

Systematically Searching for New Physics at the LHC

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UC Davis HEP Seminar, April 2014



Outline

- I. Dark Matter
- II. Topological Models
- III. Deep networks

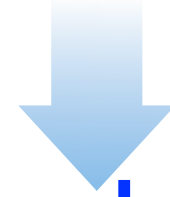
What do we know?



unknown unknown

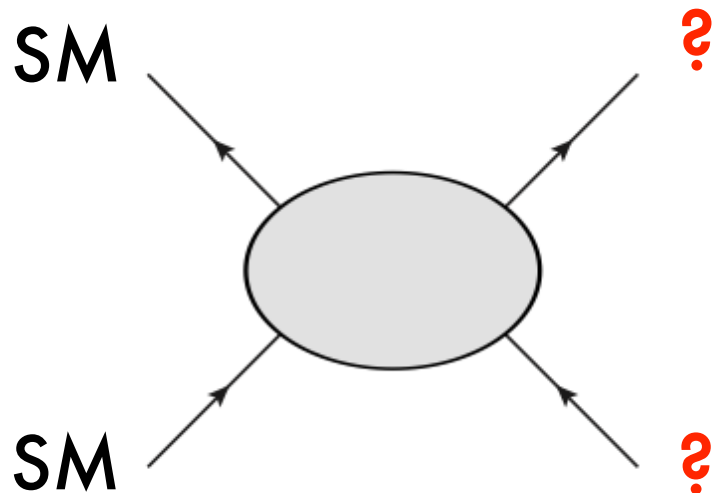


known unknown

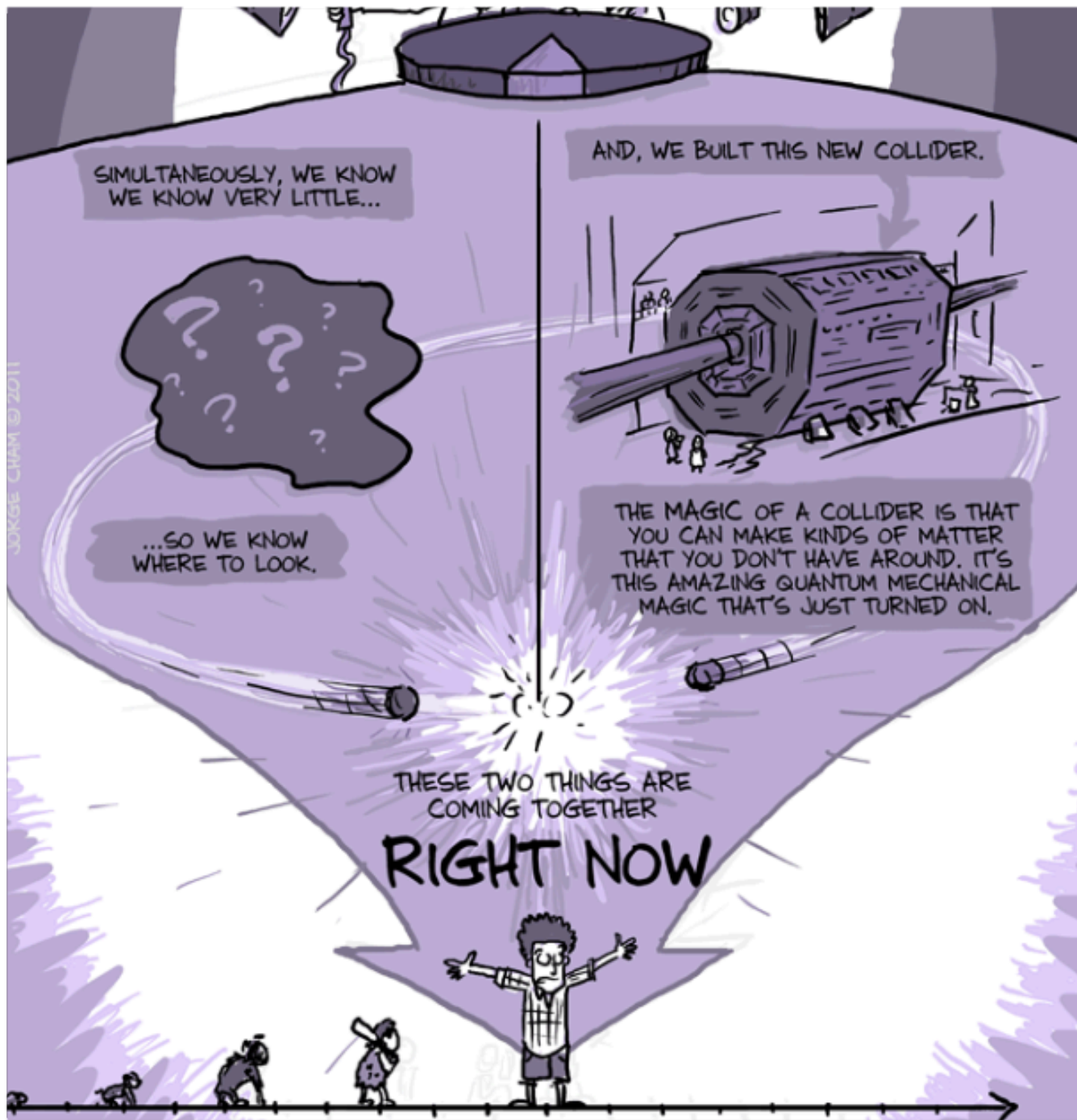


known known

Colliders: true alchemy



We can create new forms of matter, even if we have *little or no idea* of what we are looking for!



SIMULTANEOUSLY, WE KNOW WE KNOW VERY LITTLE...

AND, WE BUILT THIS NEW COLLIDER.

...SO WE KNOW WHERE TO LOOK.

THE MAGIC OF A COLLIDER IS THAT YOU CAN MAKE KINDS OF MATTER THAT YOU DON'T HAVE AROUND. IT'S THIS AMAZING QUANTUM MECHANICAL MAGIC THAT'S JUST TURNED ON.

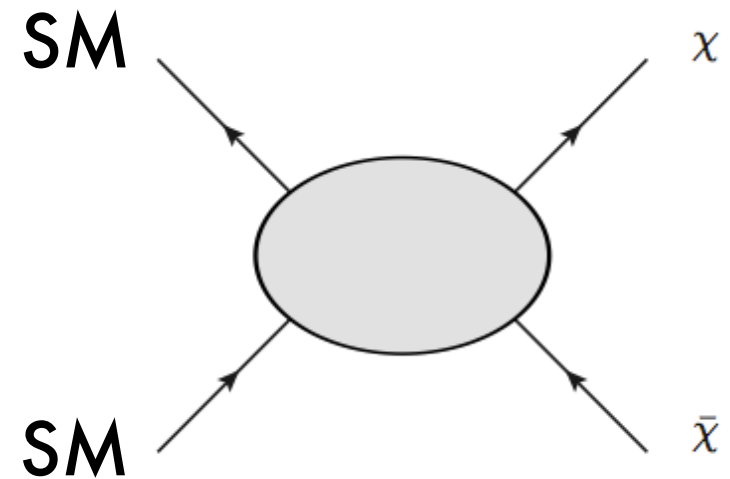
THESE TWO THINGS ARE COMING TOGETHER
RIGHT NOW

JORGE CHAM © 2011

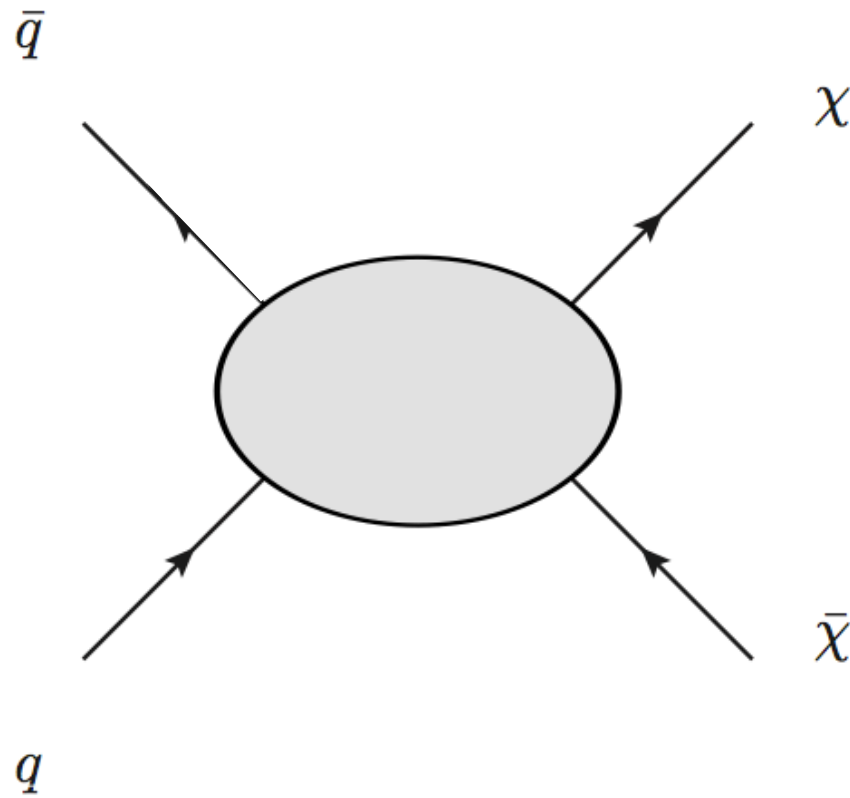
Interactions



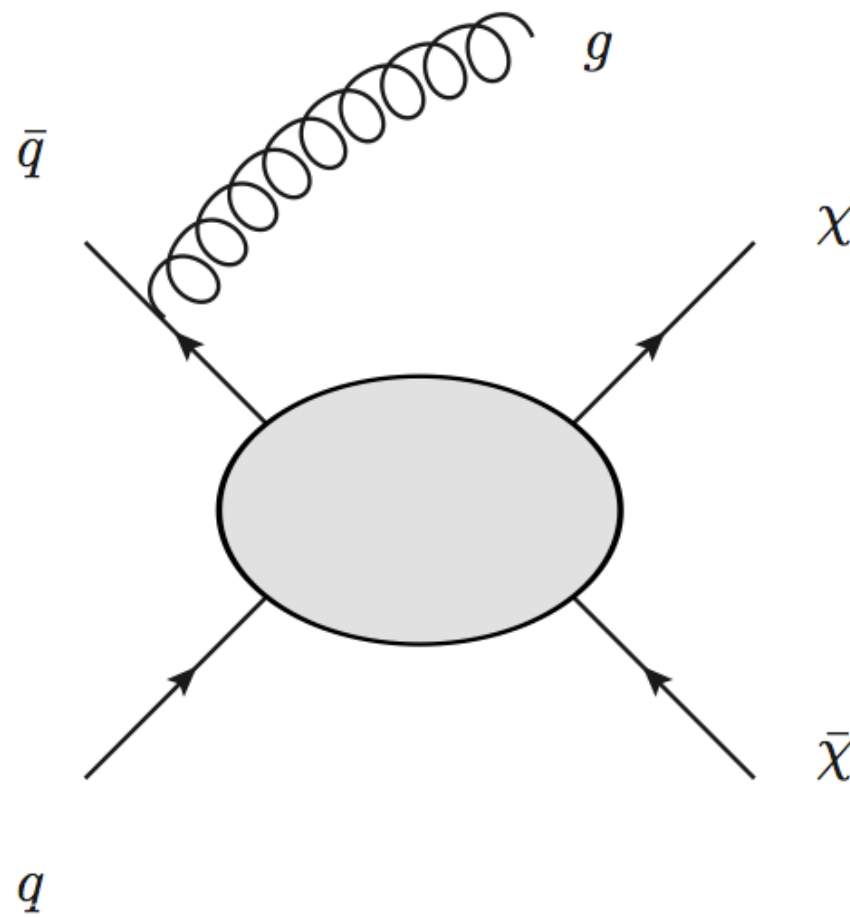
Important assumption:
Requires **some**
interaction with SM



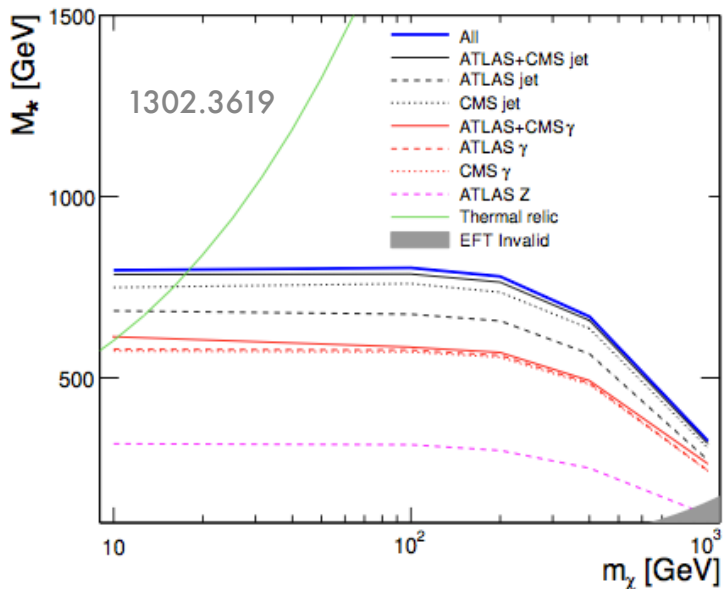
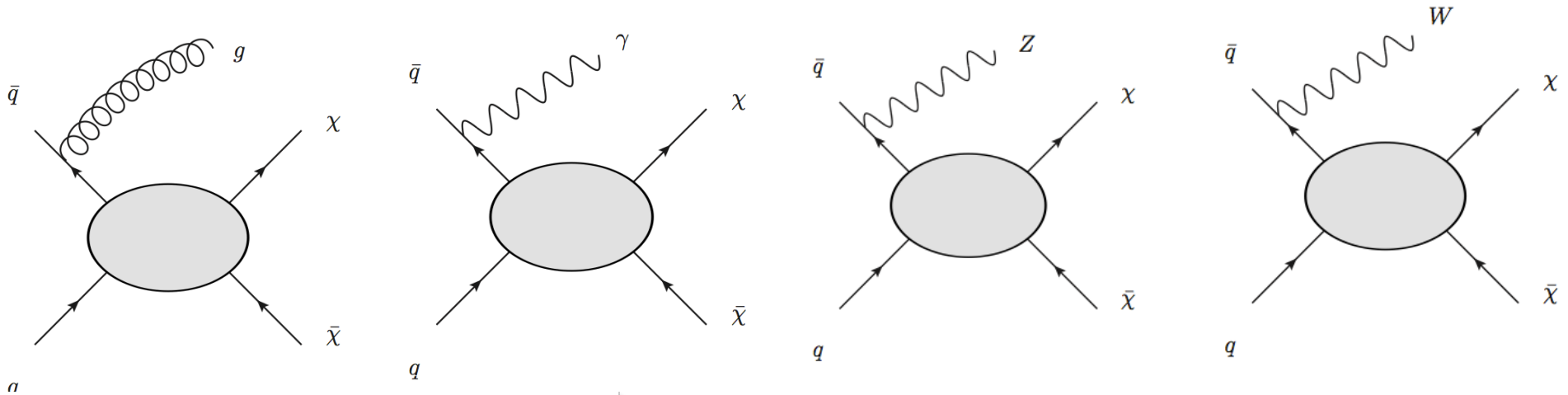
DM @ Colliders



DM @ Colliders

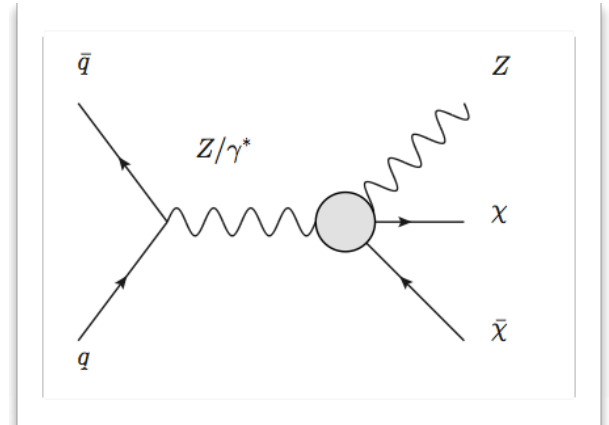


Look everywhere



Mono-jet most powerful for $qqXX$

Each mode has unique models where it is a **discovery mode.**



Outline

A. Mono-W

B. Mono-Z

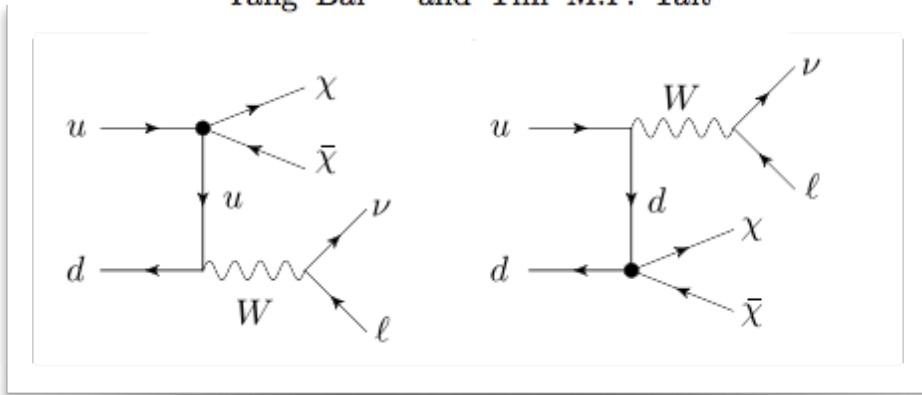
C. Mono-Higgs

Mono-W

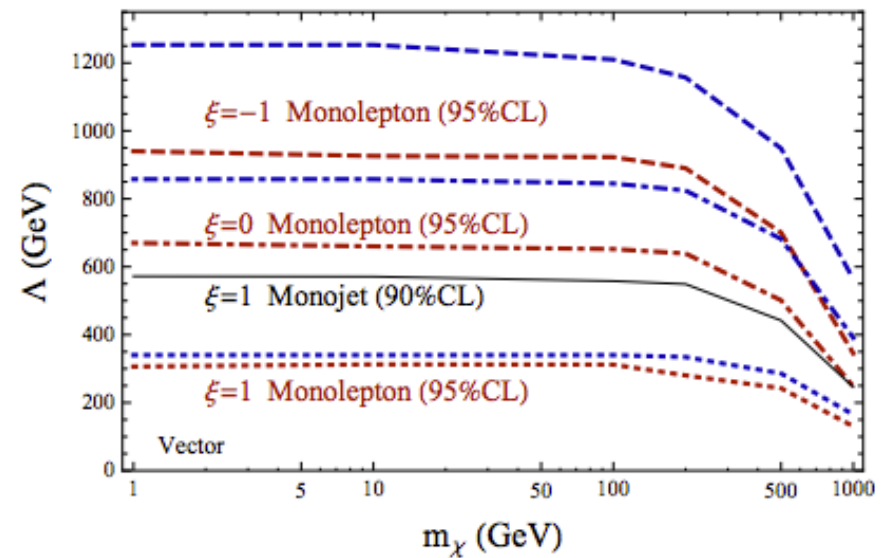
Mono-W theory

Searches with Mono-Leptons

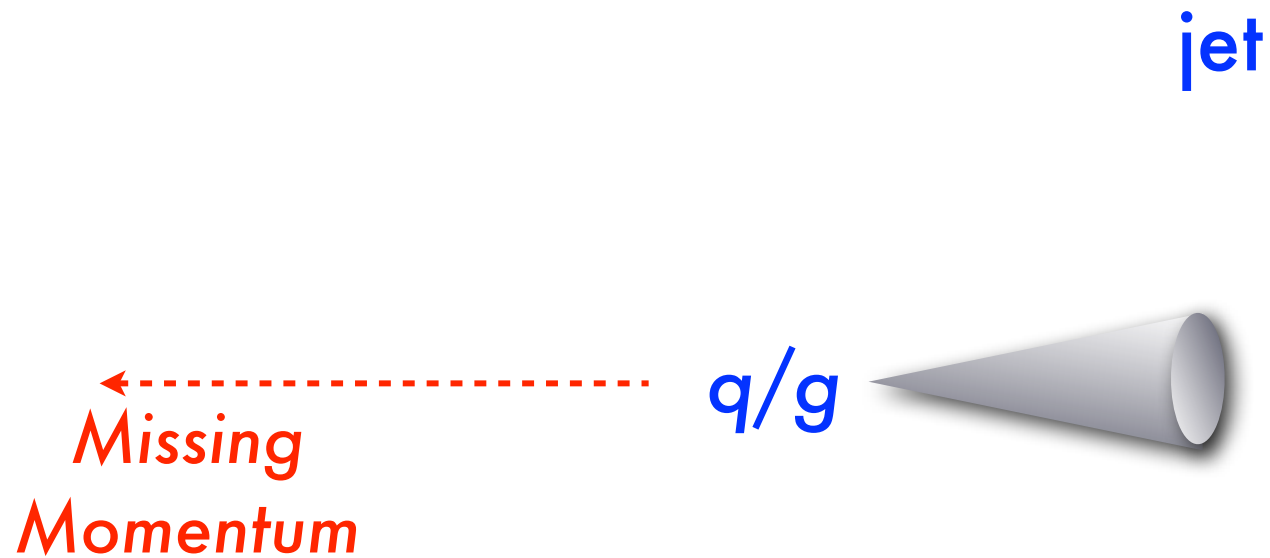
Yang Bai^{a,b} and Tim M.P. Tait^c



1208.4361

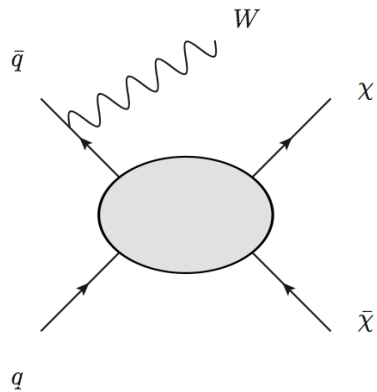


Mono-jet



Mono-heavy jet

1309.4017 (PRL)



**Missing
Momentum**

W/Z

fat jet

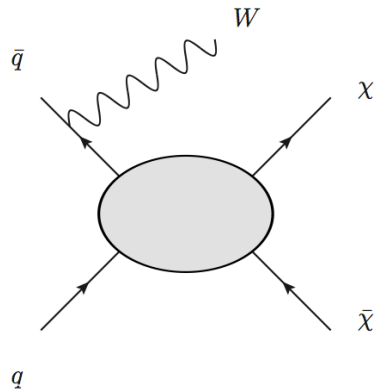
sub-jet

sub-jet

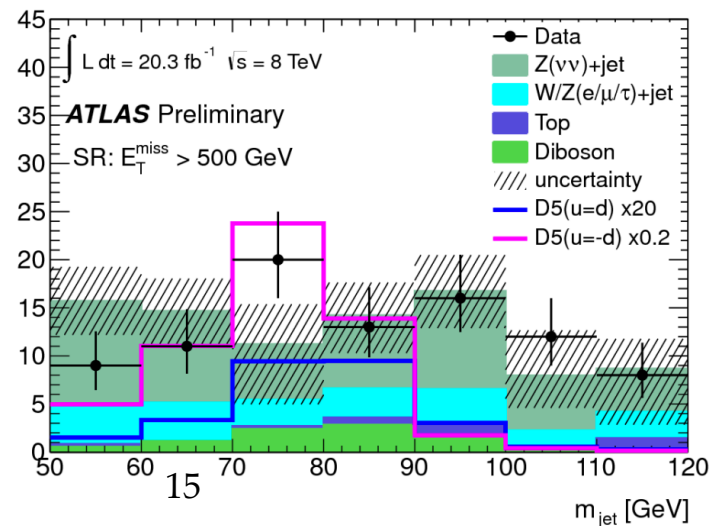
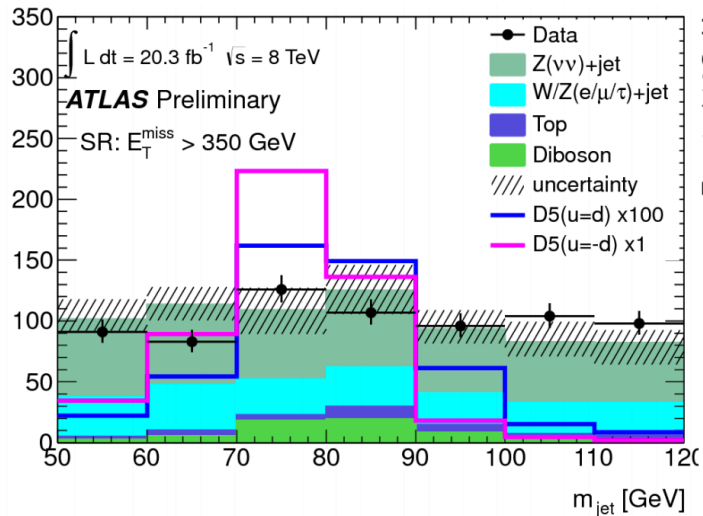


