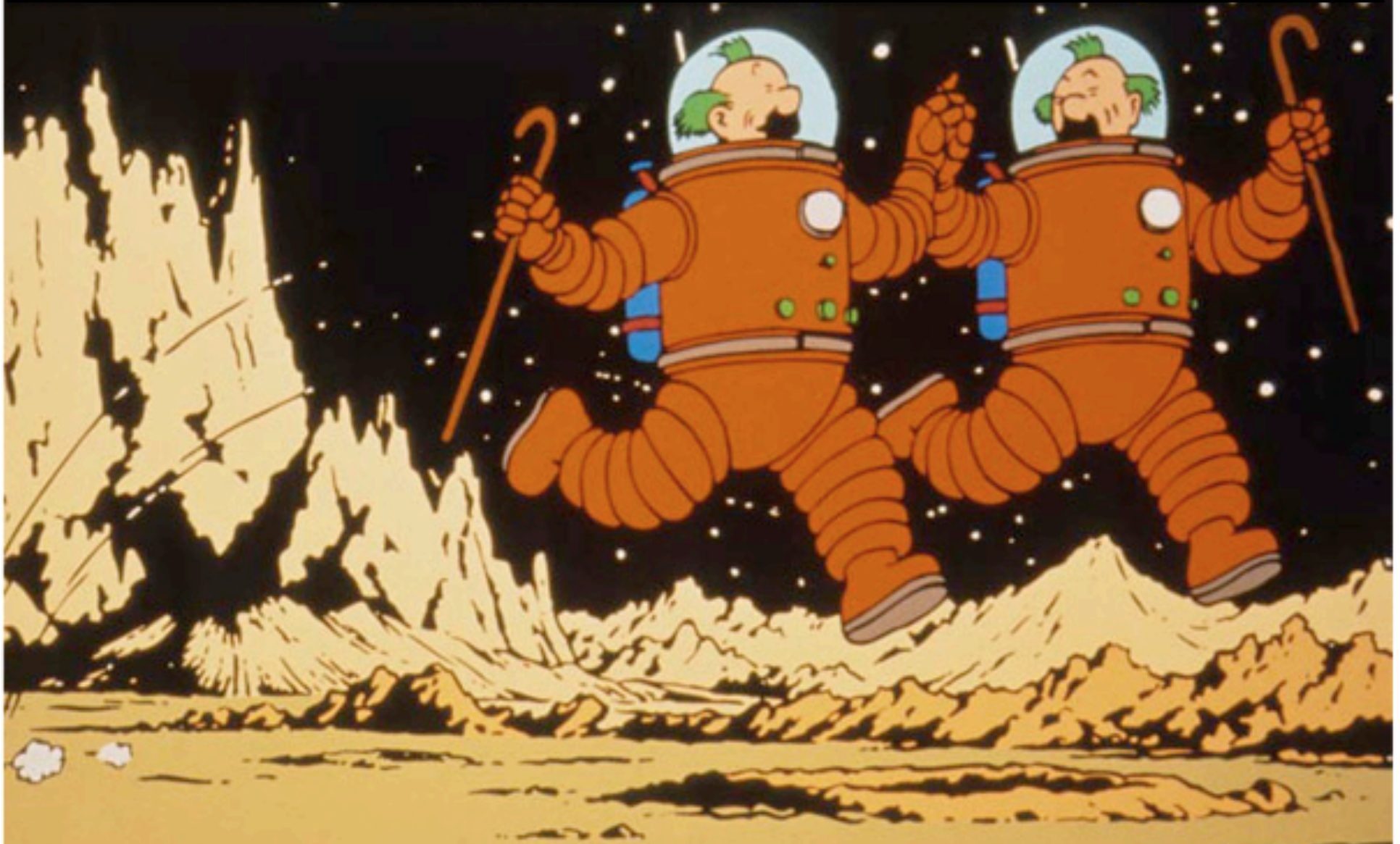
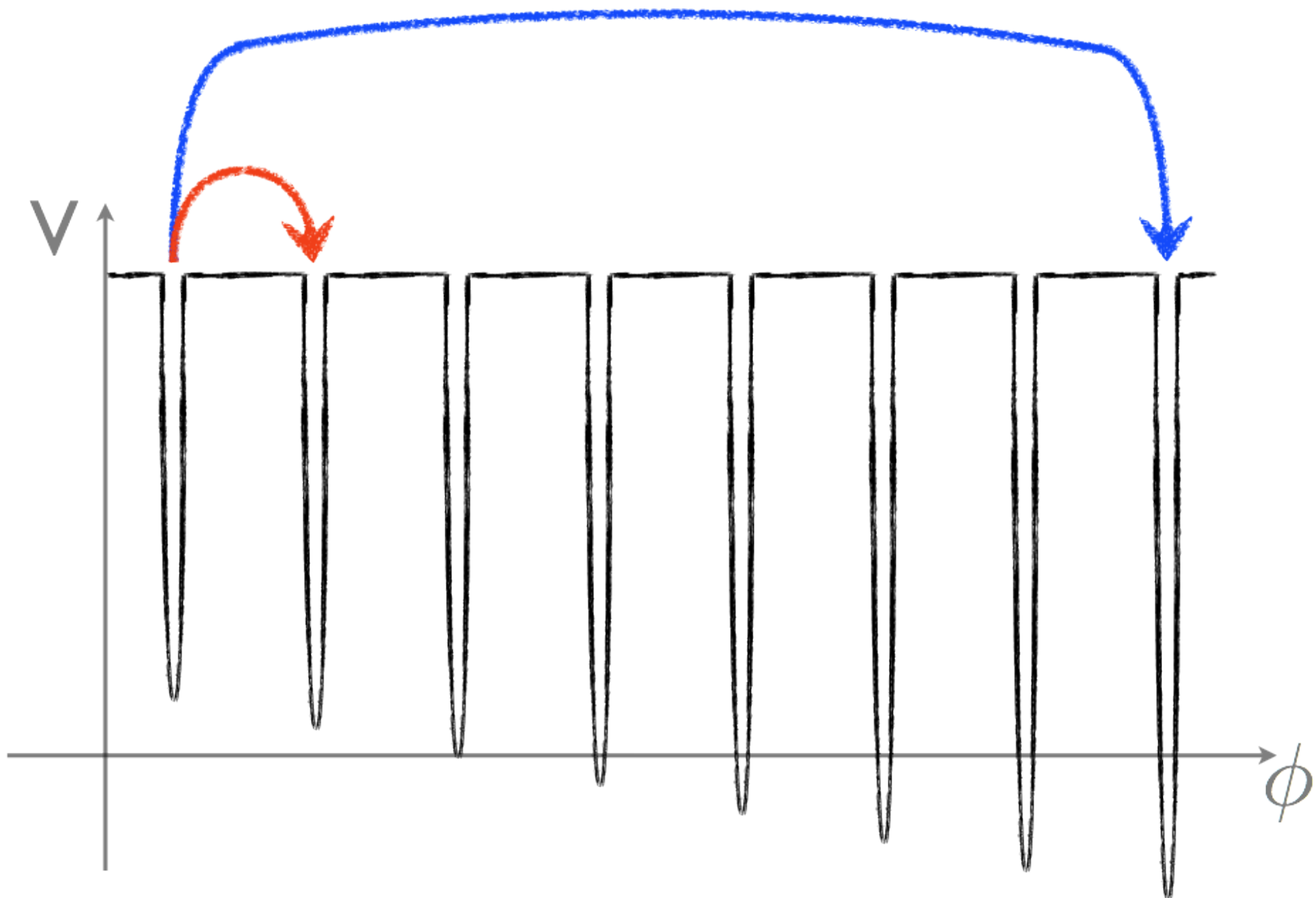


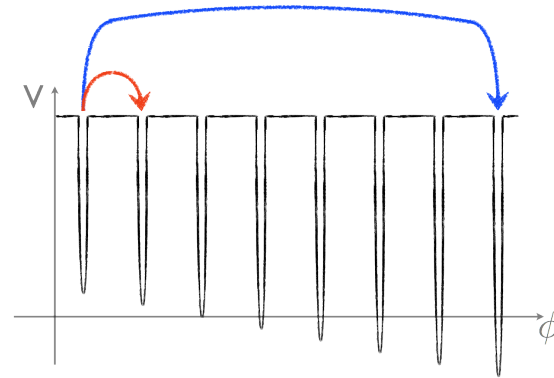
The Fastest Decay in the Landscape



Based on work with Adam Brown



1. This potential gives small steps



2. Flux compactifications with many fluxes give giant leaps

3. Monkey branes

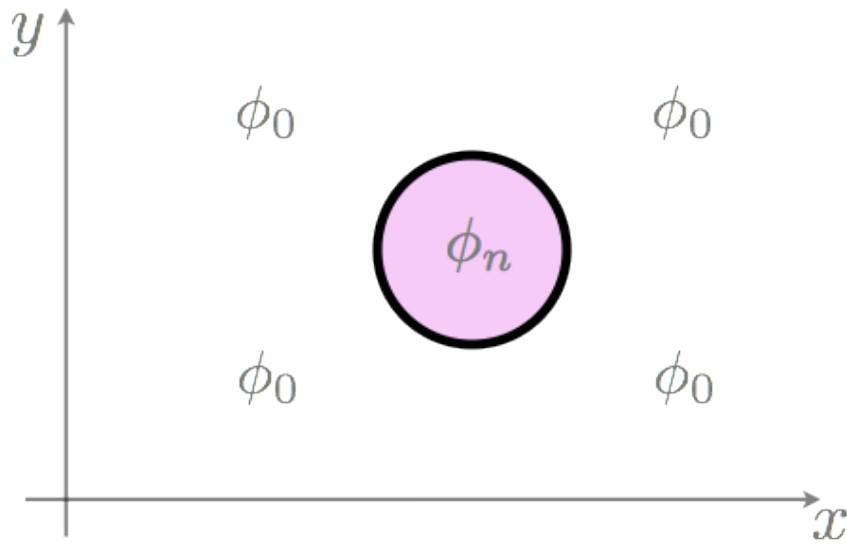
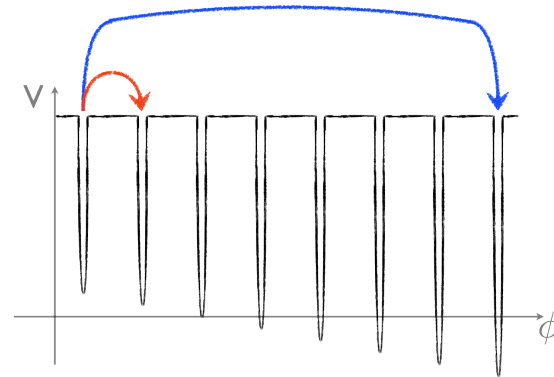


4. Adding back-reaction

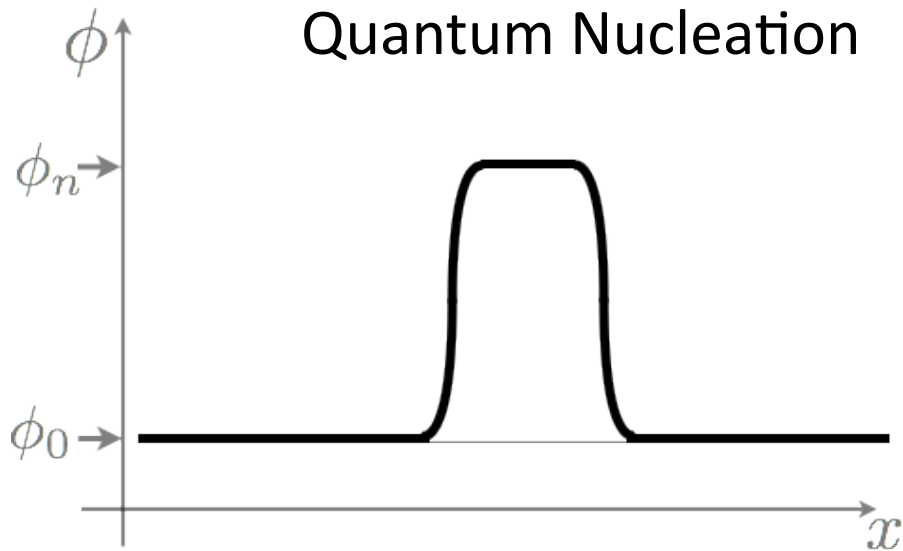
5. The giantest leap of all is a bubble of nothing



