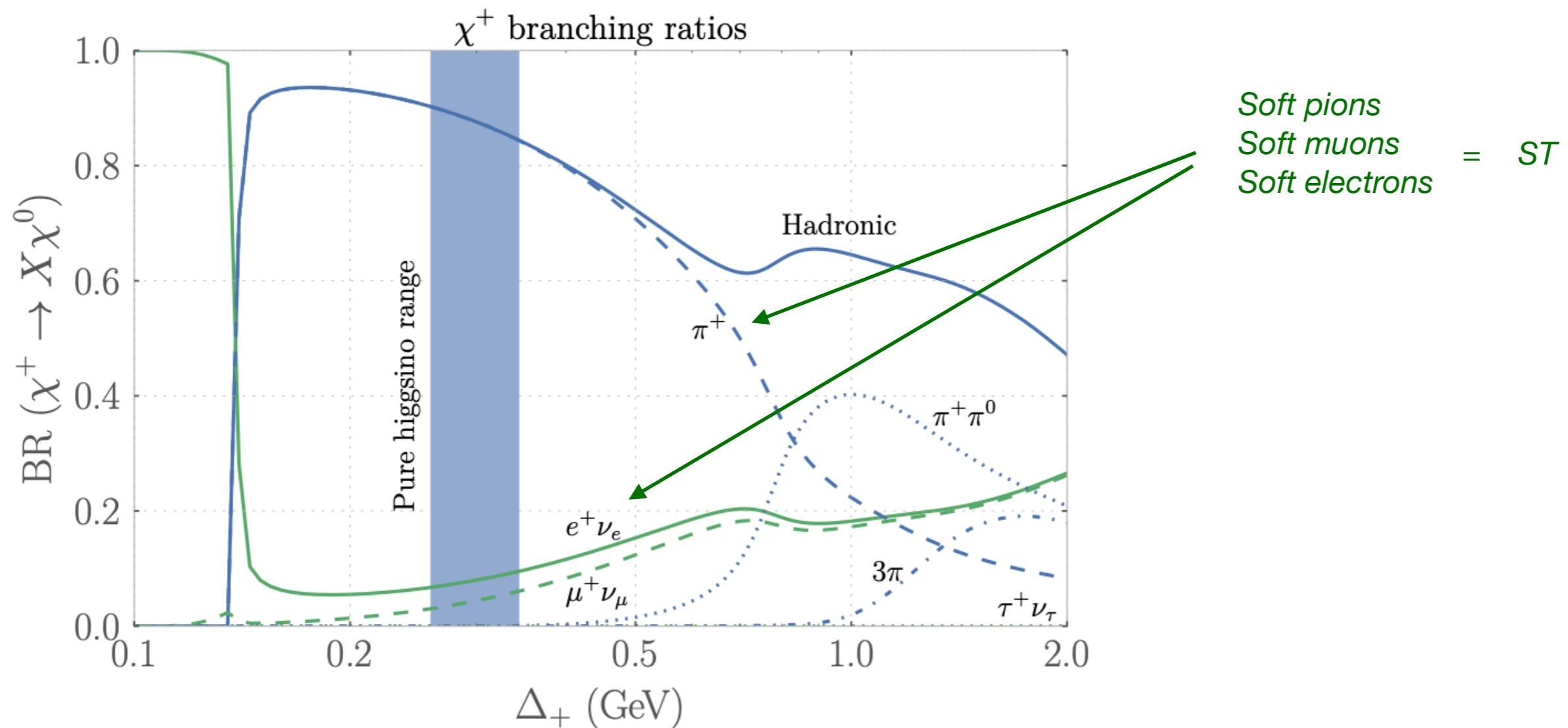
RC, Federico Meloni, José Zurita, arXiv:2405.08858



- Focus on Higgsinos:

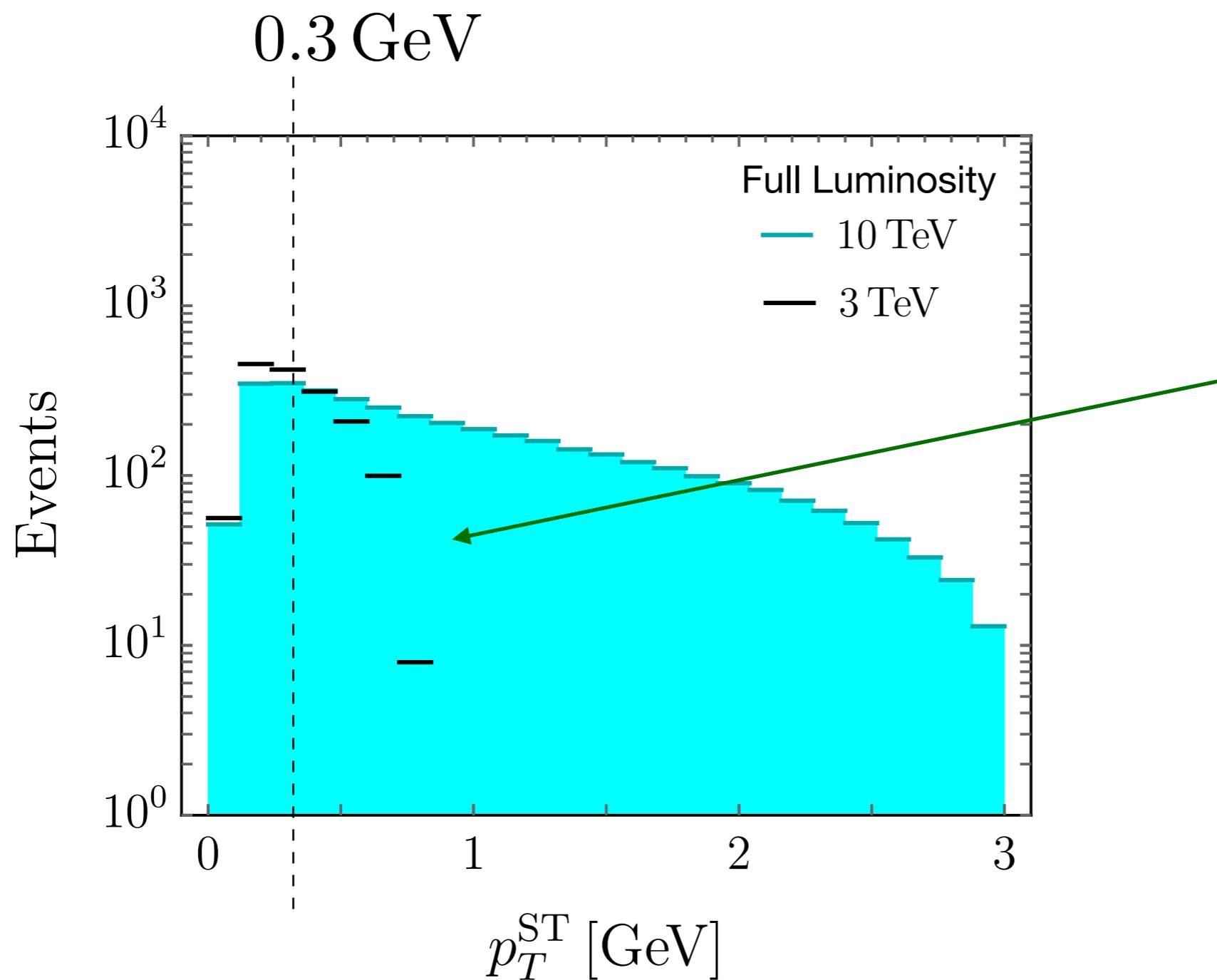
$$m_\chi = 1.1 \text{ TeV}$$

$$\Delta m \sim 0.3 \text{ GeV}$$



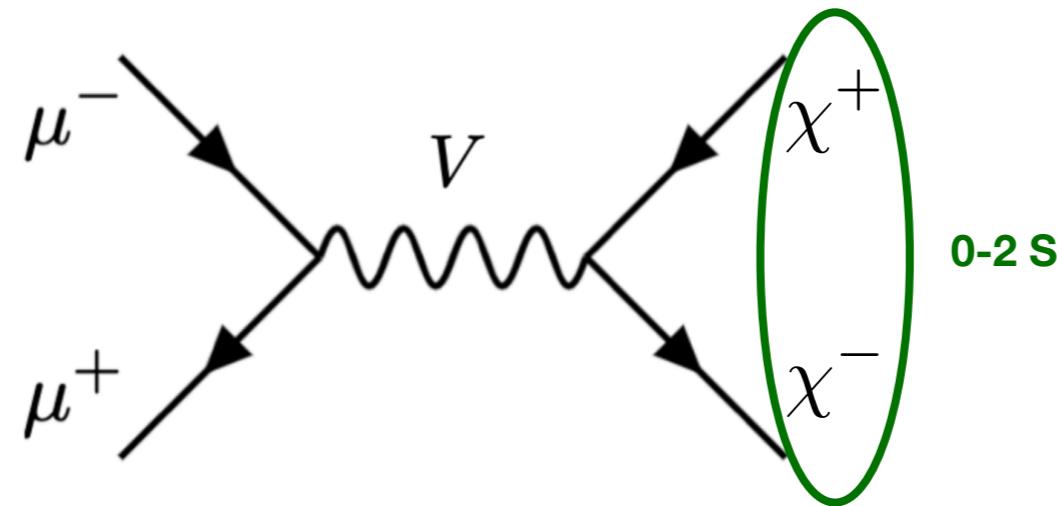
R. Mahbubani, P. Schwaller, J. Zurita, JHEP 06 (2017) 119

- Focus on the 3 TeV Collider: $\mathcal{L} = 1 \text{ ab}^{-1}$

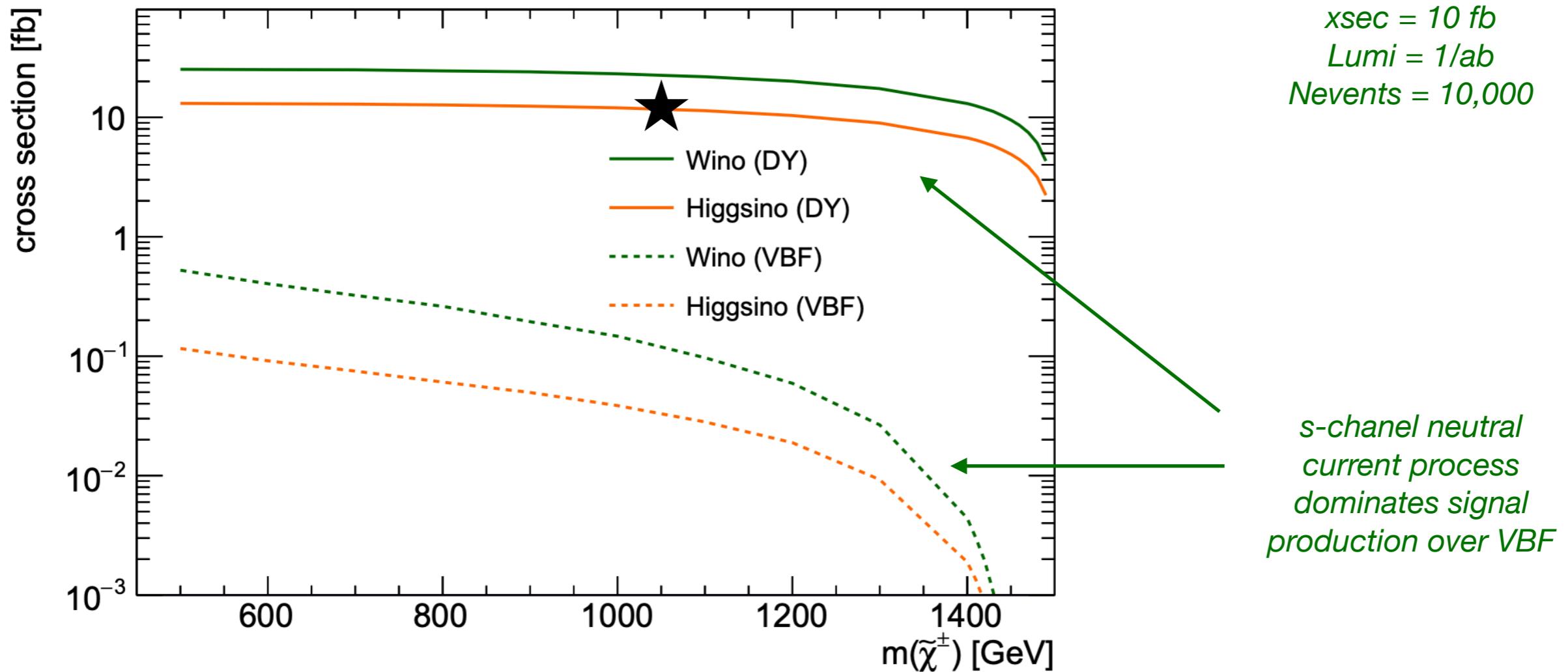


Very soft spectrum:
 $p_T^{\text{ST}} \leq 0.75 \text{ GeV}$

- Signal Events:



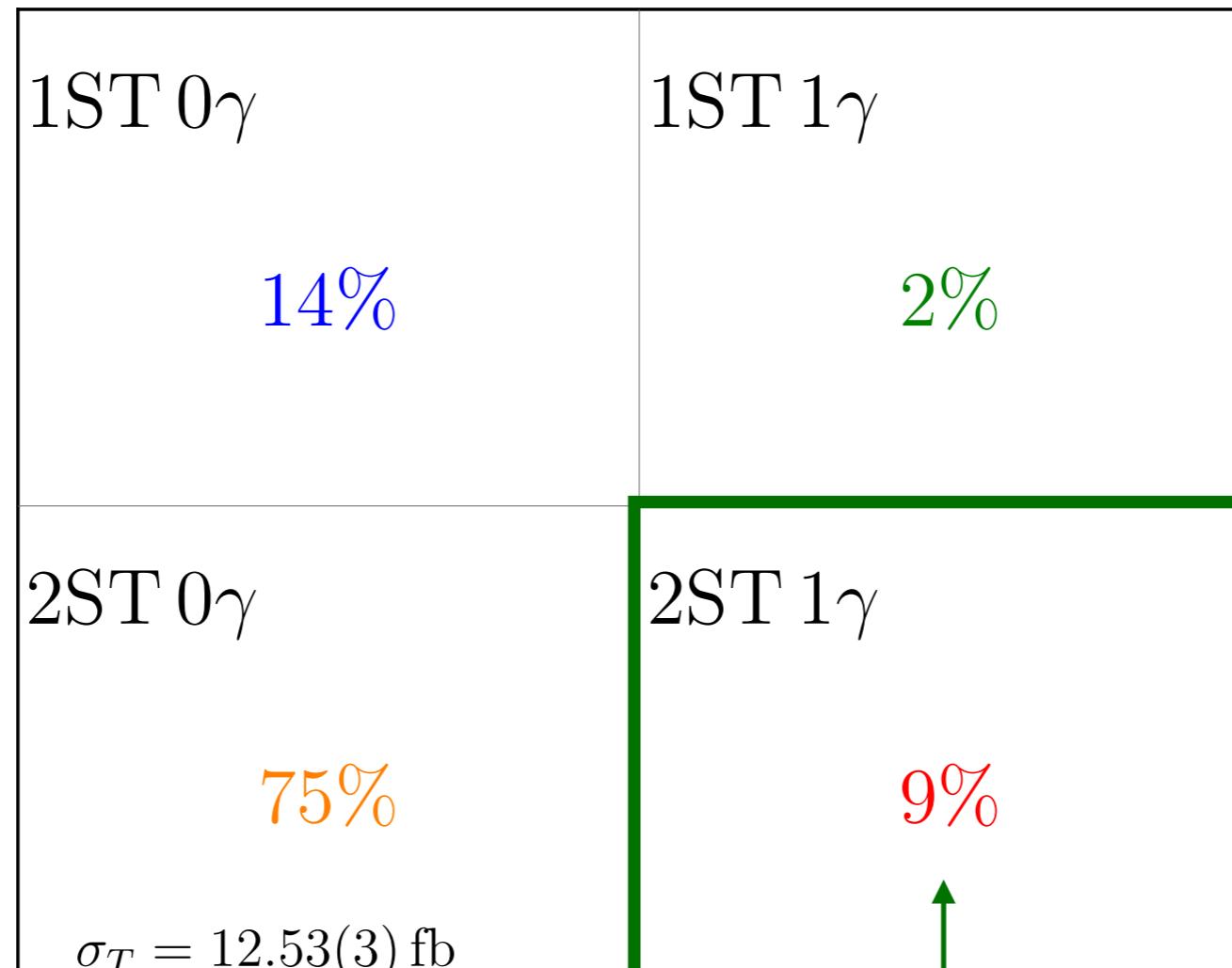
MuC 3 TeV



RC, Meloni, Simoniello, Zurita, JHEP 06 (2021) 133

- **Signal Events:**

MuC 3 TeV

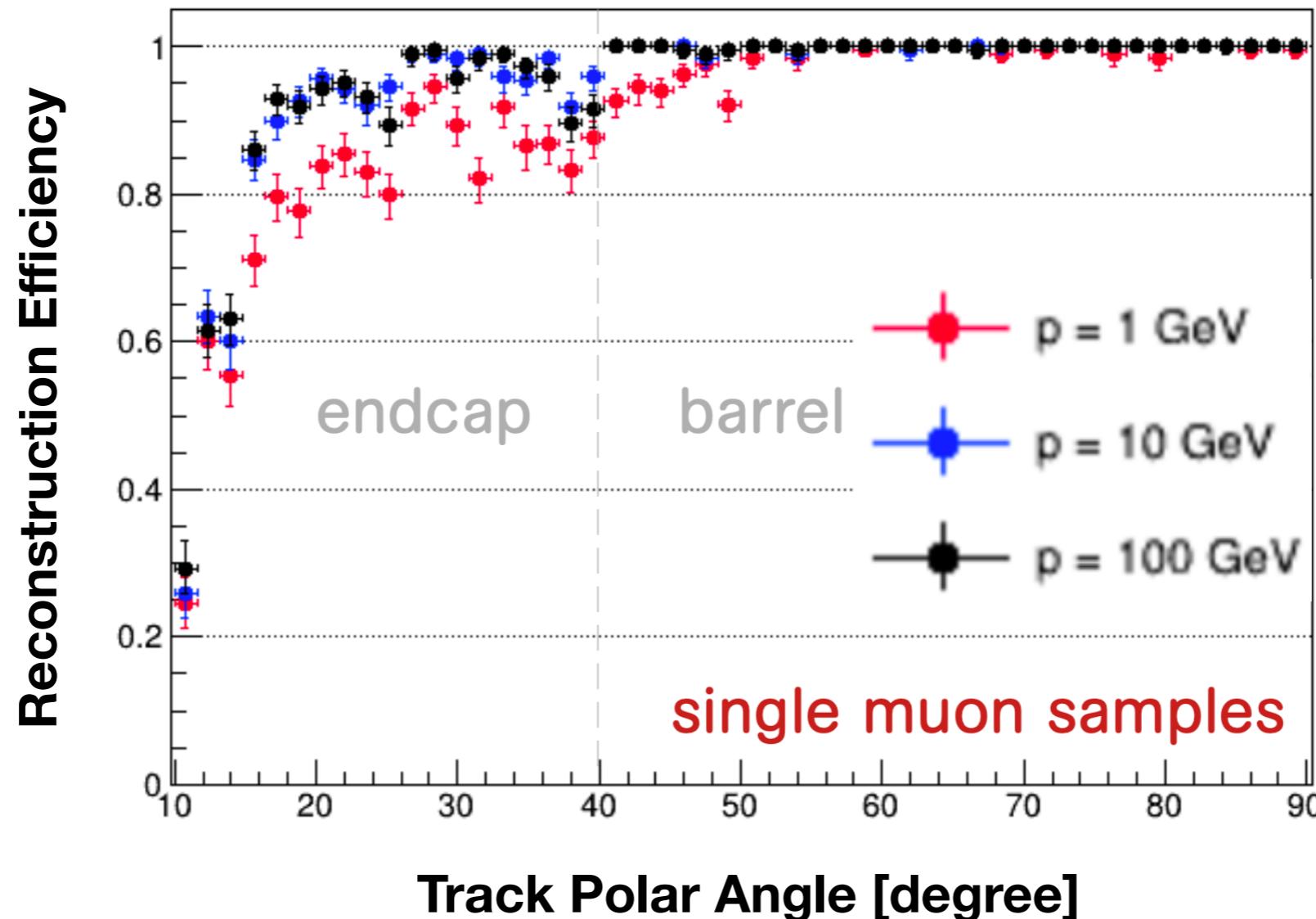


About **1,000** signal events in
this signal region

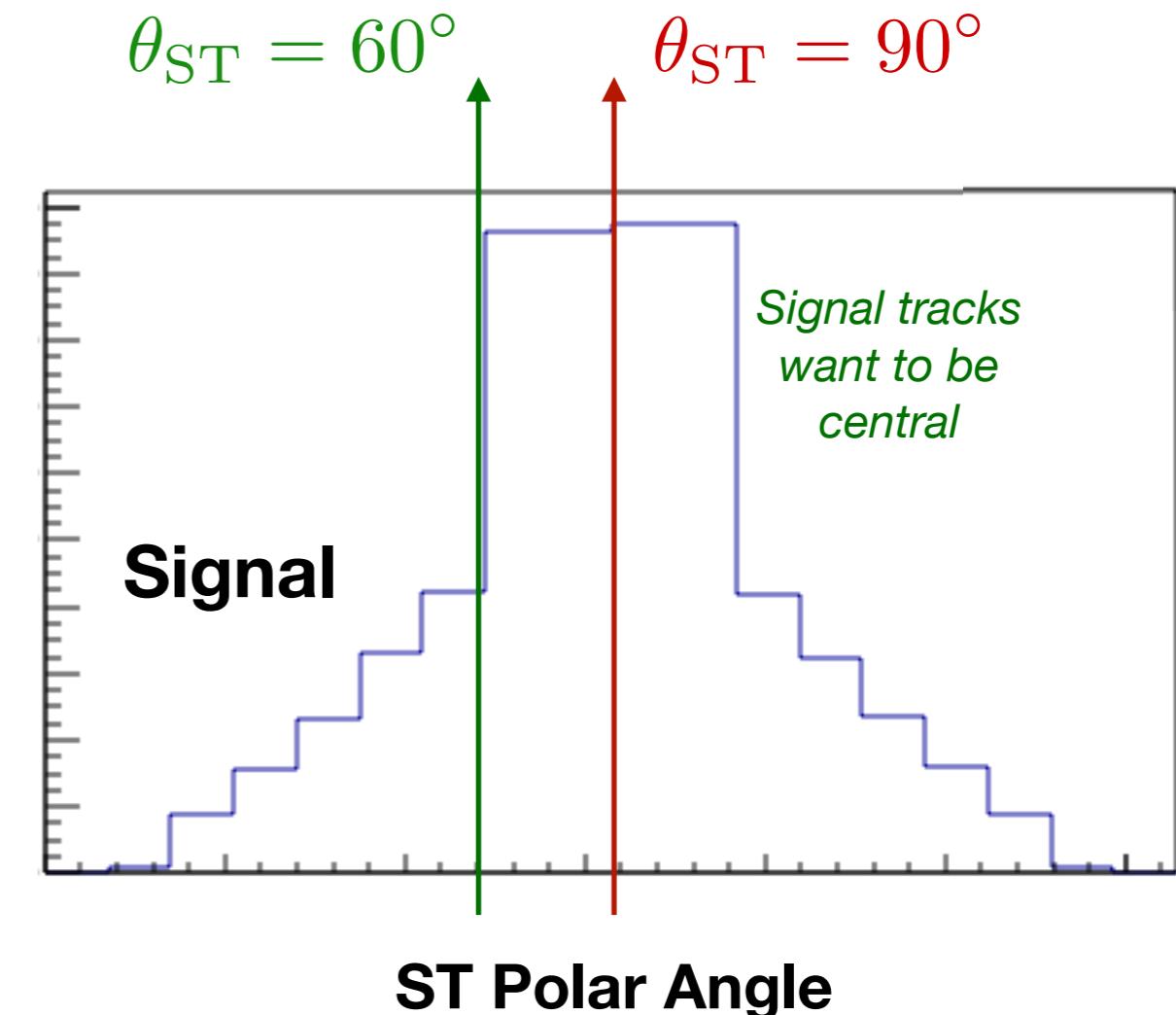
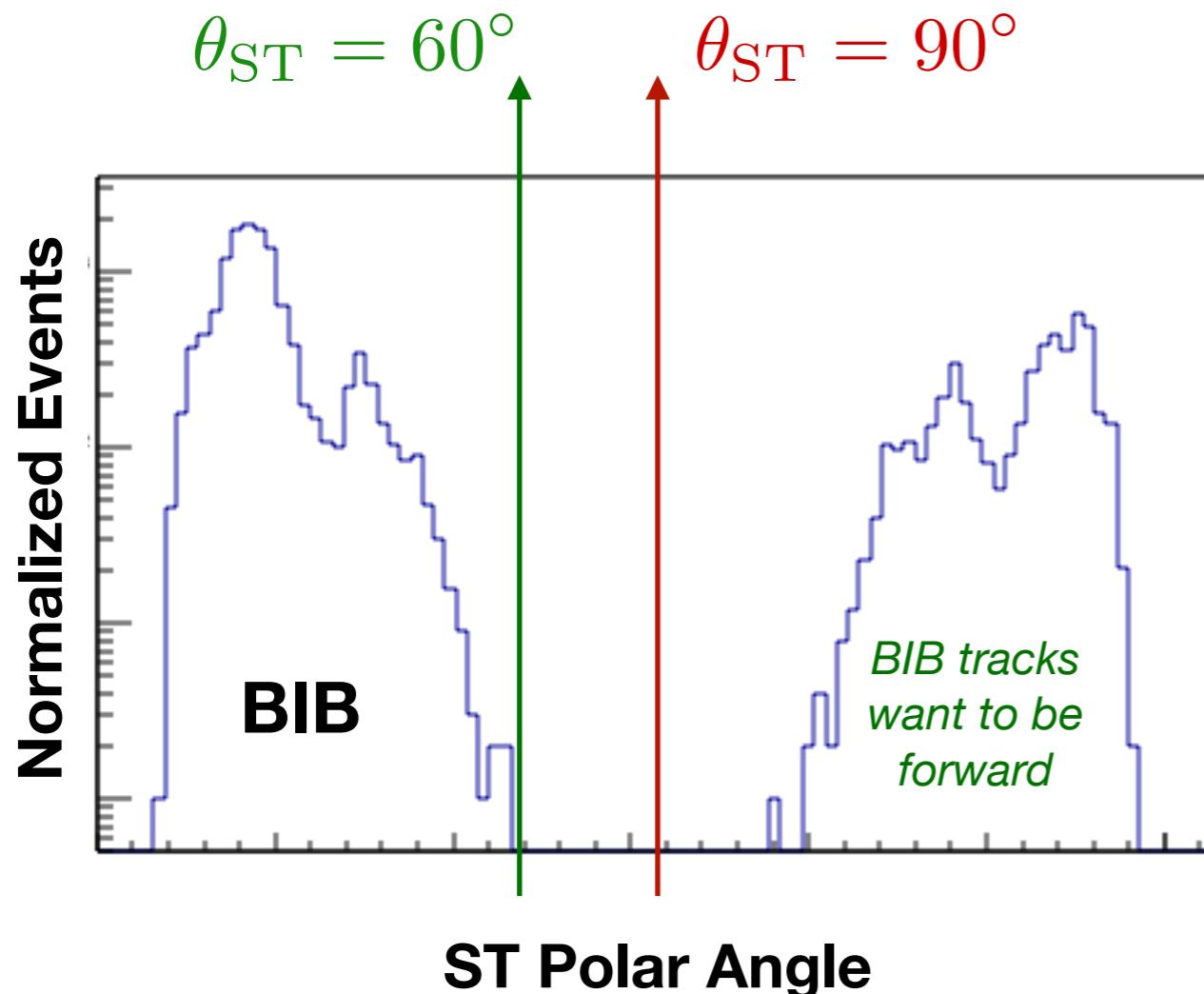
- **Backgrounds: BIB (efficiency)**