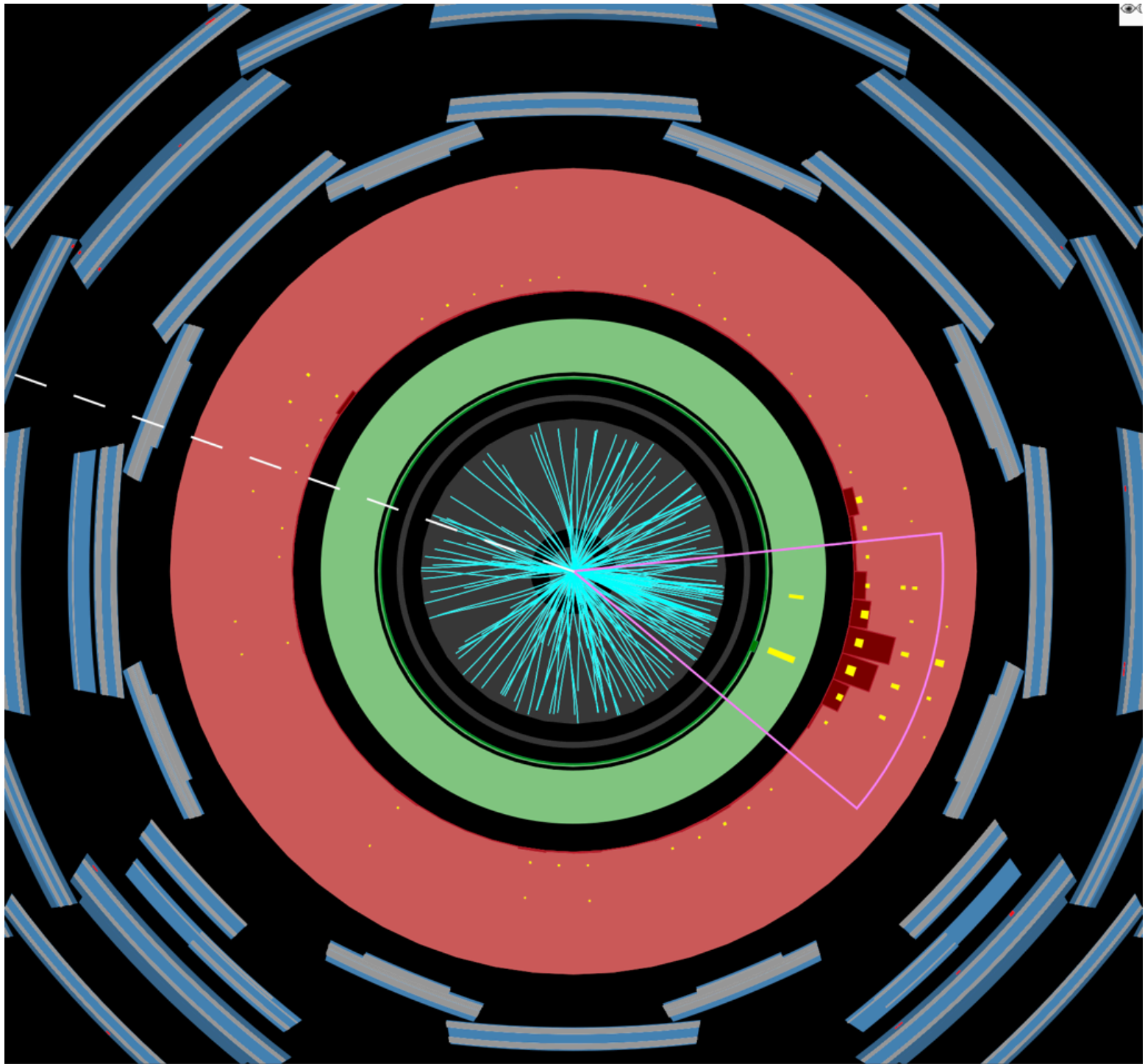
14 Ning Zhou, UCI

mono-W, etc



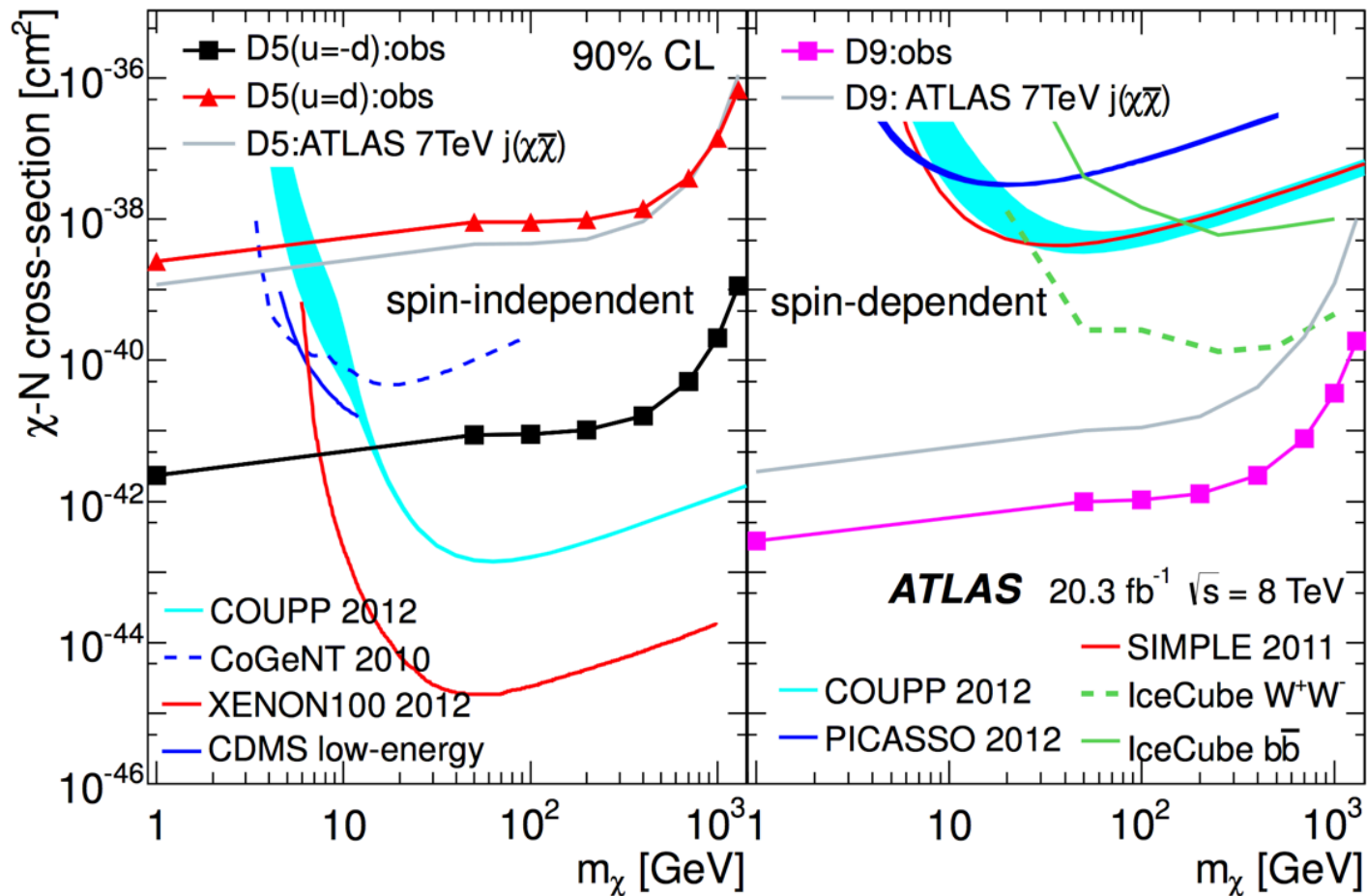
Fat jet $p_T > 250$ 1309.4017 (PRL)
 two subjets giving $m_{\text{jet}} = [50, 120]$
 No e, mu, gamma
 ≤ 1 additional narrow jets
 MET > 350 or 500



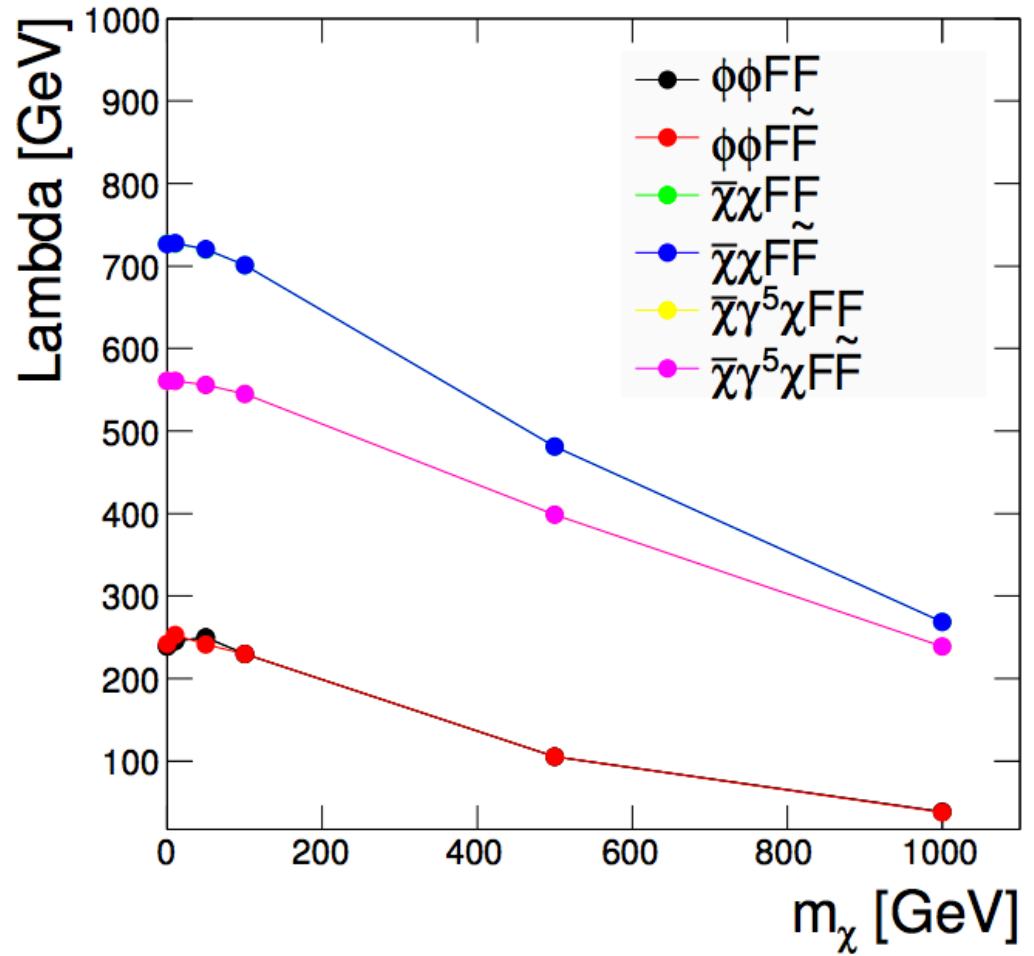
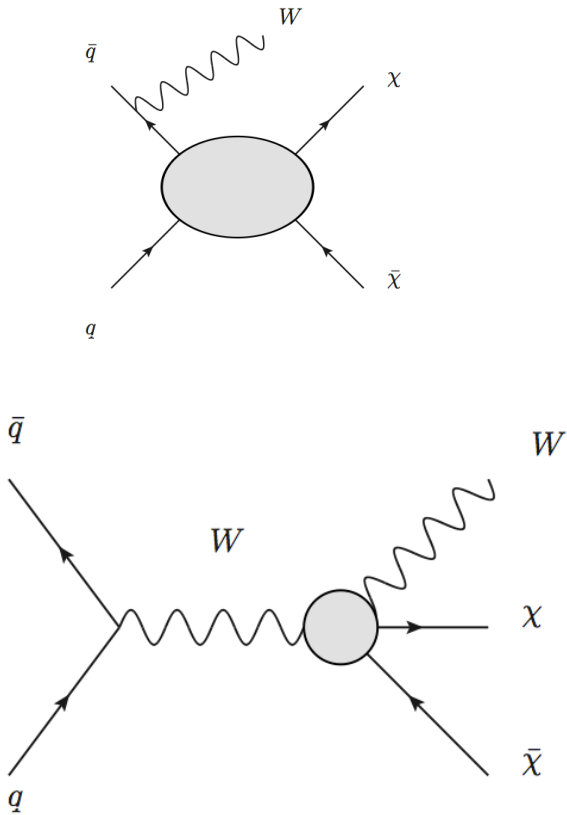


Limits

1309.4017 (PRL)



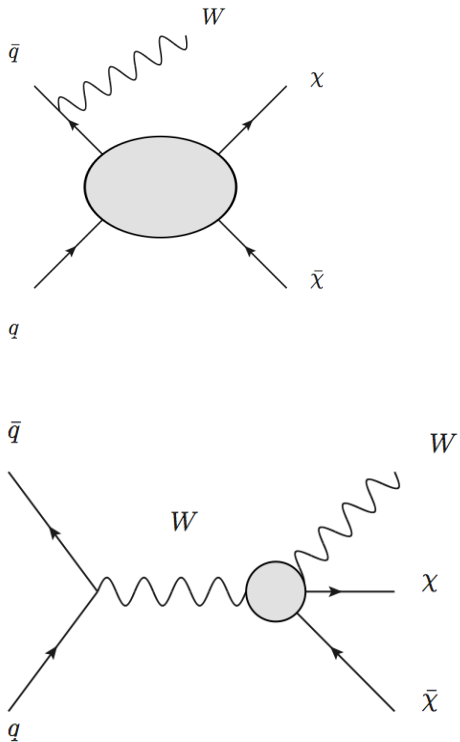
XX \rightarrow WW



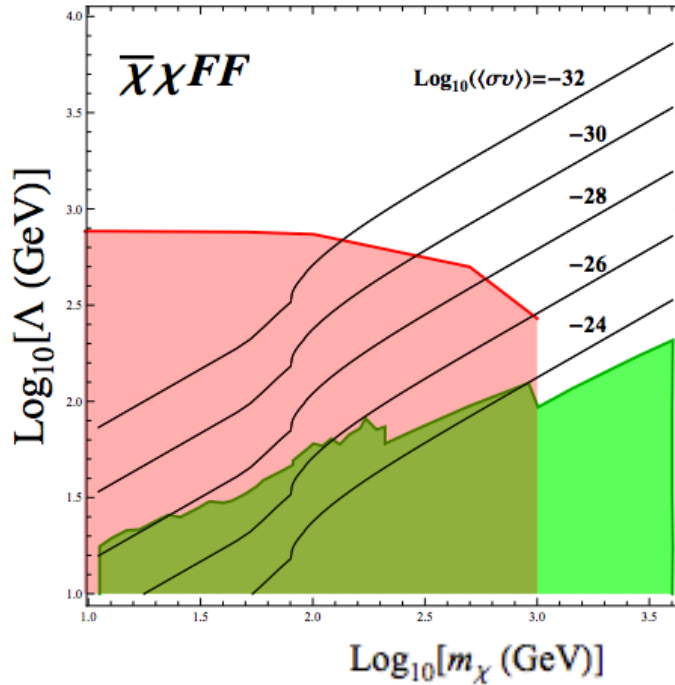
1403.6734

FIG. 4: Limits on Λ as a function of m_χ .

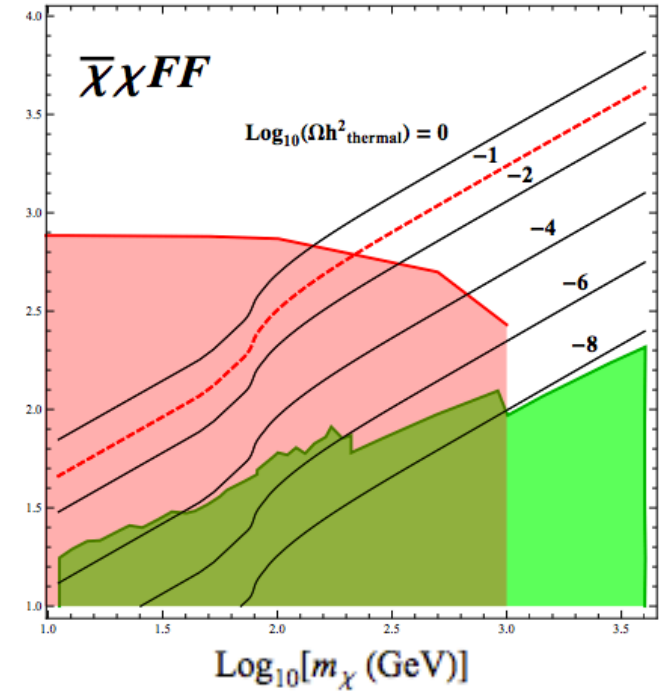
XX- \rightarrow WW



MonoW and Indirect



MonoW and Indirect



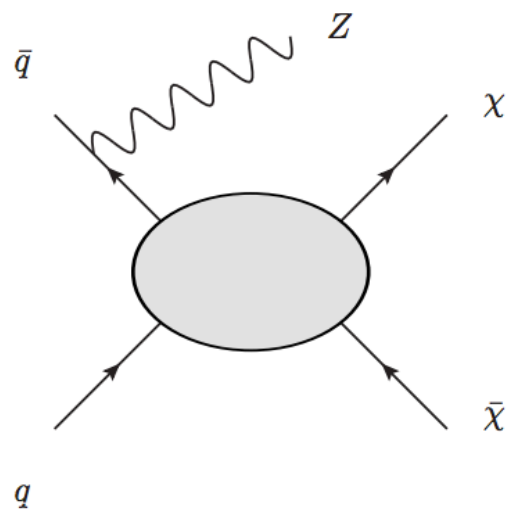
"Indirect" is an excluded region which is a combination of exclusions from the LAT line search, the LAT dwarf bounds and (at higher m_χ) the VERITAS Segue bounds. It is assumed that this DM makes up 100% of cosmological DM, no matter what its annihilation cross section is.

1403.6734

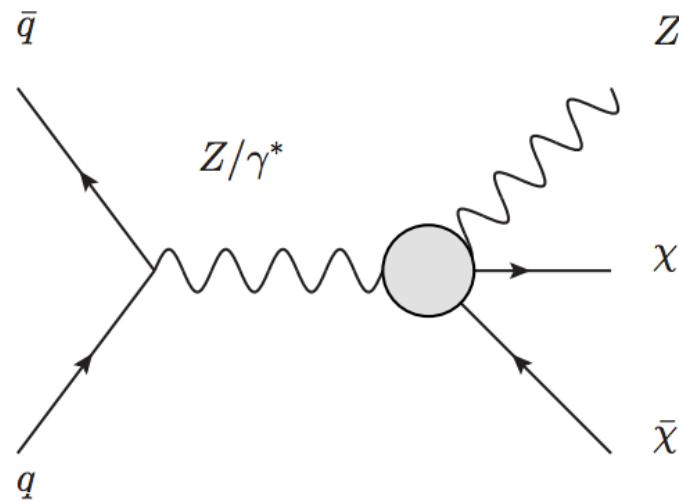
Mono-Z

1404.0051

EFTs



(a) Feynman diagram showing an ISR operator.

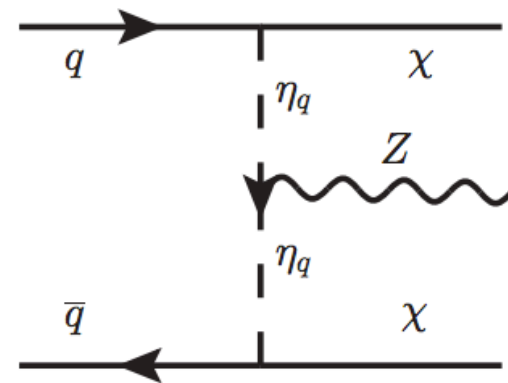
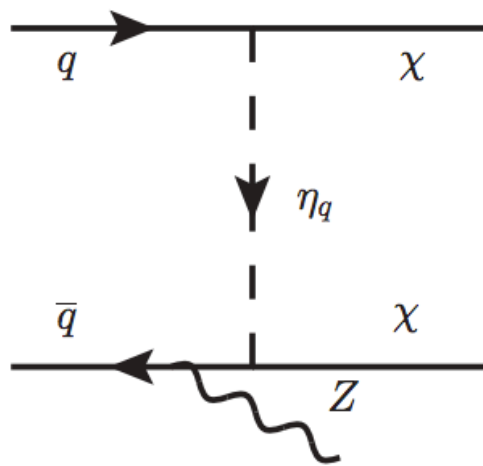
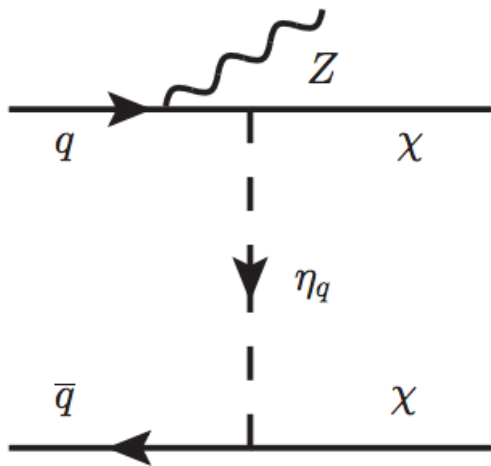


(b) Feynman diagram showing a $ZZ\chi\chi$ operator.

1404.0051



Simplified models



1404.0051



Selection

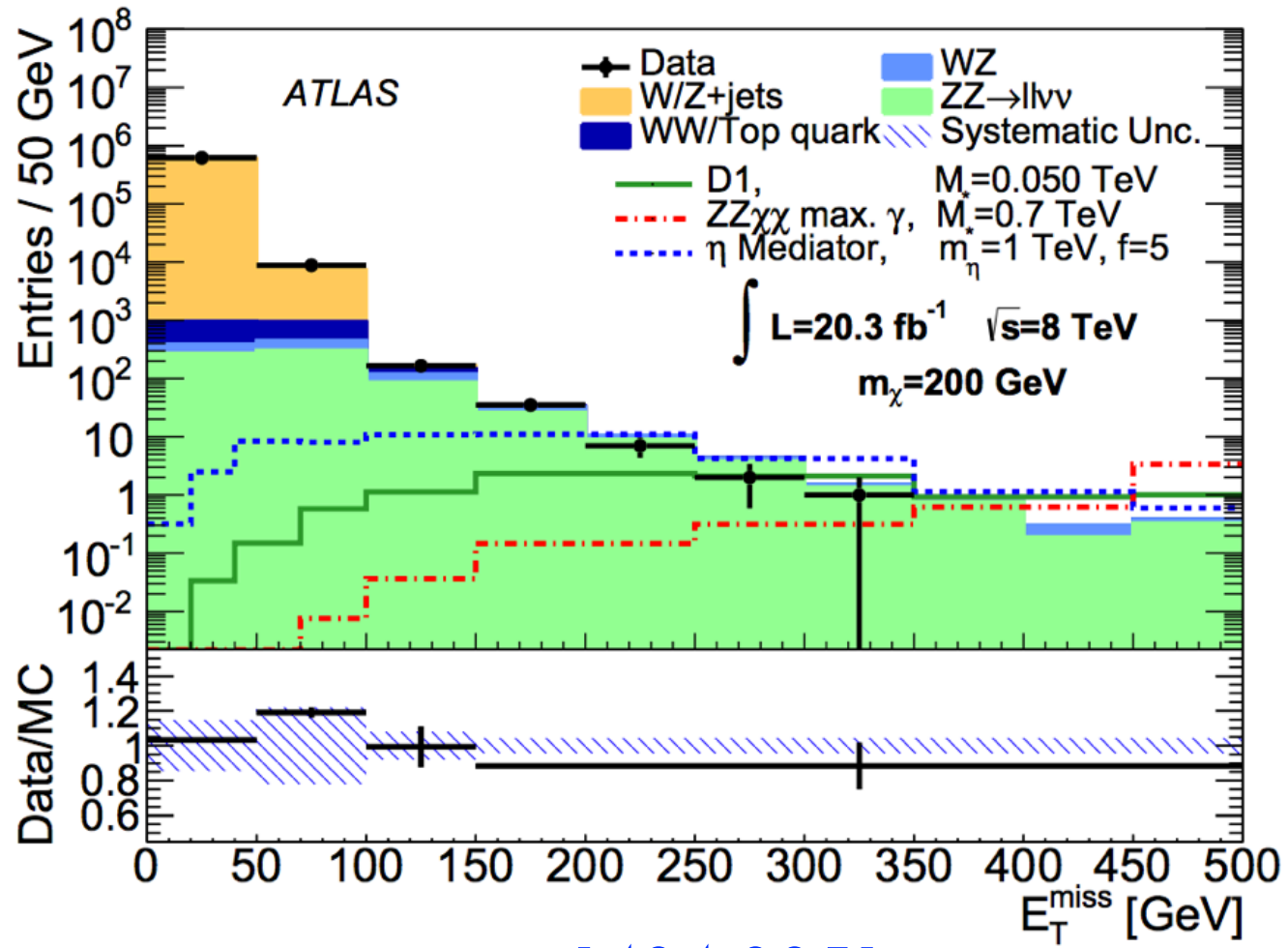
Process	E_T^{miss} threshold [GeV]			
	150	250	350	450
ZZ	41 ± 15	6.4 ± 2.4	1.3 ± 0.5	0.3 ± 0.1
WZ	8.0 ± 3.1	0.8 ± 0.4	0.2 ± 0.1	0.1 ± 0.1
$WW, t\bar{t}, Z \rightarrow \tau^+\tau^-$	1.9 ± 1.4	$0_{-0.0}^{+0.7}$	$0_{-0.0}^{+0.7}$	$0_{-0.0}^{+0.7}$
$Z+\text{jets}$	0.1 ± 0.1	–	–	–
$W+\text{jets}$	0.5 ± 0.3	–	–	–
Total	52 ± 18	7.2 ± 2.8	1.4 ± 0.9	$0.4_{-0.4}^{+0.7}$
Data	45	3	0	0

Selection

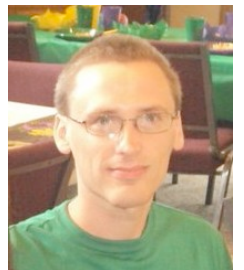
Two OS SF leps
 m_{ll} in [76,106]
veto jets, 3rd lep
MET angle cuts

$$|p_T^{ll} - E_T^{\text{miss}}|/p_T^{ll} < 0.5$$

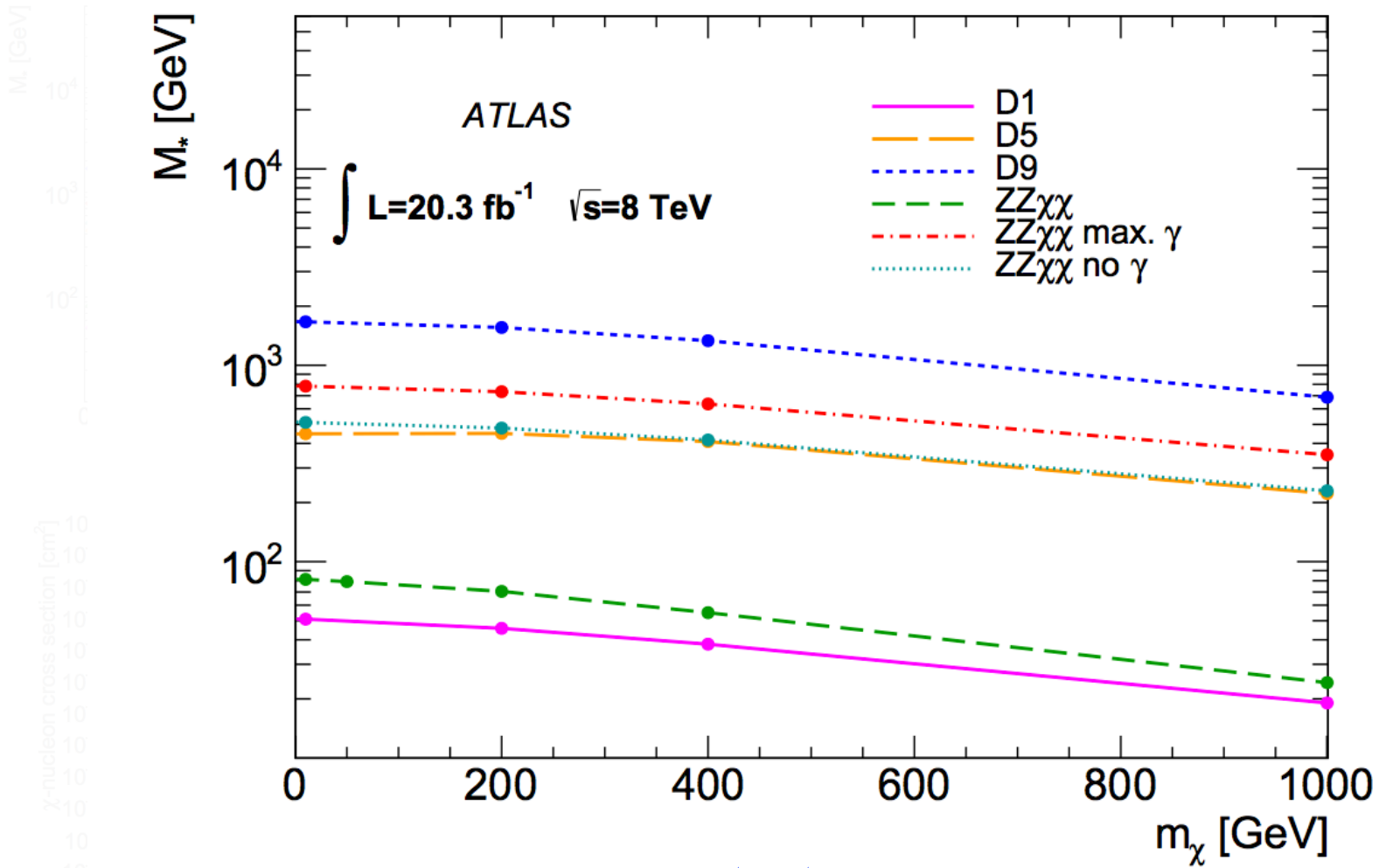
Data



1404.0051



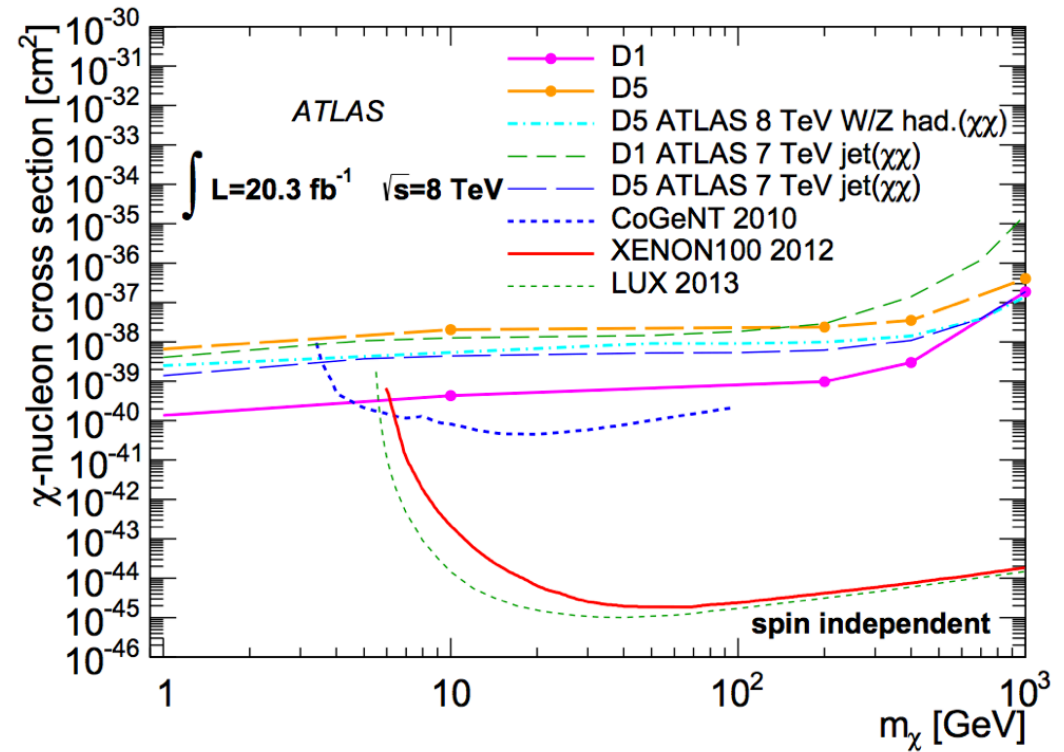
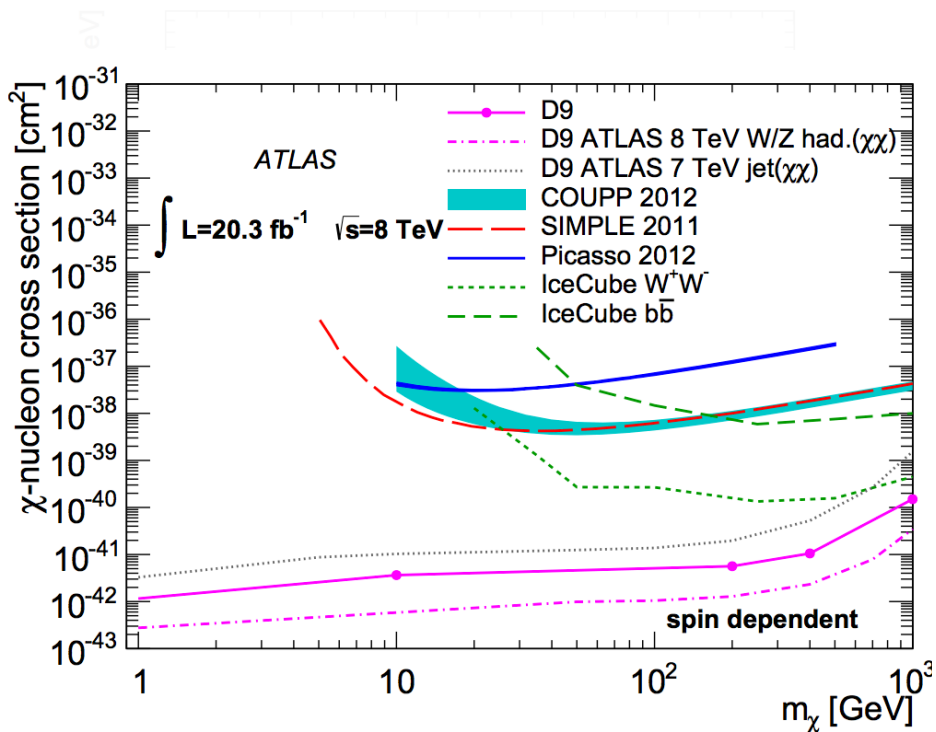
Limits....