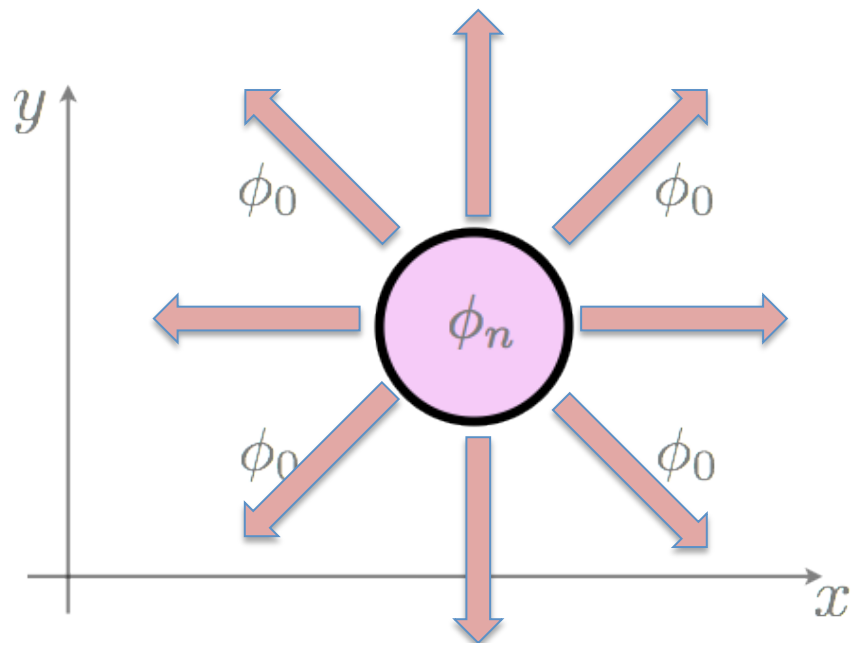
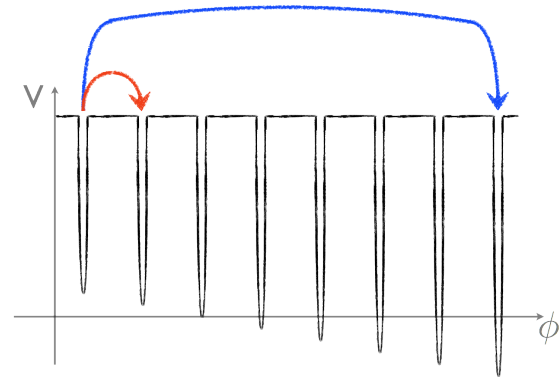
1. This potential gives small steps



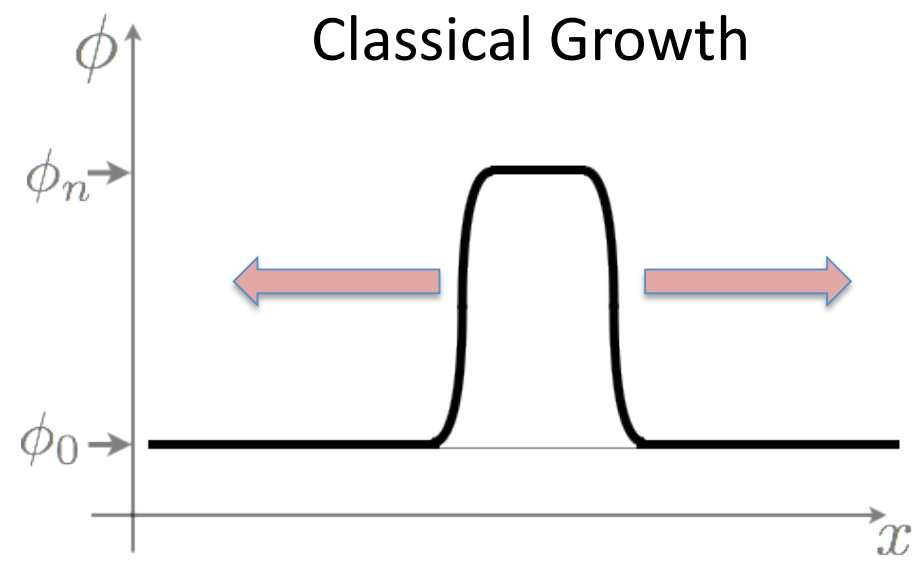
Quantum Nucleation



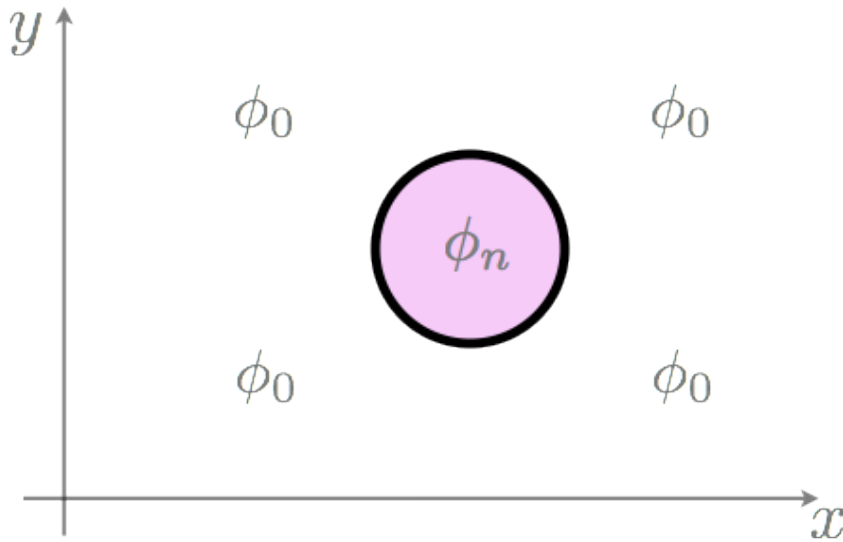
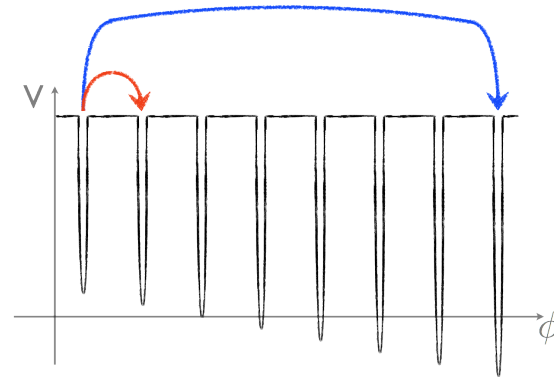
1. This potential gives small steps



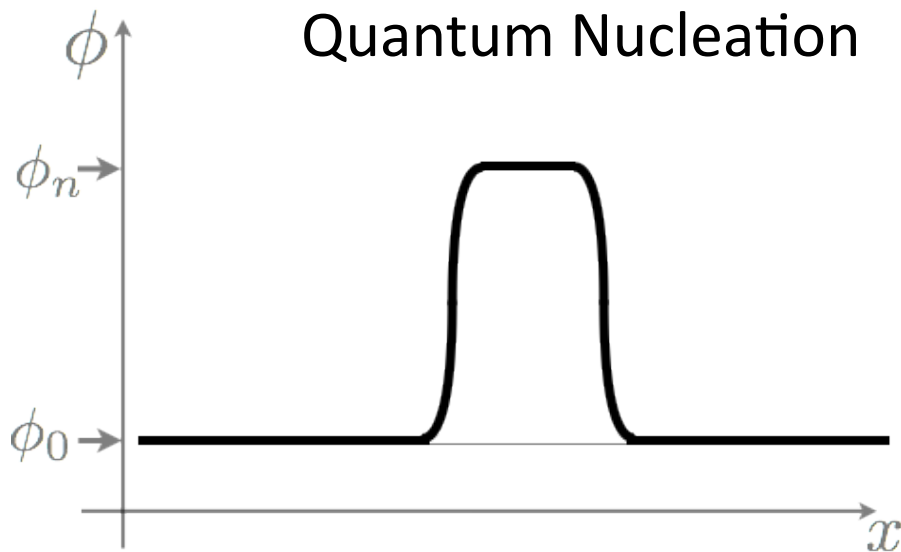
Classical Growth



1. This potential gives small steps



- thin wall
- flat spacetime



$$\Gamma_n \sim \exp\left[-\frac{\#}{\hbar} \frac{\sigma_n^4}{\epsilon_n^3}\right]$$

Coleman (1977)

$$\sigma_n \equiv \int_{\phi_0}^{\phi_n} d\phi \sqrt{2[V(\phi) - V(\phi_0)]}$$

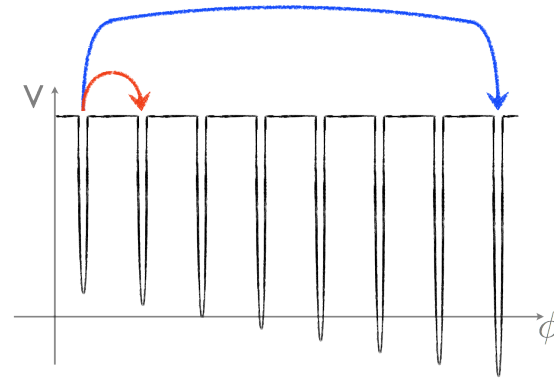
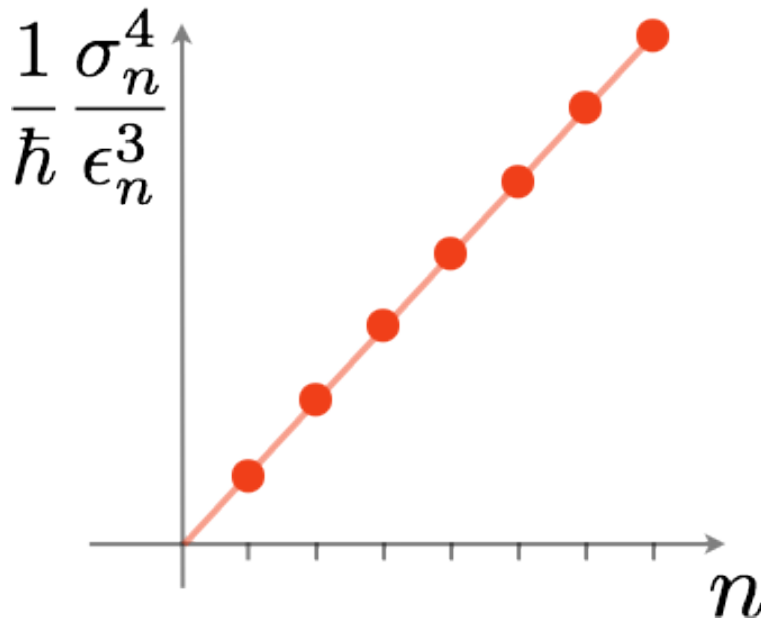
$$\epsilon_n \equiv V(\phi_0) - V(\phi_n)$$

1. This potential gives small steps

$$\sigma_n = n\sigma_1$$

$$\epsilon_n = n\epsilon_1$$

$$\frac{\sigma_n^4}{\epsilon_n^3} = n \frac{\sigma_1^4}{\epsilon_1^3}$$



- thin wall
- flat spacetime

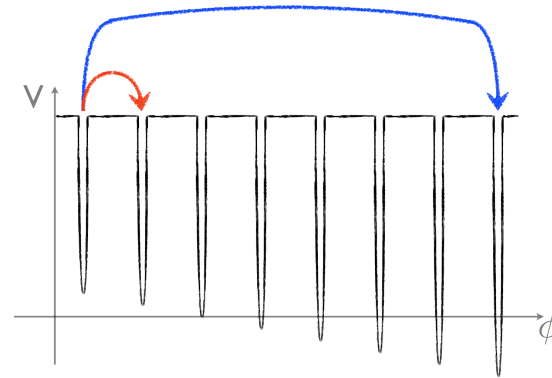
$$\Gamma_n \sim \exp\left[-\frac{\#}{\hbar} \frac{\sigma_n^4}{\epsilon_n^3}\right]$$

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$$\epsilon_n \equiv V(\phi_0) - V(\phi_n)$$