BIB Data from the
MAP collaboration

C. Accettura et al., Eur. Phys. J. C 83 (2023) 9, 864

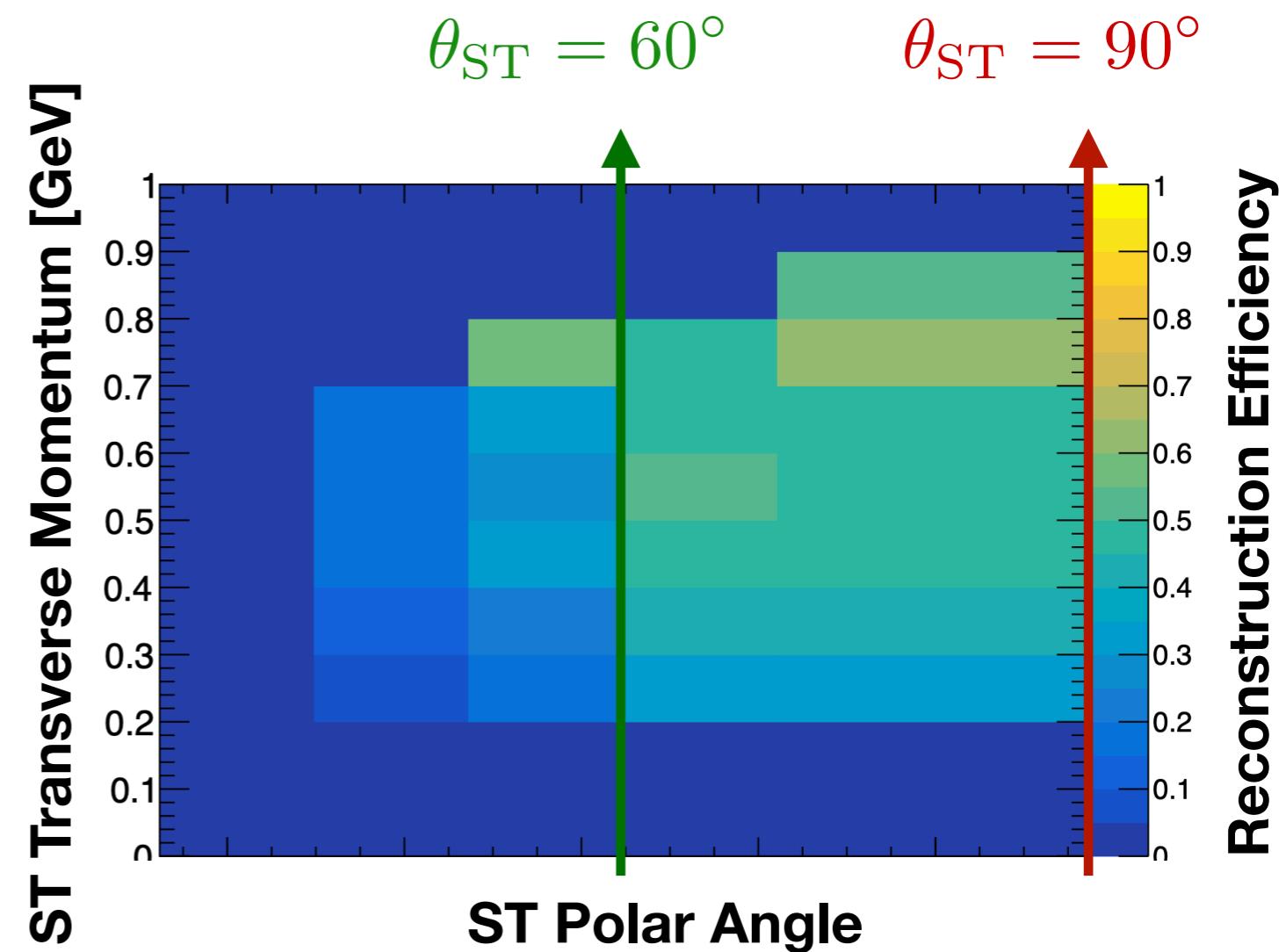
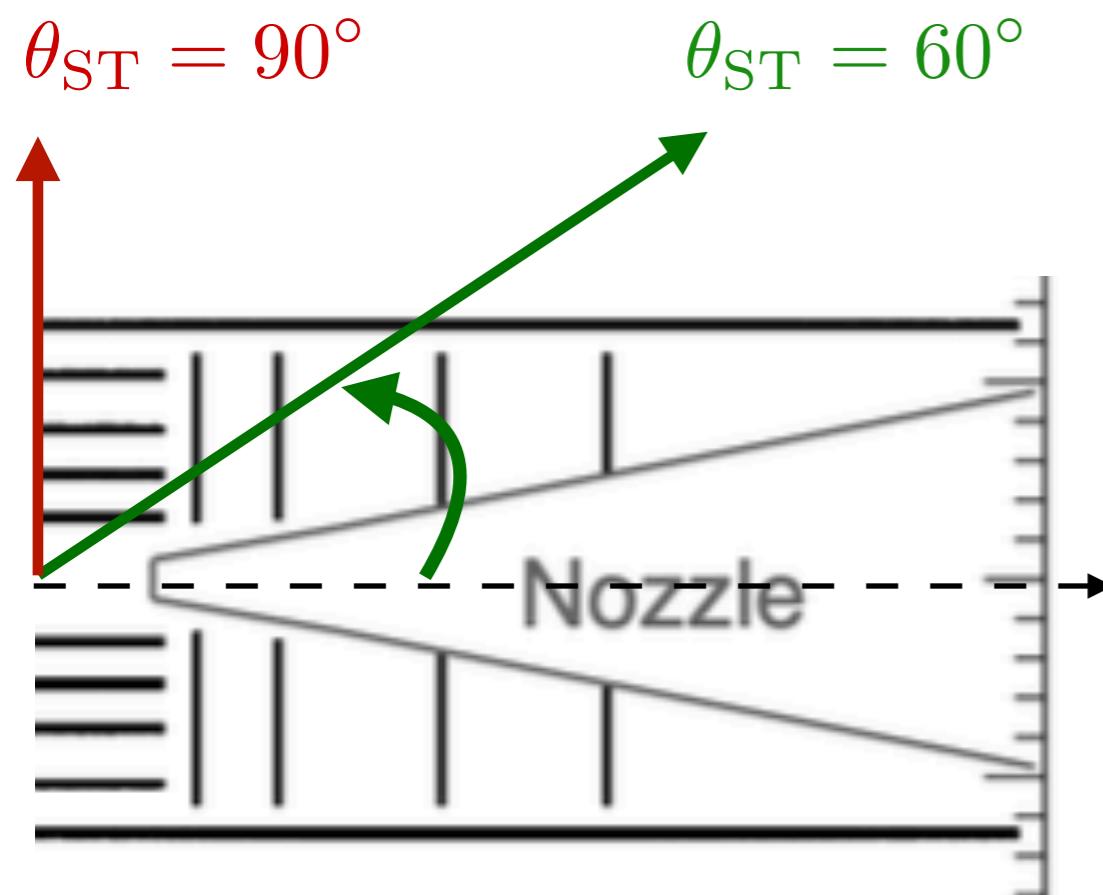


- **Backgrounds: BIB (fake tracks)**



- **Efficiency:**

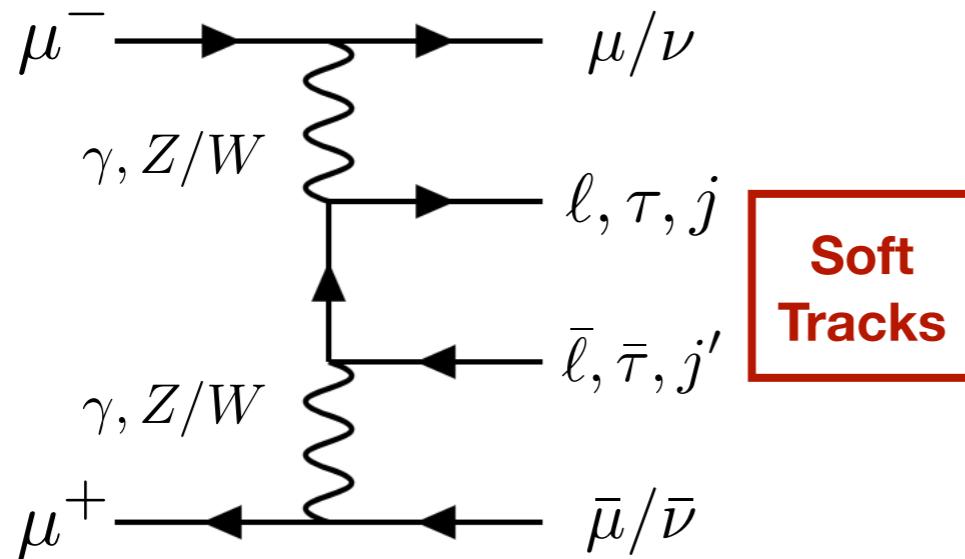
Decent reconstruction efficiency ~30-60%



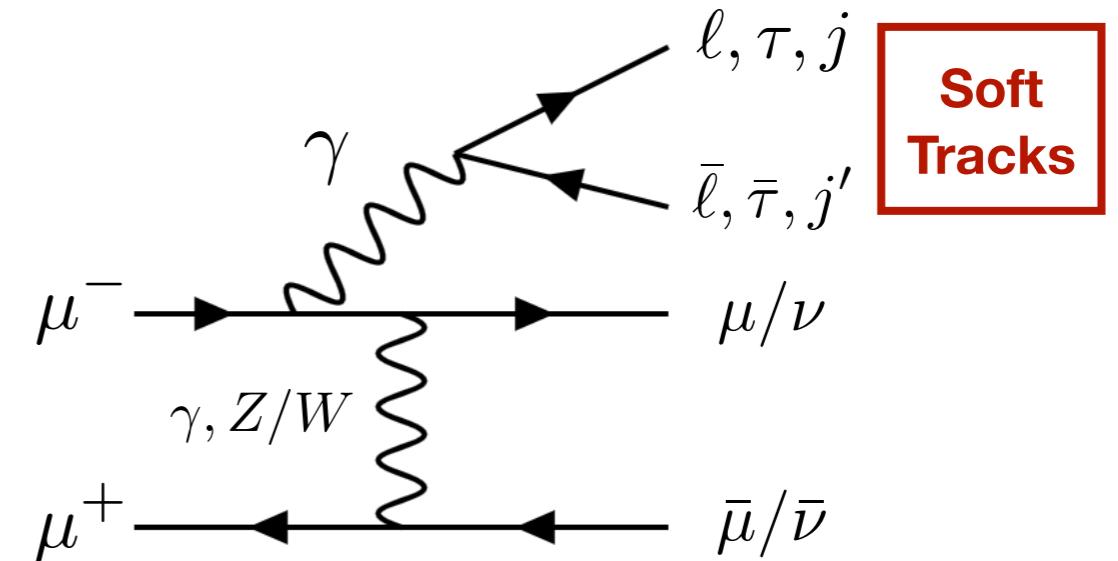
- **Background Processes:**

*VBF & Bhabha-like processes
dominate background production*

**Vector Boson
Fusion**

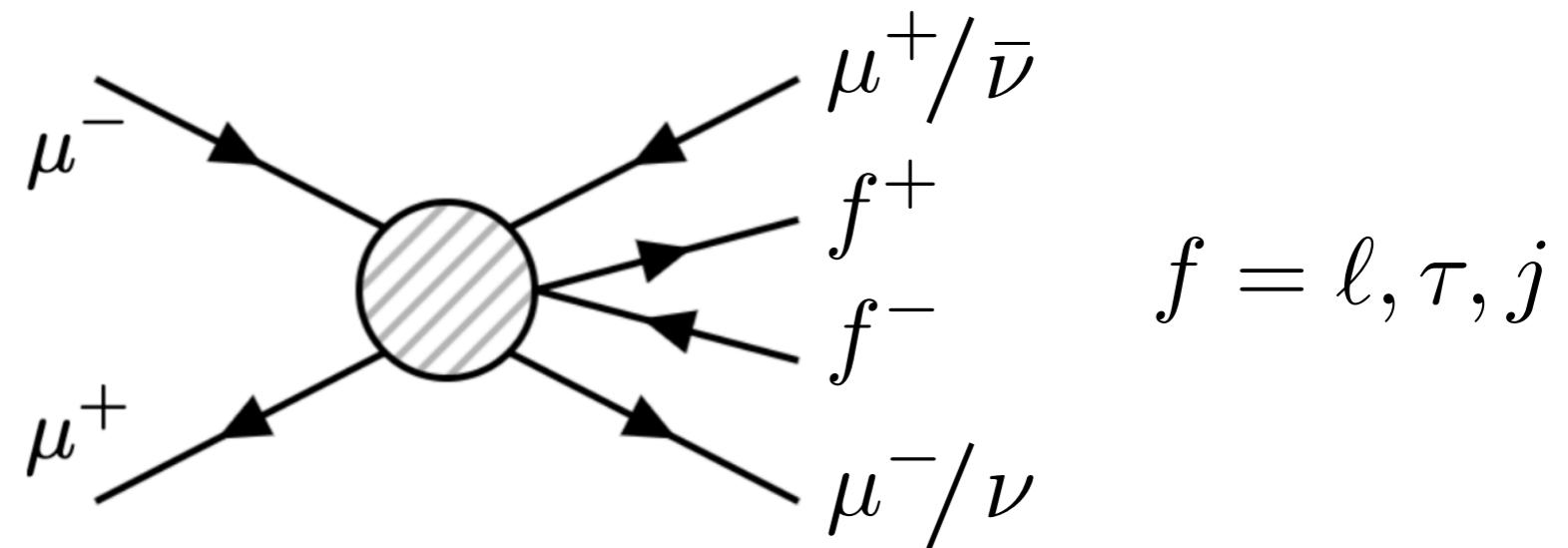


**Bhabha-Like
Scattering**



- Background Processes:

VBF & Bhabha-like processes dominate background production



$\mu^+ \mu^- \rightarrow \gamma + X (+ Z \rightarrow \nu \nu)$		
X	$\sigma(\gamma X)$ [fb]	$\sigma(\gamma X Z)$ [fb]
$\ell^+ \ell^- \nu_\ell \bar{\nu}_\ell$	242.0	2.828
$\ell^+ \ell^- \mu^+ \mu^-$	60.45	0.012
$e^+ \nu_e \mu^- \bar{\nu}_\mu + \text{CP}$	226.6	2.710
$\tau^+ \tau^- \nu_\ell \bar{\nu}_\ell$	6.493	0.058
$\tau^+ \tau^- \mu^+ \mu^-$	30.86	0.006
$\tau^+ \nu_\tau \mu^- \bar{\nu}_\mu + \text{CP}$	226.2	2.722
$jj \nu_\ell \bar{\nu}_\ell$	104.5	0.904
$jj \mu^+ \mu^-$	30.63	0.019
$jj \mu^- \bar{\nu}_\mu + \text{CP}$	1215.	11.57

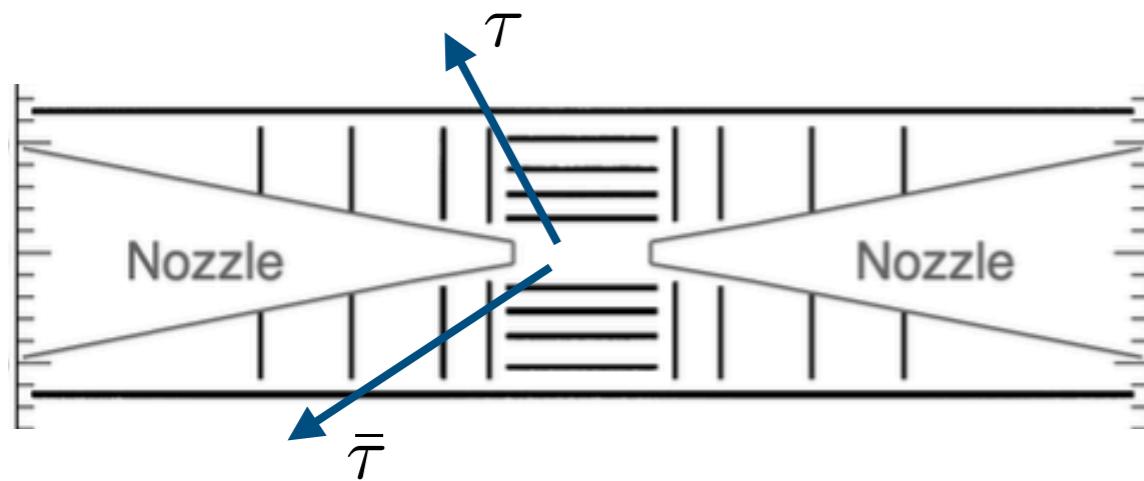
$p_T^\gamma \geq 20 \text{ GeV}$
 $|\eta_\gamma| < 2.44$