1404.0051



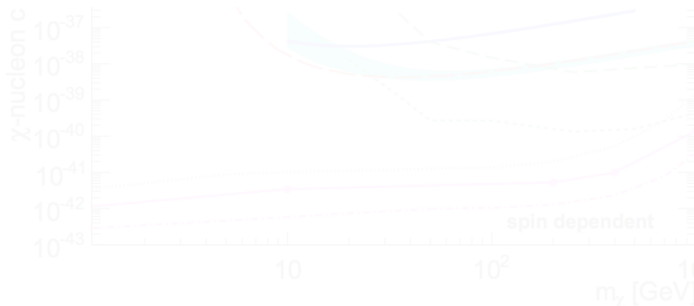
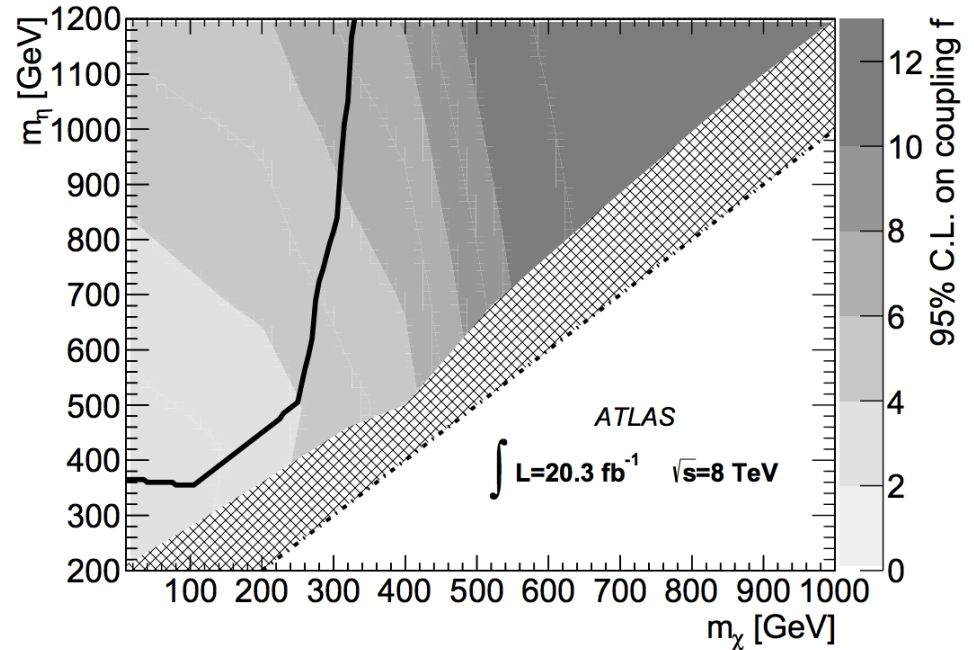
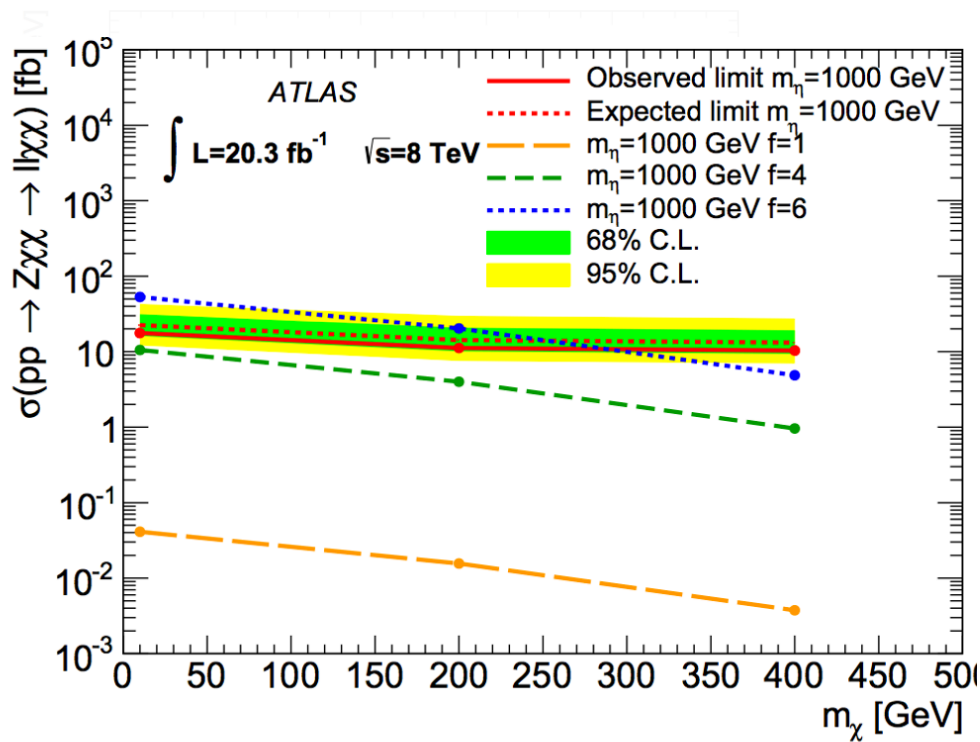
Limits....



1404.0051



Limits....



1404.0051



Mono-Higgs

1312.2592

Models

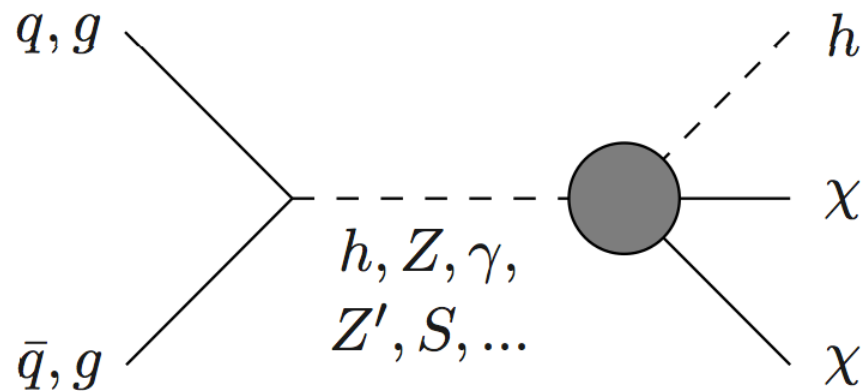


FIG. 1: Schematic diagram for mono-Higgs production in pp collisions mediated by electroweak bosons (h, Z, γ) or new mediator particles such as a Z' or scalar singlet S . The gray circle denotes an effective interaction between DM, the Higgs boson, and other states.

Models: EFT

$$\lambda |H|^2 |\chi|^2$$

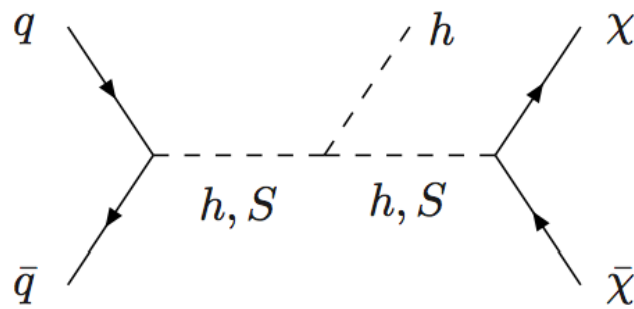
Scalar wimp

$$\frac{1}{\Lambda} |H|^2 \bar{\chi} \chi, \quad \frac{1}{\Lambda} |H|^2 \bar{\chi} i \gamma_5 \chi$$

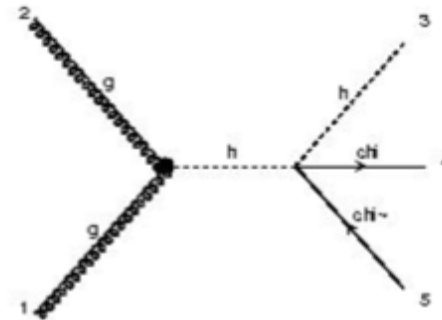
Fermion wimp

Vertices

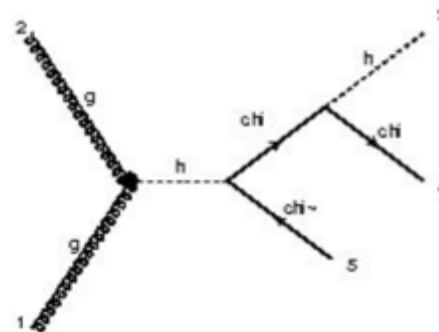
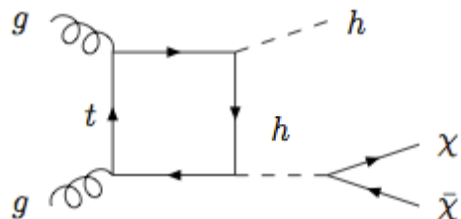
di-Higgs



4-point vertex



Off-shell s-channel Higgs



- (1) $h \rightarrow XX$ limited by invisible Higgs for $m_X < m_h/2$
- (2) For large coupling, $h \rightarrow XX$ grows, suppresses SM H decays!

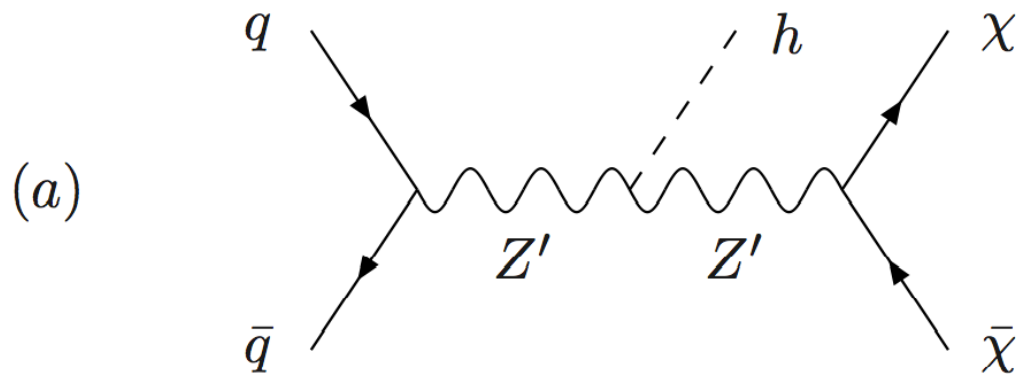
Other EFTs

Allow ZhXX-like vertices

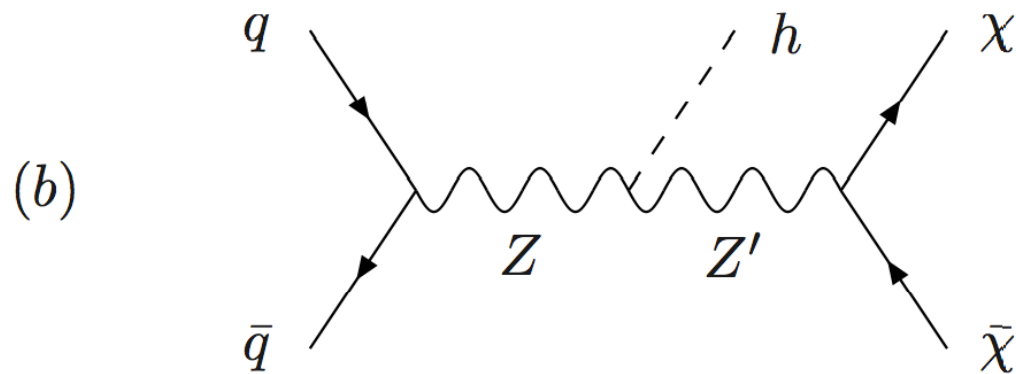
$$\frac{1}{\Lambda^2} \chi^\dagger i \overleftrightarrow{\partial}^\mu \chi H^\dagger i D_\mu H \quad \text{Scalar wimp}$$

$$\frac{1}{\Lambda^4} \bar{\chi} \gamma^\mu \chi B_{\mu\nu} H^\dagger D^\nu H. \quad \text{Fermion wimp}$$

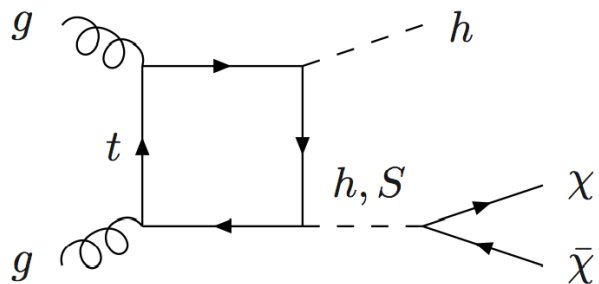
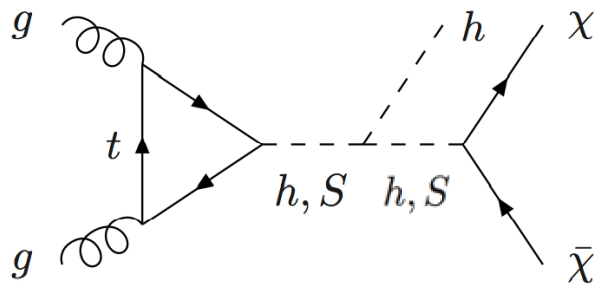
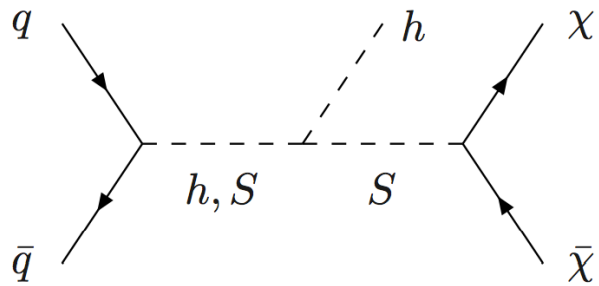
Simplified models: vector



with and
without
 Z - Z' mixing



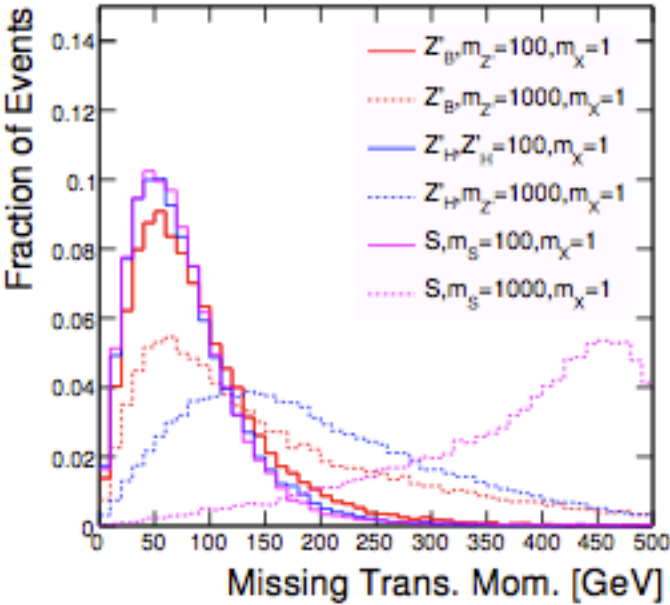
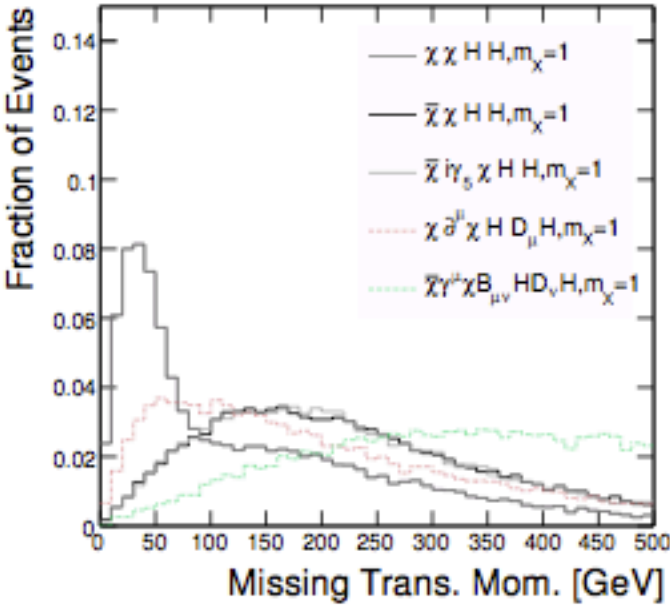
Simplified models: scalar



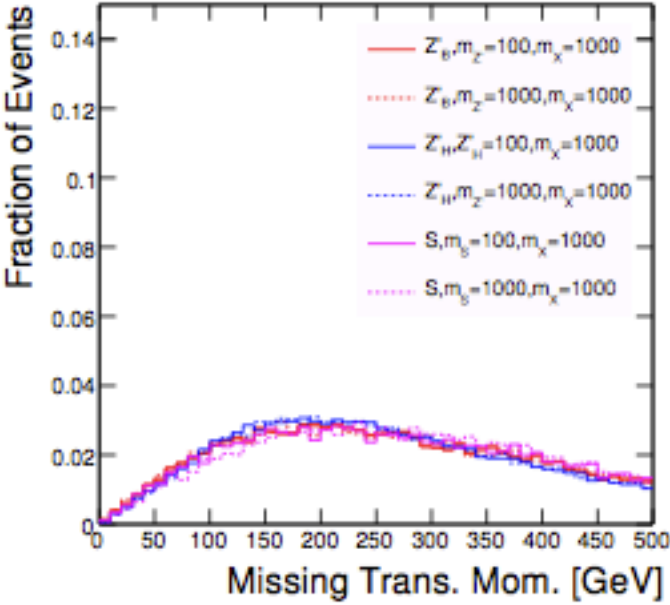
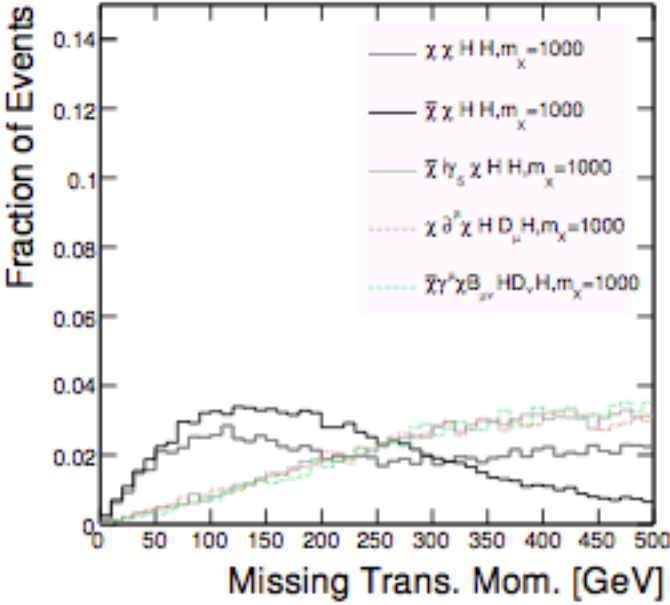
Box implemented as effective vertex in madgraph

MET

$m_\chi = 1$ GeV



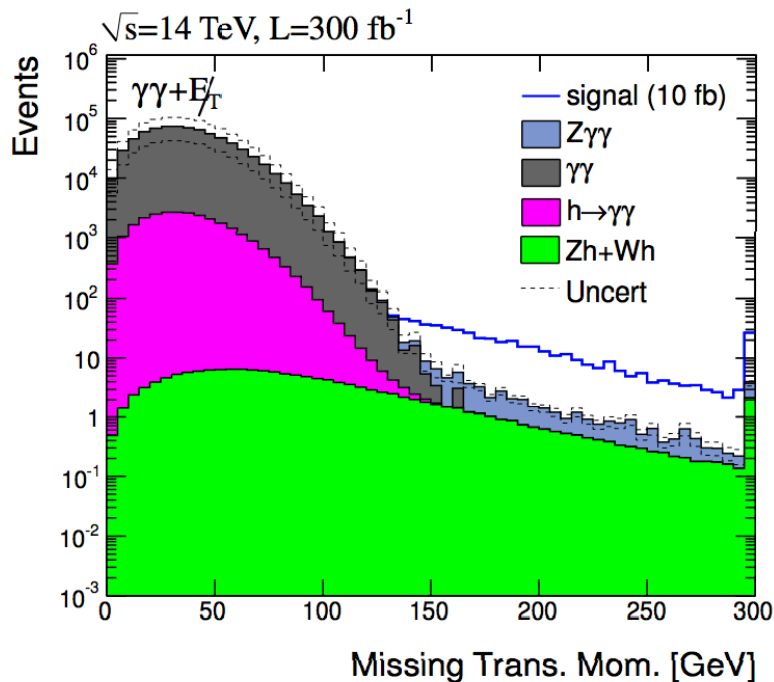
$m_\chi = 1$ TeV



EFTs

Simp. models

Gamma-gamma



Selection

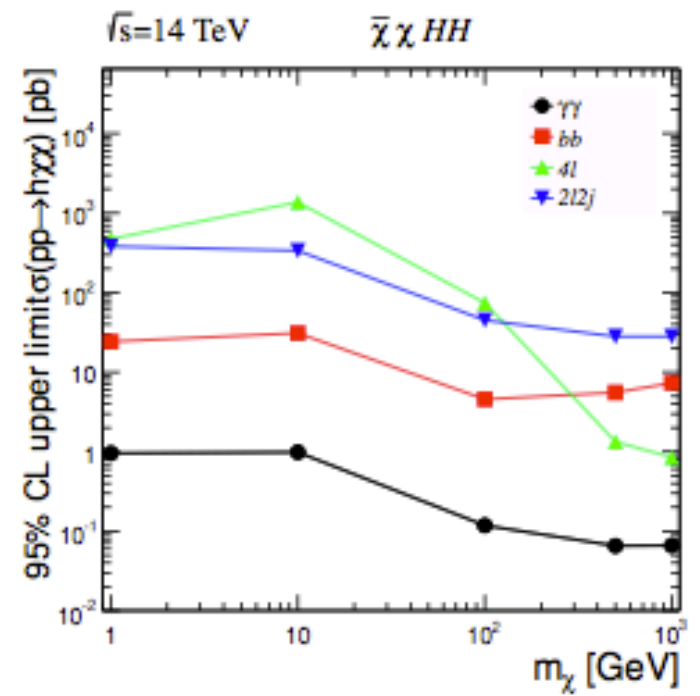
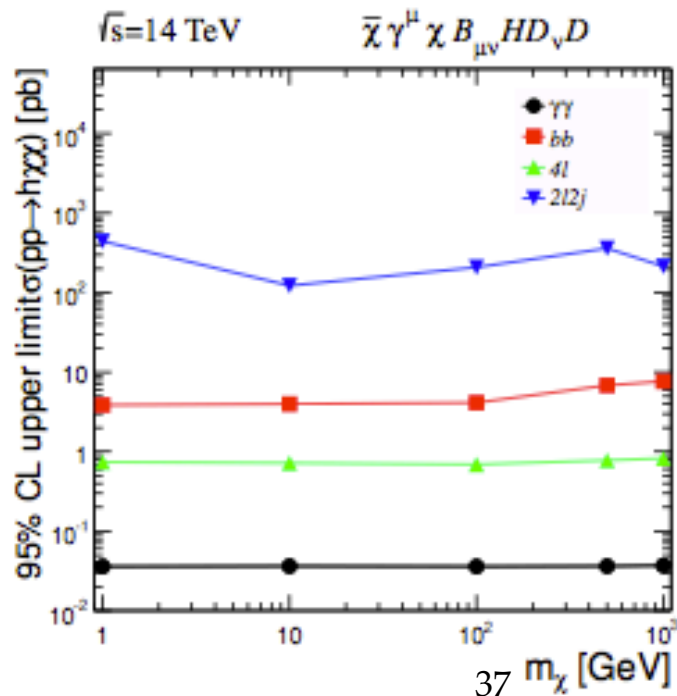
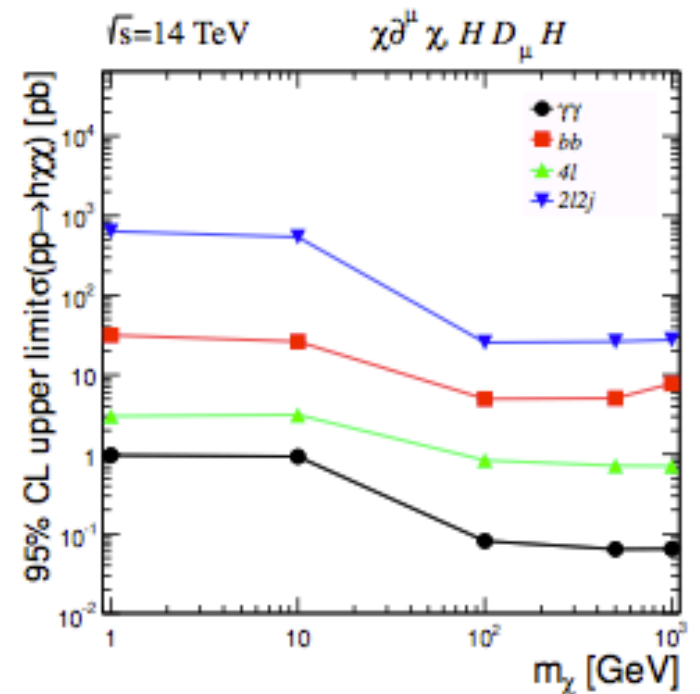
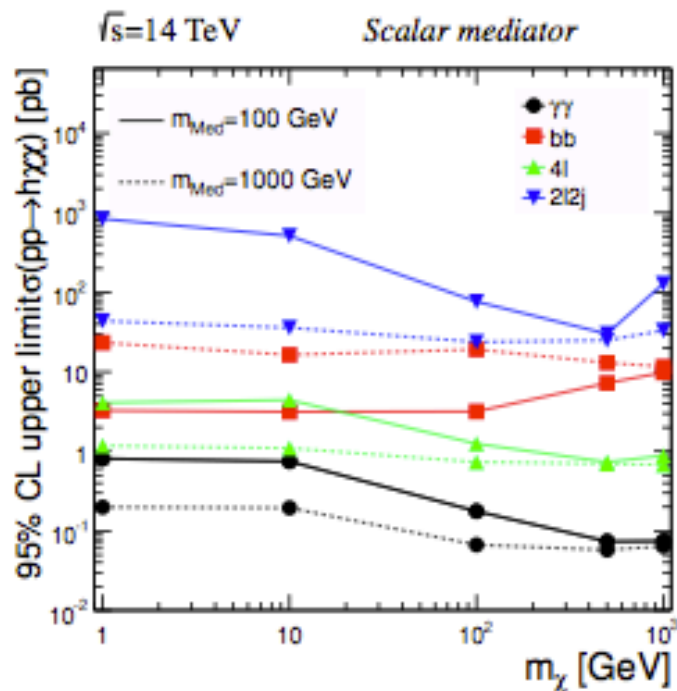
- two photons
- $m_{\gamma\gamma}$ in [110-130]
- $MET > 100, 250$ (8,14 TeV)