1. This potential gives small steps



2. Flux compactifications with many fluxes give giant leaps

3. Monkey branes



4. Adding back-reaction

5. The giantest leap of all is a bubble of nothing



Flux Landscapes

There are many non-cosmological reasons to consider extra dimensions

Extra dimensions and flux naturally give rise to a landscape of vacua

Flux shifts cosmological constant

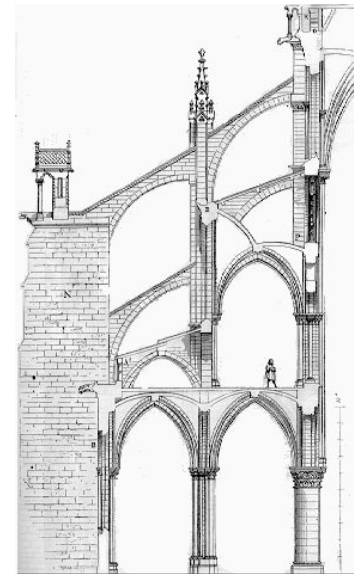
$$\Lambda_{\text{eff}} = \Lambda_0 + \frac{1}{2} F^2$$

Flux stabilizes extra dimensions

If you build a compactification, it will flux decay

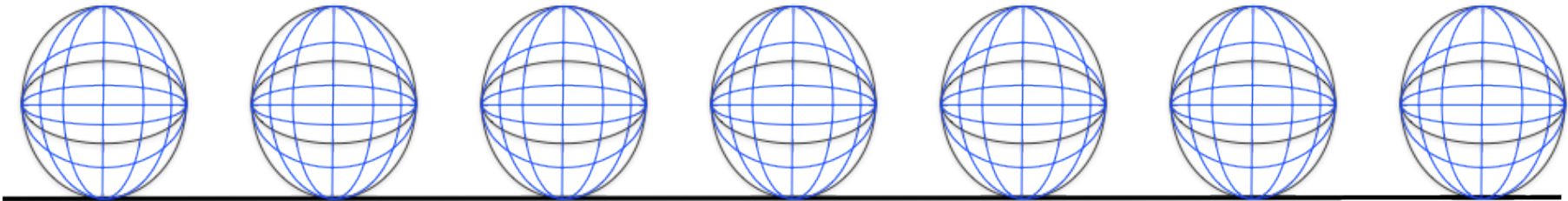
What will our Universe decay to?

Where did it come from?



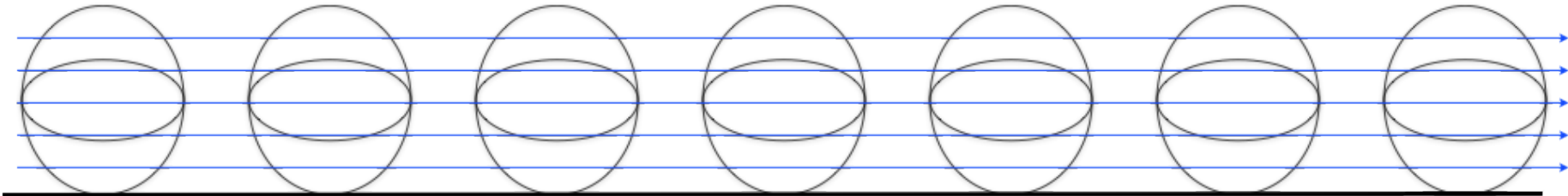
Consider flux tunneling in 1+1+2 dimensions

space time sphere



F_{ab} wraps

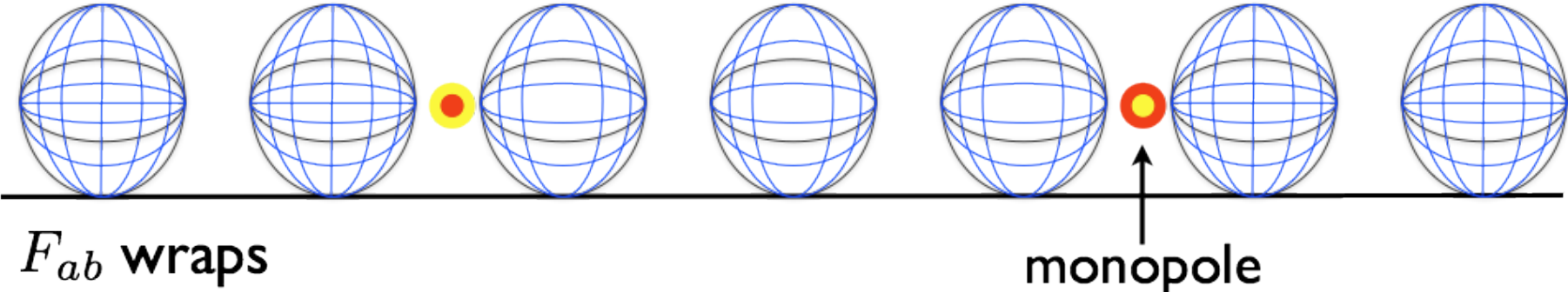
N units of flux N units of flux N units of flux