$p_T^\ell \geq 0.1 \text{ GeV}$
 $p_T^j \geq 0.1 \text{ GeV}$

Compare with the
~1fb signal

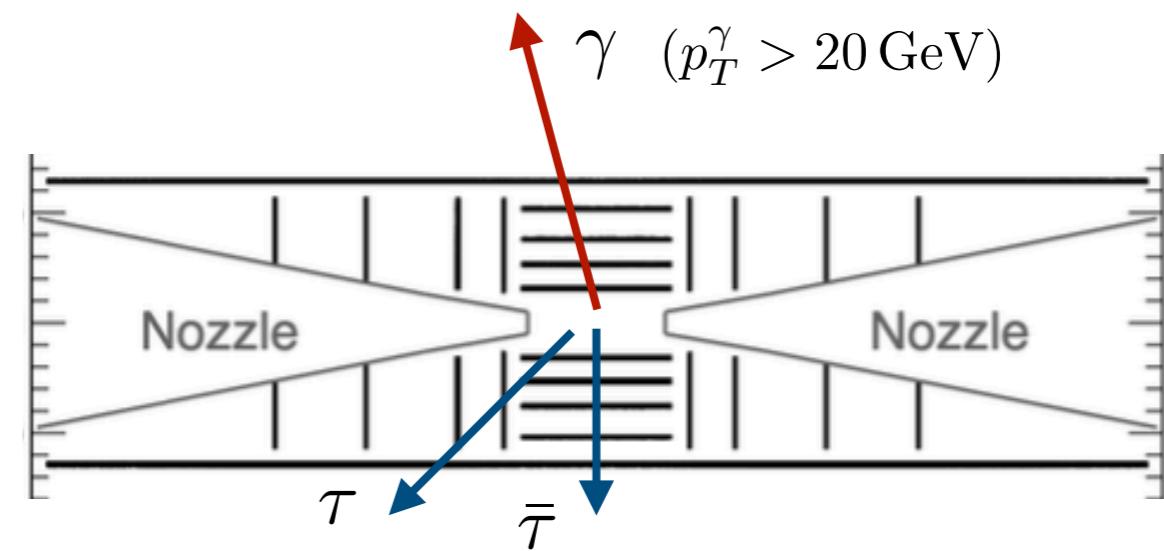
- The Importance of the Photon:**

Signal region:
 $2 \text{ ST} + 1 \text{ gamma}$
 $0.2 < p_T < 0.75 \text{ GeV}$



$$N_{\text{signal}} \sim 9000$$

$$N_{\text{background}} \sim 10^3 \times N_{\text{signal}}$$

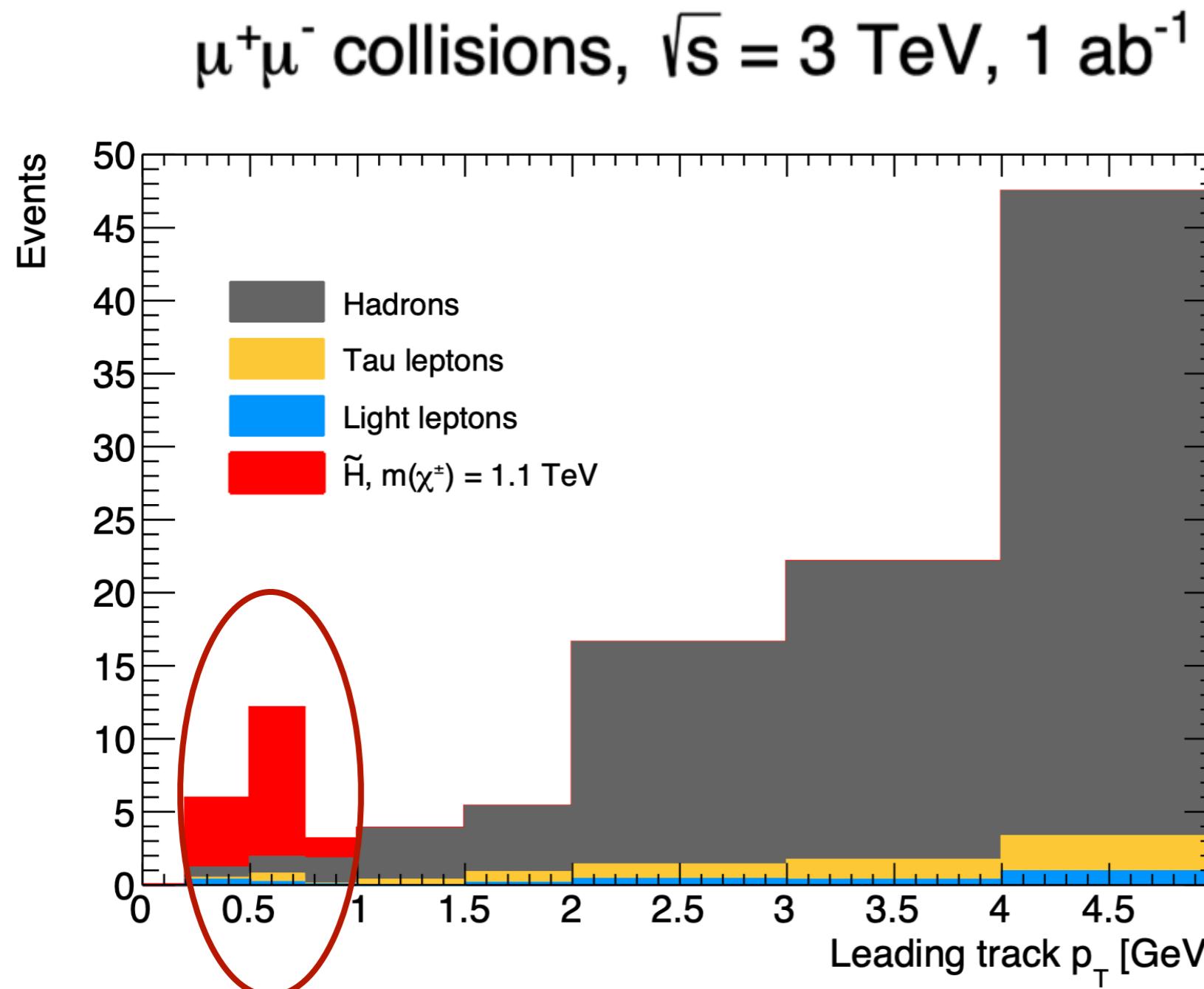


$$N_{\text{signal}} \sim 2000$$

$$N_{\text{background}} \sim 6000$$

Note: Considering di-tau production using the improved Weizsäcker-Williams approximation for the photon PDF.

- **Projection:**



Five sigma+ for the
Thermal Higgsino
(Doublet MDM)

Signal region:
2 ST + 1 gamma
 $0.2 < pT < 0.75 \text{ GeV}$

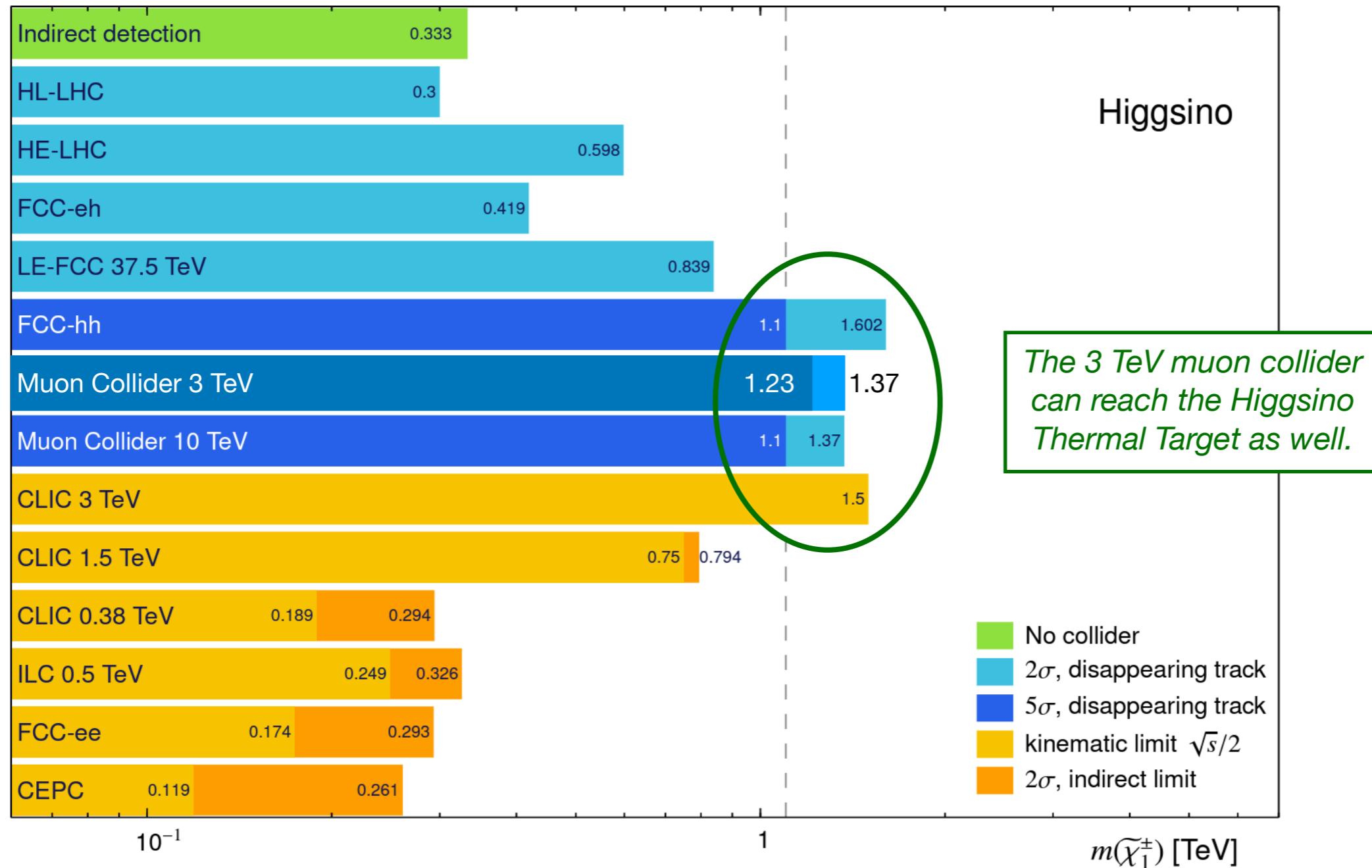
Photon pT above
 40 GeV

Fakes:

Pairs of ST along with
an uncorrelated photon

Random ECAL hits
from the BIB that can
mimic a photon

- **Projection:**



Outline

1. Introduction

2. Minimal WIMP Searches

- Direct Detection
- Indirect Detection
- LHC

3. Future p/e Colliders

4. Muon Collider

- Disappearing Tracks
- Soft Tracks

Summary

- Thermal Targets

$$\chi_2 \equiv \chi_{\tilde{H}} = \begin{pmatrix} \chi_{\tilde{H}}^+ \\ \chi_{\tilde{H}}^0 \end{pmatrix},$$

*Doublets
(Higgsino-like)*

$$\chi_3 \equiv \chi_{\tilde{W}} = \begin{pmatrix} \chi_{\tilde{W}}^+ \\ \chi_{\tilde{W}}^0 \\ \chi_{\tilde{W}}^- \end{pmatrix} \dots \quad \chi_5 \dots \quad \chi_9 \dots$$

*Triplets
(Wino-like)*

Fiveplets... Nineplets...