Backgrounds

- $h\rightarrow\gamma\gamma$ + fake MET
- $\gamma\gamma$ + fake MET
- $Z\gamma\gamma, Z\rightarrow\nu\nu$
- $Zh, Z\rightarrow\nu\nu$ + $Wh, W\rightarrow l\nu$

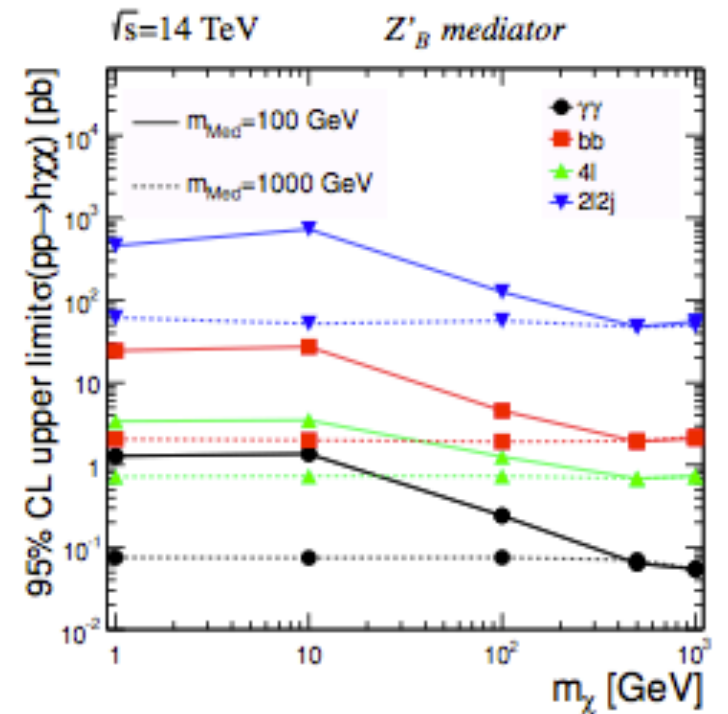
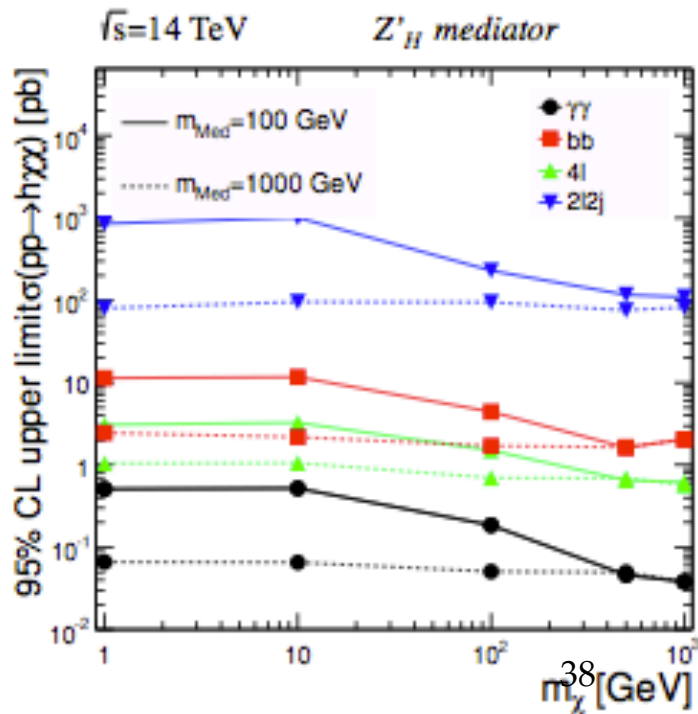
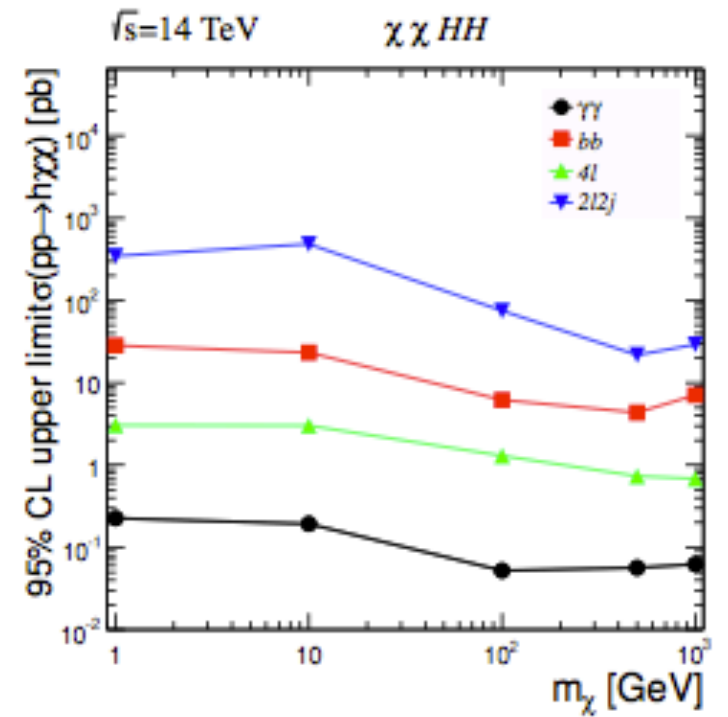
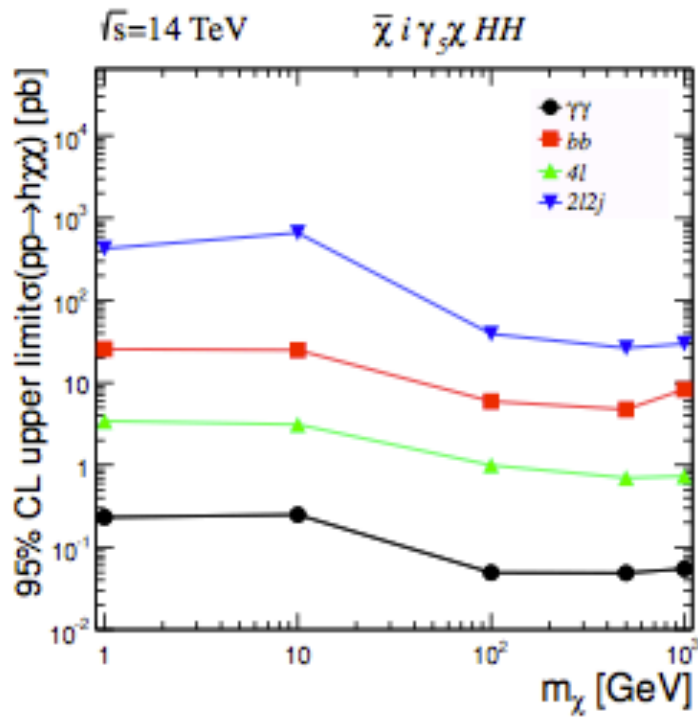


Assuming
 $h \rightarrow \text{SM}$
rates are
unchanged





Assuming
 $h \rightarrow \text{SM}$
rates are
unchanged



Parameter limits

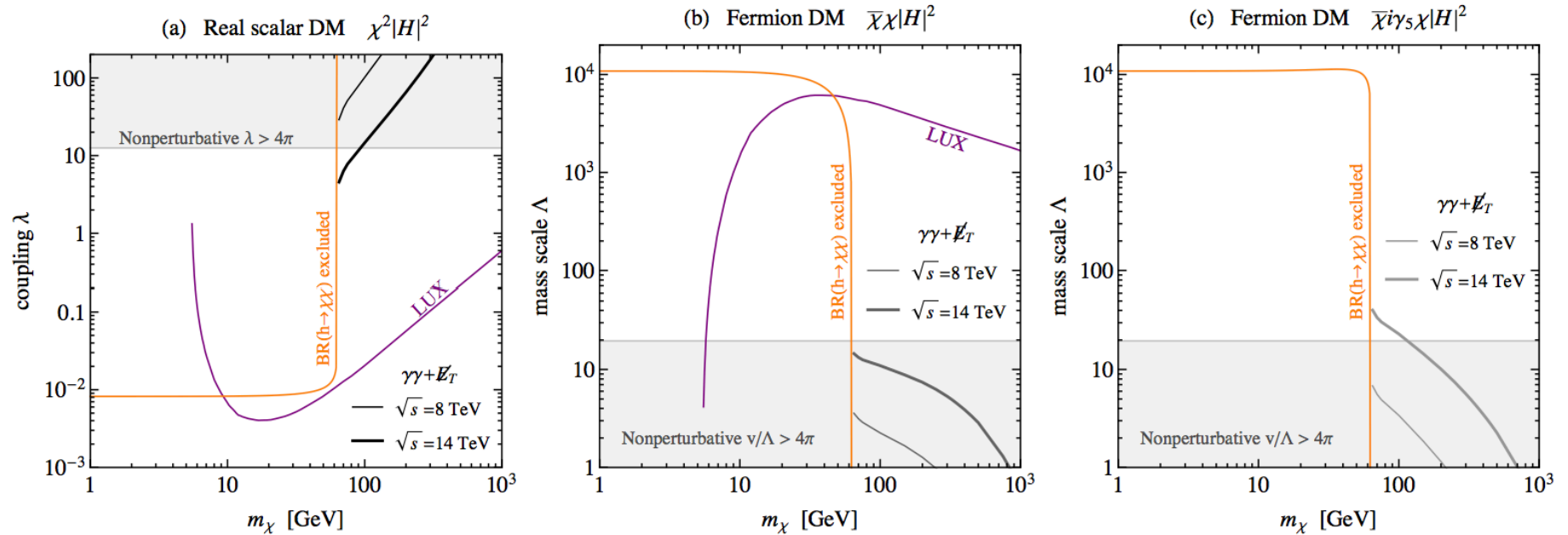


FIG. 20: Projected LHC mono-Higgs sensitivities at $\sqrt{s} = 8$ TeV (20 fb^{-1}) and 14 TeV (300 fb^{-1}), with $\gamma\gamma + \cancel{E}_T$ final states, on Higgs portal effective operators. All constraint contours exclude larger coupling λ or smaller mass scale Λ . Shaded region is excluded based on perturbativity arguments; orange contours denote limits from invisible h decays; purple contours are exclusion limits from LUX.

Note:

for $m_\chi < m_h/2$, no valid limits.

Large Lambda **boosts** $h \rightarrow \chi\chi$, **suppresses** $h \rightarrow \text{visible}$

Parameter limits

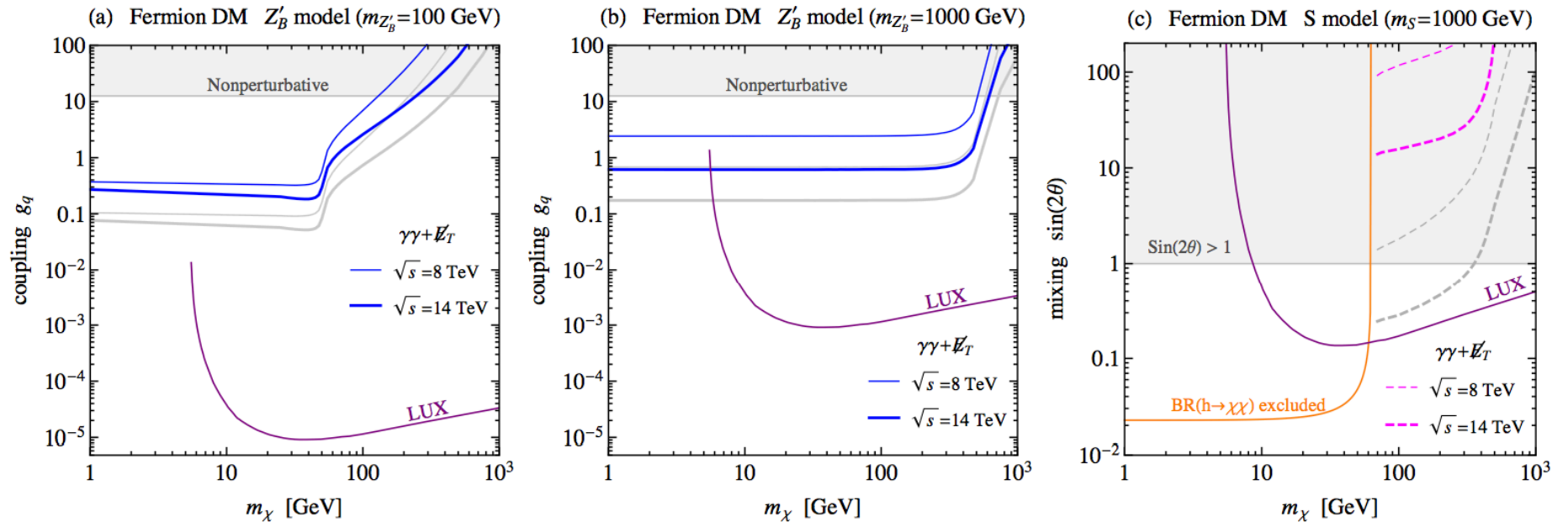


FIG. 22: Projected LHC mono-Higgs sensitivities at $\sqrt{s} = 8$ TeV (20 fb^{-1}) and 14 TeV (300 fb^{-1}), with $\gamma\gamma + \cancel{E}_T$ final states, on simplified models. All constraint contours exclude larger couplings or mixing angles. Shaded region is excluded based on perturbativity arguments or requiring $\sin\theta \leq 1$; orange contour denotes limit from invisible h decays; purple contours are exclusion limits from LUX.

DM References + Plans

ATLAS

7 TeV γ +MET (1209.4625)
W \rightarrow jj +MET (1309.4017)
Invisible Higgs (1402.3244)
Z+MET (1404.0051)

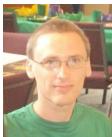
W \rightarrow lv +MET (soon)

VBF Invisible Higgs (forthcoming)

8 TeV γ +MET (forthcoming)

dijets (forthcoming)

Higgs+MET (forthcoming)

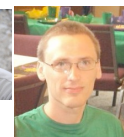
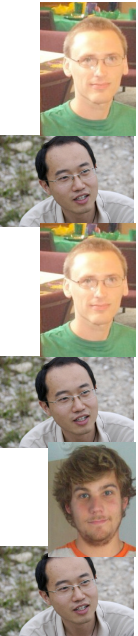


Pheno

monoZ (1212.3352)
DM combo (1302.3619)
Fermi/LHC (1307.5064)
DM future (1307.5327)
H+MET (1312.2592)
Indirect WW (1403.6734)

Compressed spectra (forthcoming)

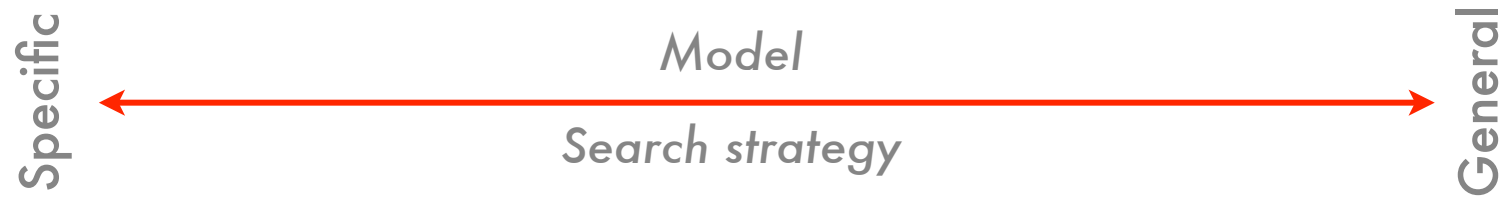
mono-Z' (forthcoming)



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Searching for new physics

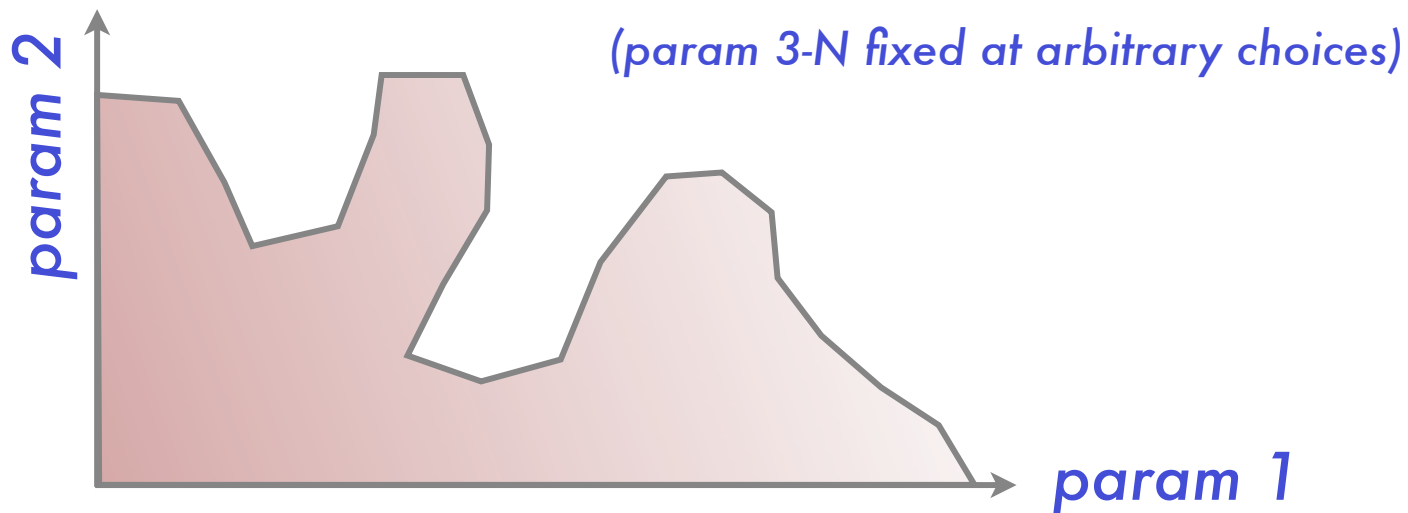


Traditional approach

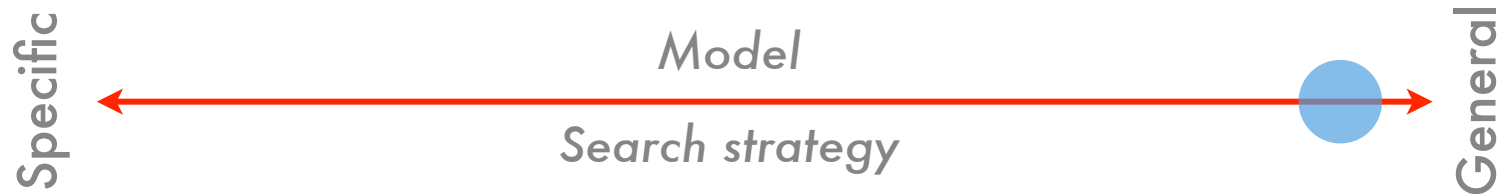


Bet on a specific theory

Optimize analysis to squeeze out maximal sensitivity to new physics.

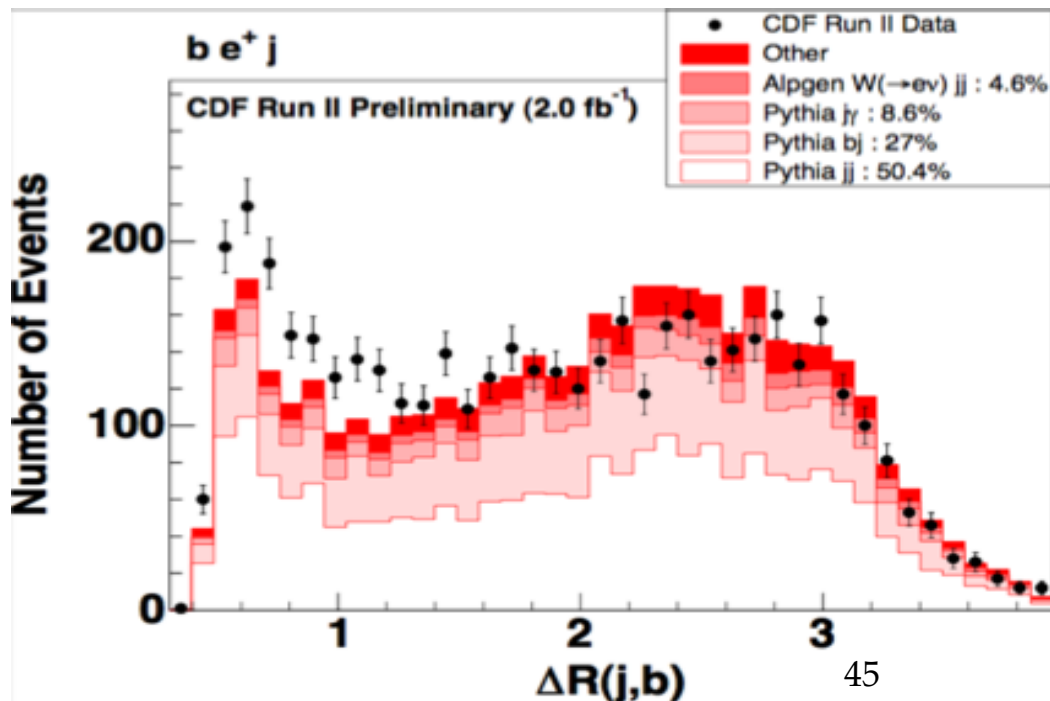


Model independent search



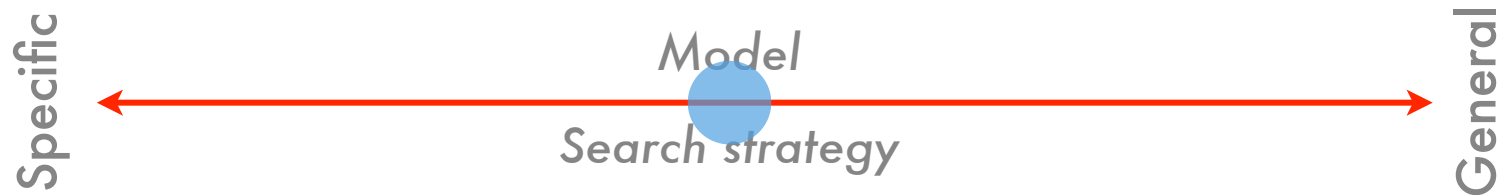
Discard the model

compare data to standard model



“Never listen to theorists.
Just go look for it.”
–A. Pierce, 2010

Compromise



Admit the need for a model

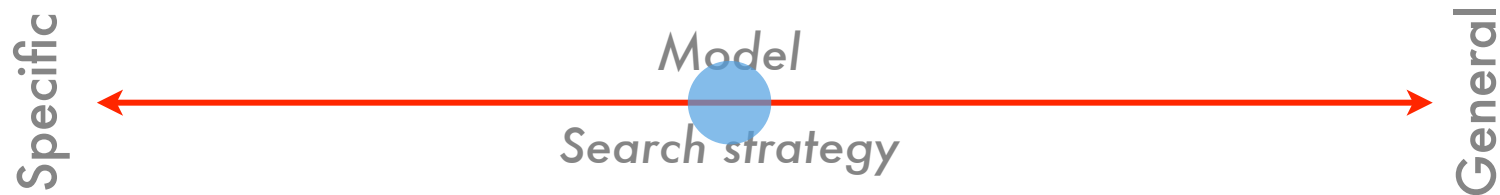
New signal requires a coherent physical explanation,
even trivial or effective

Generalize your model

Construct simple models that describe classes of new physics **which can be discovered at the LHC.**

What are we good at discovering?

Compromise



Admit the need for a model

New signal requires a coherent physical explanation,
even trivial or effective

Generalize your model

Construct simple models that describe classes of new physics **which can be discovered at the LHC.**

What are we good at discovering? **Resonances!**

Is this being done?