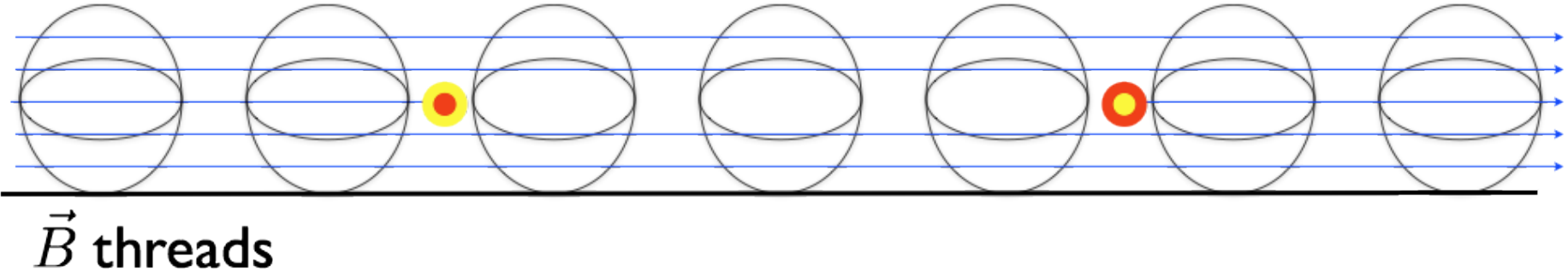
\vec{B} threads

Consider flux tunneling in 1+1+2 dimensions

space time sphere

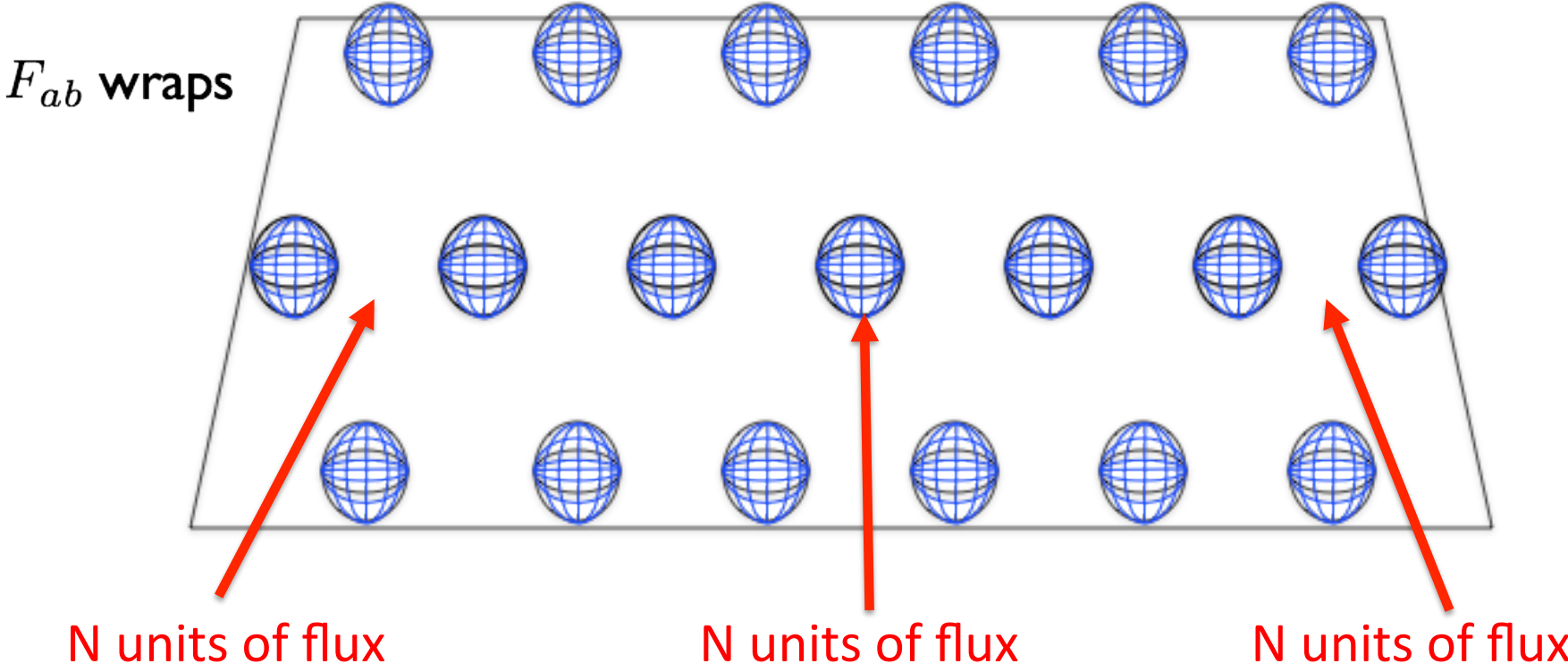


N units of flux || N-1 units of flux || N units of flux



Consider flux tunneling in 3+1+2 dimensions

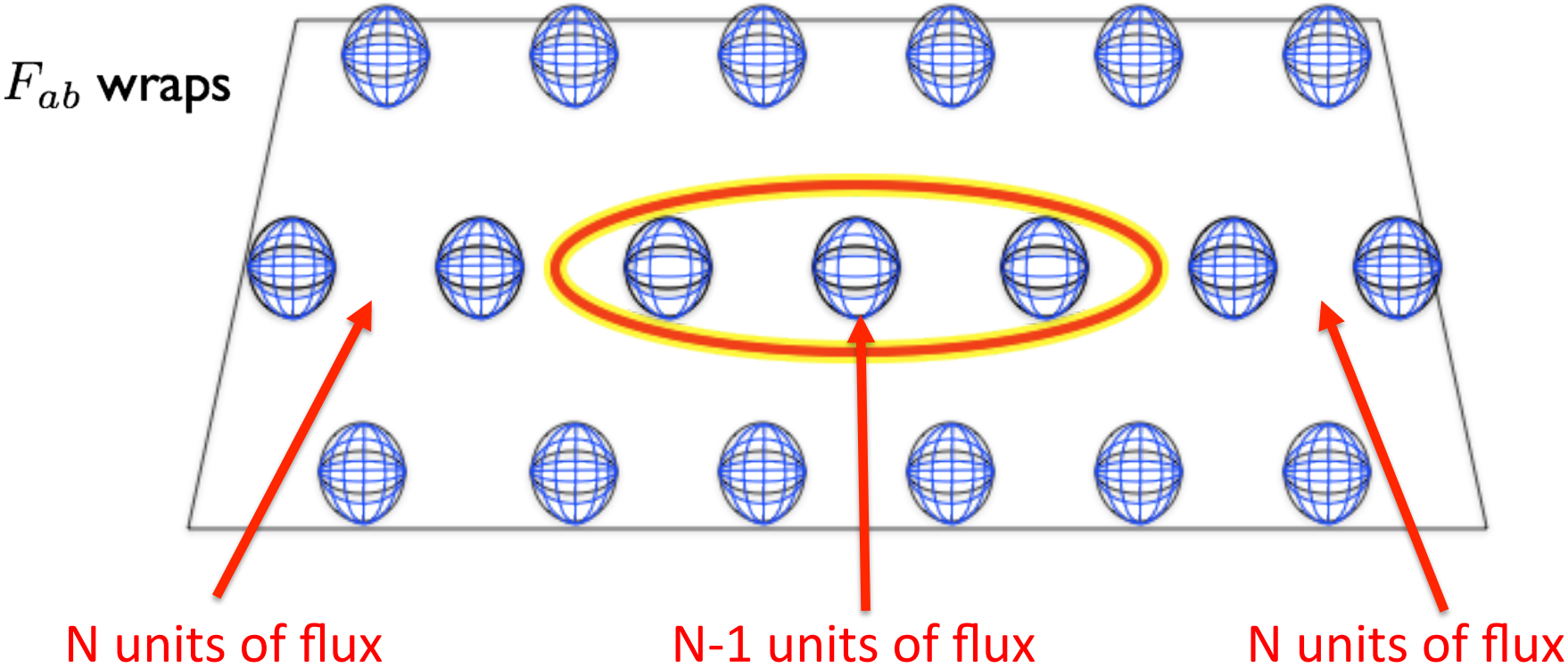
space time sphere



Consider flux tunneling in 3+1+2 dimensions

space time sphere

Nucleate an extremal black 2-branes

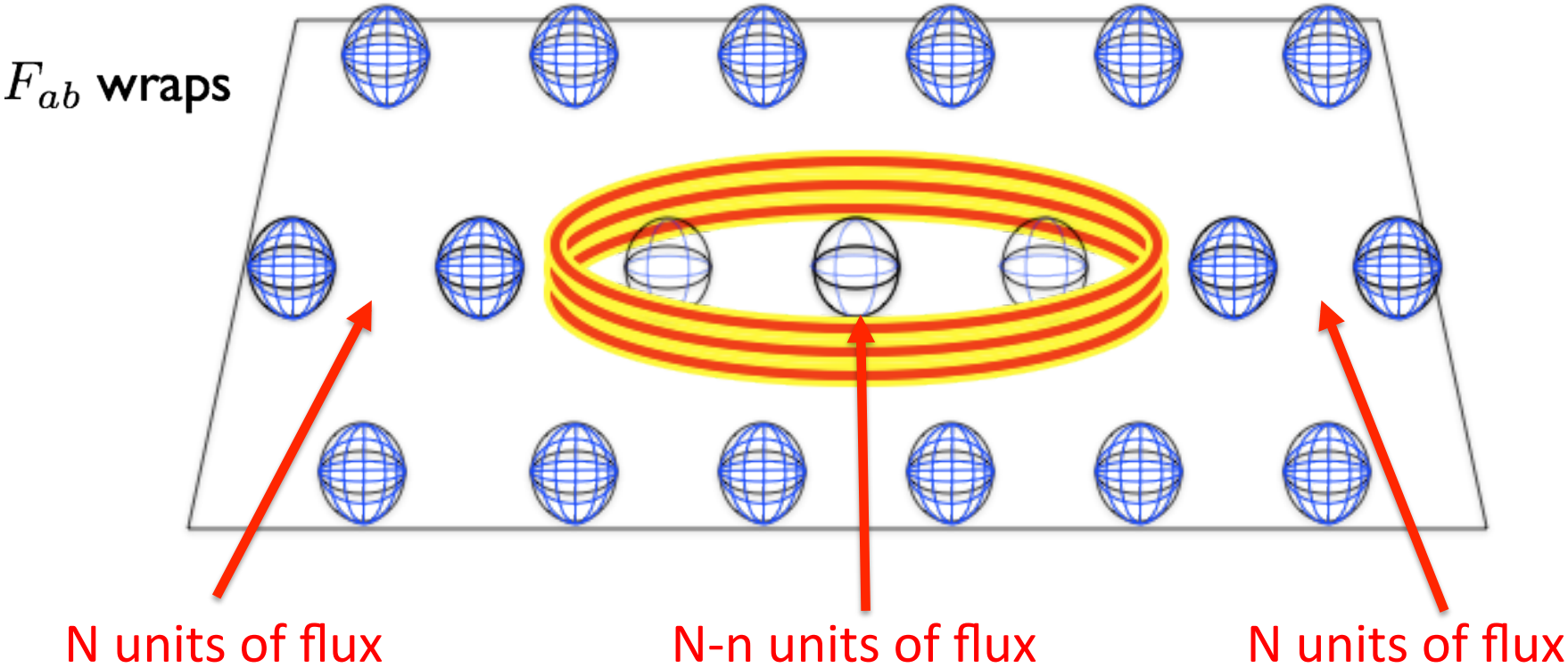


Consider flux tunneling in 3+1+2 dimensions

Giant Leaps?

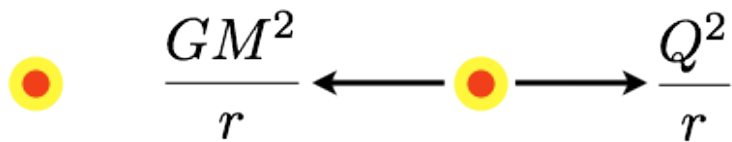


Nucleate a stack of n 2-branes



What's the tension of the brane?

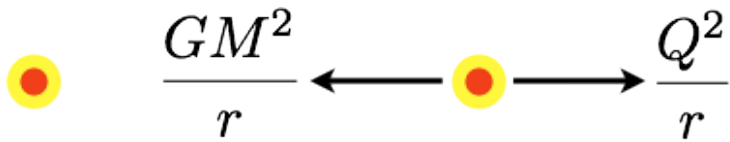
as light as possible, given its charge



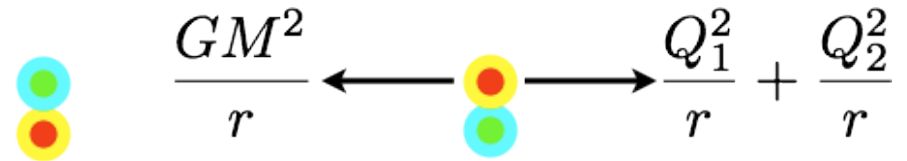
 $M = Q$


What's the tension of the brane?

as light as possible, given its charge



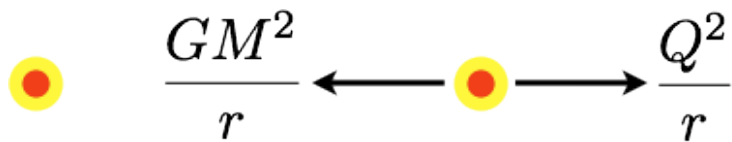
 $M = Q$




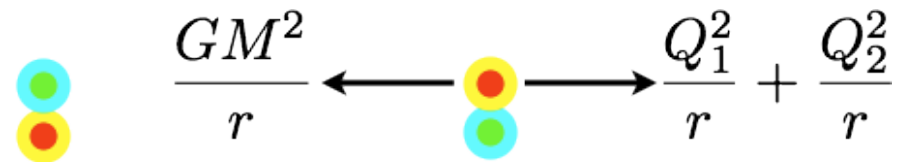
 $M = \sqrt{Q_1^2 + Q_2^2}$


What's the tension of the brane?

as light as possible, given its charge



 $M = Q$



 $M = \sqrt{Q_1^2 + Q_2^2}$

$$T \sim \left(\sum_{i=1}^{\mathfrak{n}} g_i^2 n_i^2 \right)^{1/2}$$

Large landscape: many different types of fluxes

$$F^2 = \sum_{i=1}^{\mathfrak{N}} g_i^2 N_i^2 \quad T \sim \left(\sum_{i=1}^{\mathfrak{N}} g_i^2 n_i^2 \right)^{1/2}$$

Large landscape: many different types of fluxes

$$F^2 = \sum_{i=1}^{\mathfrak{N}} g_i^2 N_i^2 \quad T \sim \left(\sum_{i=1}^{\mathfrak{N}} g_i^2 n_i^2 \right)^{1/2}$$

Two decay directions:

MONOFLUX decay

MULTIFLUX decay

Large landscape: many different types of fluxes

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Two decay directions:

MONOFLUX decay

1. $T_n \sim n$

MULTIFLUX decay

1. $T_n \sim \sqrt{n}$

Large landscape: many different types of fluxes

$$F^2 = \sum_{i=1}^{\mathfrak{n}} g_i^2 N_i^2.$$

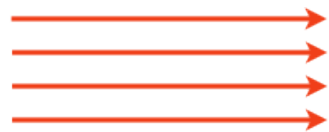
$$T \sim \left(\sum_{i=1}^{\mathfrak{n}} g_i^2 n_i^2 \right)^{1/2}$$

Two decay directions:

MONOFLUX decay

I. $T_n \sim n$

II. $\Delta(F^2) \sim (2N - n)n$



like flux lines repel

MULTIFLUX decay

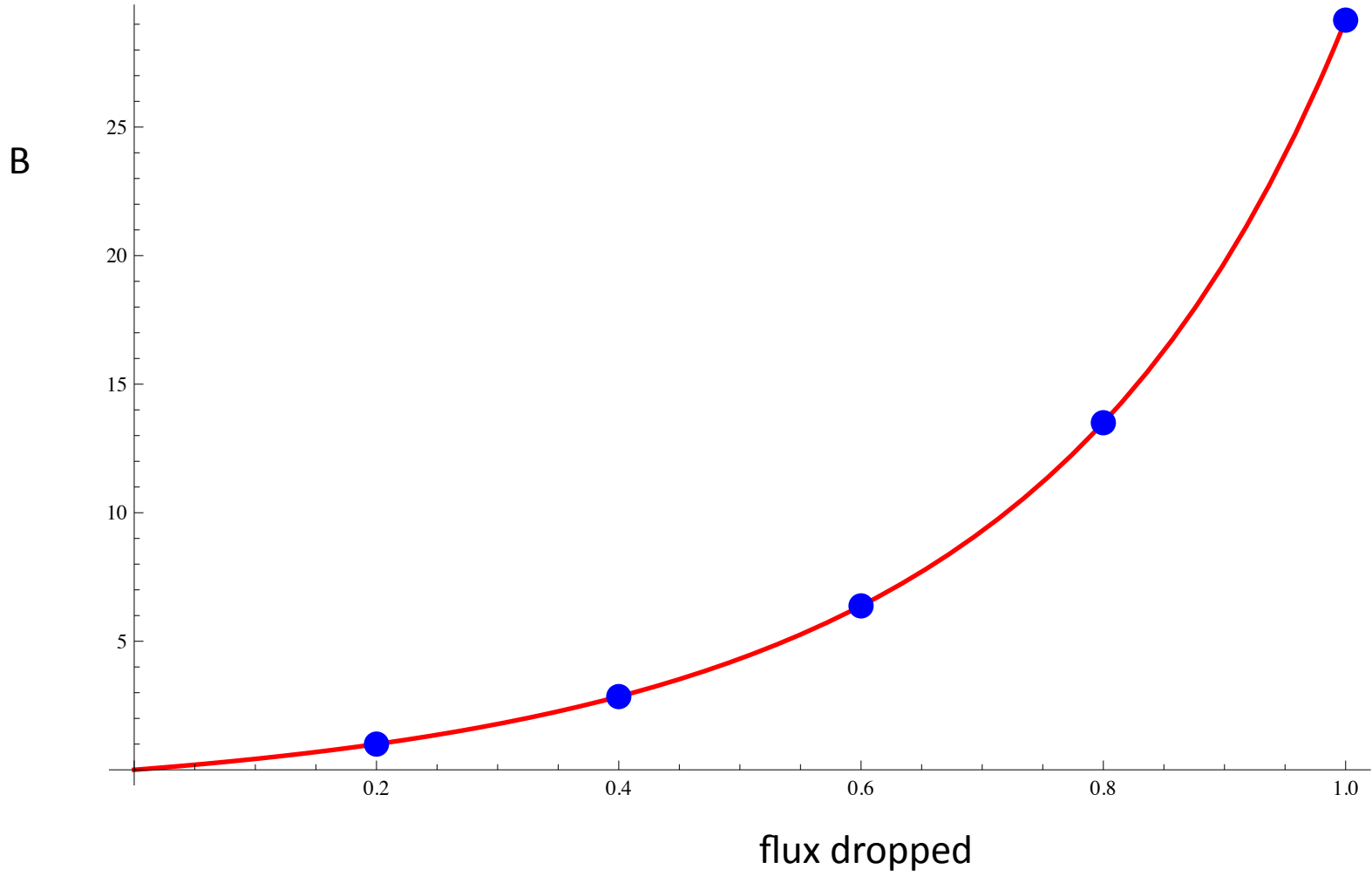
I. $T_n \sim \sqrt{n}$

II. $\Delta(F^2) \sim n$

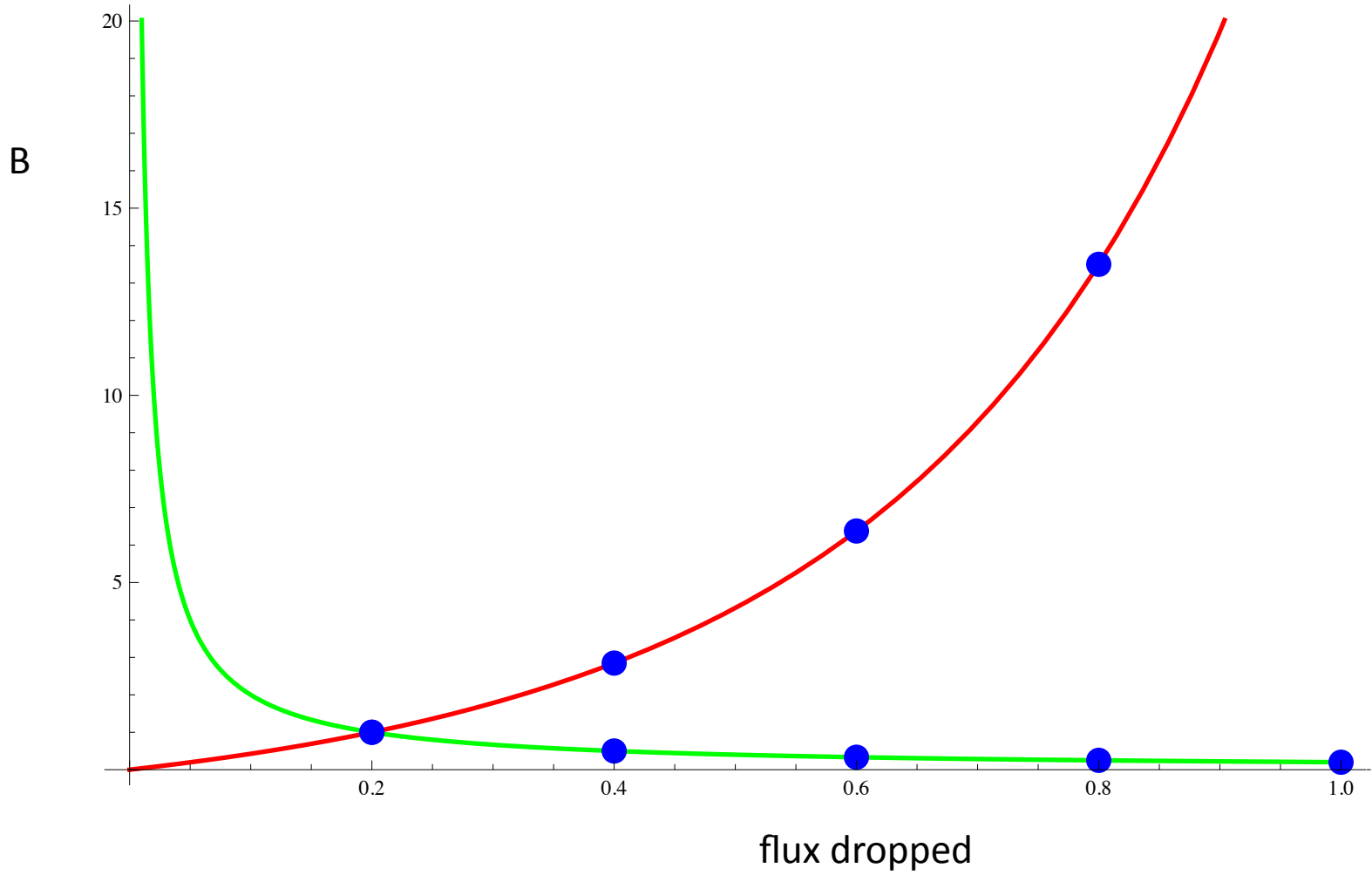


unlike flux lines don't

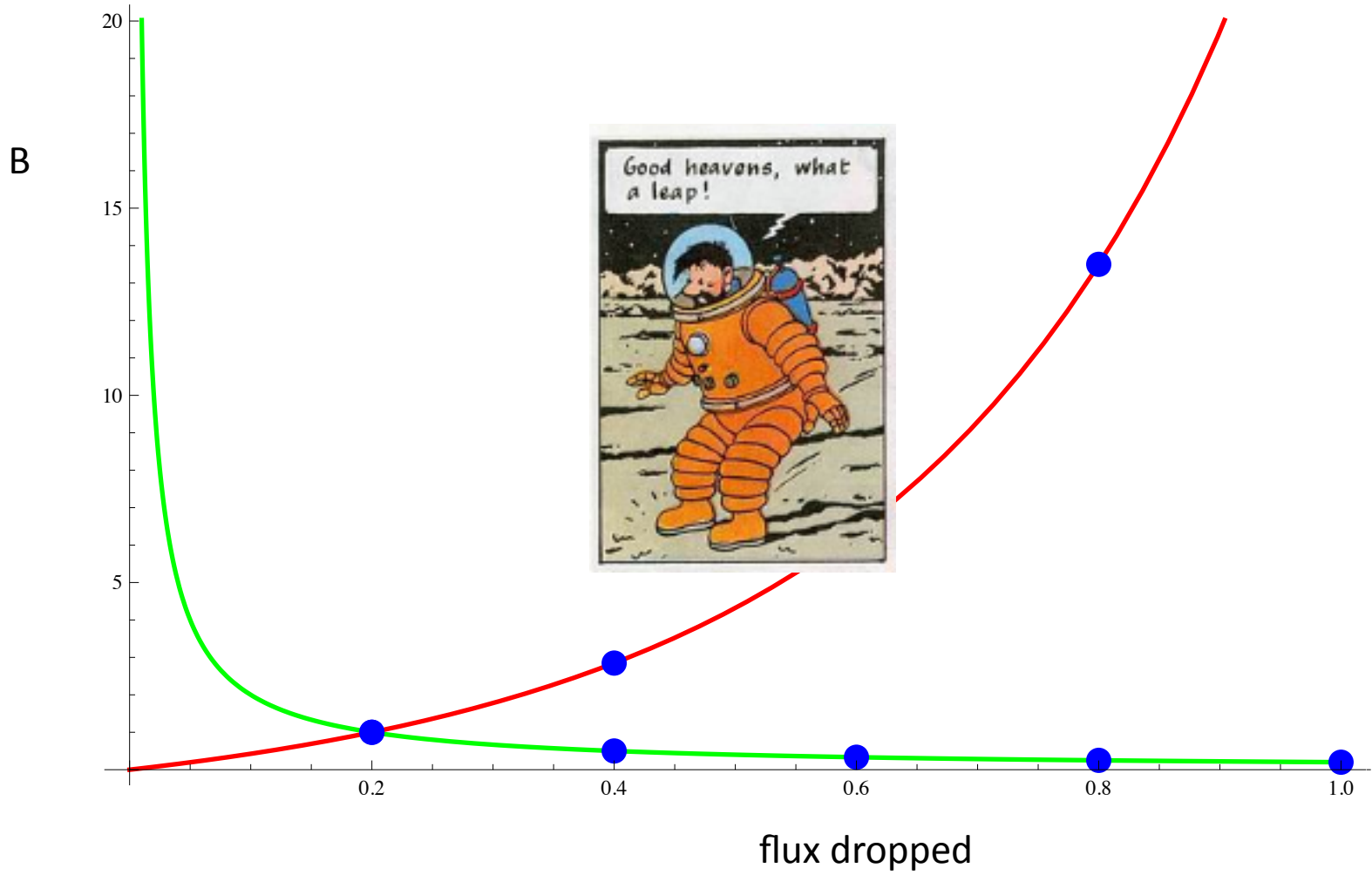
MONOFLUX decay



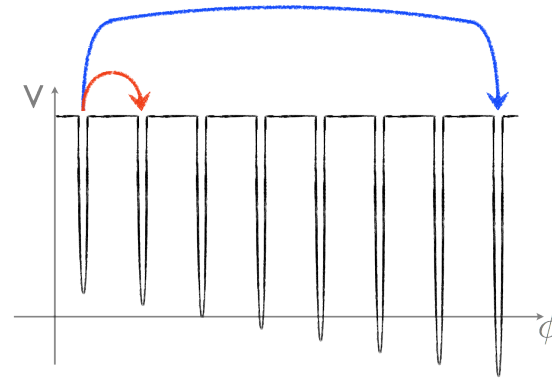
MULTIFLUX decay



MULTIFLUX decay



1. This potential gives small steps



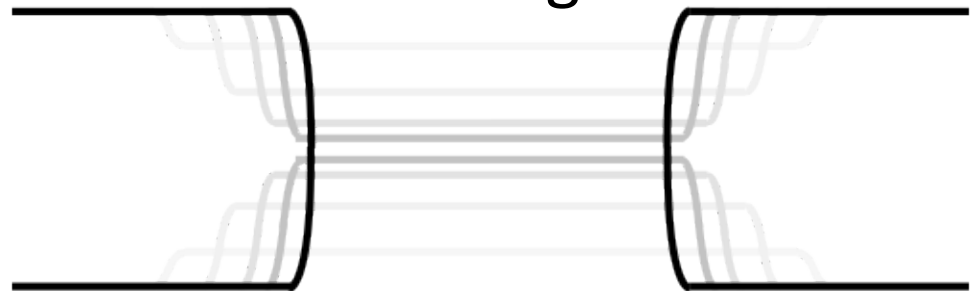
2. Flux compactifications with many fluxes give giant leaps

3. Monkey branes



4. Adding back-reaction

5. The giantest leap of all is a bubble of nothing



Lifting to Higher Dimensions

A single higher dimensional flux that wraps many distinct cycles in the internal manifold.



Many fluxes

Bousso-Polchinski model

To what extent does the story stay true?

(fix the radion)

$$F^2 = \sum_{i=1}^{\mathfrak{n}} g_i^2 N_i^2 \text{ ?}$$

$$T \sim \left(\sum_{i=1}^{\mathfrak{n}} g_i^2 n_i^2 \right)^{1/2} \text{ ?}$$

Lifting to Higher Dimensions

q -form \mathbf{F}

flat m -torus T^m

$$\mathbf{F} = \sum_{b=1}^{\mathfrak{n}} F_b dw_{i_1} \wedge \cdots \wedge dw_{i_q}$$

\mathbf{F} must point ALL legs down the extra dimensions

| | t | r | θ | ψ | w_1 | \dots | w_q | \dots | w_m |
|-------------------|-----|-----|----------|--------|-------|---------|-------|---------|-------|
| \mathbf{F} | | | | | X | X | X | | |
| $\star\mathbf{F}$ | X | X | X | X | | | | X | X |

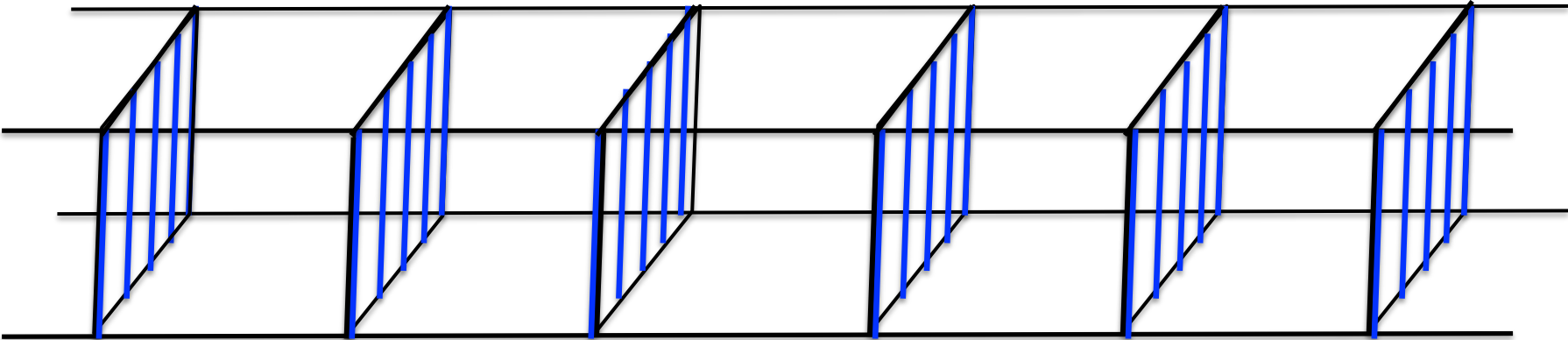
$$\mathfrak{n} = \binom{m}{q}$$

$$\frac{1}{2}\mathbf{F}^2 \rightarrow \sum_{b=1}^{\mathfrak{n}} F_b^2$$

How to discharge?