Mass: 1.1 TeV 2.7 TeV 14 TeV

Lifetime: 0.02 ns 0.2 ns 0.2 ns

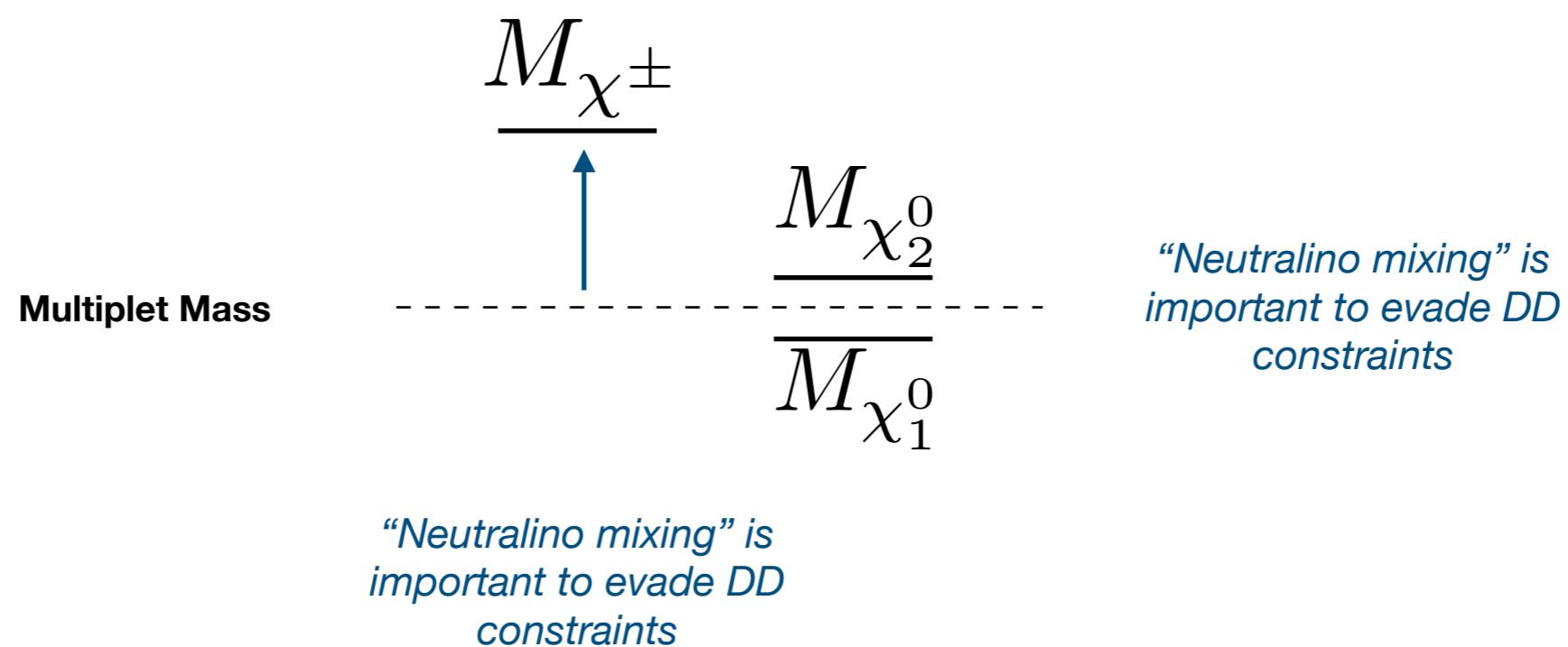
- Lifetime

For a given multiplet

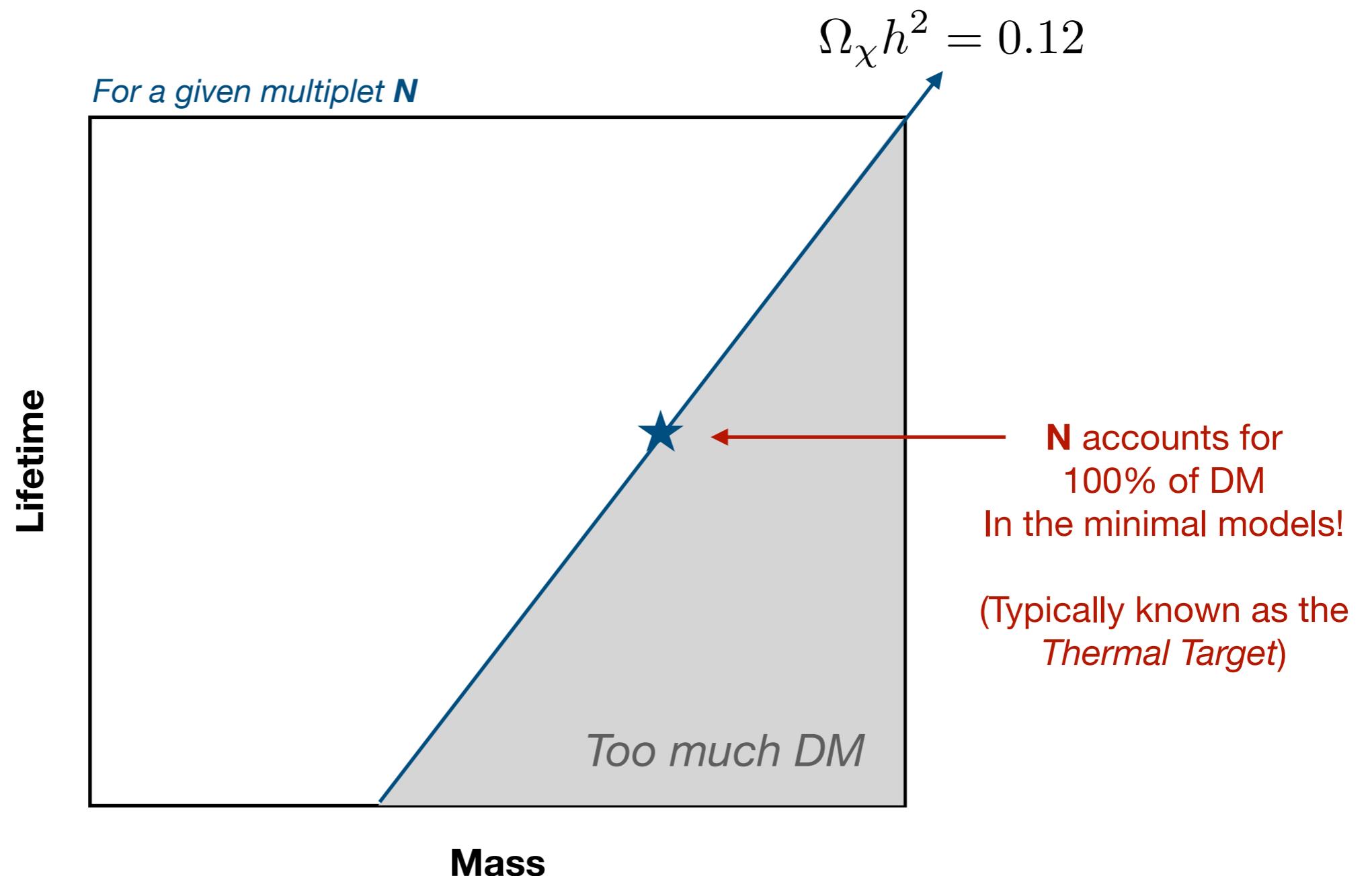
$$\chi_{\tilde{H}} = \begin{pmatrix} \chi_{\tilde{H}}^+ \\ \chi_0^0 \\ \chi_{\tilde{H}}^- \end{pmatrix}$$

*Loop effects generate
chargino-neutralino splitting*

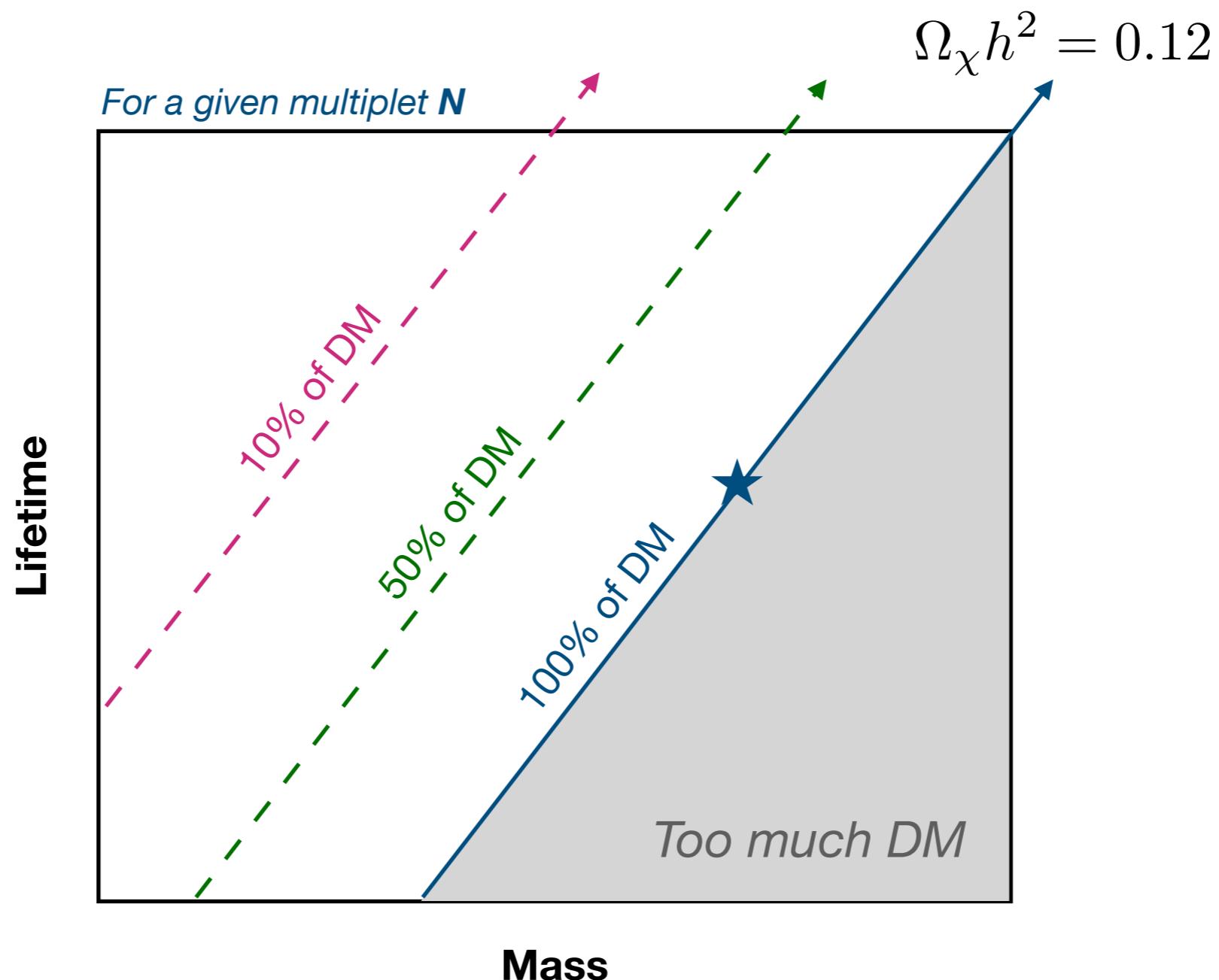
$$\Delta M = M_{\chi^\pm} - M_{\chi^0} > 0$$



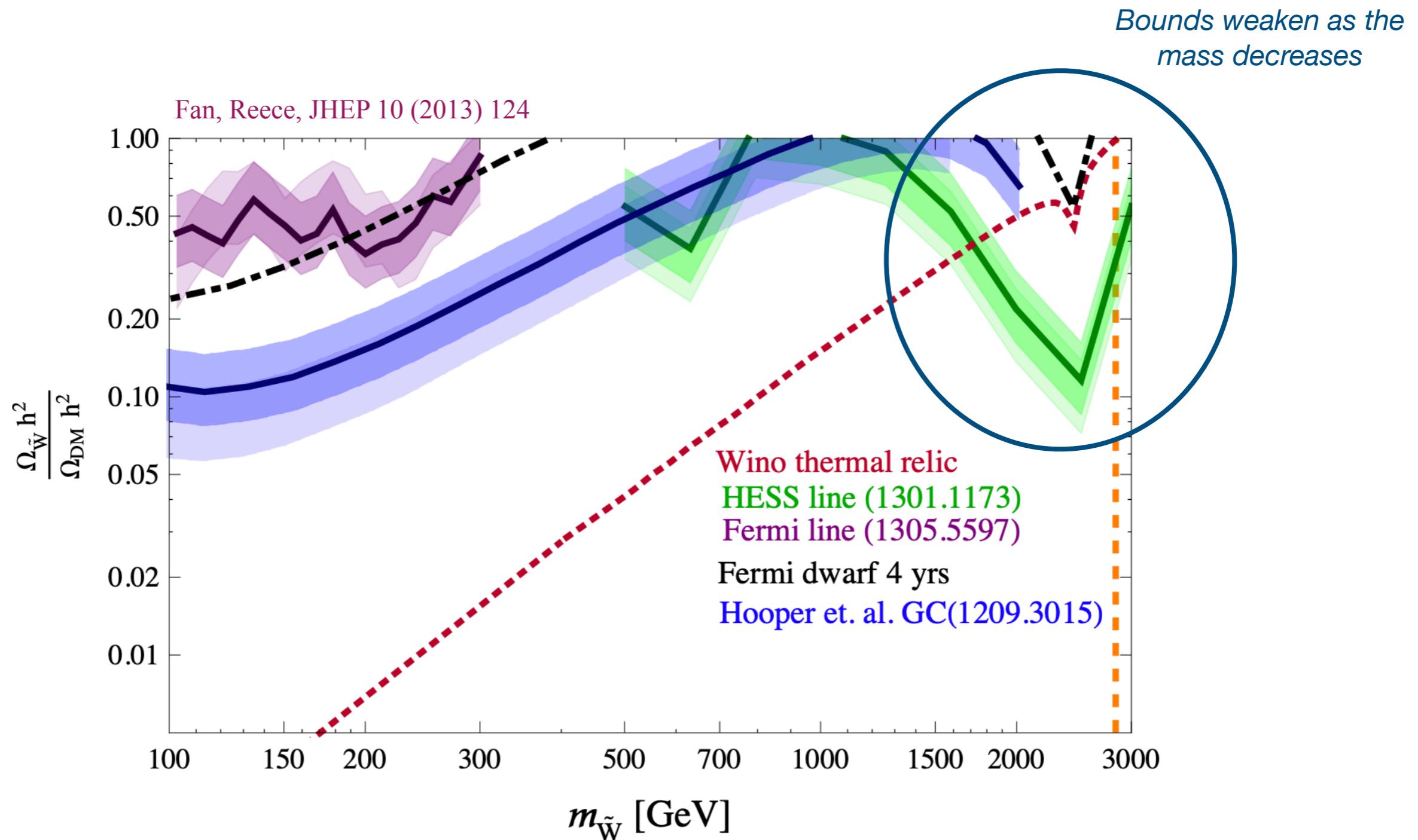
- Target Parameter Space:



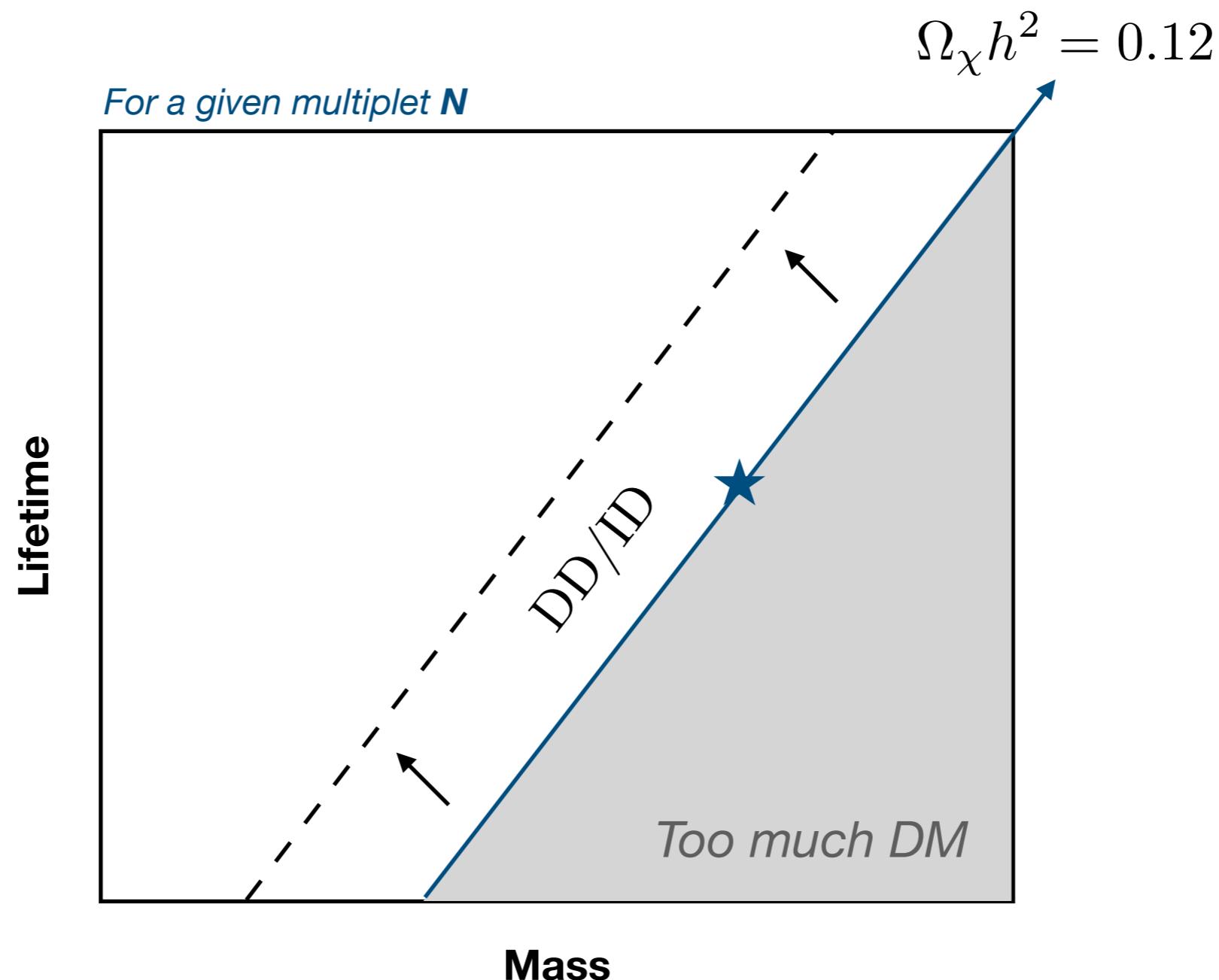
- Target Parameter Space:



- Indirect Detection (ID)