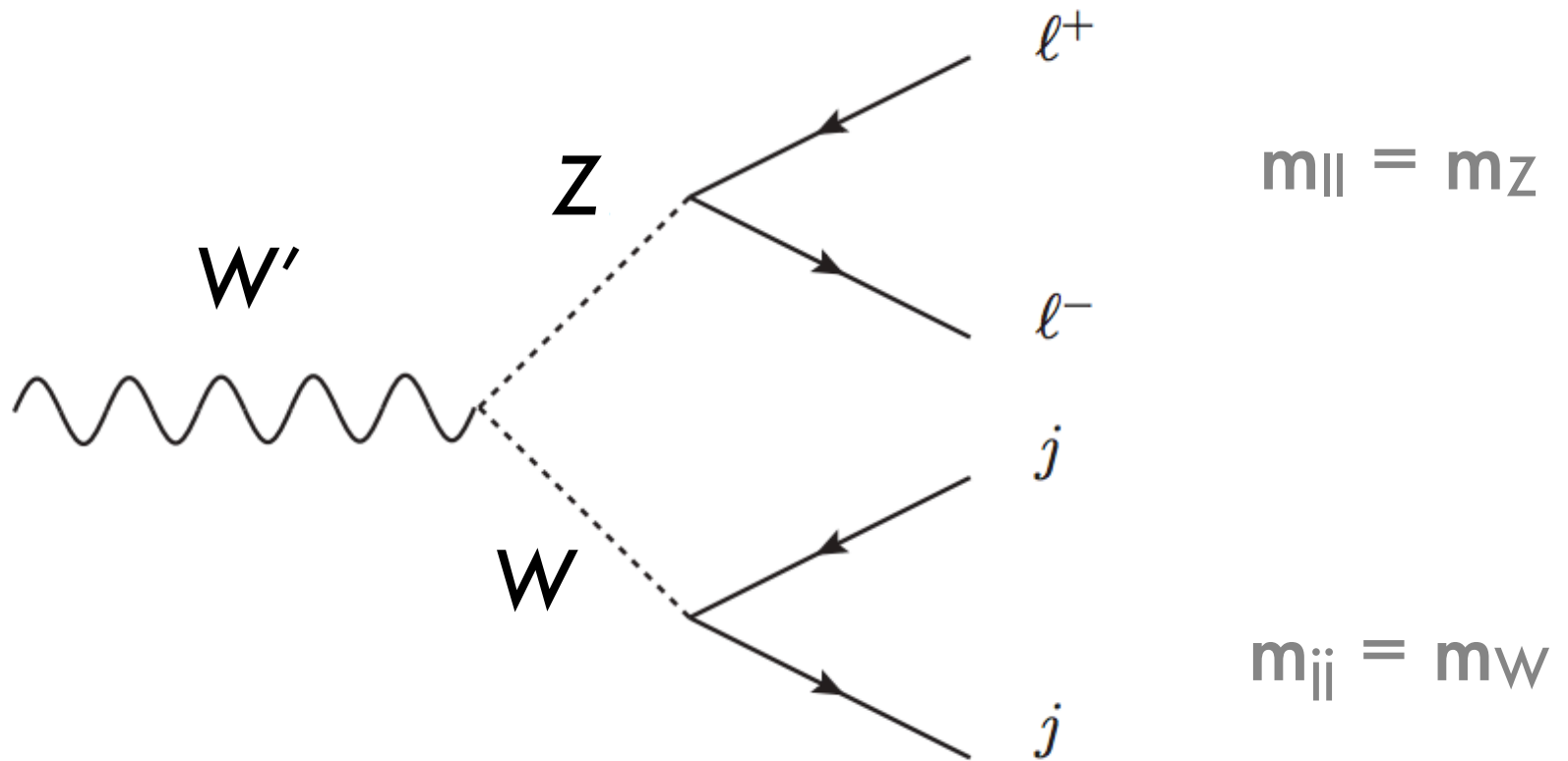
l^+

l^-

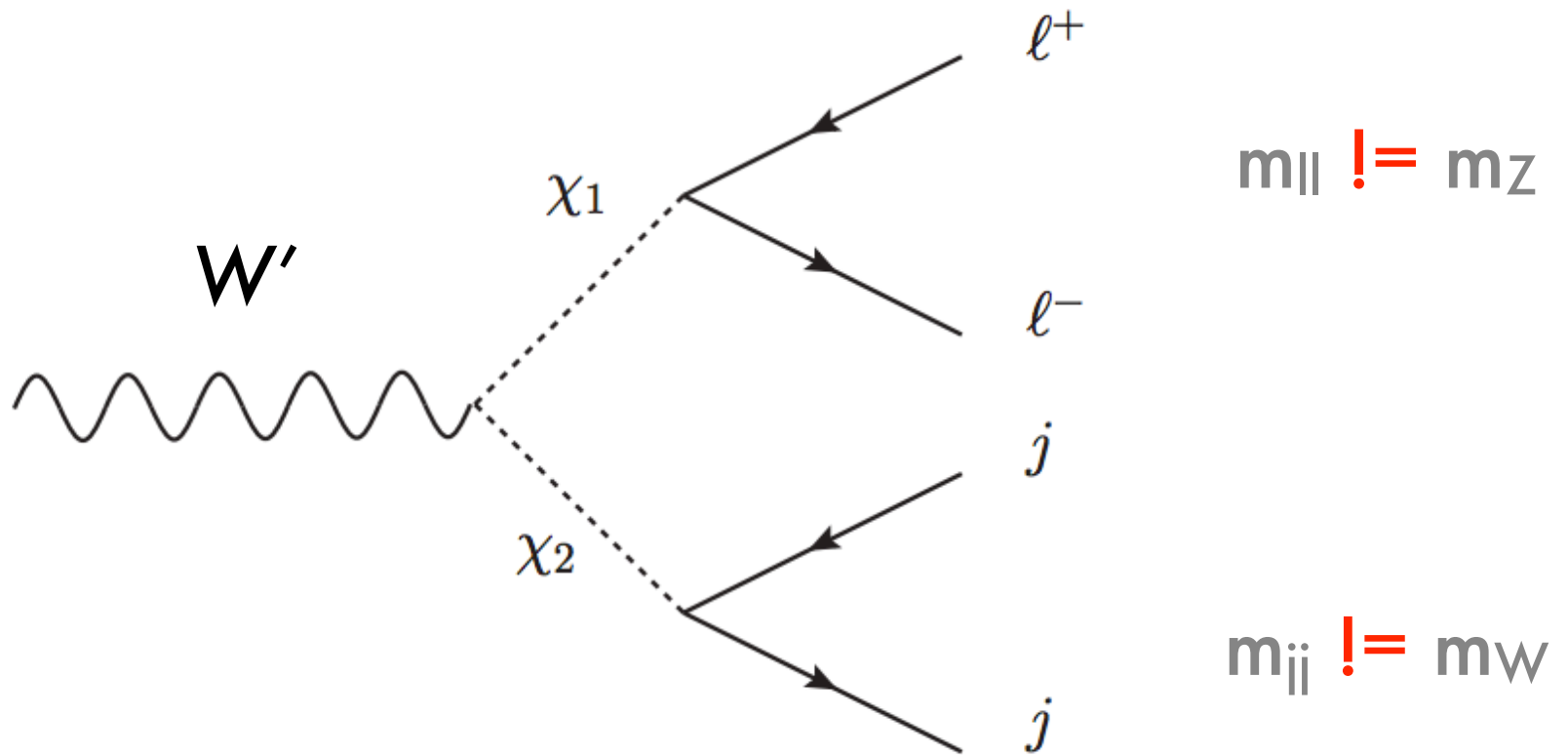
j

j

Is this being done?



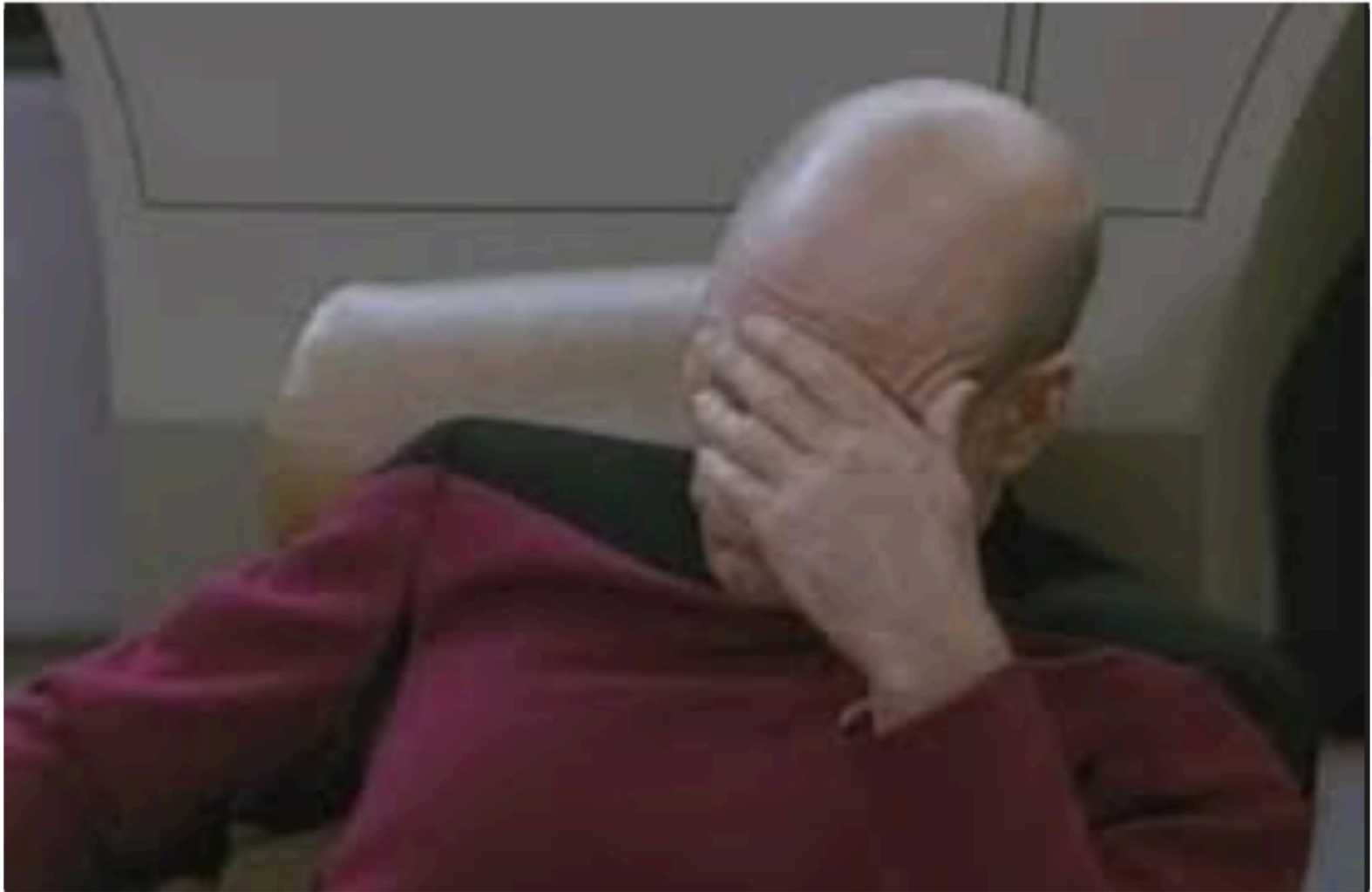
What about this?



Missed resonances?

Easy-to-find resonances
may exist in our data and
nobody has looked!

Missed resonances?



Topological models

UCI Physics 247
Final project
arXiv: 1401.1462

FERMILAB-PUB-13-529-T

Systematically Searching for New Resonances at the Energy Frontier using Topological Models

Mohammad Abdullah,¹ Eric Albin,¹ Anthony DiFranzo,¹ Meghan Frate,¹ Craig Pitcher,¹ Chase Shimmin,¹ Suneet Upadhyay,¹ James Walker,¹ Pierce Weatherly,¹ Patrick J. Fox,² and Daniel Whiteson¹

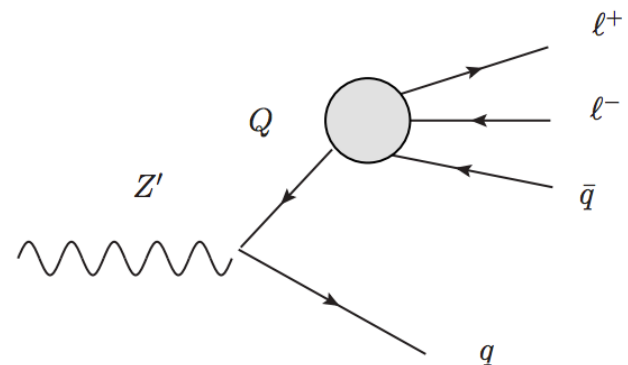
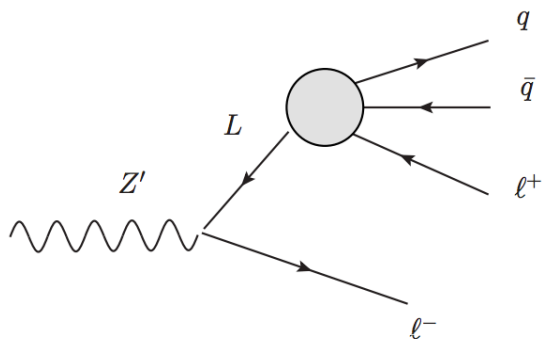
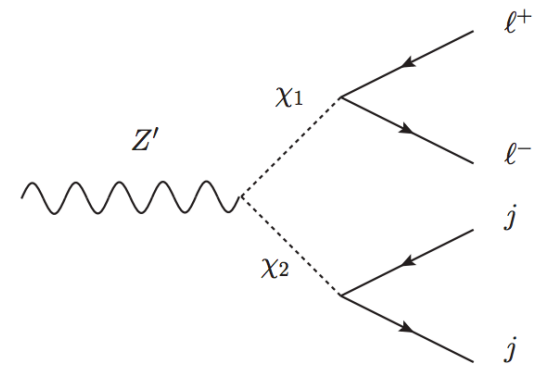
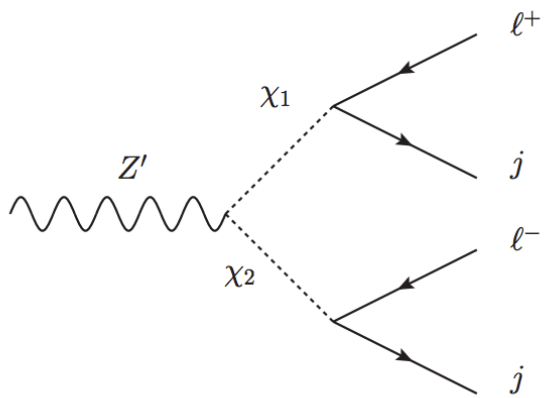
¹*Department of Physics and Astronomy, University of California, Irvine, CA 92697*

²*Fermi National Accelerator Laboratory, Batavia, IL 60615*

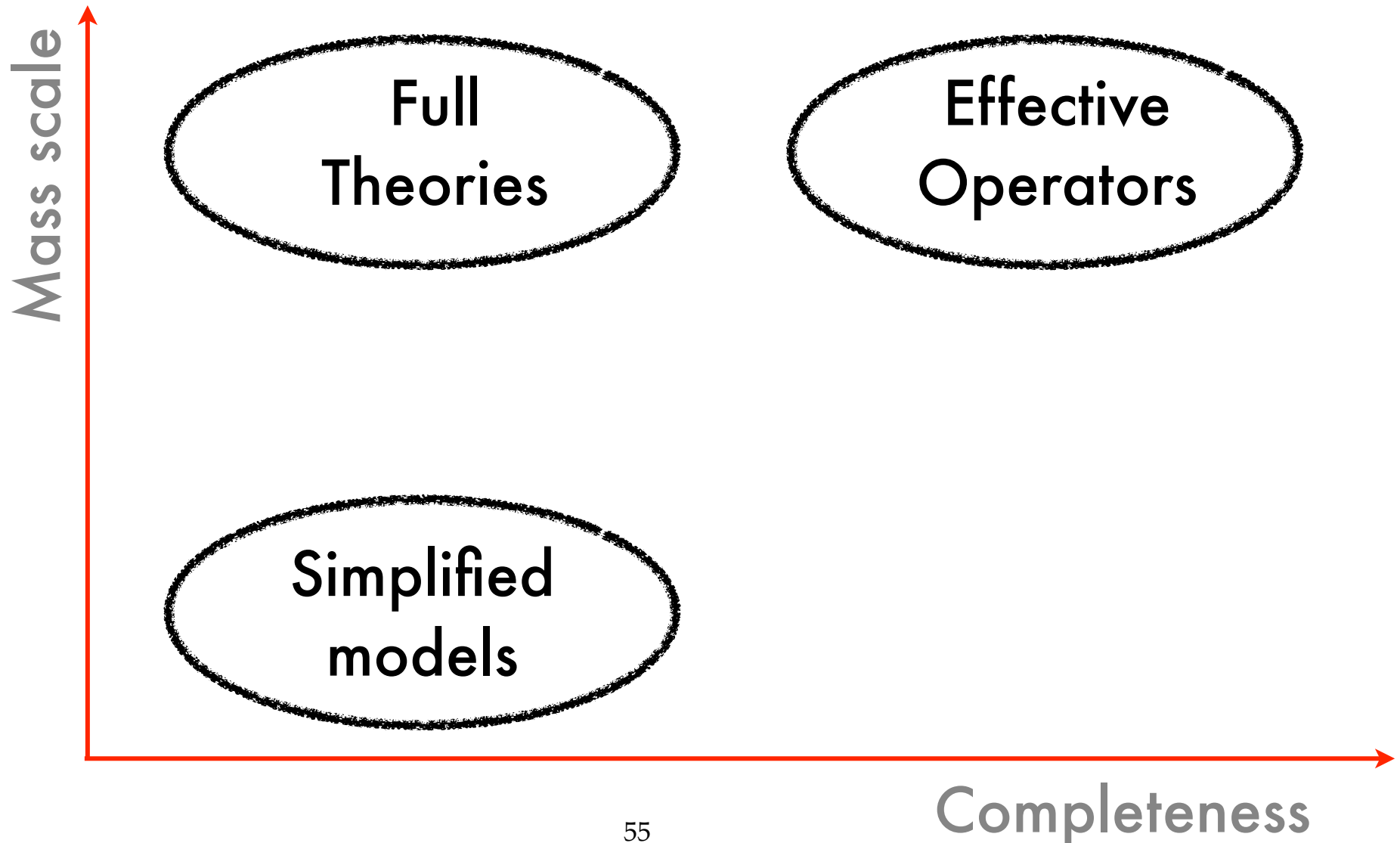
We propose a new strategy to systematically search for new physics processes in particle collisions at the energy frontier. An examination of all possible topologies which give identifiable resonant features in a specific final state leads to a tractable number of ‘topological models’ per final state and gives specific guidance for their discovery. Using one specific final state, $\ell\ell jj$, as an example, we find that the number of possibilities is reasonable and reveals simple, but as-yet-unexplored, topologies which contain significant discovery potential. We propose analysis techniques and estimate the sensitivity for pp collisions with $\sqrt{s} = 14$ TeV and $\mathcal{L} = 300$ fb⁻¹.

Topological models

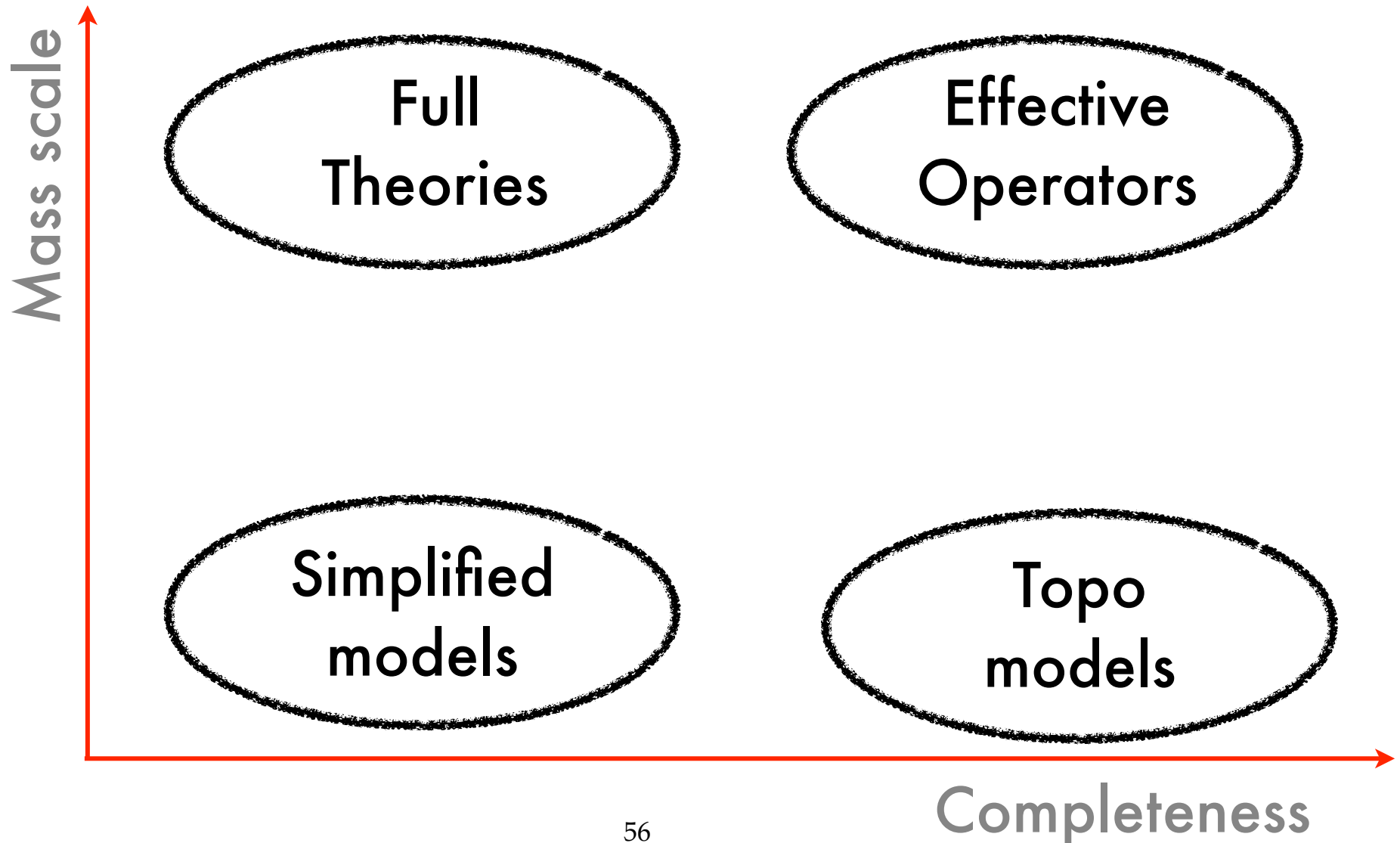
For a given final state (eg $lljj$) construct all models with resonances. Then look for them!



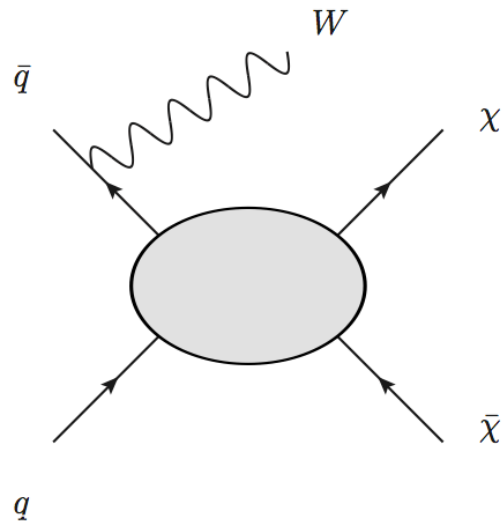
Connections to EFT, Simp. Models



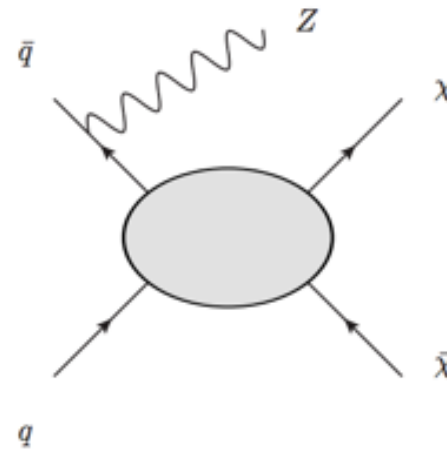
Connections to EFT, Simp. Models



Mono- Z'

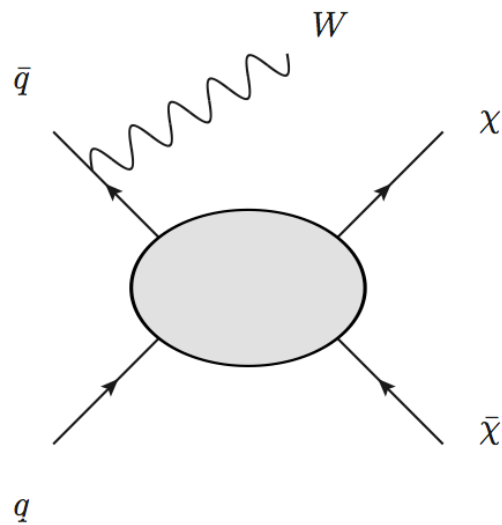


$$m_{ij} = m_W \text{ or } m_Z$$

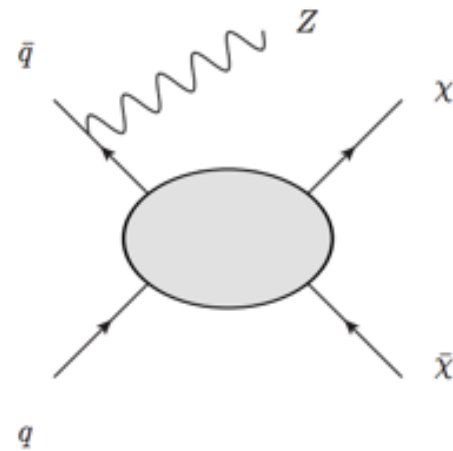


$$m_{ij} = m_Z$$

Mono-Z'



$$m_{ij} = m_W \text{ or } m_Z$$



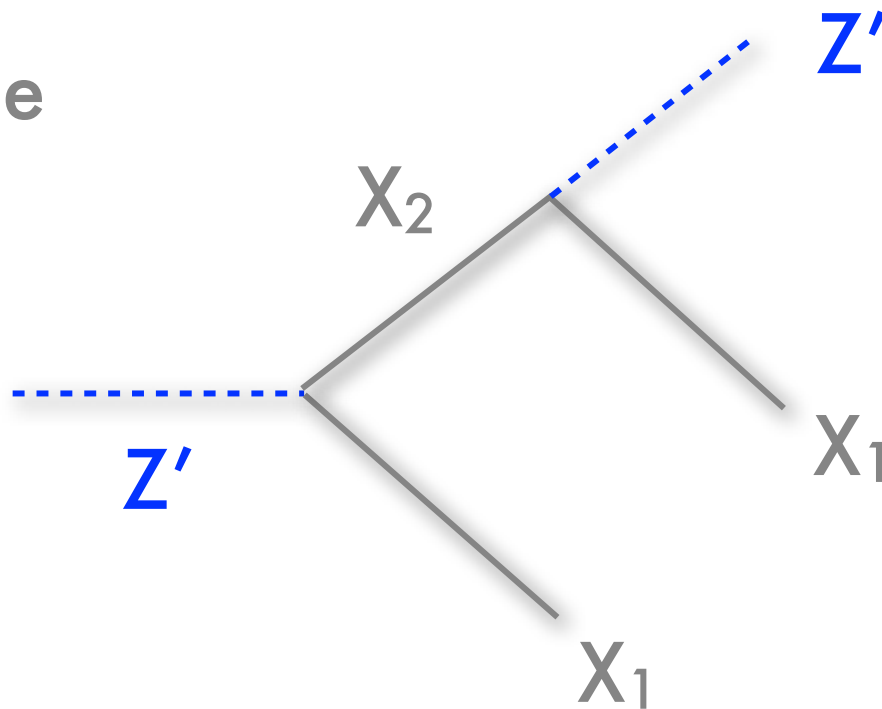
$$m_{ij} = m_Z$$

What about other values?

Mono-....