Consider flux tunneling in 1+1+2 dimensions

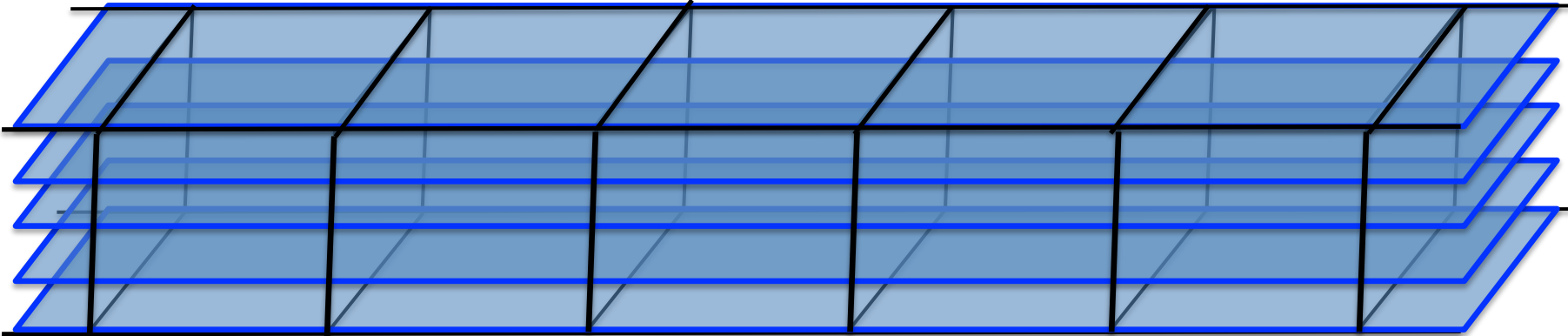
space time sphere



N units of flux

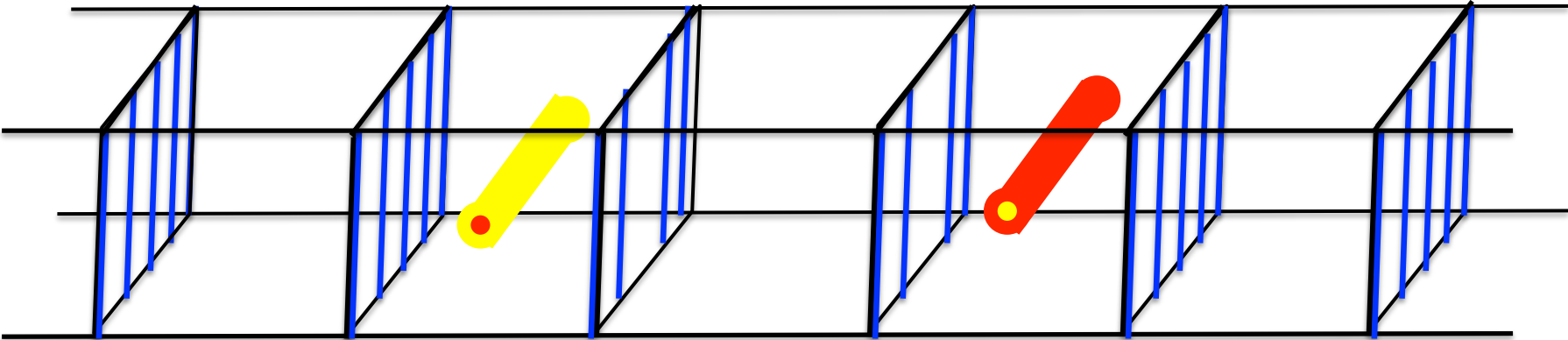
N units of flux

N units of flux

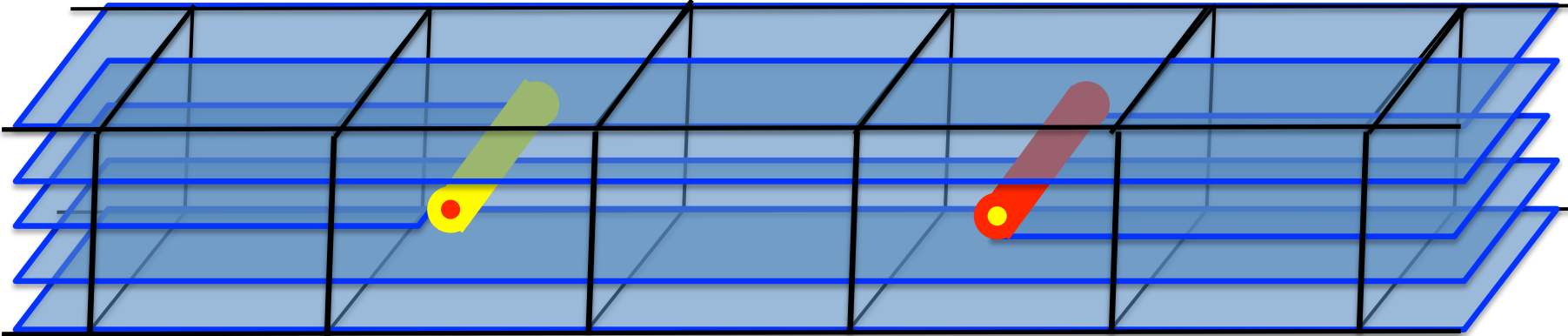


Consider flux tunneling in 1+1+2 dimensions

space time sphere



N units of flux N-1 units of flux N units of flux



Lifting to Higher Dimensions

q -form \mathbf{F}

flat m -torus T^m

$$\mathbf{F} = \sum_{b=1}^n F_b dw_{i_1} \wedge \cdots \wedge dw_{i_q}$$

\mathbf{F} must point ALL legs down the extra dimensions

| | t | r | θ | ψ | w_1 | \dots | w_q | \dots | w_m |
|-------------------|-----|-----|----------|--------|-------|---------|-------|---------|-------|
| \mathbf{F} | | | | | X | X | X | | |
| $\star\mathbf{F}$ | X | X | X | X | | | | X | X |
| brane | X | | X | X | | | | X | X |

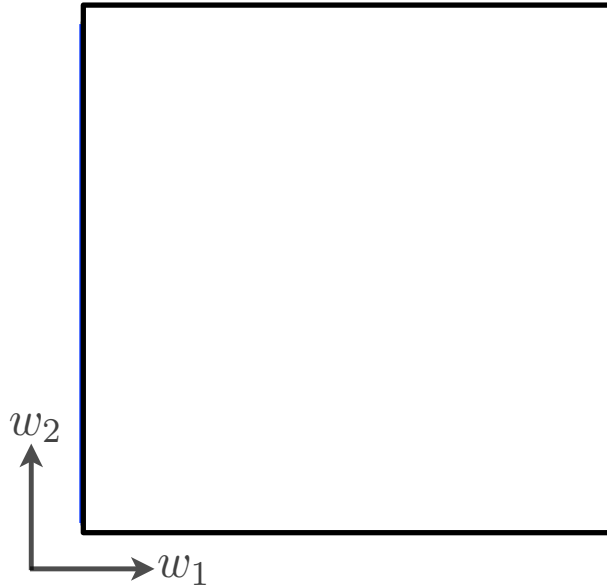
Brane makes a bubble in the extended directions
wraps the dual cycle

Let's see how this works in some specific cases

Lifting to Higher Dimensions

1-form \mathbf{F}
flat 2-torus T^2

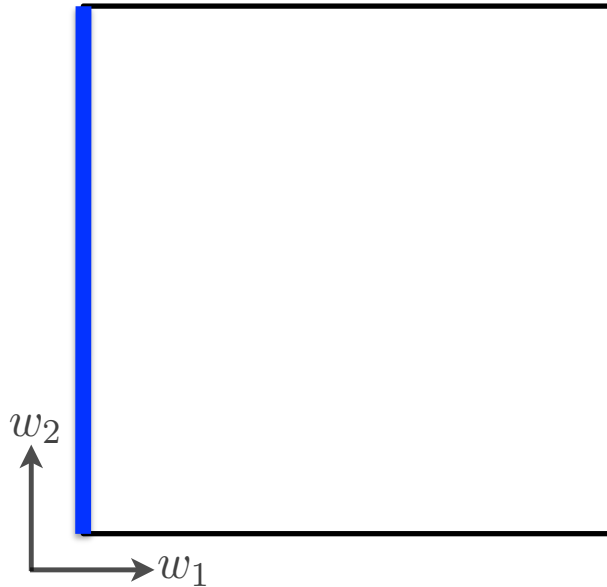
$$\mathfrak{N} = \begin{pmatrix} 2 \\ 1 \end{pmatrix} = 2$$



Lifting to Higher Dimensions

1-form \mathbf{F}
flat 2-torus T^2

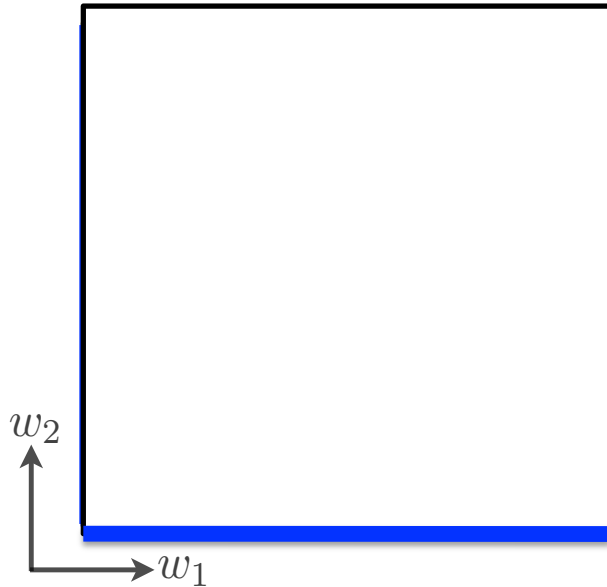
$$\mathfrak{N} = \begin{pmatrix} 2 \\ 1 \end{pmatrix} = 2$$



$$\Delta \mathbf{F} = dw_1$$

Lifting to Higher Dimensions

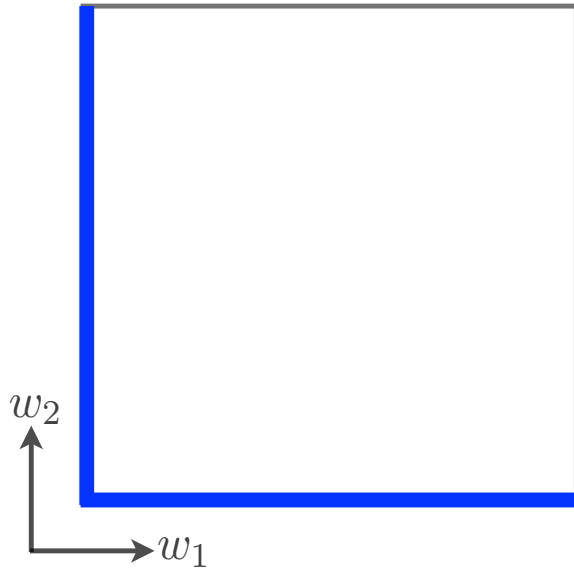
1-form \mathbf{F}
flat 2-torus T^2 $\mathfrak{N} = \begin{pmatrix} 2 \\ 1 \end{pmatrix} = 2$