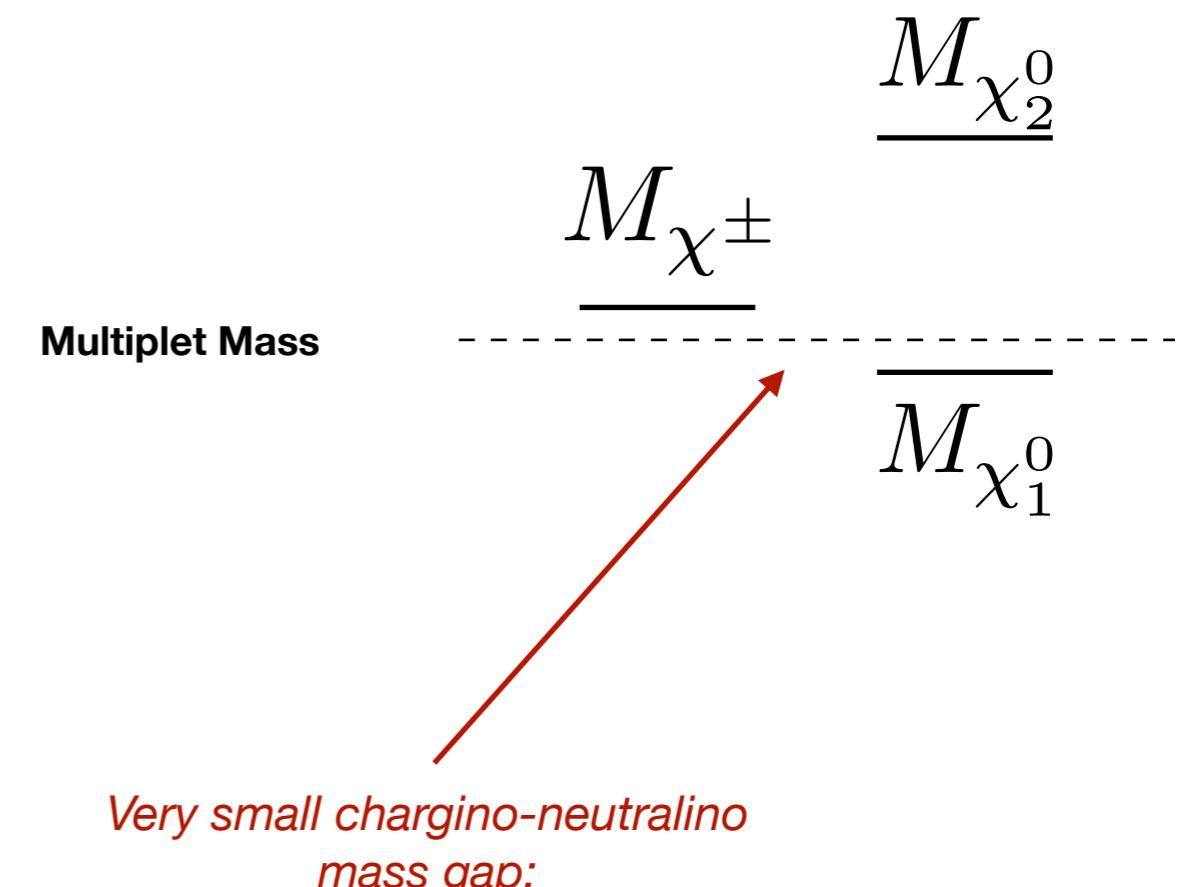
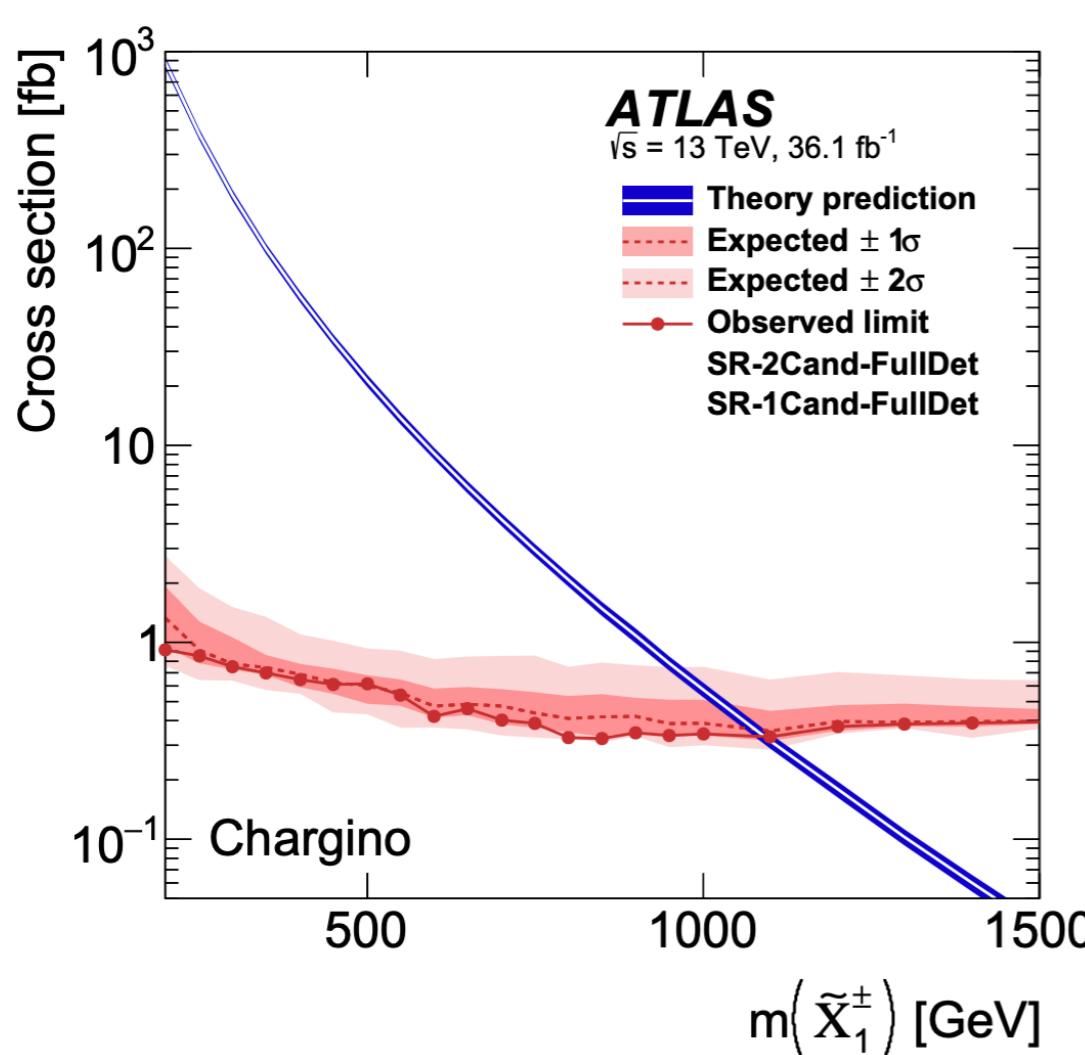
- Target Parameter Space:



- LHC:

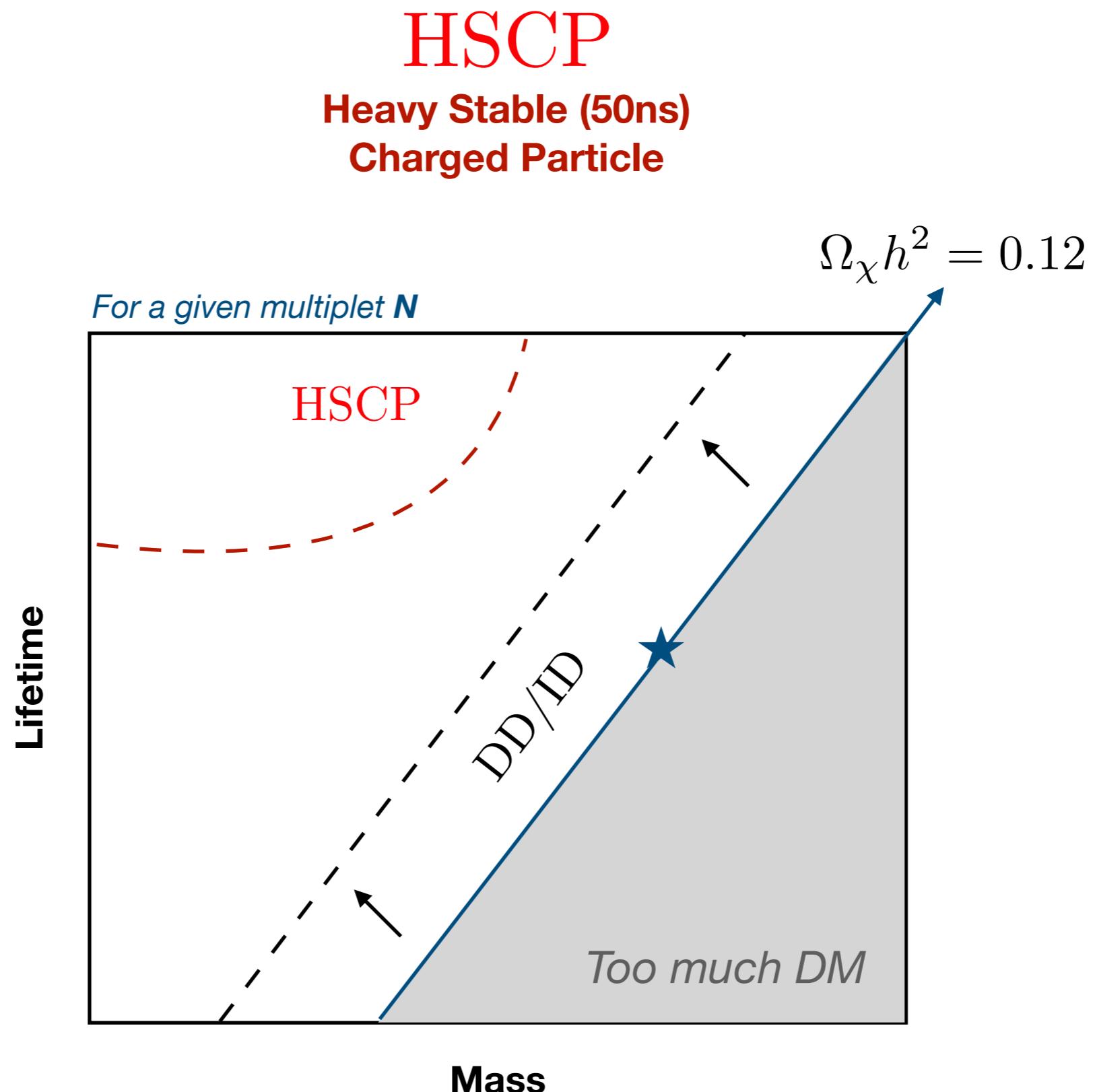
HSCP

Heavy Stable (50ns)
Charged Particle



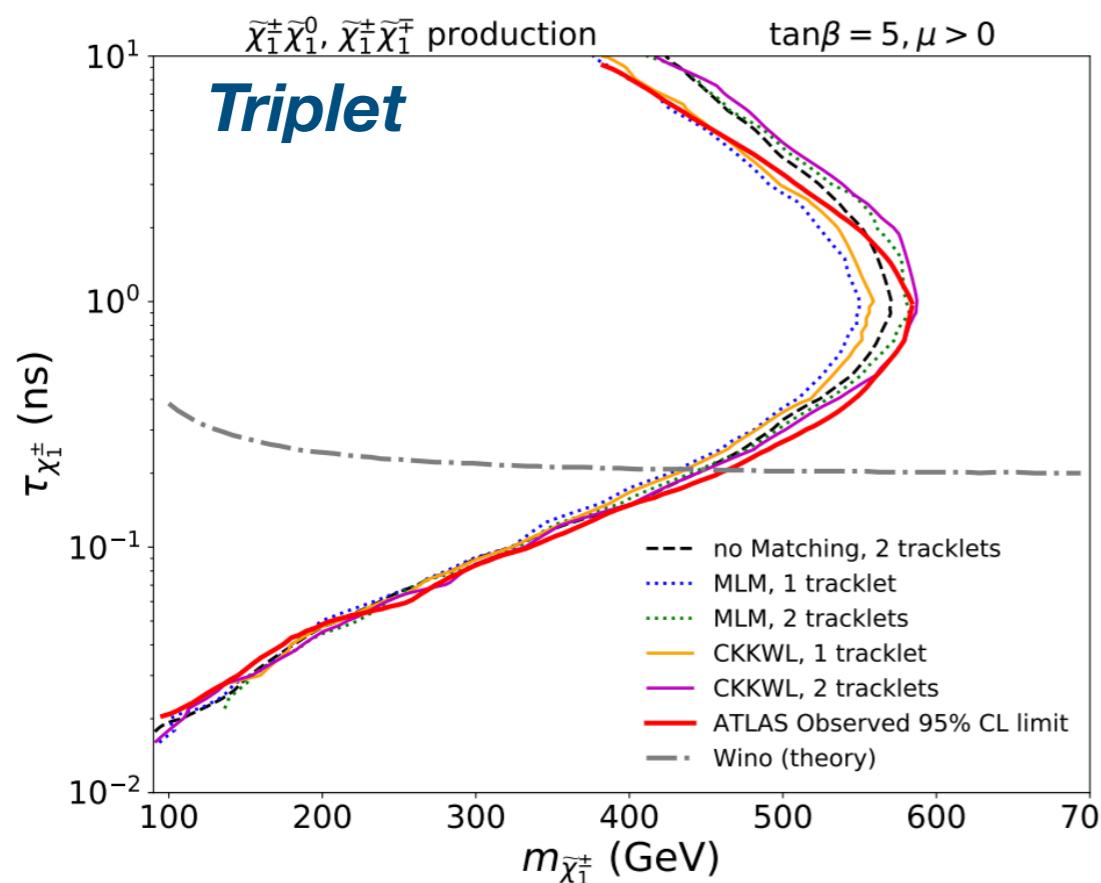
ATLAS Collaboration, Phys. Rev. D **99** (2019) 9, 092007

- LHC:

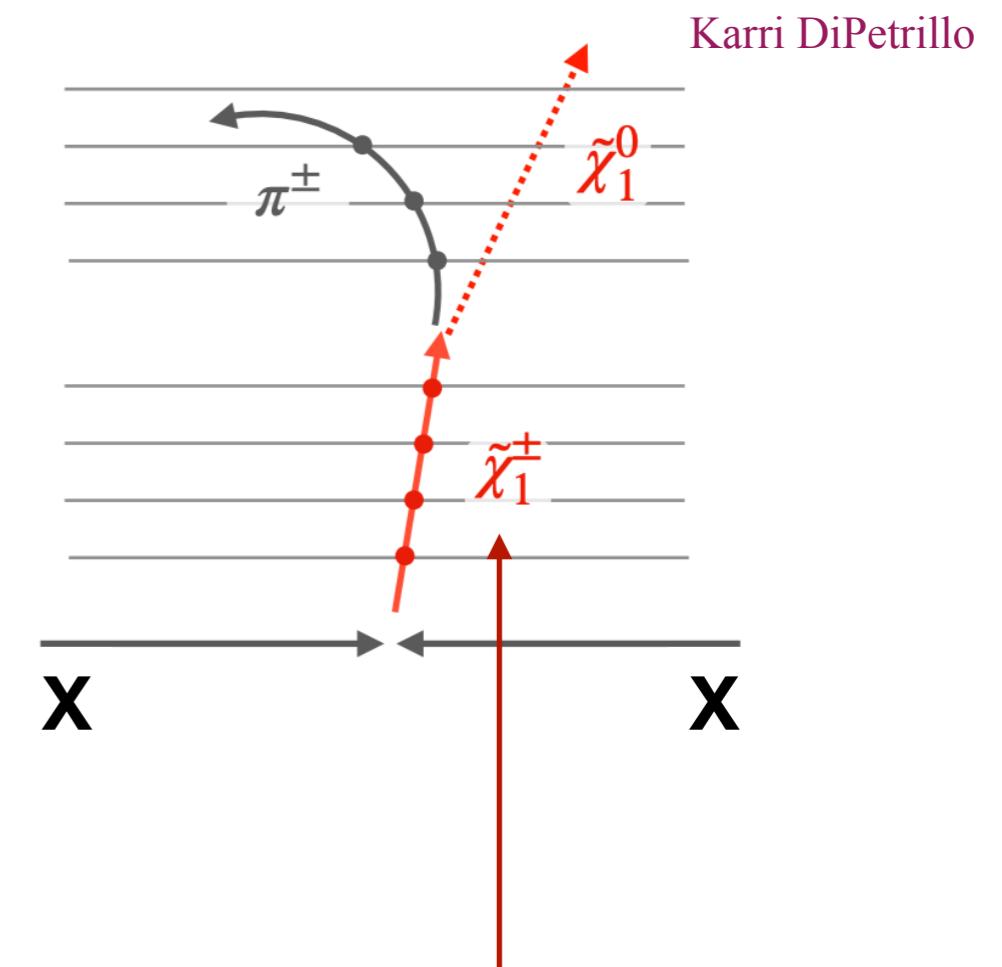


- LHC:

DT Disappearing Tracks

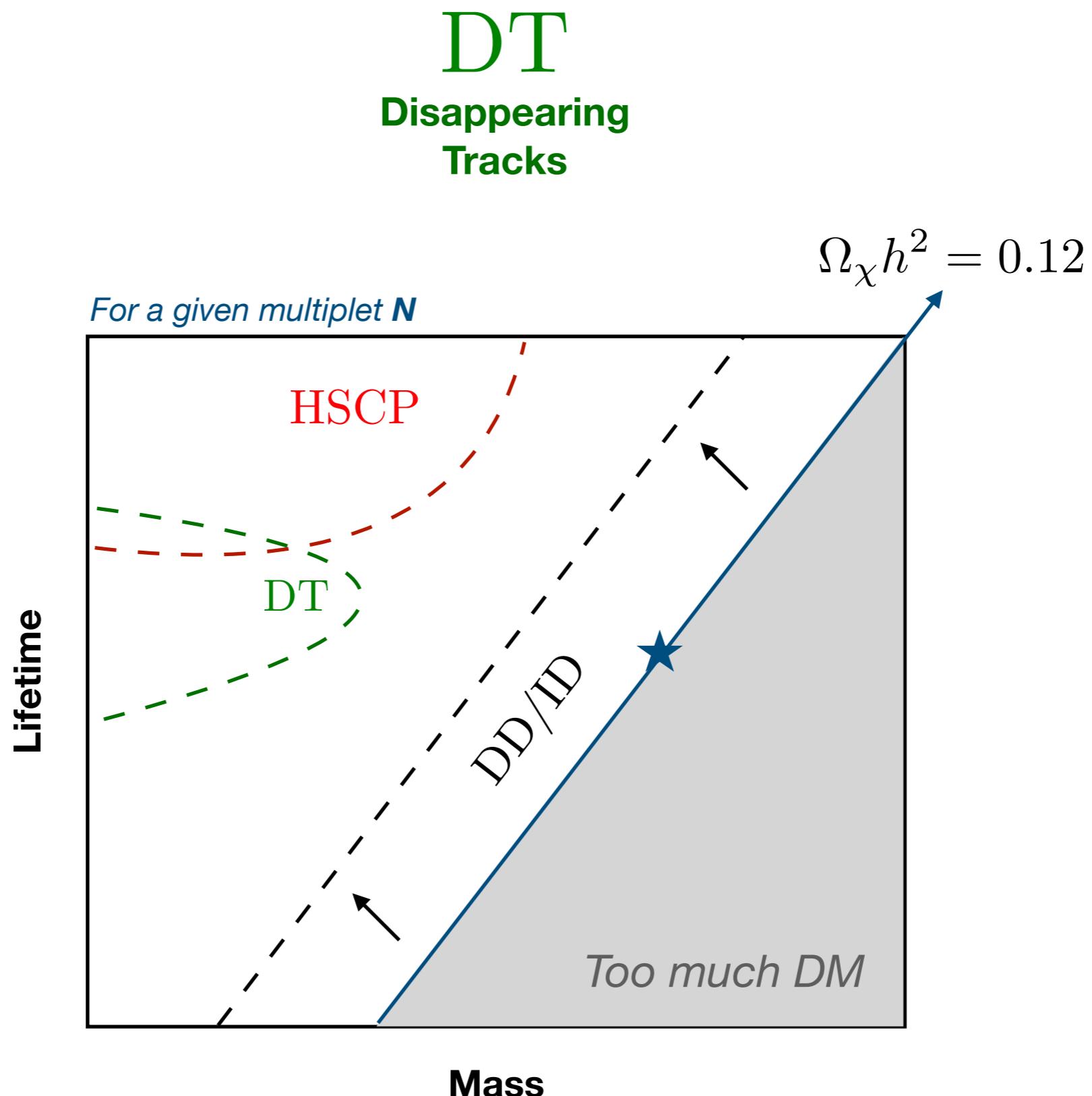


A. Belyaev et al., Phys. Rev. D **103** (2021) 9, 095006



Chargino has a long lifetime due to small mass gap!

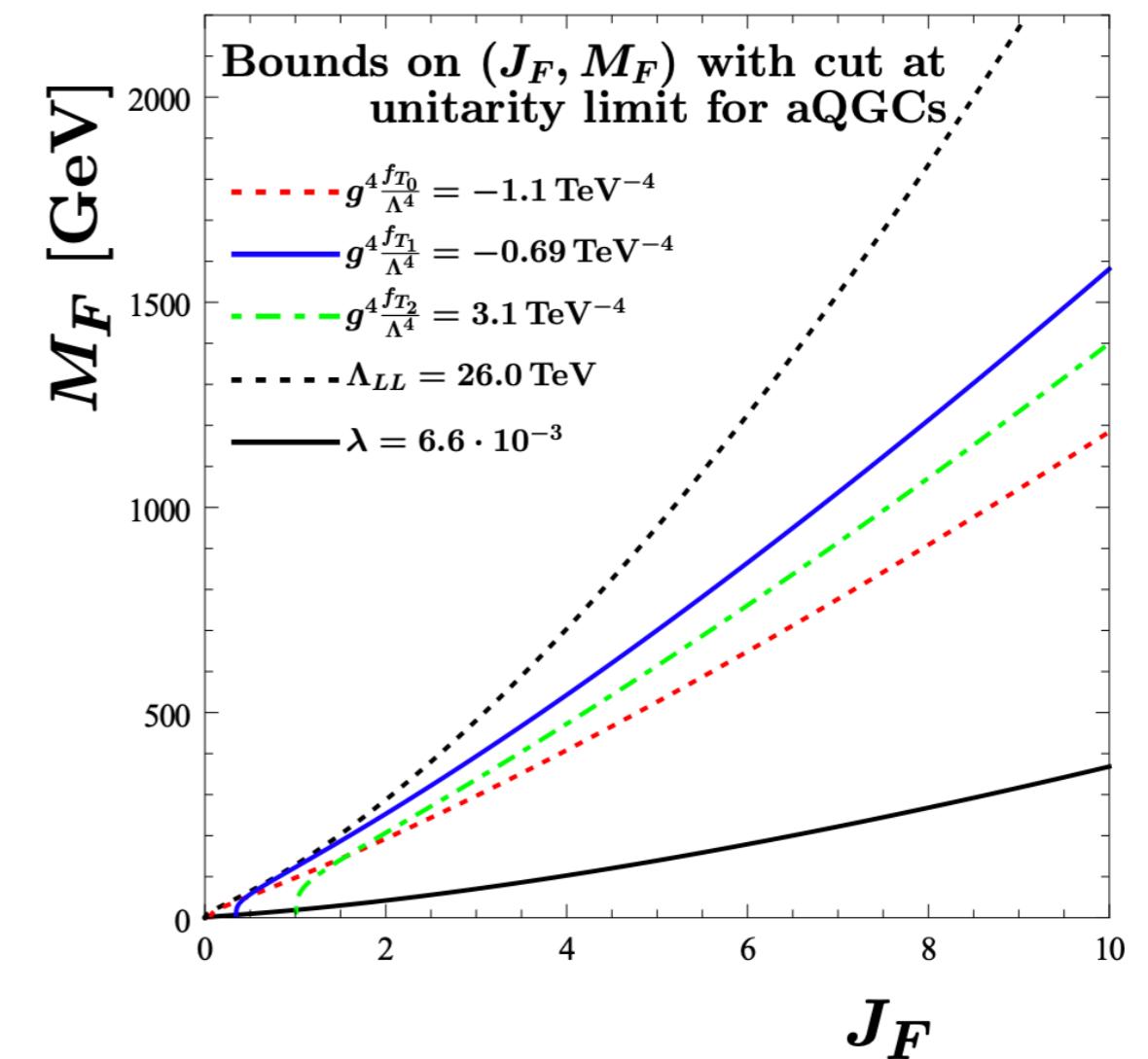
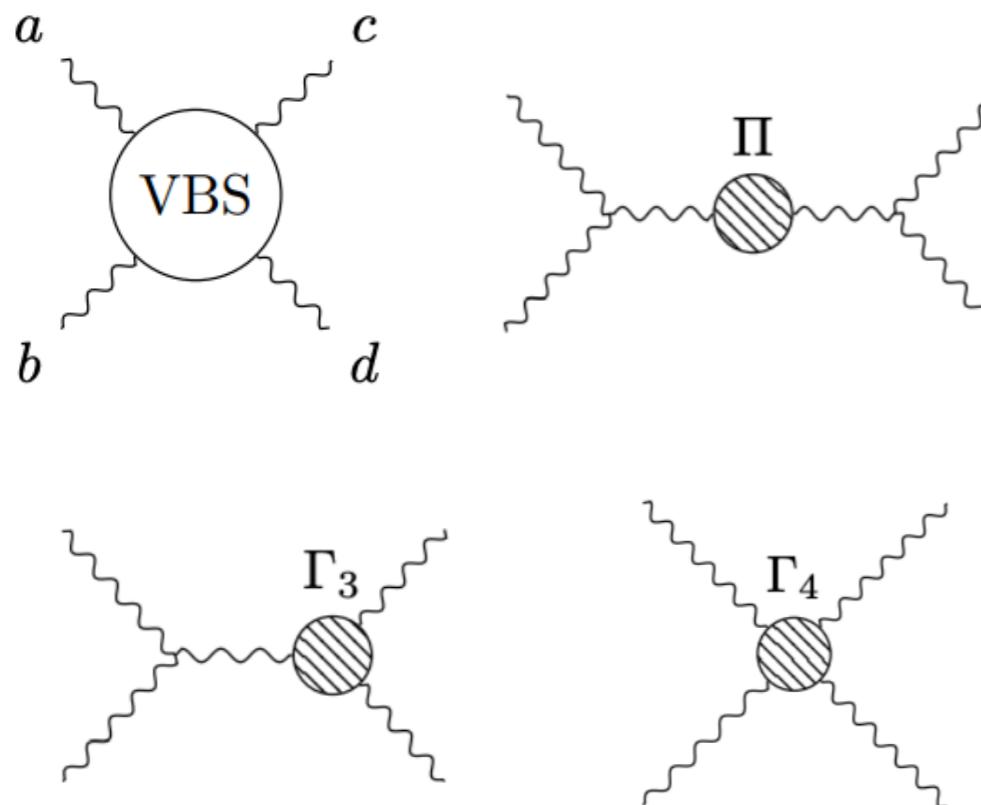
- LHC:



- LHC:

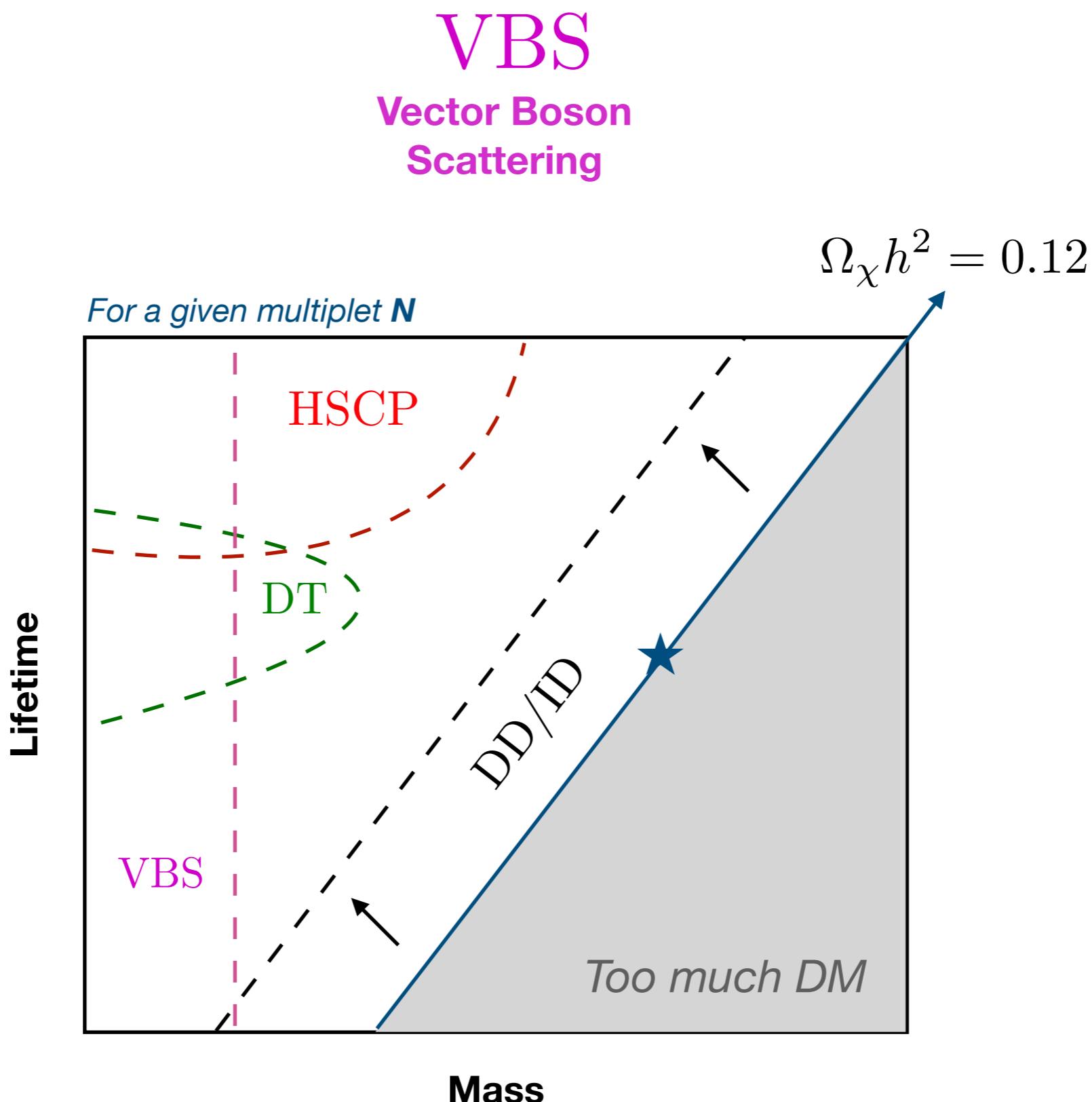
VBS

Vector Boson
Scattering

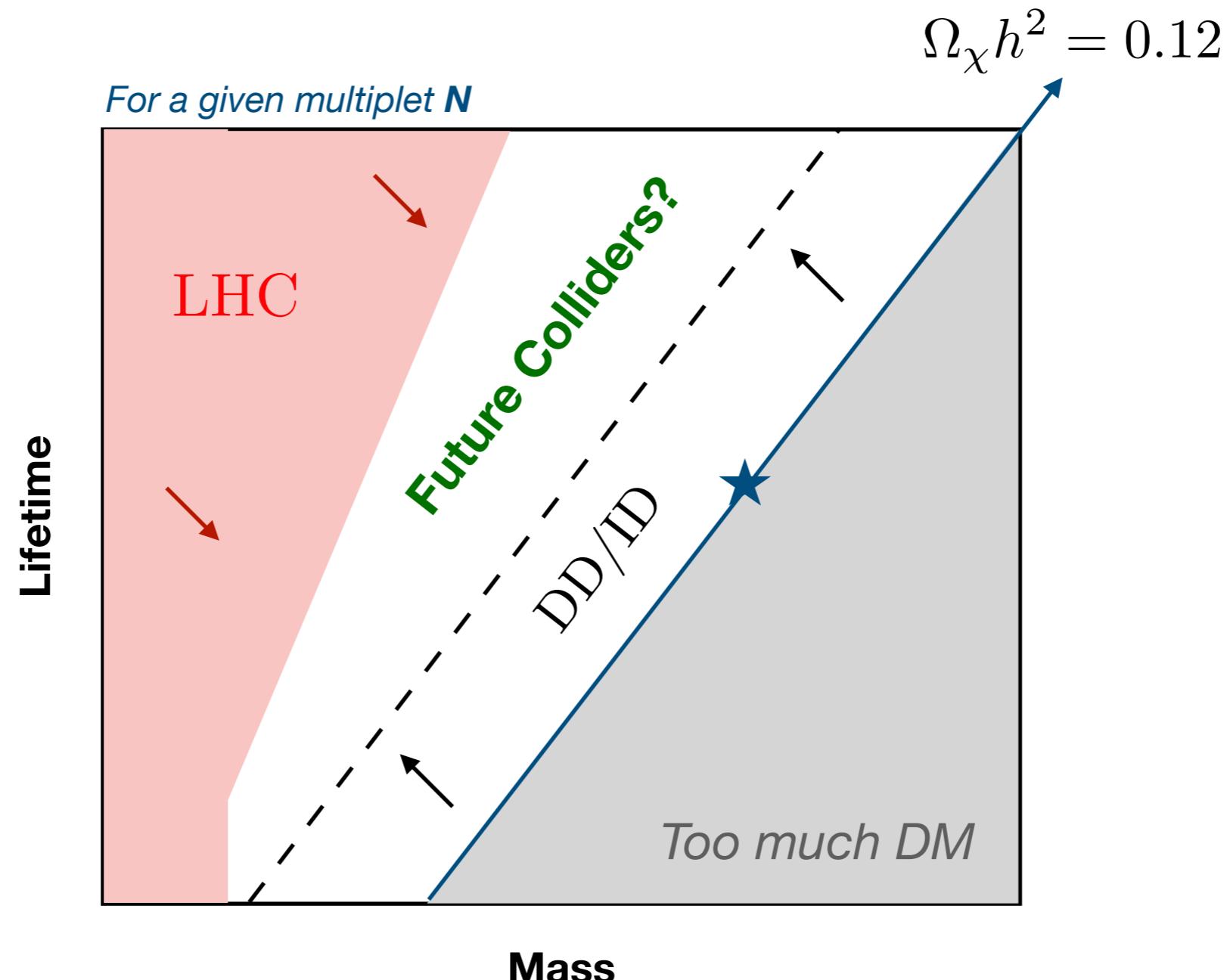


J. Lang et al., Eur. Phys. J. C 81 (2021) 7, 659

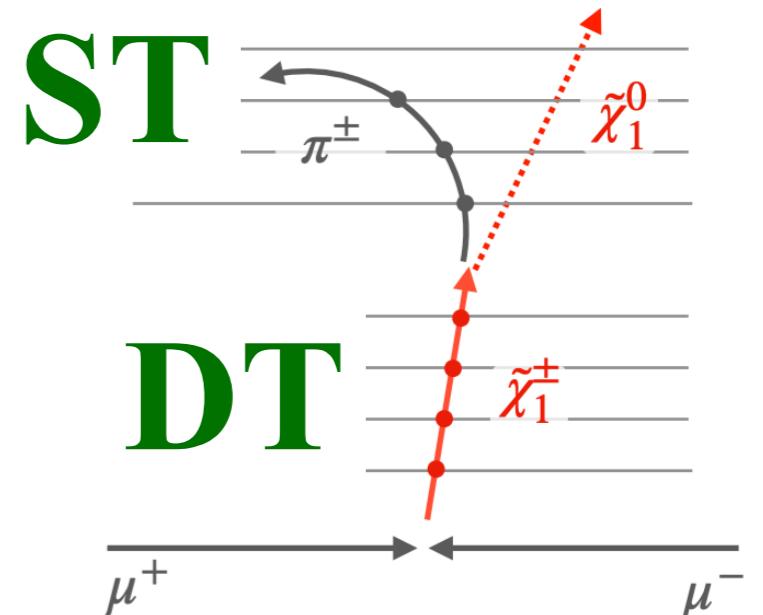
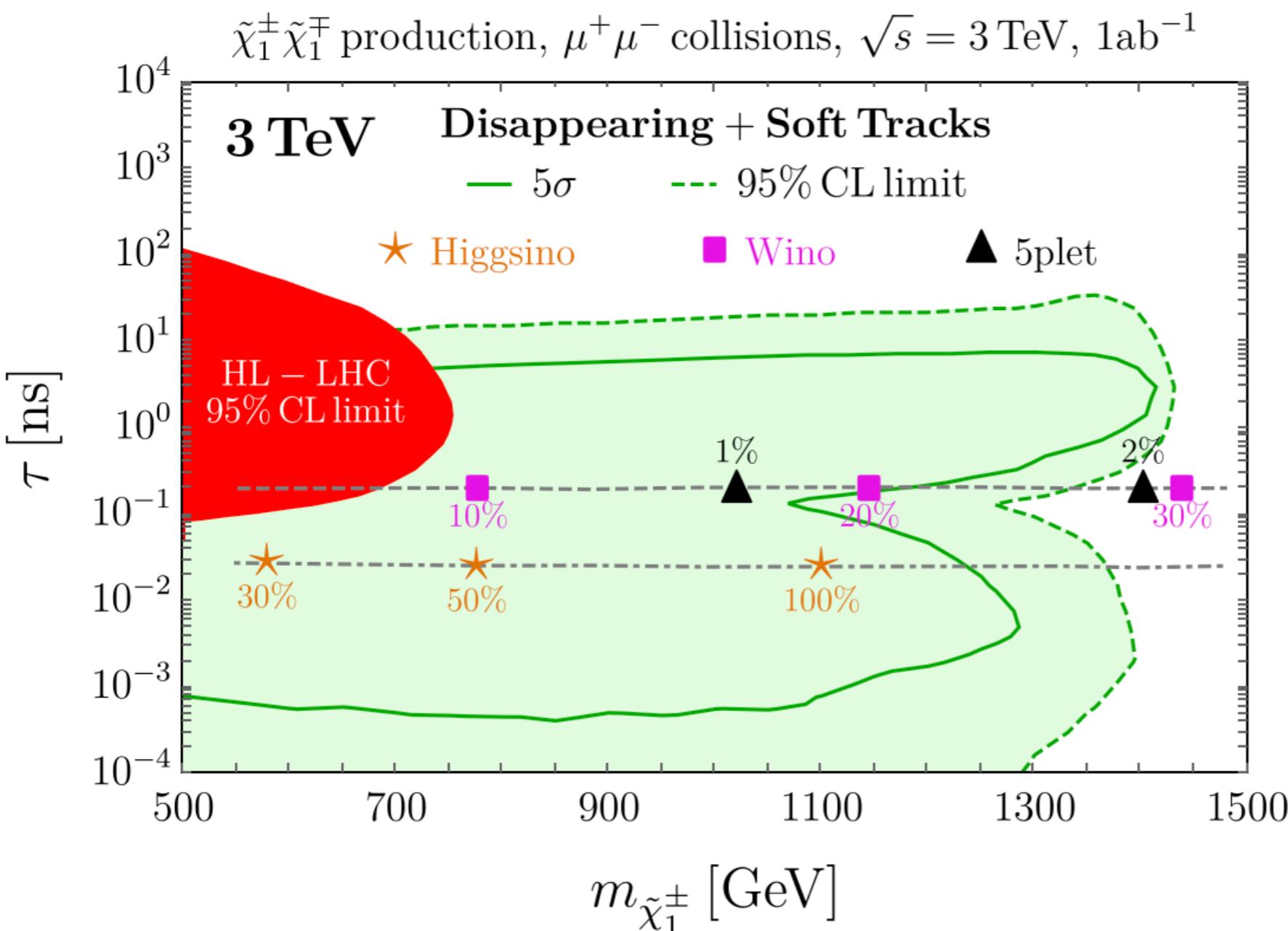
- LHC:



- Road to Minimal WIMPs?



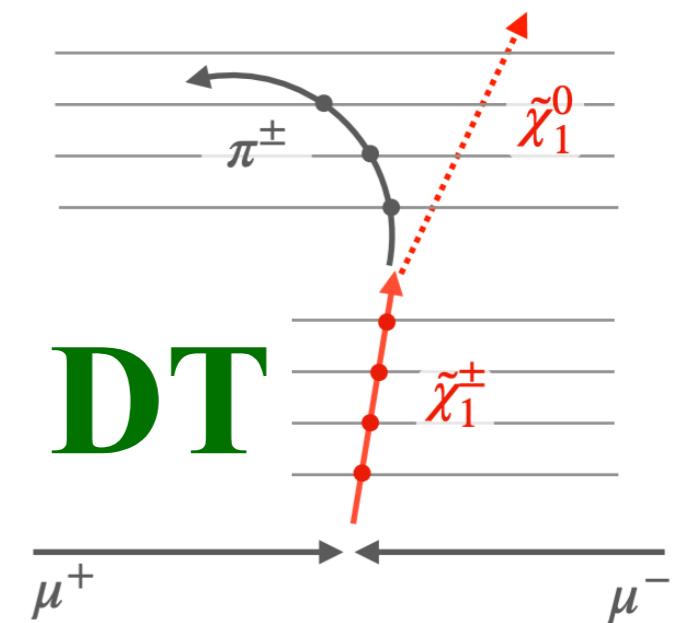
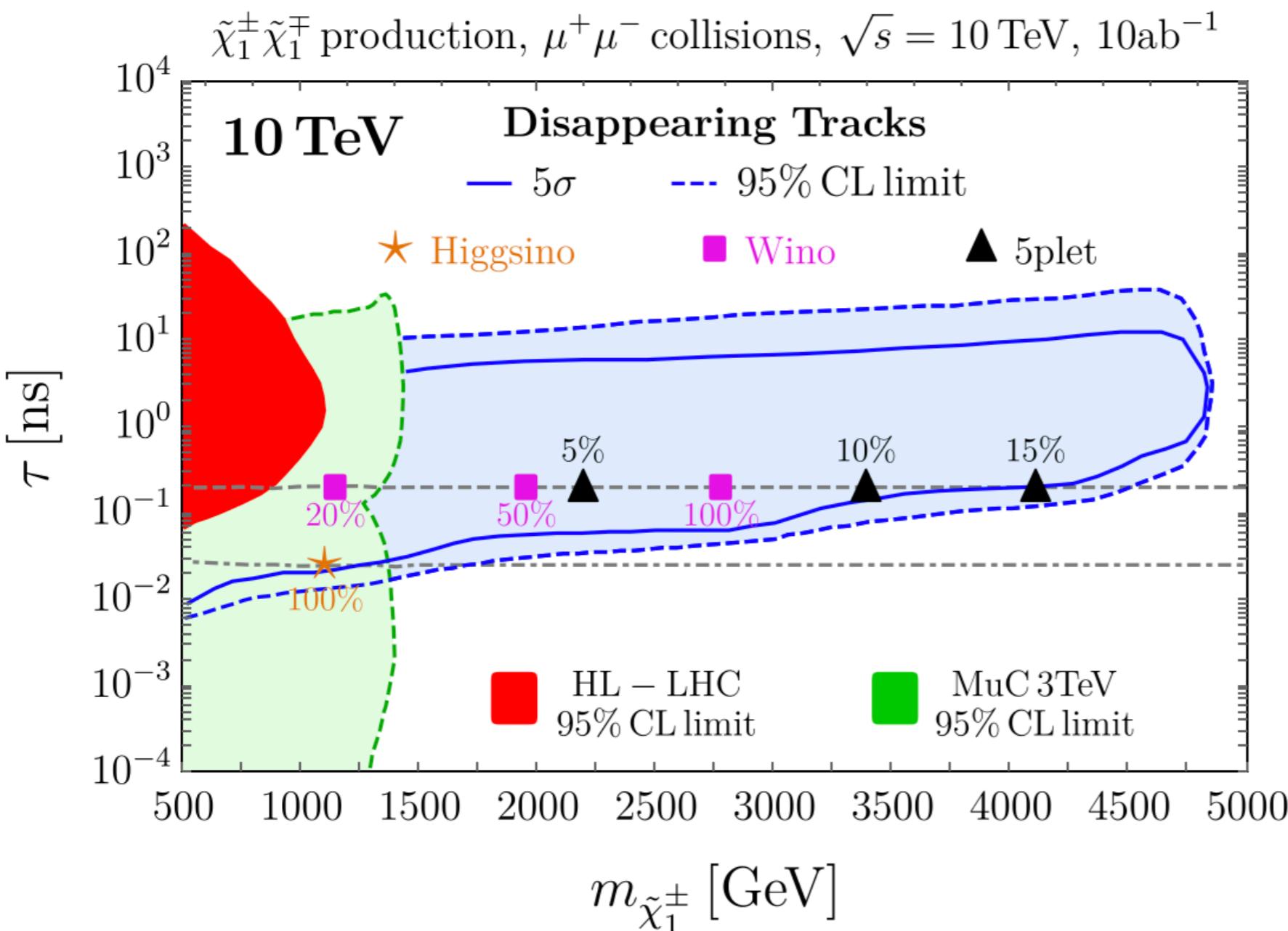
- Road to Minimal WIMPs:



The 3 TeV collider:

Doublets	100% of DM
Triplet	20% of DM
Fiveplet	1% of DM

- Road to Minimal WIMPs:



The 3 TeV collider:

Doublets	100% of DM
Triplet	100% of DM
Fiveplet	10% of DM

Summary

1. Minimal WIMPs constitute highly motivated targets for BSM physics. **Douplets/Triplets** that account for **100% of DM** in the Universe as well as **5-plets** that can account for **~10% of DM** fall into the multi-TeV rage, reachable at future colliders.
2. In our ***Road to Minimal WIMPs***, the LHC plays an important role. Searches for HSCP, DT, VBS cover important portions of the lifetime/mass parameter space.
3. The **Muon Collider** program ($3 \rightarrow 10$ TeV) could discover and characterize minimal WIMPs. **Disappearing Track** and **Soft Track** searches will allow us to determine the **mass** of the thermal relic, as well as the **mass gap** between this particle and its companion charged state.
4. These studies with full detector simulation and BIB overlay require the collaboration between **theorists and experimentalists**. Preliminary studies with new detector designs and new BIB data show improvements with respect to the results shown in this talk.

Thank You!