Signature

Heavy resonance
+ MET

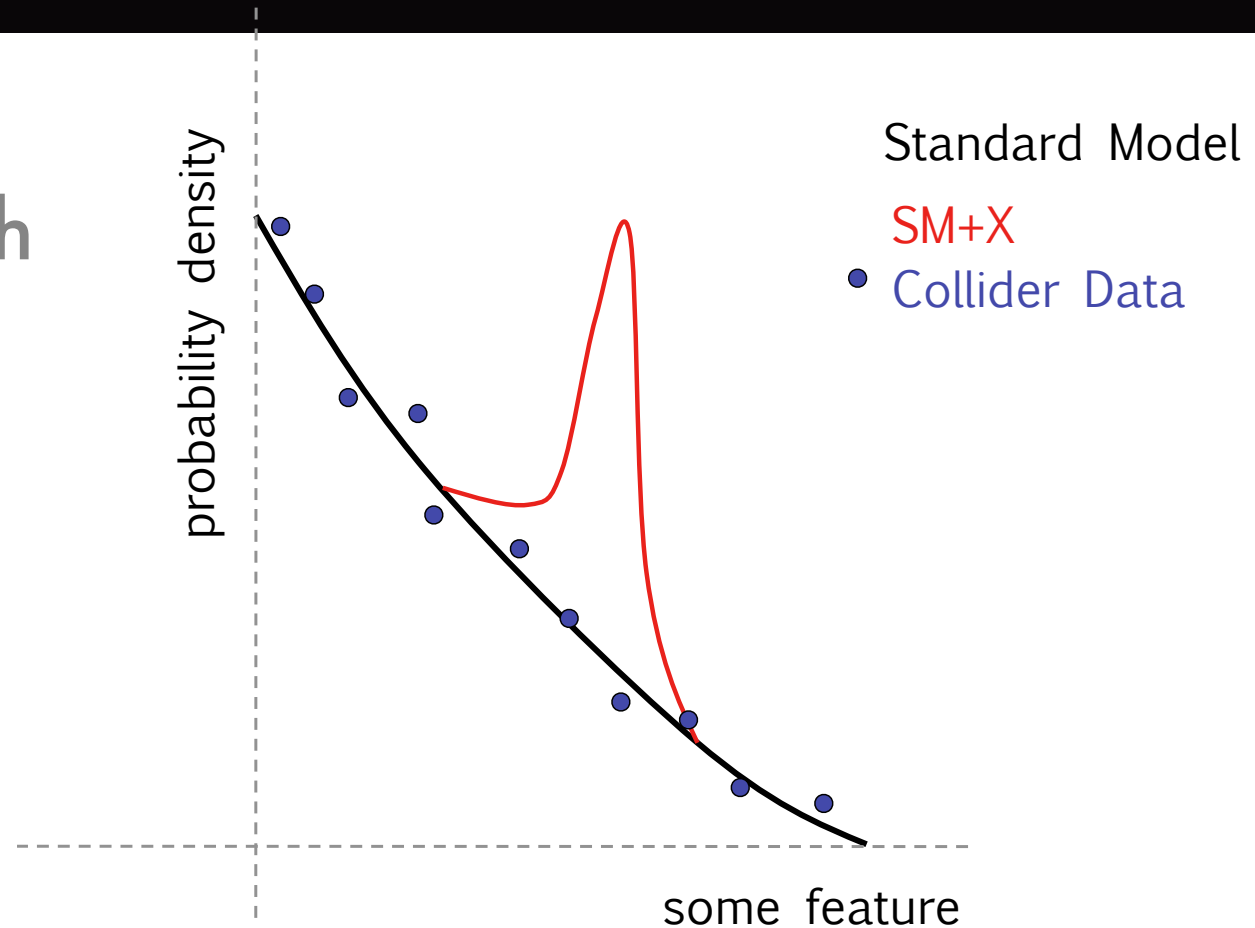


Outline

- I. Dark Matter
- II. Topological Models
- III. Deep networks

How to find NP

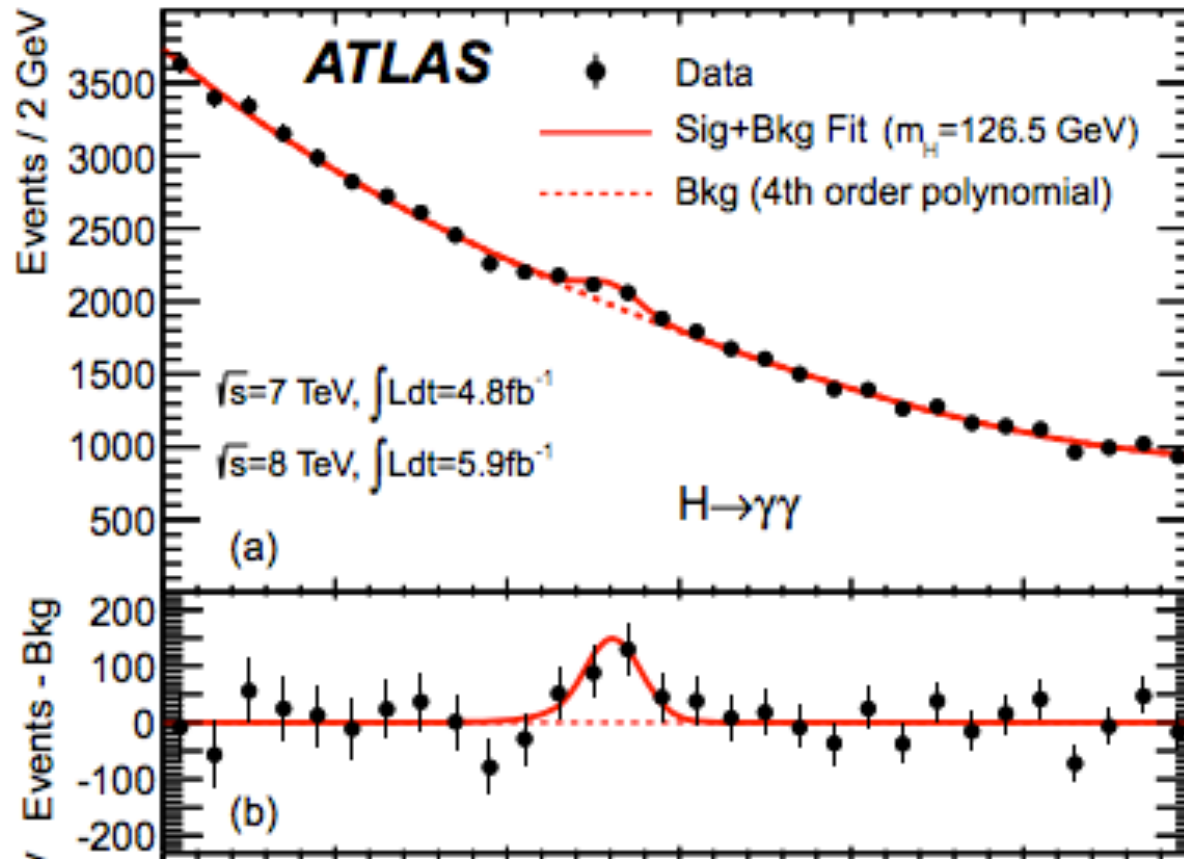
Isolate some feature in which two theories **SM**, **SM+X** can be best distinguished.



The data can tell us which hypothesis is preferred via a likelihood ratio:

$$\frac{L_{SM+X}}{L_{SM}} = \frac{P(\text{data} \mid \text{SM+X})}{P(\text{data} \mid \text{SM})}$$

e.g.

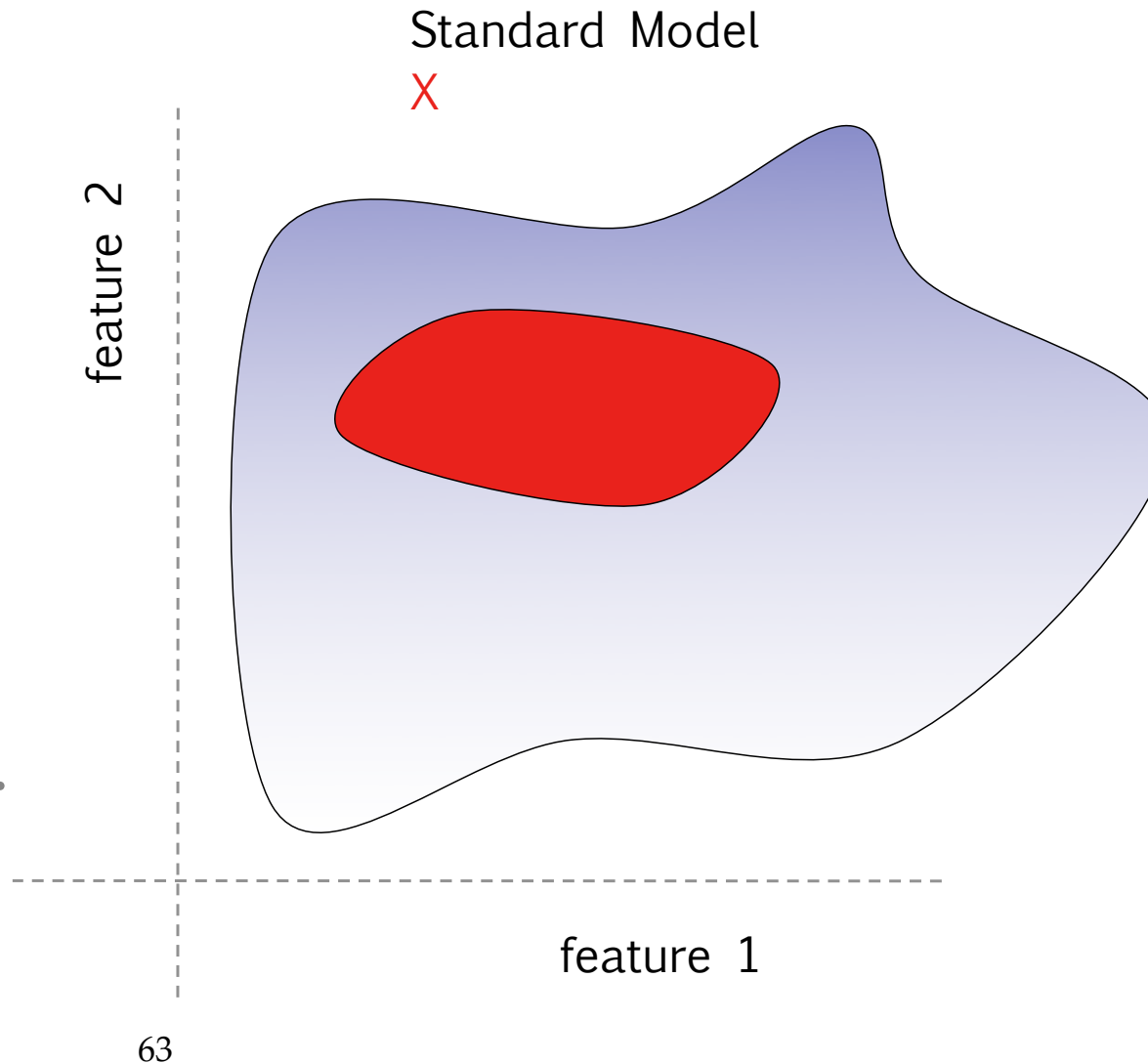


But...

Reality is more complicated.

The full space can be very high dimensional.

Calculating likelihood in d -dimensional space requires $\sim 100^d$ MC events.



ML tools

Standard Model

X

feature 2

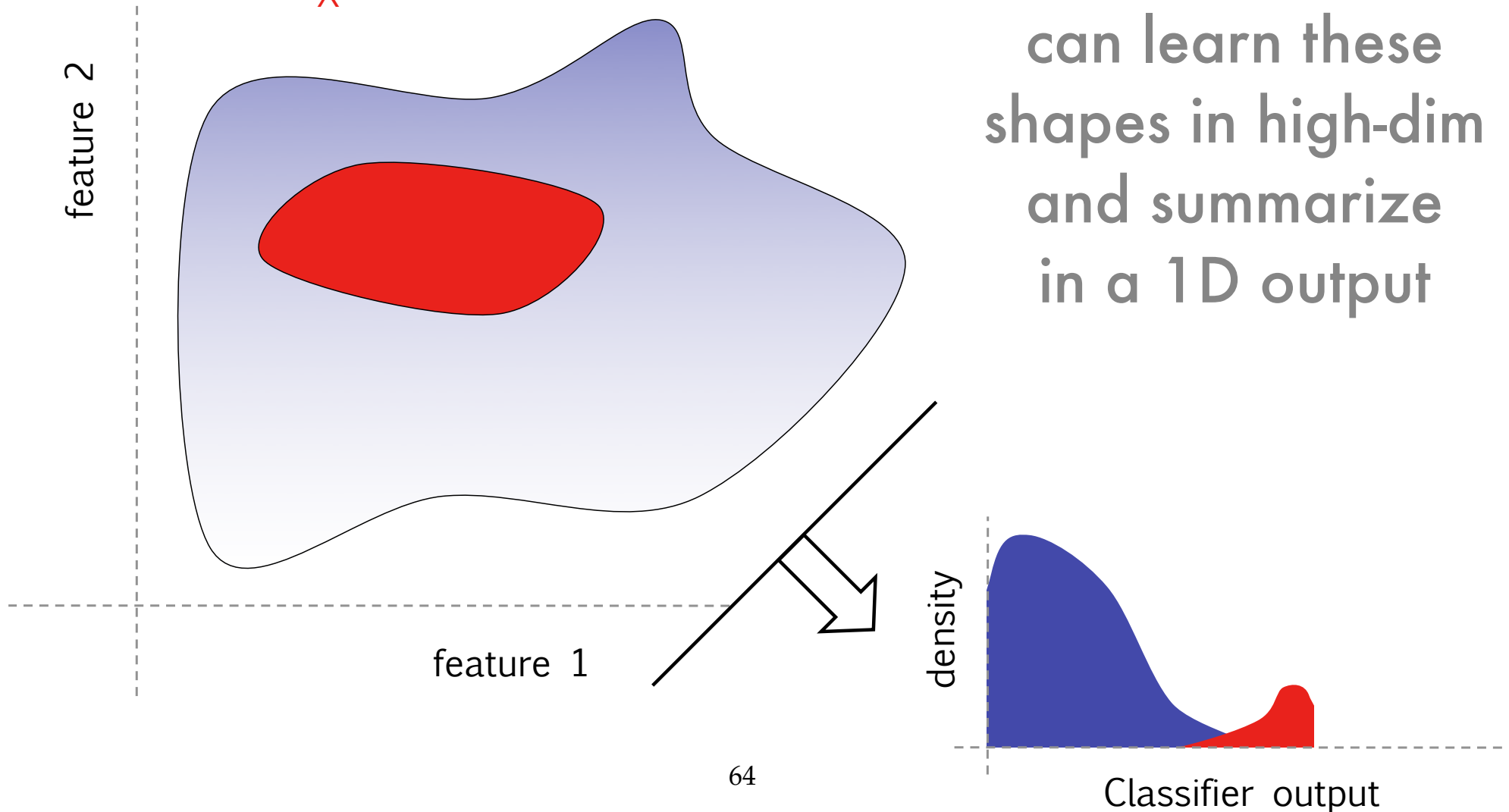
feature 1

64

density

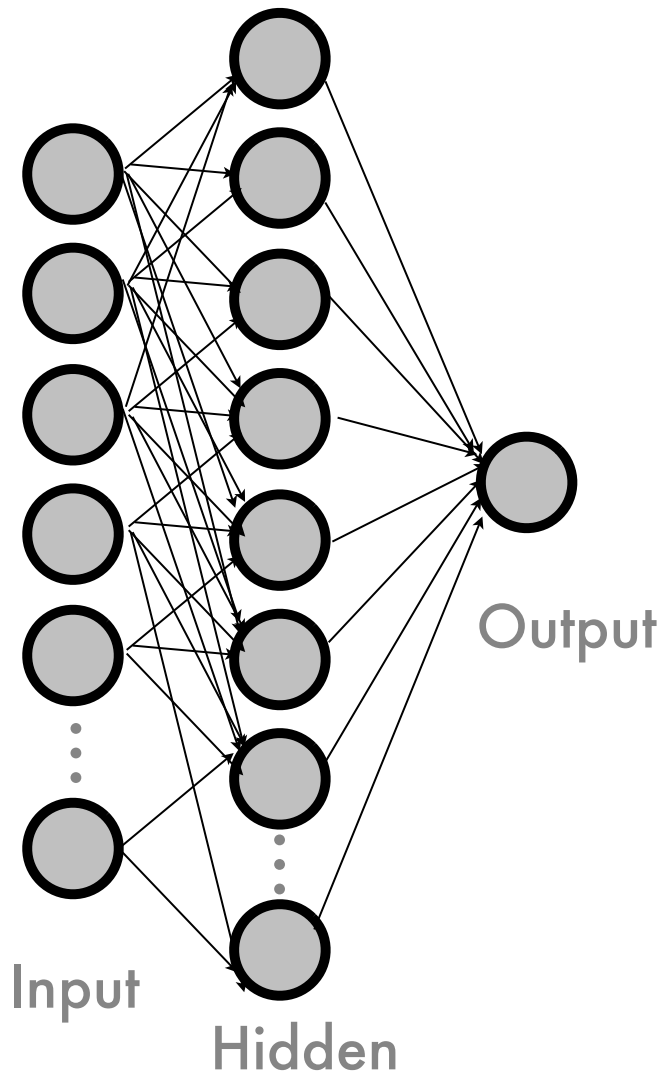
Classifier output

Neural networks
can learn these
shapes in high-dim
and summarize
in a 1D output



Neural Networks

Essentially a functional fit with many parameters



Function

Each neuron's output is a function of the weighted sum of inputs.

Goal

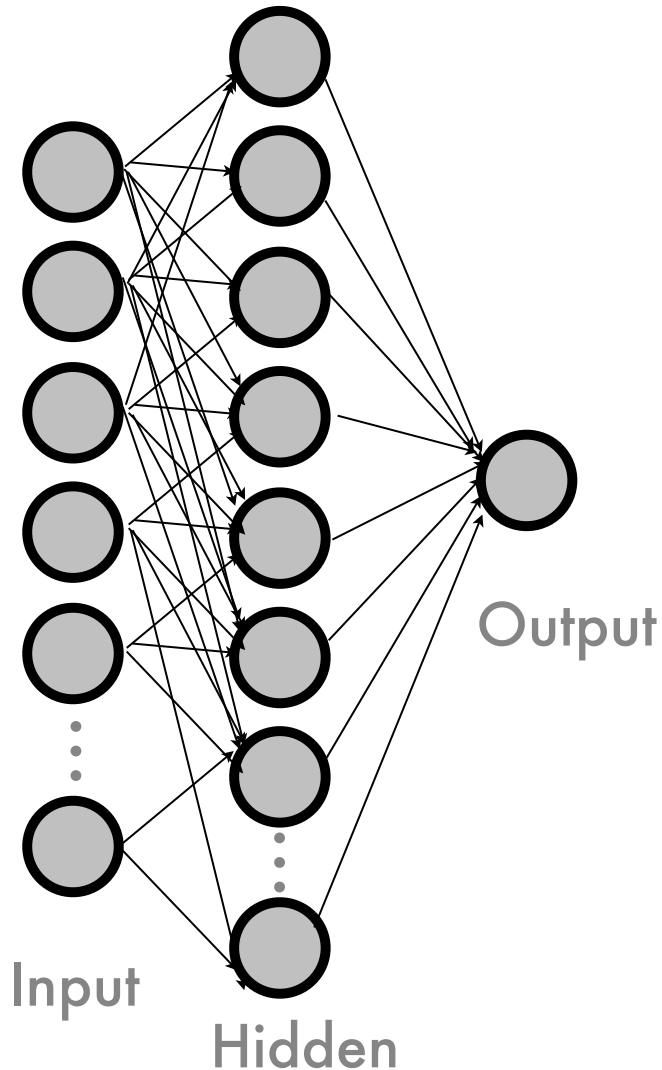
find set of weights which give most useful function

Learning

give examples, back-propagate error to adjust weights

Neural Networks

Essentially a functional fit with many parameters



Problem:

Networks with > 1 layer are very difficult to train.

Consequence:

Networks are not good at learning non-linear functions.
(like invariant masses!)

In short:

Can't just throw 4-vectors at NN.

Search for Input

ATLAS-CONF-2013-108

Can't just use $4v$

Can't give it too many inputs

Painstaking search through input feature space.

Variable	VBF			Boosted		
	$\tau_{lep}\tau_{lep}$	$\tau_{lep}\tau_{had}$	$\tau_{had}\tau_{had}$	$\tau_{lep}\tau_{lep}$	$\tau_{lep}\tau_{had}$	$\tau_{had}\tau_{had}$
$m_{\tau\tau}^{MMC}$	•	•	•	•	•	•
$\Delta R(\tau, \tau)$	•	•	•		•	•
$\Delta\eta(j_1, j_2)$	•	•	•			
m_{j_1, j_2}	•	•	•			
$\eta_{j_1} \times \eta_{j_2}$		•	•			
p_T^{total}		•	•			
sum p_T					•	•
$p_T(\tau_1)/p_T(\tau_2)$					•	•
$E_T^{miss} \phi$ centrality		•	•	•	•	•
$x_{\tau 1}$ and $x_{\tau 2}$						•
$m_{\tau\tau, j_1}$				•		
m_{ℓ_1, ℓ_2}				•		
$\Delta\phi_{\ell_1, \ell_2}$				•		
sphericity				•		
$p_T^{\ell_1}$				•		
$p_T^{j_1}$				•		
$E_T^{miss}/p_T^{\ell_2}$				•		
m_T		•			•	
$\min(\Delta\eta_{\ell_1, \ell_2, jets})$	•					
$j_3 \eta$ centrality	•					
$\ell_1 \times \ell_2 \eta$ centrality	•					
$\ell \eta$ centrality		•				
$\tau_{1,2} \eta$ centrality			•			

Table 3: Discriminating variables used for each channel and category. The filled circles identify which variables are used in each decay mode. Note that variables such as $\Delta R(\tau, \tau)$ are defined either between the two leptons, between the lepton and τ_{had} , or between the two τ_{had} candidates, depending on the decay mode.

Search for Input

ATLAS-CONF-2013-108

Can't just use 4v

Can't give it +
many inp

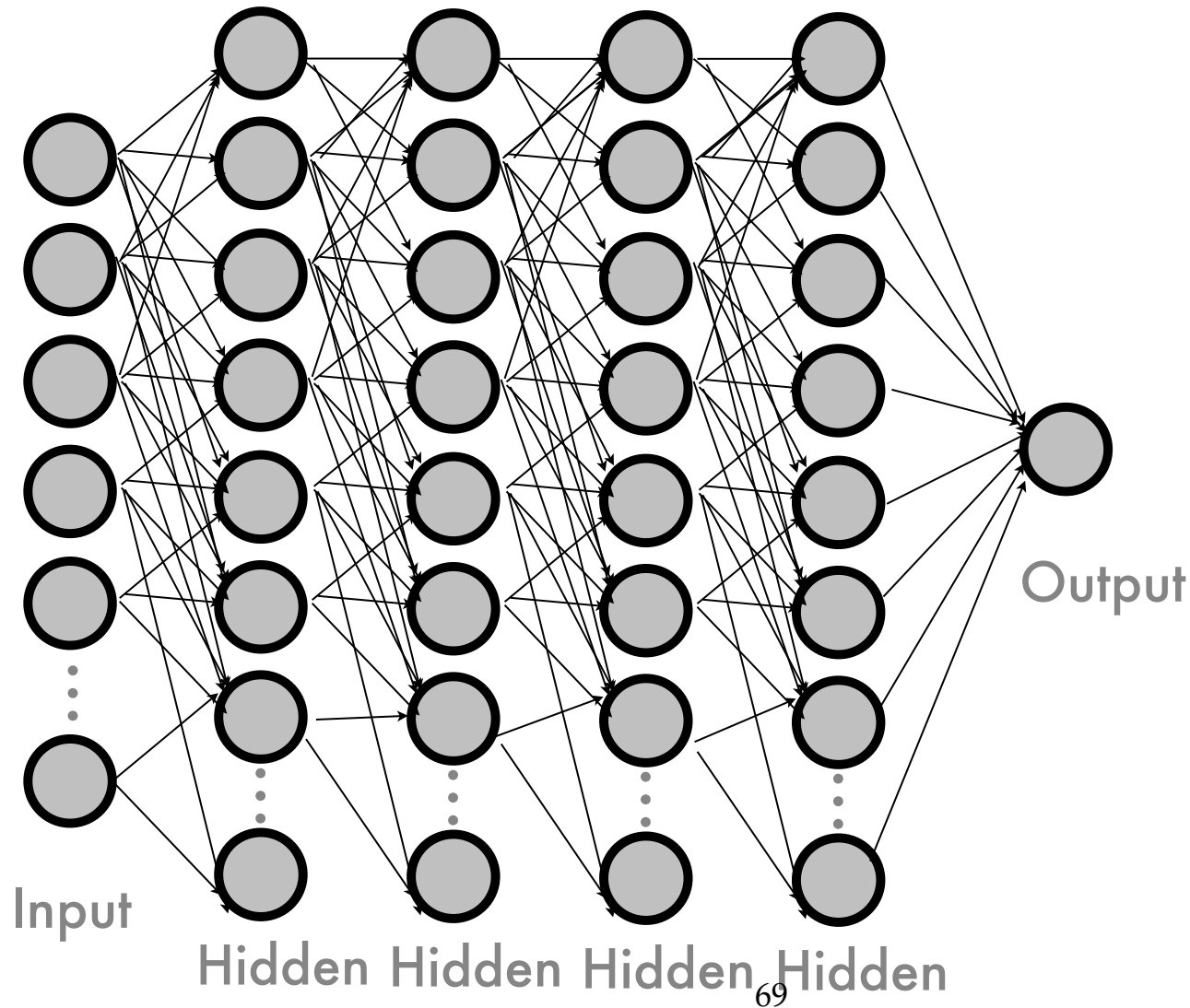
Painstaking
through inp
feature space.

**Also true for
BDTs, SVNs, etc**

Variable	VBF	Boosted	
	$\tau_{lep}\tau_{lep}$	$\tau_{lep}\tau_{had}$	$\tau_{had}\tau_{had}$
m_{TT}^{MMC}		•	•
$\Delta R(\tau, \tau)$		•	•
		•	•
		•	•
		•	•
		•	•
		•	•
		•	•
P_T^{j1}		•	
$E_T^{miss}/P_T^{\ell 2}$		•	
m_T	•		•
$\min(\Delta\eta_{\ell_1, \ell_2, jets})$	•		
$j_3 \eta$ centrality	•		
$\ell_1 \times \ell_2 \eta$ centrality	•		
$\ell \eta$ centrality	•		
$\tau_{1,2} \eta$ centrality		•	

Table 3: Discriminating variables used for each channel and category. The filled circles identify which variables are used in each decay mode. Note that variables such as $\Delta R(\tau, \tau)$ are defined either between the two leptons, between the lepton and τ_{had} , or between the two τ_{had} candidates, depending on the decay mode.

Deep networks



New tools
let us
train
deep
networks.

How well
do they work?

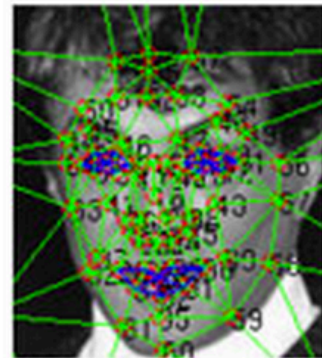
Real world applications



(a)



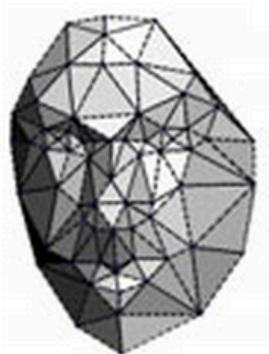
(b)



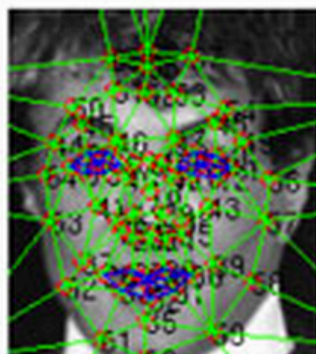
(c)



(d)



(e)



(f)



(g)



(h)

Head turn: DeepFace uses a 3-D model to rotate faces, virtually, so that they face the camera. Image (a) shows the original image, and (g) shows the final, corrected version.