$$\Delta \mathbf{F} = dw_2$$

Lifting to Higher Dimensions

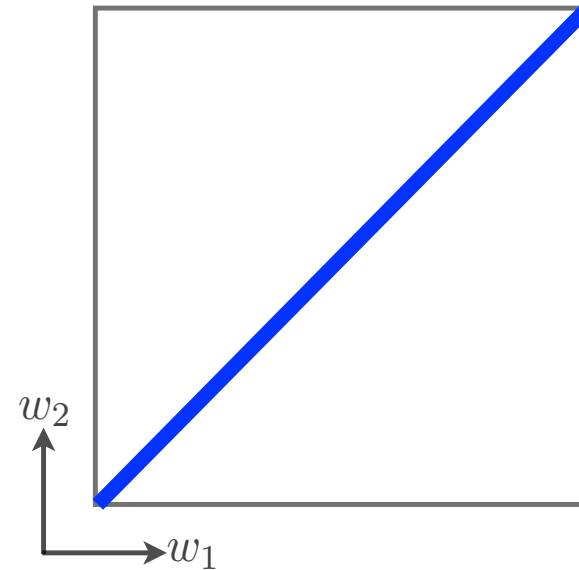
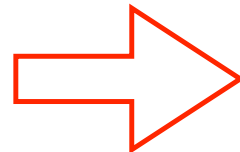
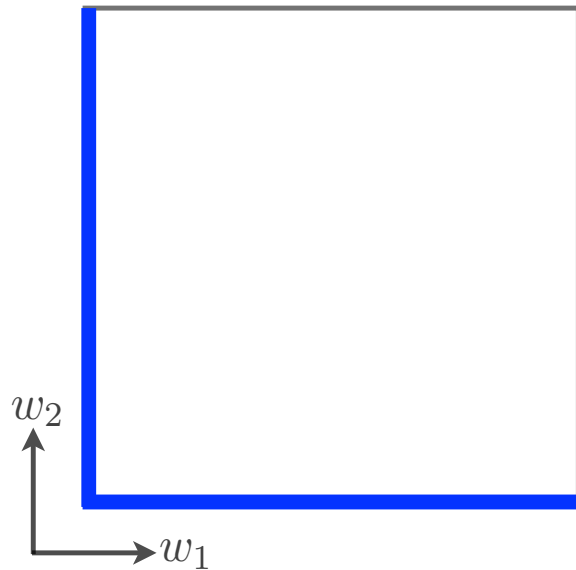
1-form \mathbf{F}
flat 2-torus T^2 $\mathfrak{n} = \begin{pmatrix} 2 \\ 1 \end{pmatrix} = 2$



Drop one unit of both $T \sim n?$

Lifting to Higher Dimensions

1-form \mathbf{F}
flat 2-torus T^2 $\mathfrak{n} = \begin{pmatrix} 2 \\ 1 \end{pmatrix} = 2$

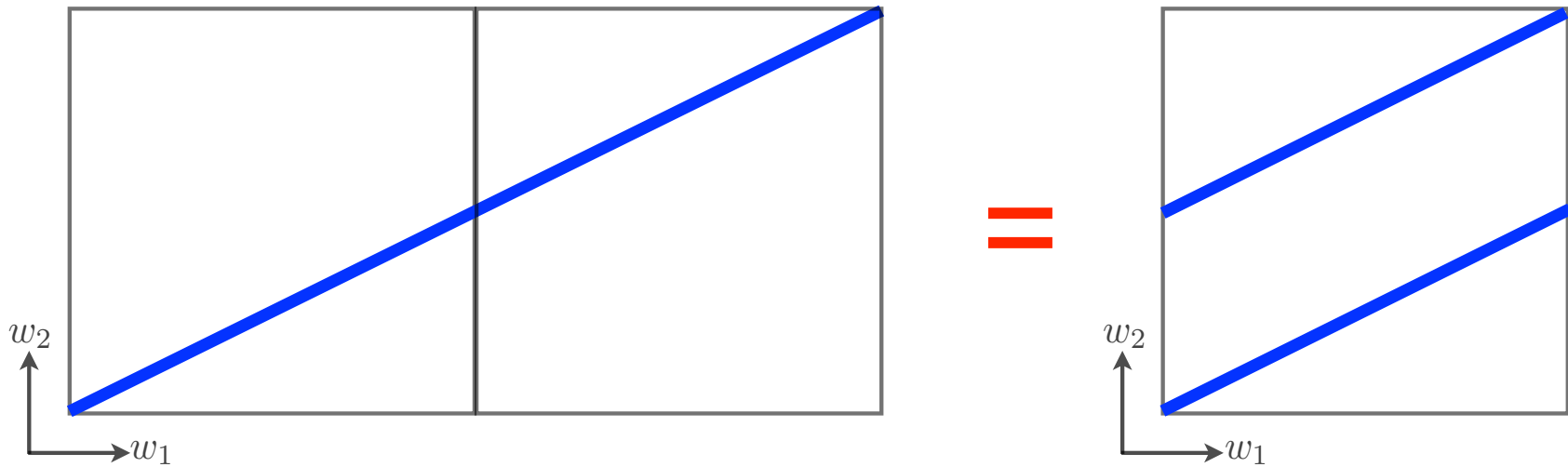


Drop one unit of both ~~$T \sim n$~~ ?

$$T \sim \sqrt{n}$$

Lifting to Higher Dimensions

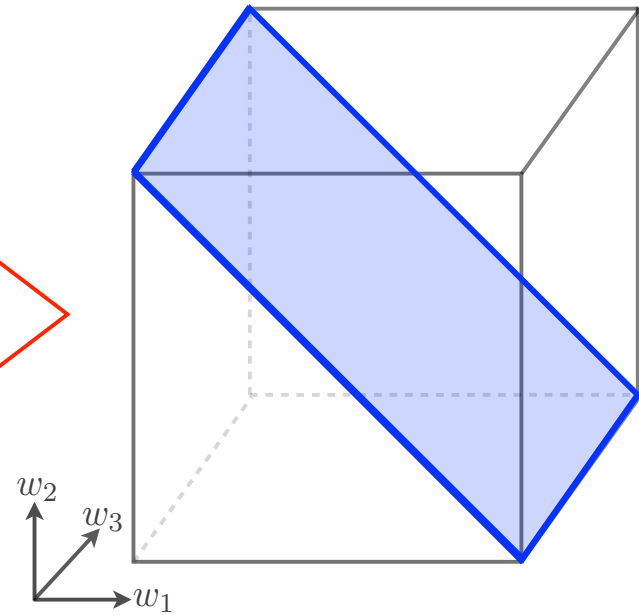
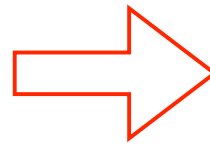
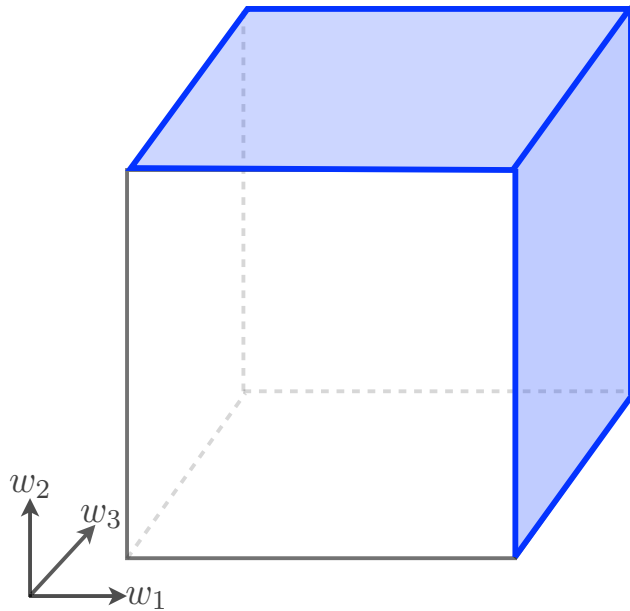
1-form \mathbf{F}
flat 2-torus T^2 $\mathfrak{n} = \begin{pmatrix} 2 \\ 1 \end{pmatrix} = 2$



$$T \sim \left(\sum_{i=1}^{\mathfrak{n}} g_i^2 n_i^2 \right)^{1/2} = \text{Pythagoras' Theorem}$$

Lifting to Higher Dimensions

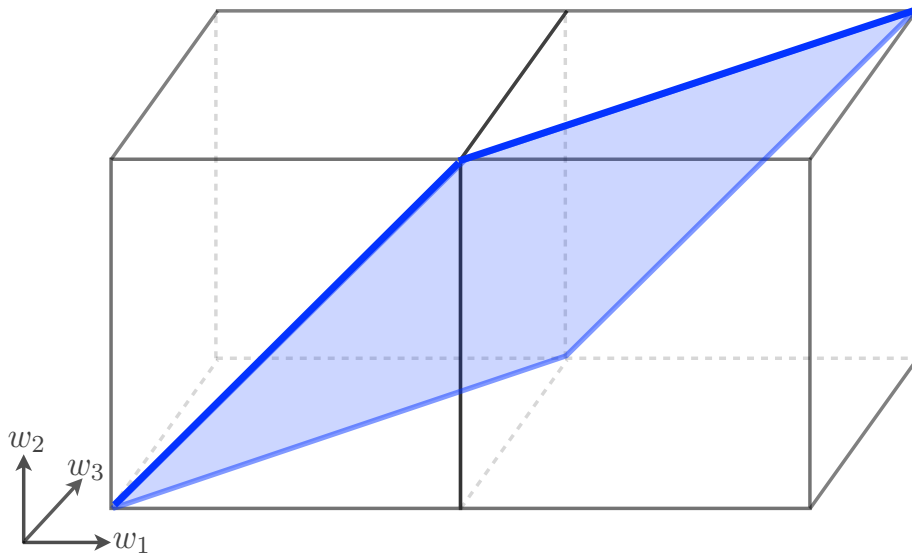
1-form \mathbf{F}
flat 3-torus T^3 $\mathfrak{n} = \begin{pmatrix} 3 \\ 1 \end{pmatrix} = 3$



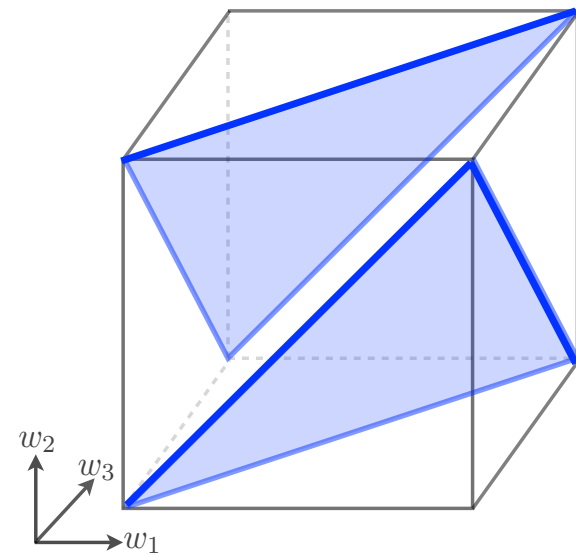
Still true

Lifting to Higher Dimensions

1-form \mathbf{F}
flat 3-torus T^3 $\mathfrak{N} = \begin{pmatrix} 3 \\ 1 \end{pmatrix} = 3$