Paper

Deep Learning in High-Energy Physics: Improving the Search for Exotic Particles

P. Baldi,¹ P. Sadowski,¹ and D. Whiteson²

¹*Dept. of Computer Science, UC Irvine, Irvine, CA 92617*

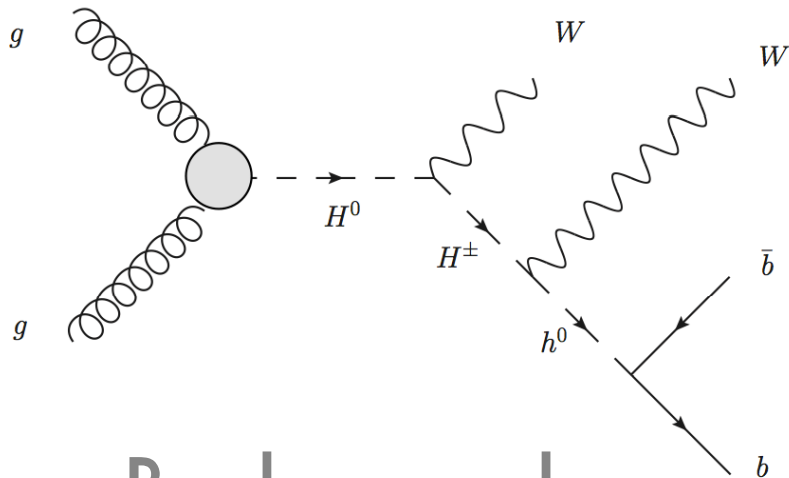
²*Dept. of Physics and Astronomy, UC Irvine, Irvine, CA 92617*

arXiv: 1402.4735

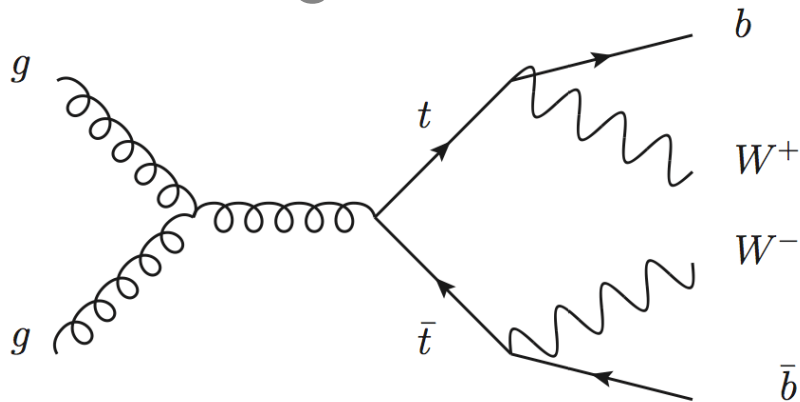
In revision at *Nature Comm.*

Benchmark problem

Signal



Background



Can deep networks automatically discover useful variables?

4-vector inputs

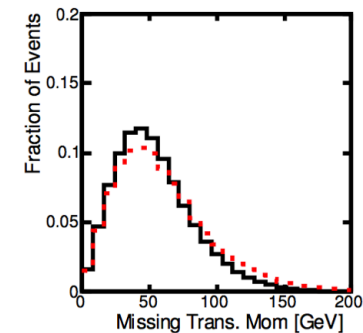
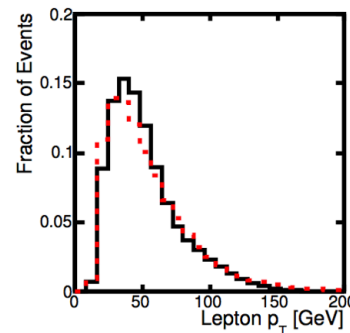
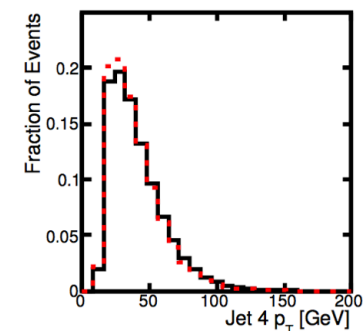
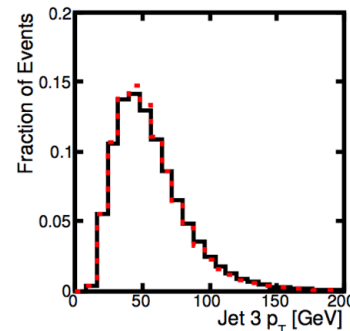
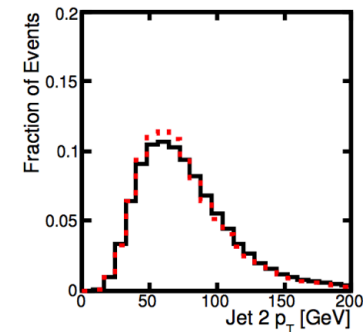
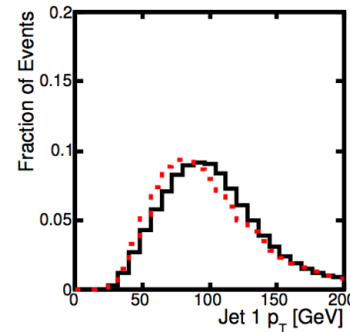
21 Low-level vars

jet+lepton mom. (3x5)

missing ET (2)

jet btags (4)

Not much
separation
visible in 1D
projections



4-vector inputs

7 High-level vars

$m(WWbb)$

$m(Wbb)$

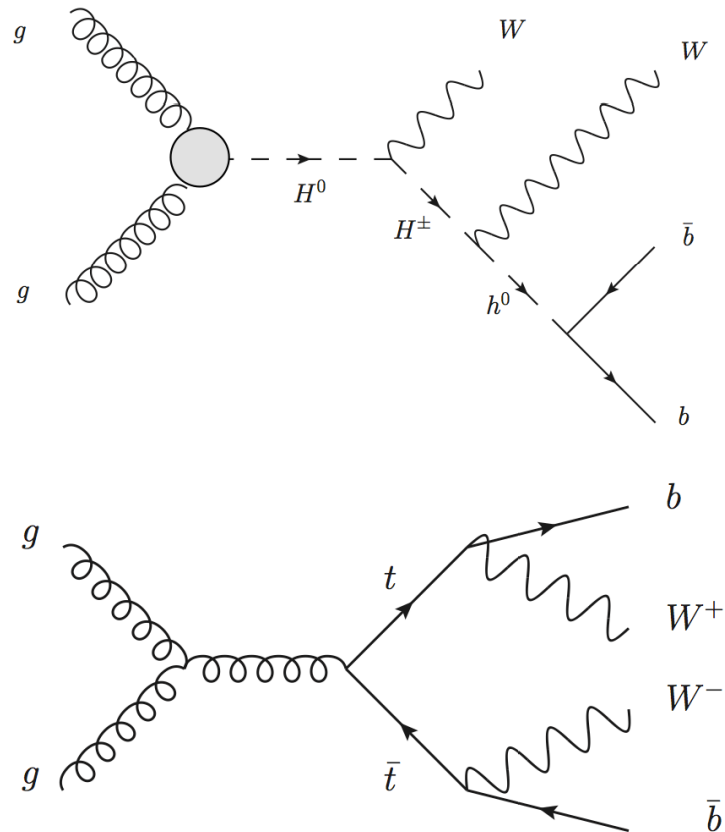
$m(bb)$

$m(bjj)$

$m(jj)$

$m(lv)$

$m(blv)$



4-vector inputs

7 High-level vars

$m(WWbb)$

$m(Wbb)$

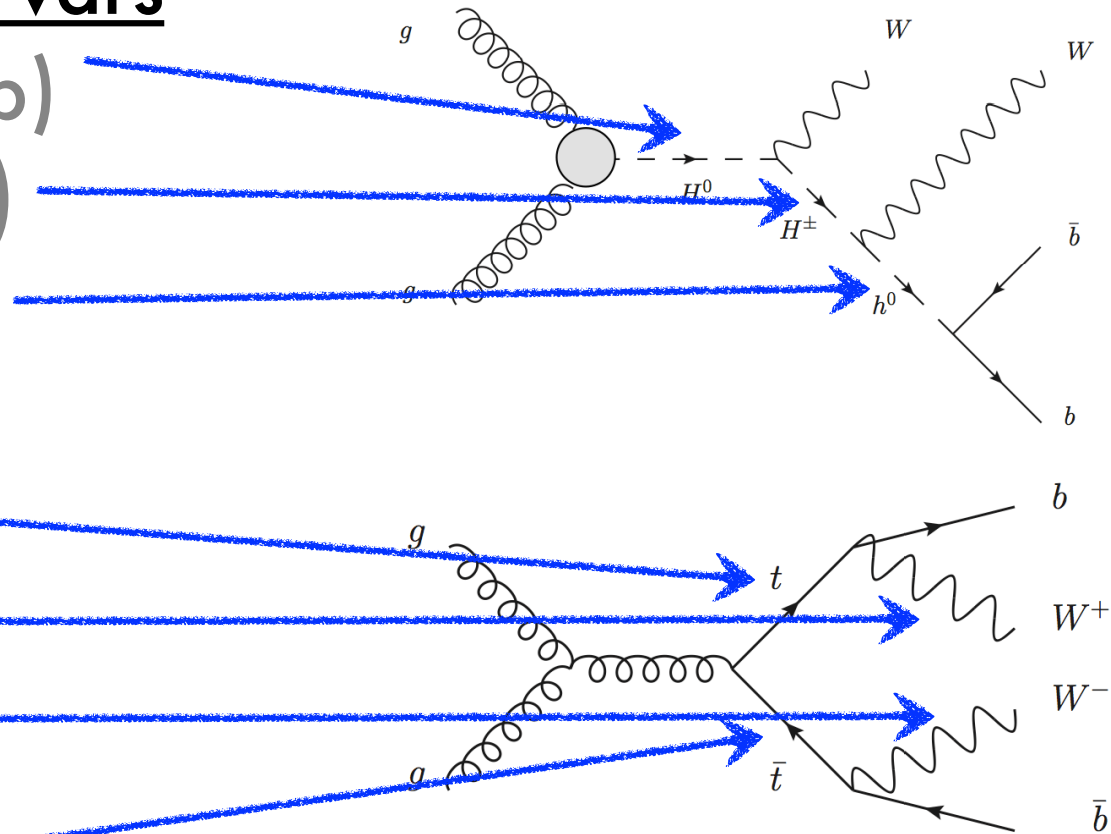
$m(bb)$

$m(bjj)$

$m(jj)$

$m(lv)$

$m(blv)$



4-vector inputs

7 High-level vars

$m(WWbb)$

$m(Wbb)$

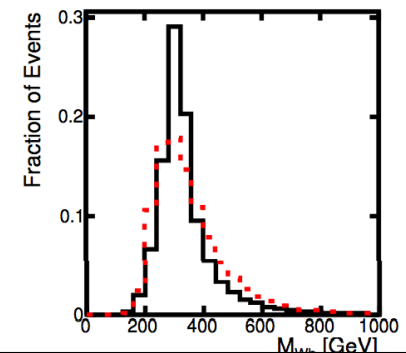
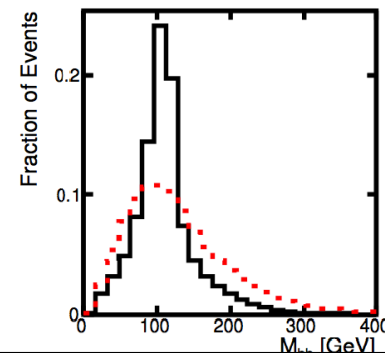
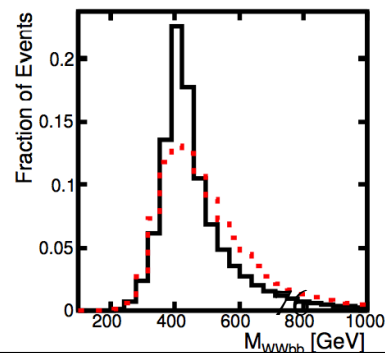
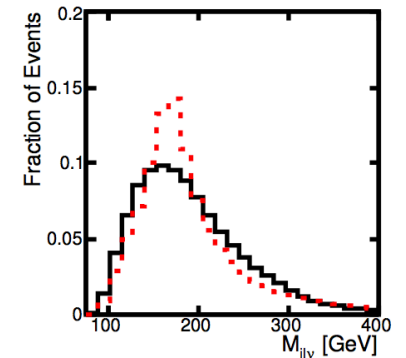
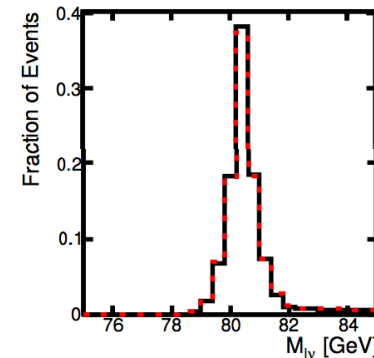
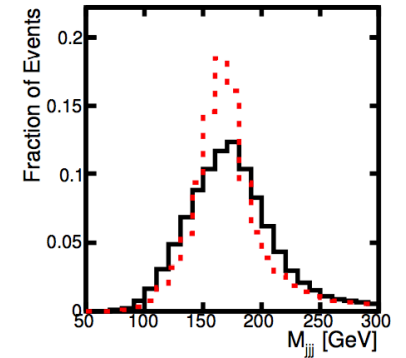
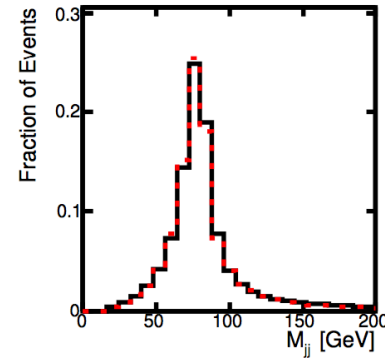
$m(bb)$

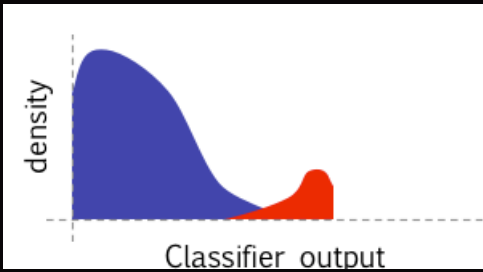
$m(bjj)$

$m(jj)$

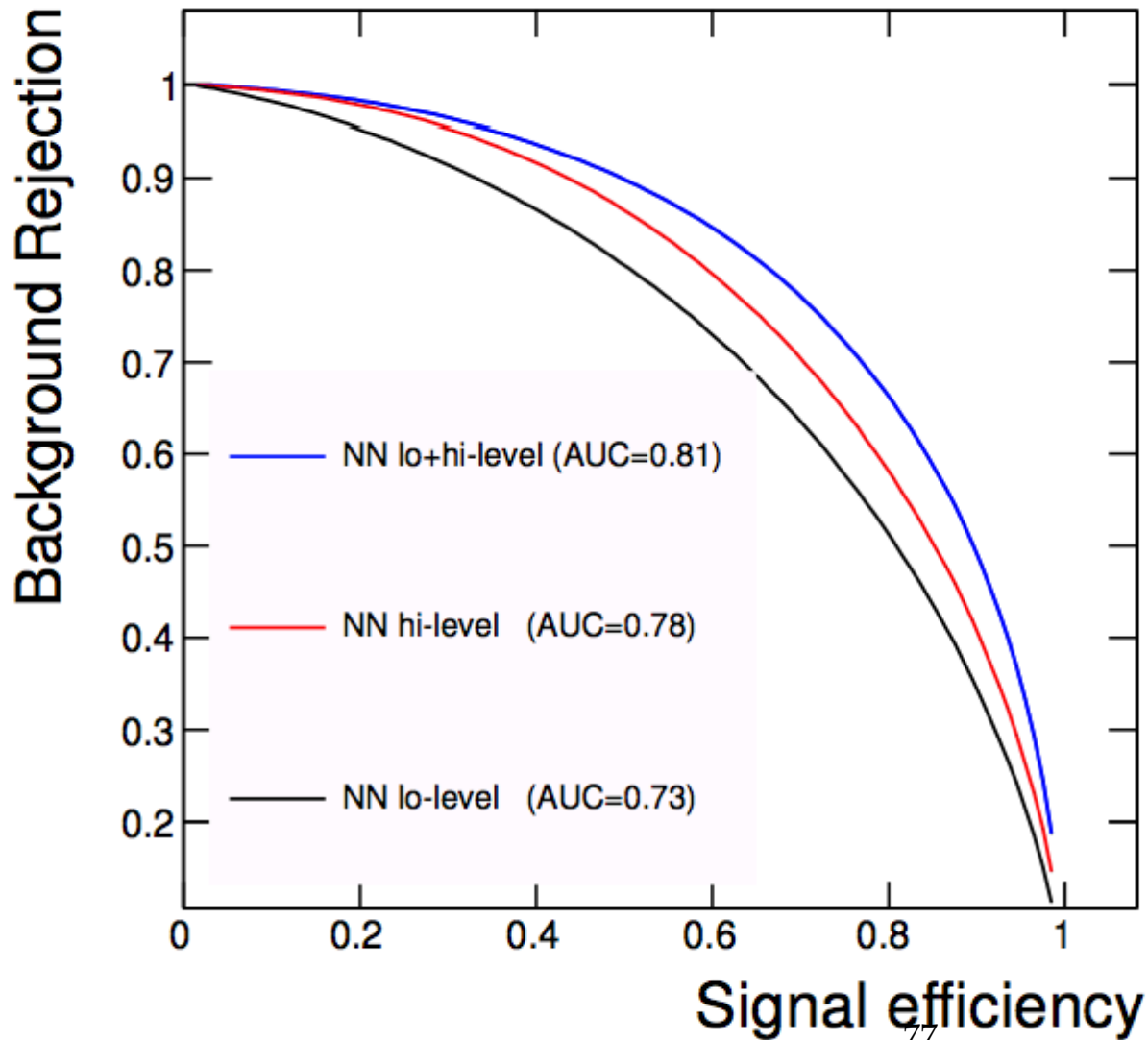
$m(lv)$

$m(blv)$





Standard NNs



Results

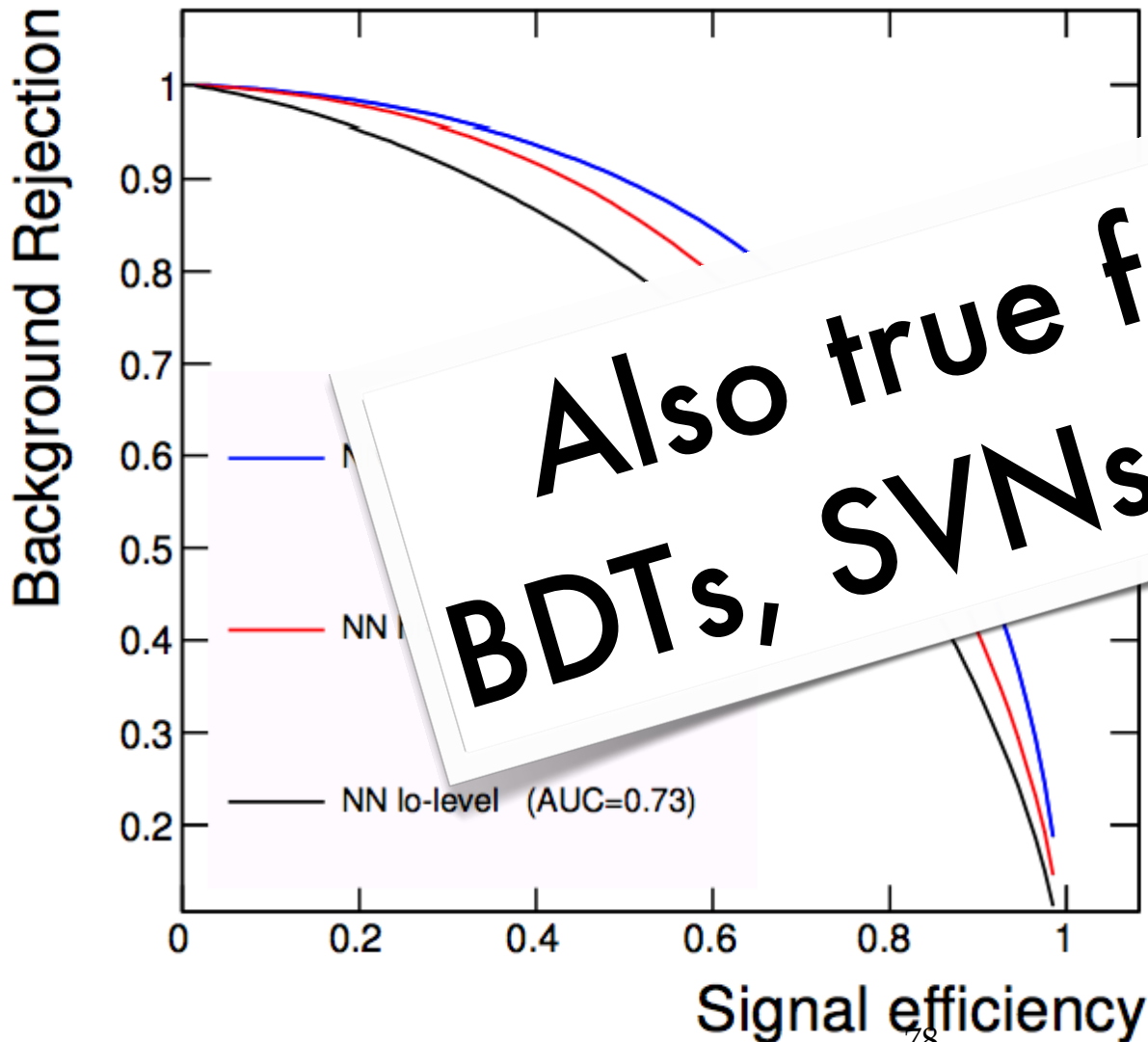
Adding hi-level
boosts performance
Better: lo+hi-level.

Conclude:

NN can't find
hi-level vars.

Hi-level vars
do not have all info

Standard NNs



Also true for
BDTs, SVNs, etc

Results

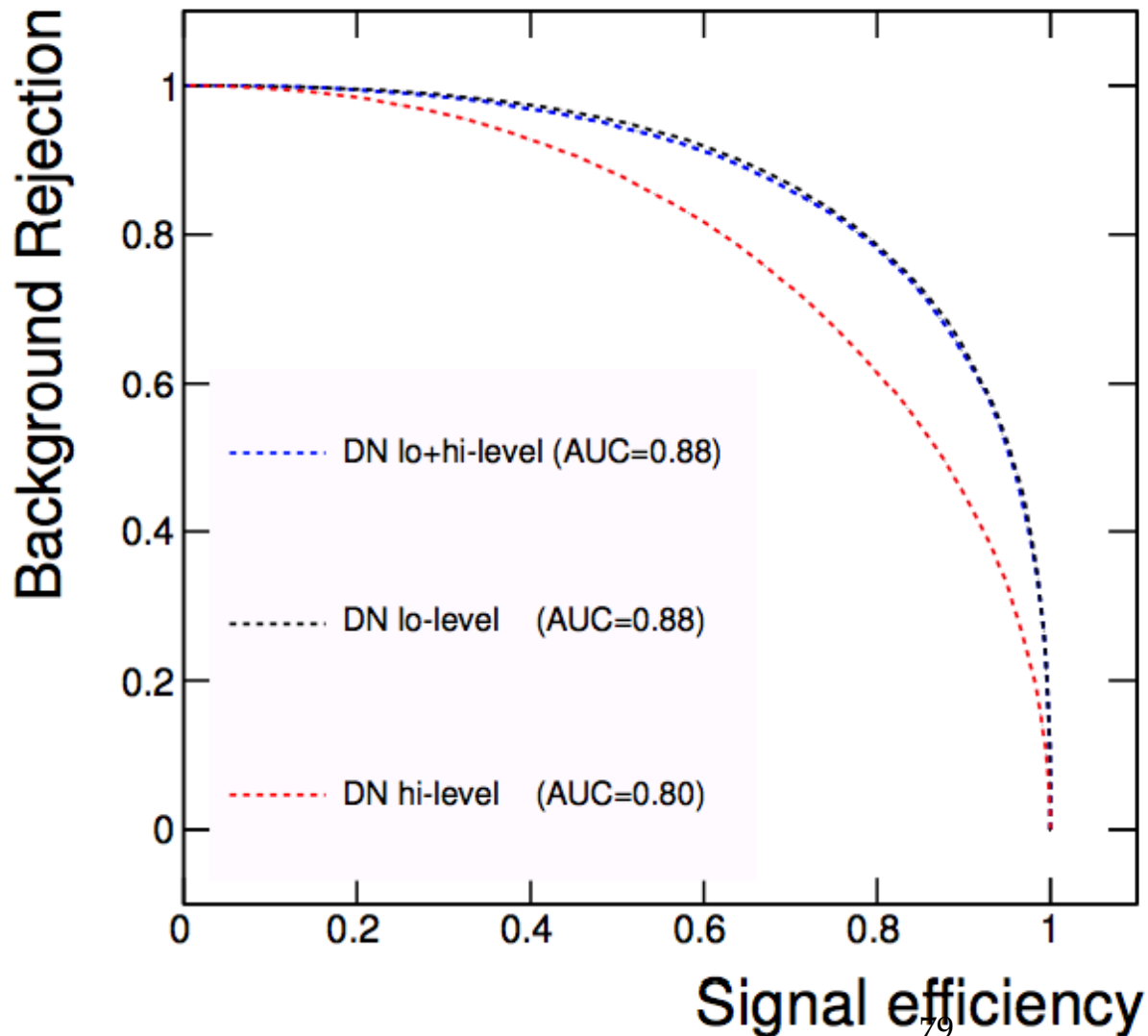
g hi-level
performance
lo+hi-level.

include:

NN can't find
hi-level vars.

Hi-level vars
do not have all info

Deep Networks



Results

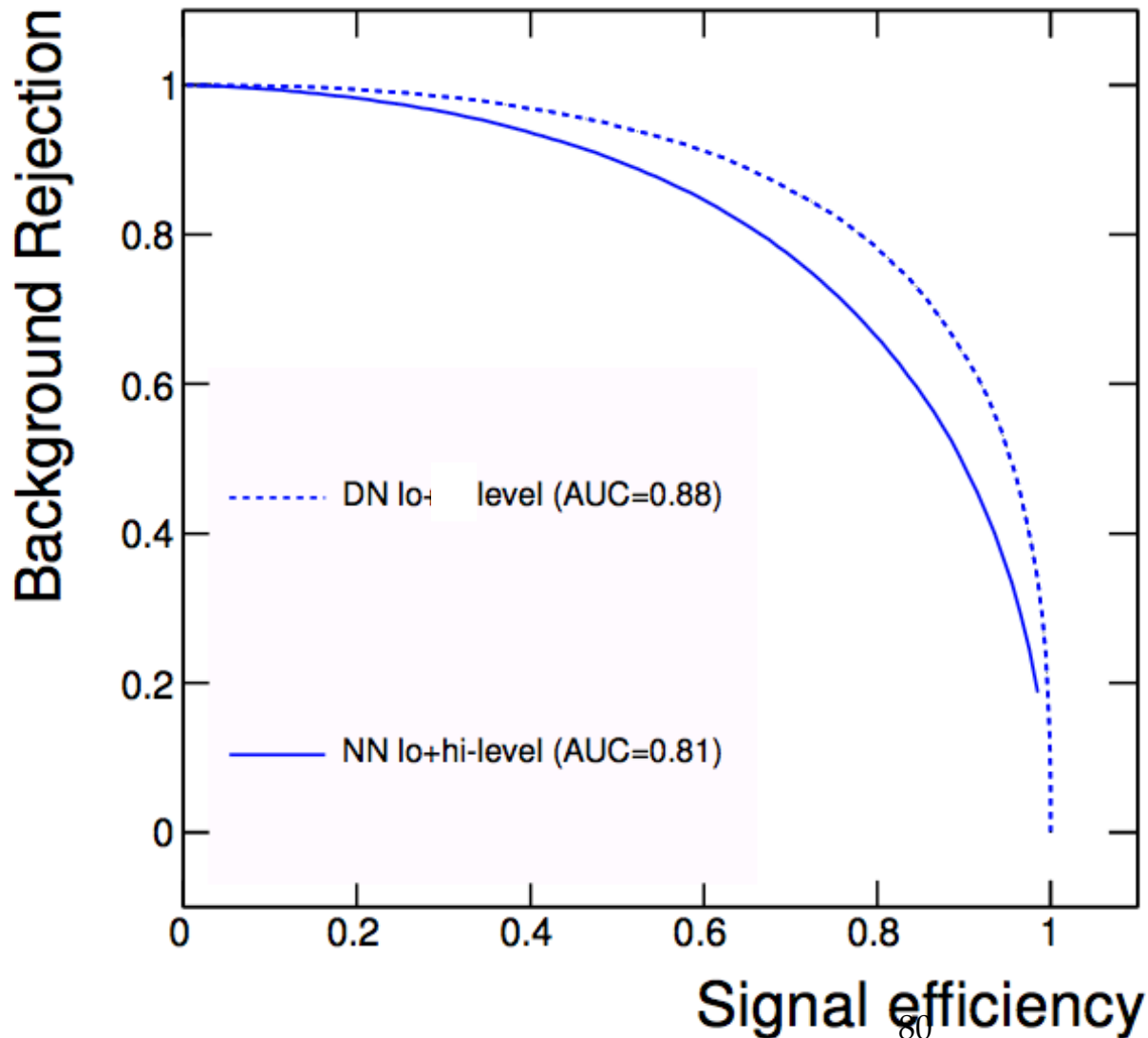
Lo+hi = lo.

Conclude:

DN can find
hi-level vars.

Hi-level vars
do not have all info
are unnecessary

Deep Networks



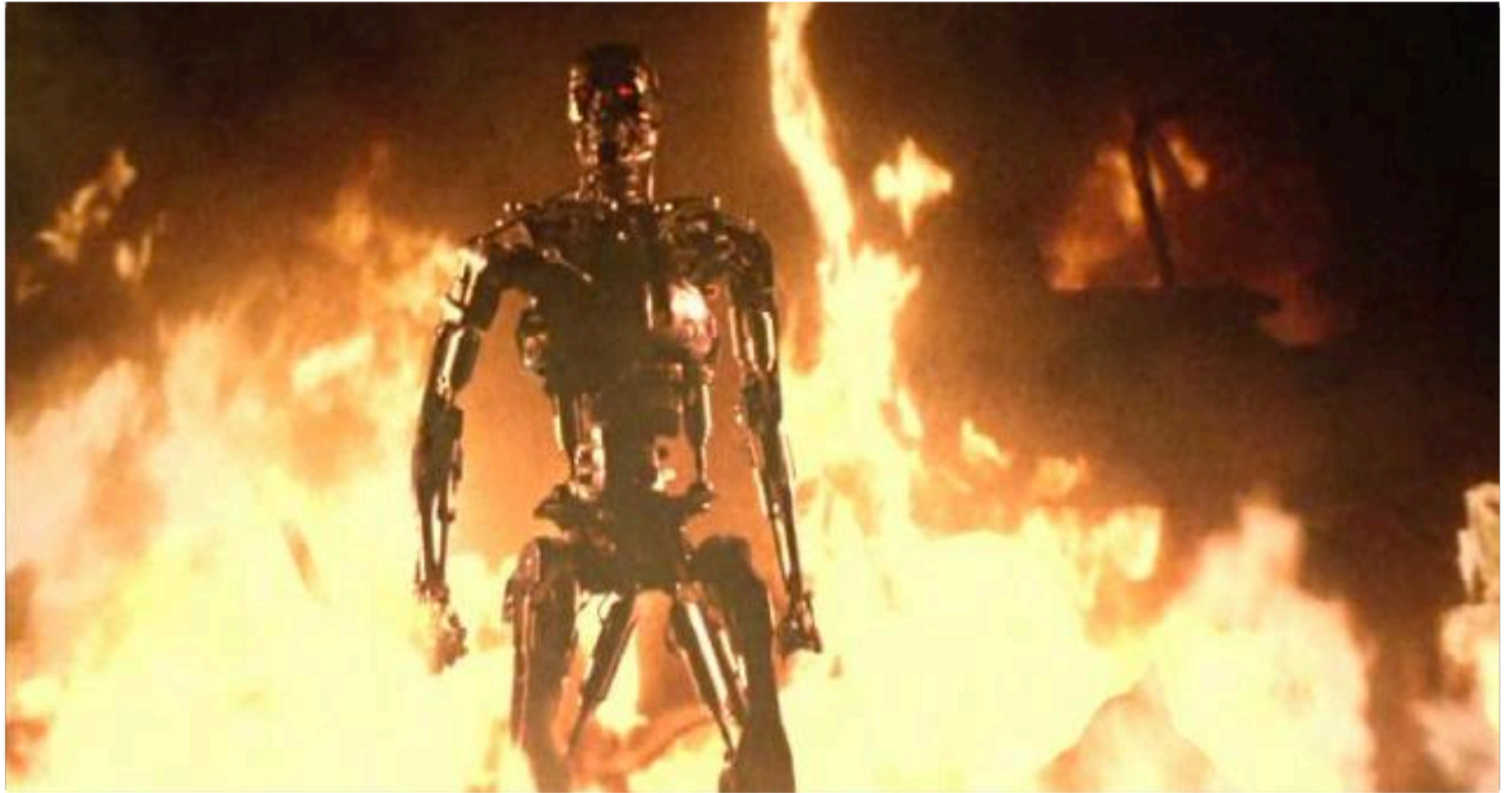
Results

DN > NN

Conclude:

DN does better
than human
assisted NN

The Als win



Results

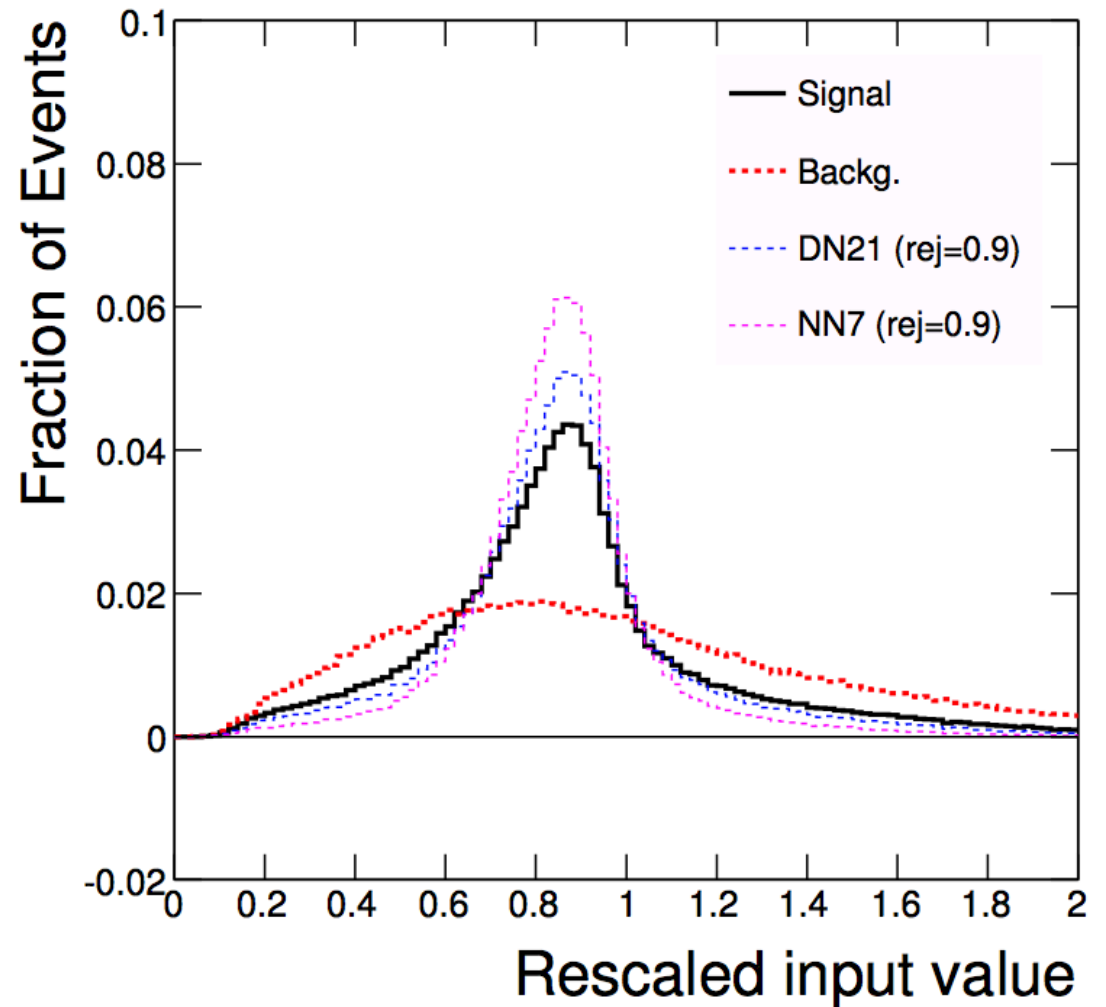
Identified example benchmark where traditional NNs fail to discover all discrimination power.

Adding human insight helps traditional NNs.

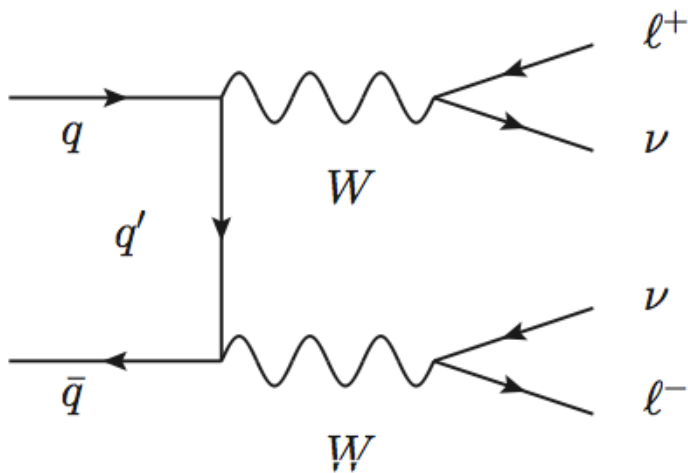
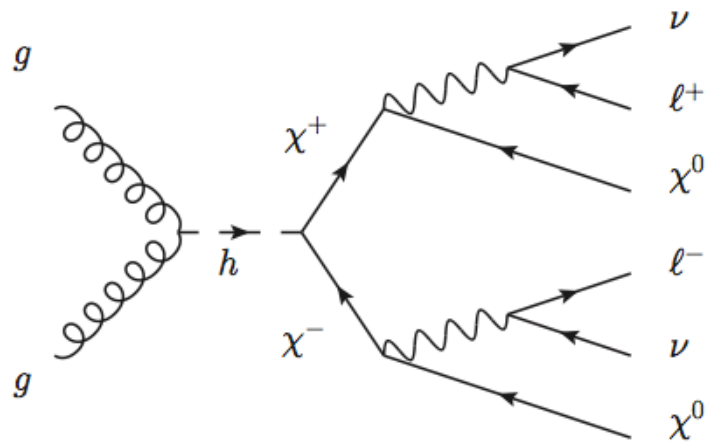
Deep networks succeed **without human insight**.
Outperform human-boosted traditional NNs.

Why?

DN not as
reliant on signal
features. Cuts into
background space.

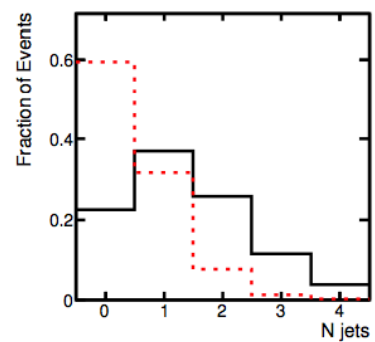
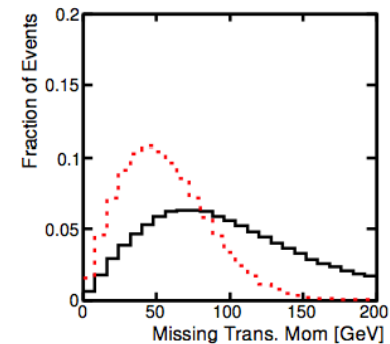
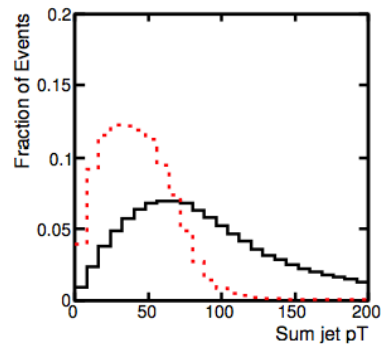
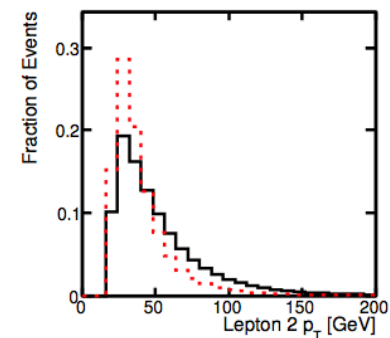
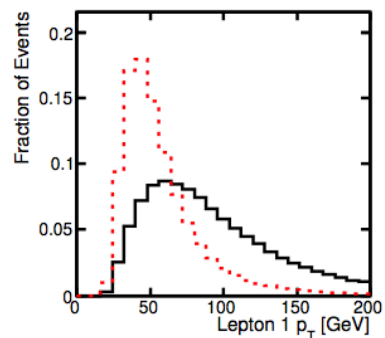


2nd case: SUSY



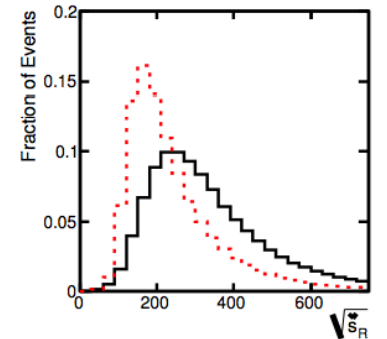
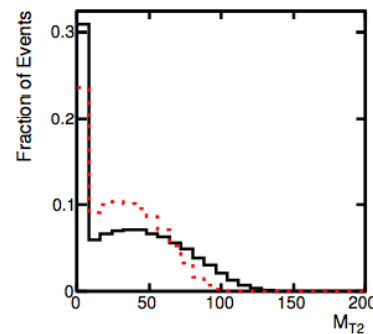
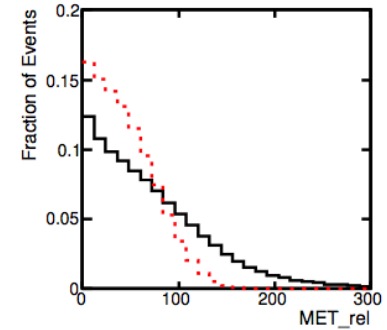
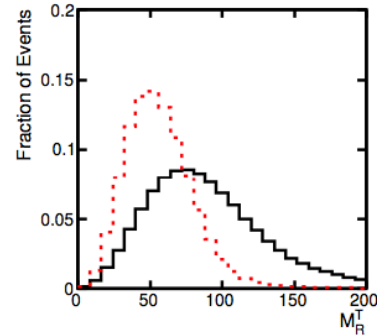
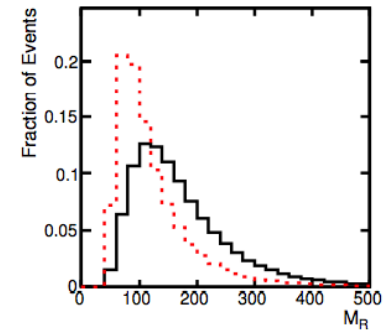
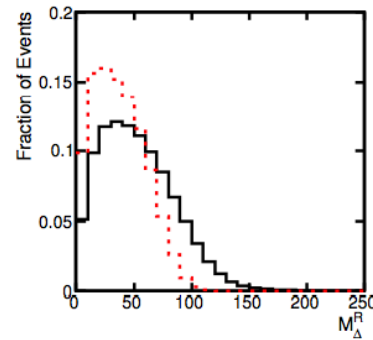
Can
Deep Networks
help us
find SUSY
in the data?

Low-level variables

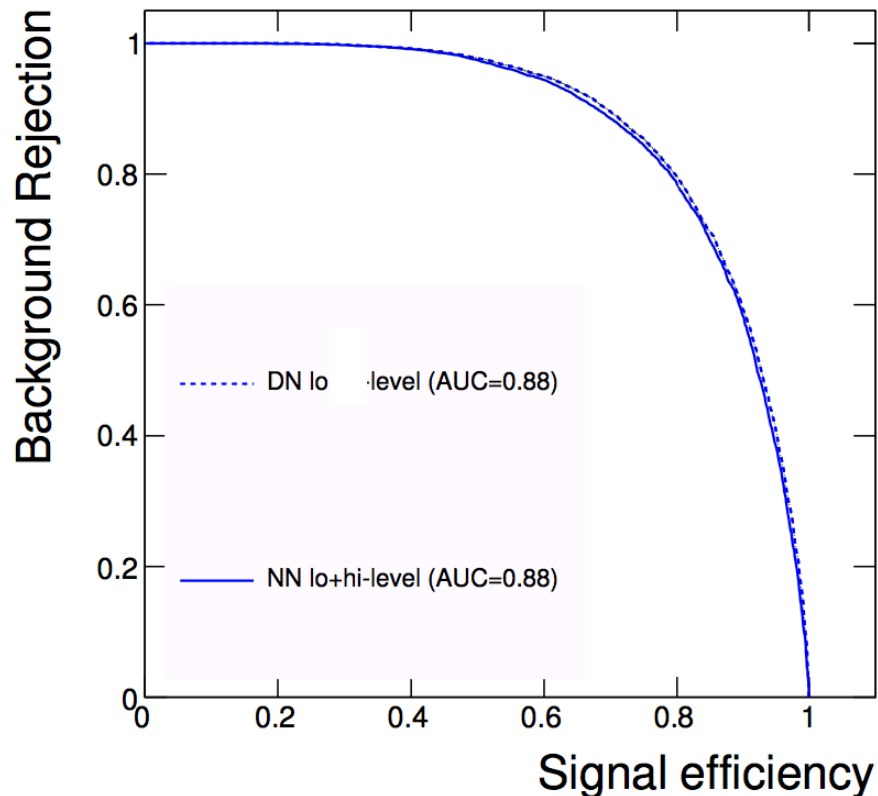


High-level variables

Axial-MET
Met-rel
MT2
Razor
Super-razor



SUSY results



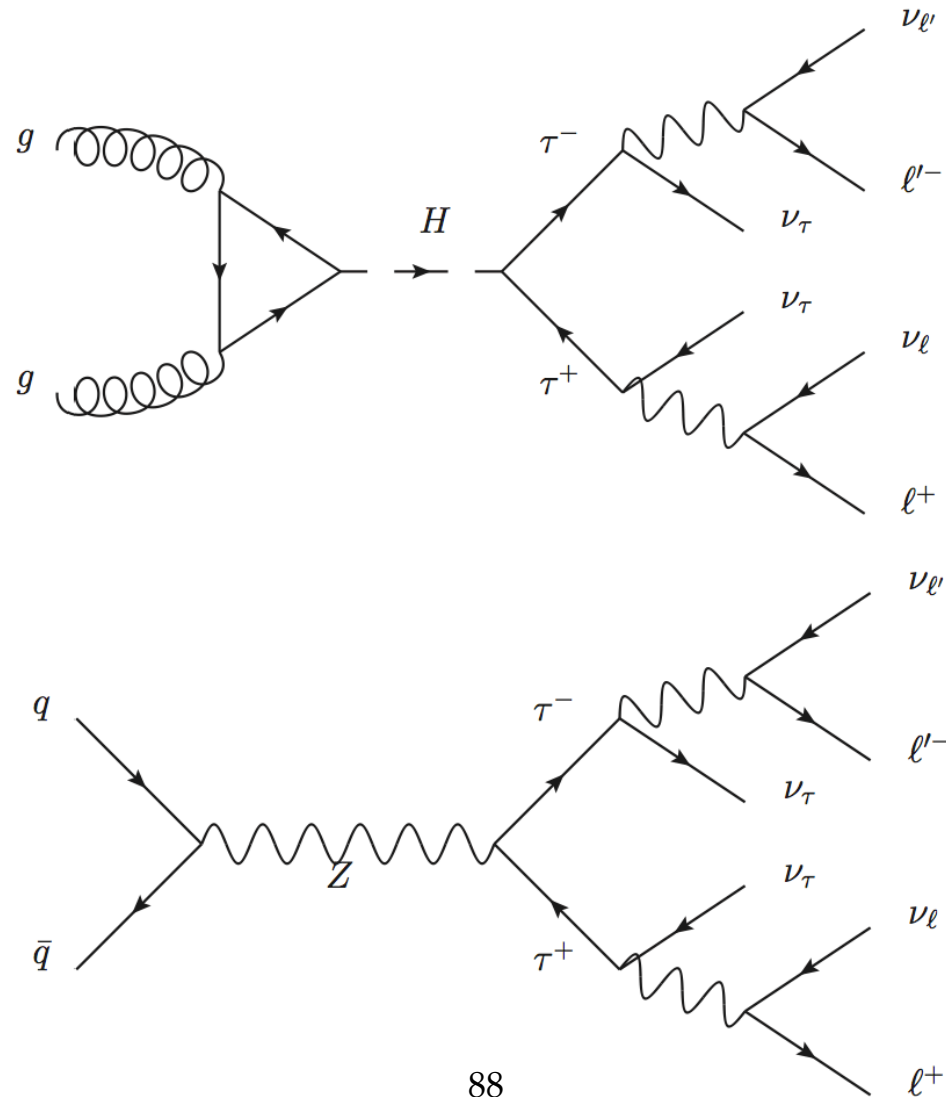
DN doesn't need help

Outperforms human assisted NN

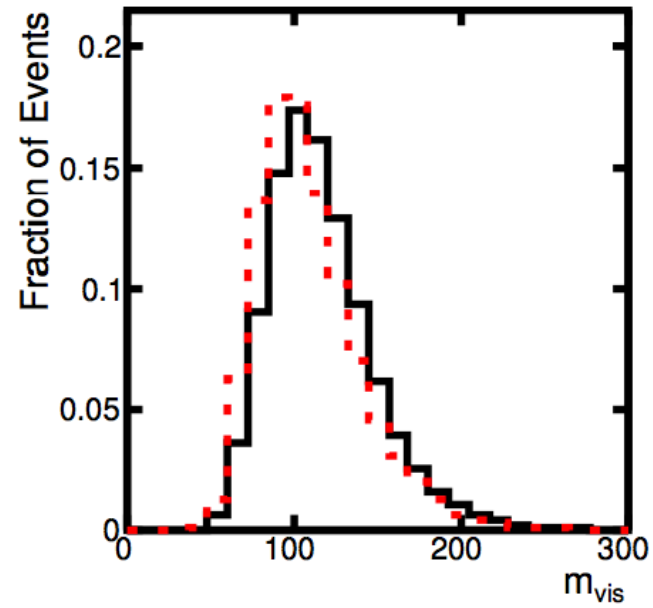
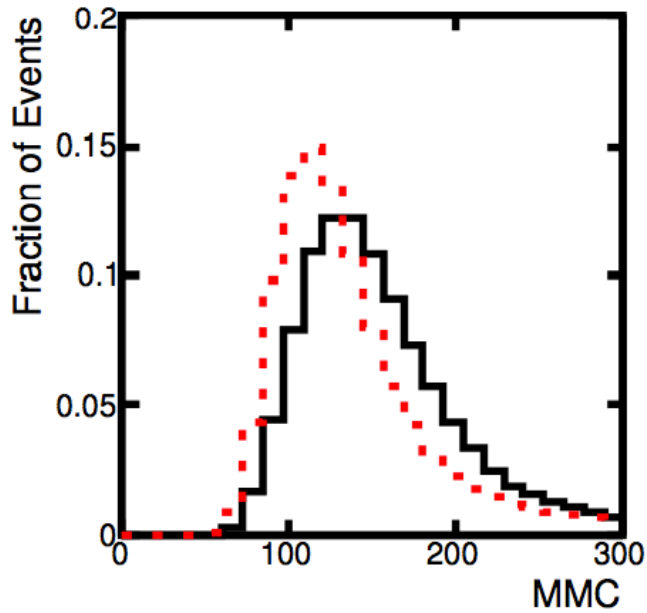
Margin is smaller
-> high level variables are **less helpful** and **less needed!**

Technique	Discovery significance		
	Low-level	High-level	Complete
NN	6.5σ	6.2σ	6.9σ
DN	7.5σ	7.3σ	7.6σ

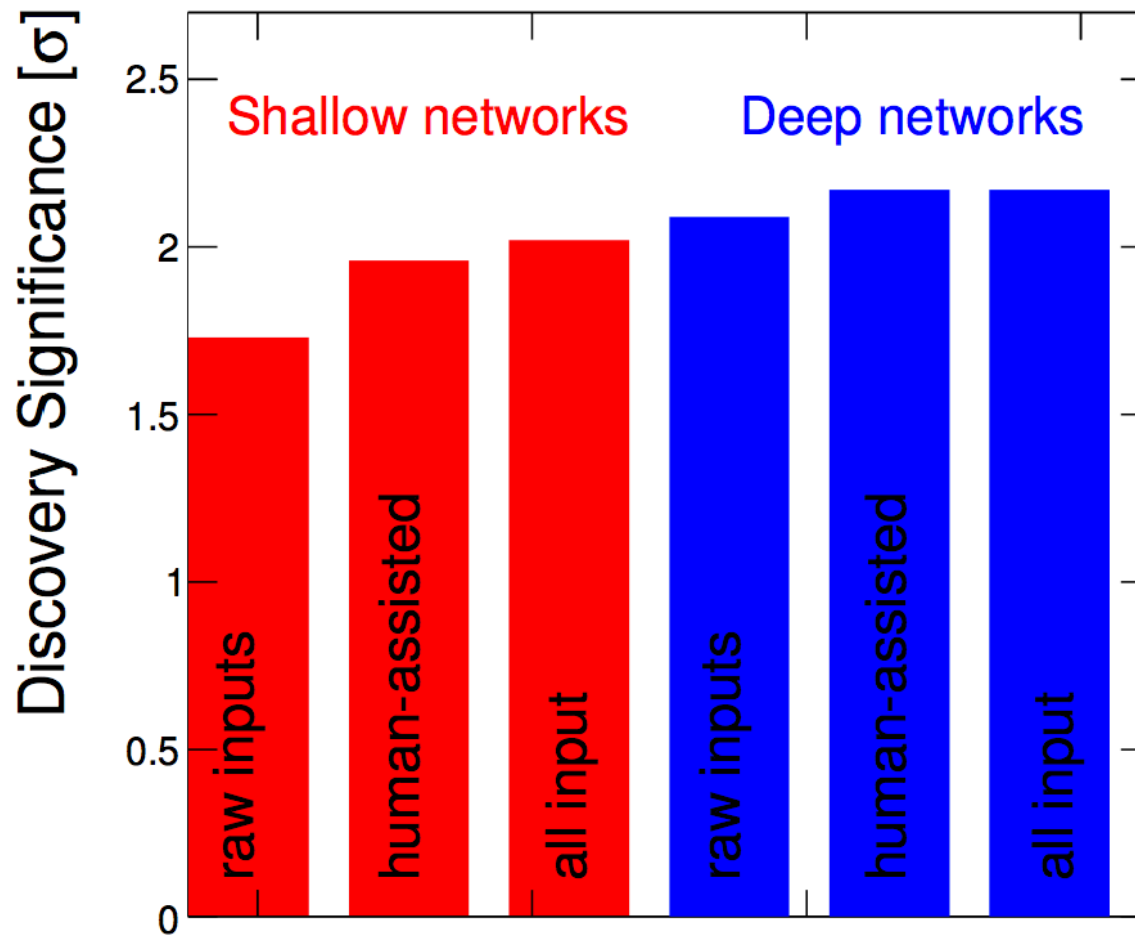
Preliminary: $h \rightarrow \tau\tau$



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Summary

Dark matter:

broad-based attack on all LHC signals

Topological models:

Strategy to build complete set of models with discoverable resonances

Deep networks:

Networks can take 4-vectors, find powerful discriminants