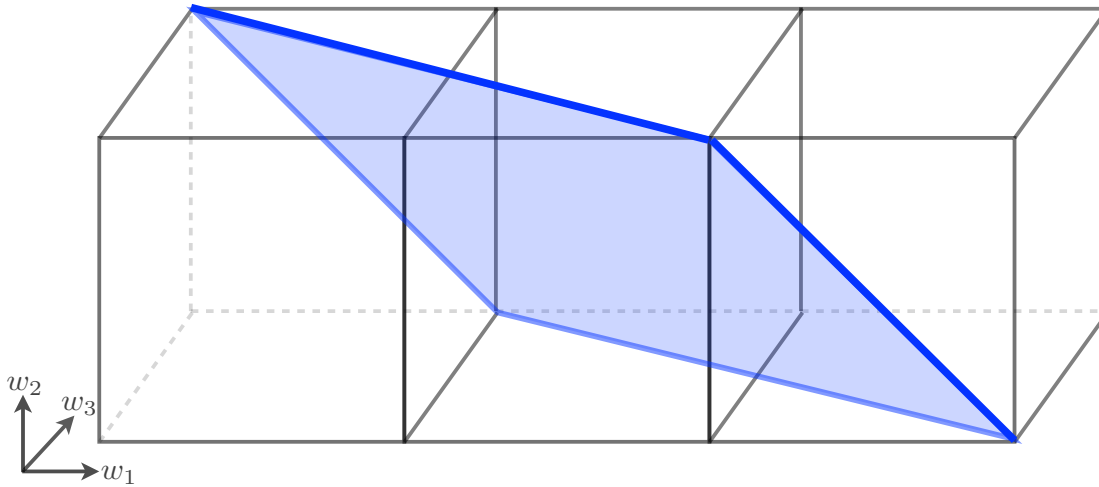
=



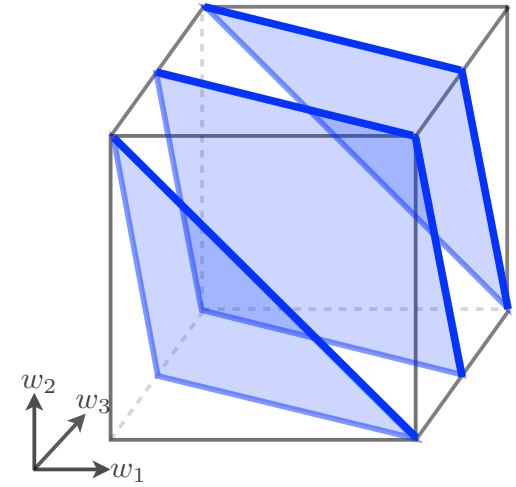
Still true

Lifting to Higher Dimensions

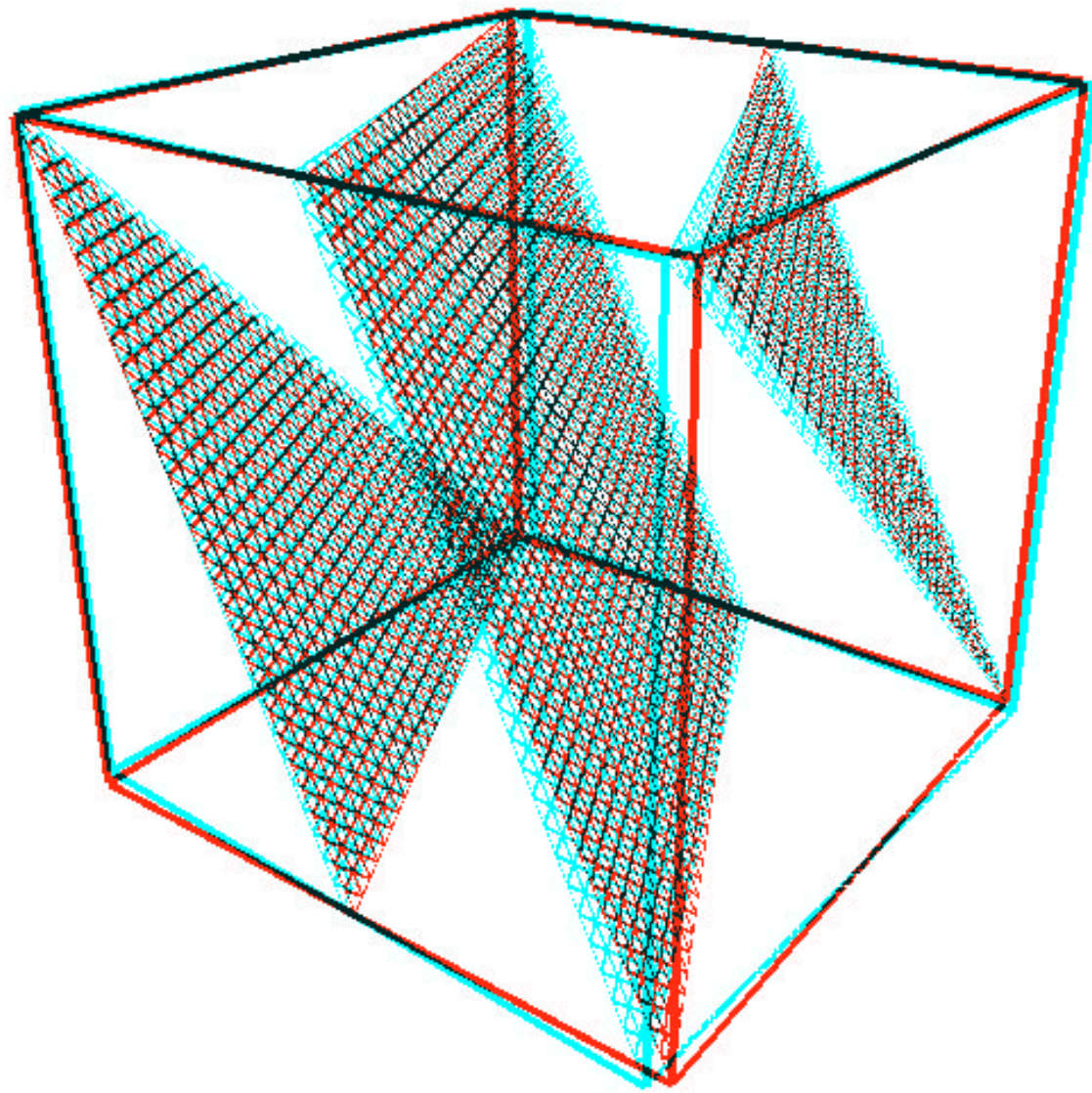
1-form \mathbf{F}
flat 3-torus T^3 $\mathfrak{N} = \begin{pmatrix} 3 \\ 1 \end{pmatrix} = 3$



=



Still true



Lifting to Higher Dimensions

2-form \mathbf{F}
flat 4-torus T^3

$$\mathfrak{n} = \binom{4}{2} = 6$$

$$\Delta \mathbf{F} = dw_1 \wedge dw_2 + dw_3 \wedge dw_4$$

Lifting to Higher Dimensions

2-form \mathbf{F}
flat 4-torus T^4

$$\mathfrak{n} = \binom{4}{2} = 6$$

$$\Delta \mathbf{F} = dw_1 \wedge dw_2 + dw_3 \wedge dw_4$$

No longer true!

Requires 2 branes

$$T \sim n$$

Lifting to Higher Dimensions

The general case

$$d\mathbf{F} = \star \mathbf{j}$$

$$\Delta \mathbf{F} = \star g \mathbf{Y}_1 \wedge \cdots \wedge \mathbf{Y}_q$$

This tells us when a single, flat brane can be found.

DECOMPOSABILITY

When it can be, $T \sim \sqrt{n}$.

Otherwise worse

Lifting to Higher Dimensions

The general case

T is intermediate

$$T \sim \sqrt{n} \quad T \sim n$$

Clustering

| | w_1 | w_2 | w_3 | w_4 |
|---|-------|-------|-------|-------|
| 1 | x | x | | |
| 2 | x | | x | |
| 3 | | x | x | |
| 4 | | | x | x |
| 5 | | x | | x |
| 6 | x | | | x |

Ideal configuration is CURVED

Breaks **rotation** symmetry of landscape

$$\Delta \mathbf{F} = dw_1 \wedge dw_2 + dw_2 \wedge dw_3$$



$$\Delta \mathbf{F} = dw_1 \wedge dw_2 + dw_3 \wedge dw_4$$



Breaks **reflection** symmetry of landscape

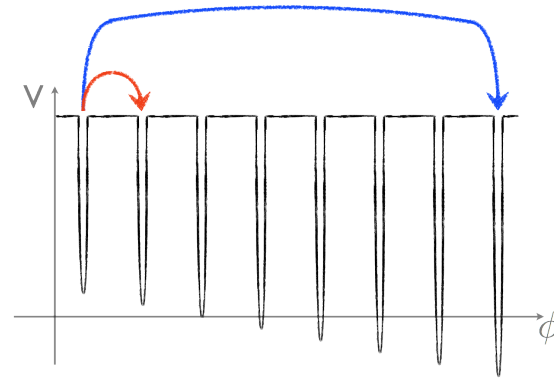
$$\Delta \mathbf{F} = dw_1 \wedge dw_3 + dw_1 \wedge dw_4 + dw_2 \wedge dw_3 + dw_2 \wedge dw_4$$



$$\Delta \mathbf{F} = dw_1 \wedge dw_3 + dw_1 \wedge dw_4 - dw_2 \wedge dw_3 + dw_2 \wedge dw_4$$



1. This potential gives small steps



2. Flux compactifications with many fluxes give giant leaps

3. Monkey branes



4. Adding back-reaction

5. The giantest leap of all is a bubble of nothing

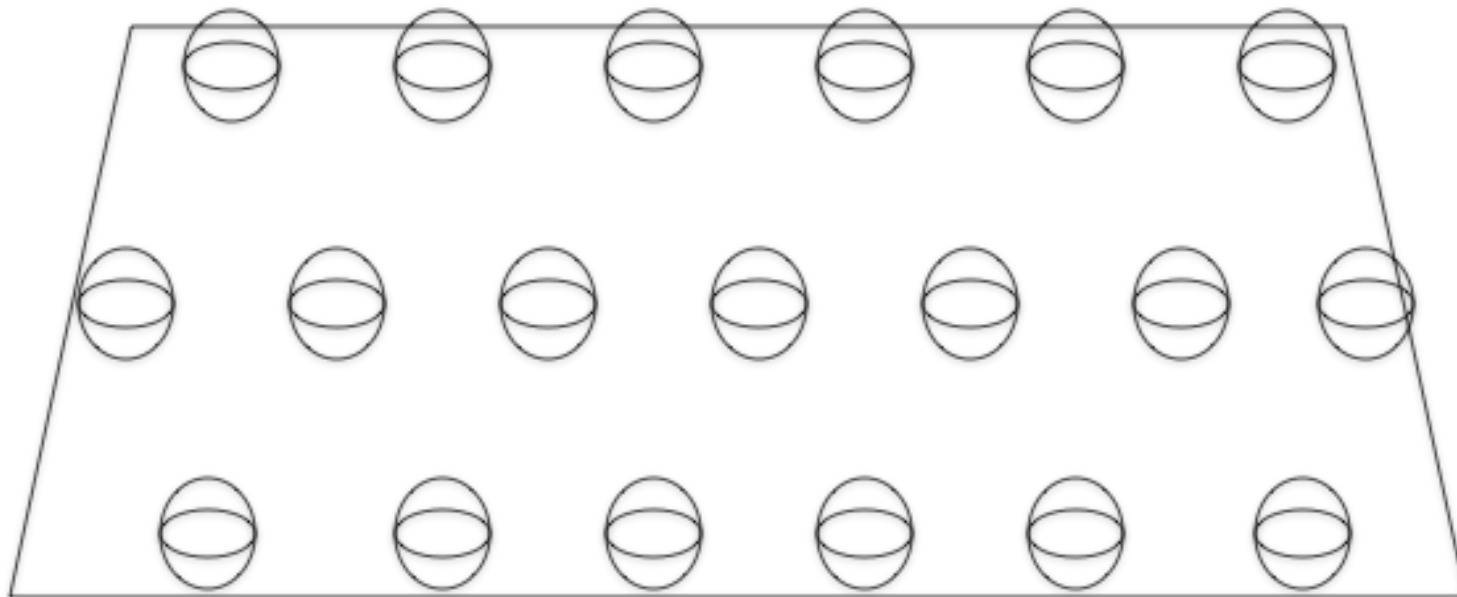


Simplest possible model with **stabilized XD**s that supports **Minkowski and de Sitter** 4D slices is **6D Einstein-Maxwell**

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu + R^2 d\Omega_2^2$$

4D dS, Min, or AdS

2D sphere



three ingredients:

- spatial **curvature** of 2-spheres
- 2-form F_{ab} **flux** wrapping 2-spheres
- 6D **cosmo-constant**

i.e. Maxwell
magnetic flux

$$S_{\text{EM}} = \int d^6x \sqrt{-G} \left(\frac{1}{2} \mathcal{R}^{(6)} - \frac{1}{4} F_{AB} F^{AB} - \Lambda_6 \right)$$

Freund Rubin (1980)

Effective 4D theory:

$$ds^2 = e^{-\psi(x)/M} g_{\mu\nu} dx^\mu dx^\nu + e^{\psi(x)/M} R^2 d\Omega_2^2$$

\mathfrak{N} different two-form fluxes, each with field strength

$$\mathbf{F}_i = \frac{g_i N_i}{4\pi} \sin \theta d\theta \wedge d\phi,$$

$$F^2 = \sum_{i=1}^{\mathfrak{N}} g_i^2 N_i^2.$$

$$S = \int d^4x \sqrt{-g} \left(\frac{1}{2} M^2 \mathcal{R}^{(4)} - \frac{1}{2} \partial_\mu \psi \partial^\mu \psi - V(\psi) \right)$$

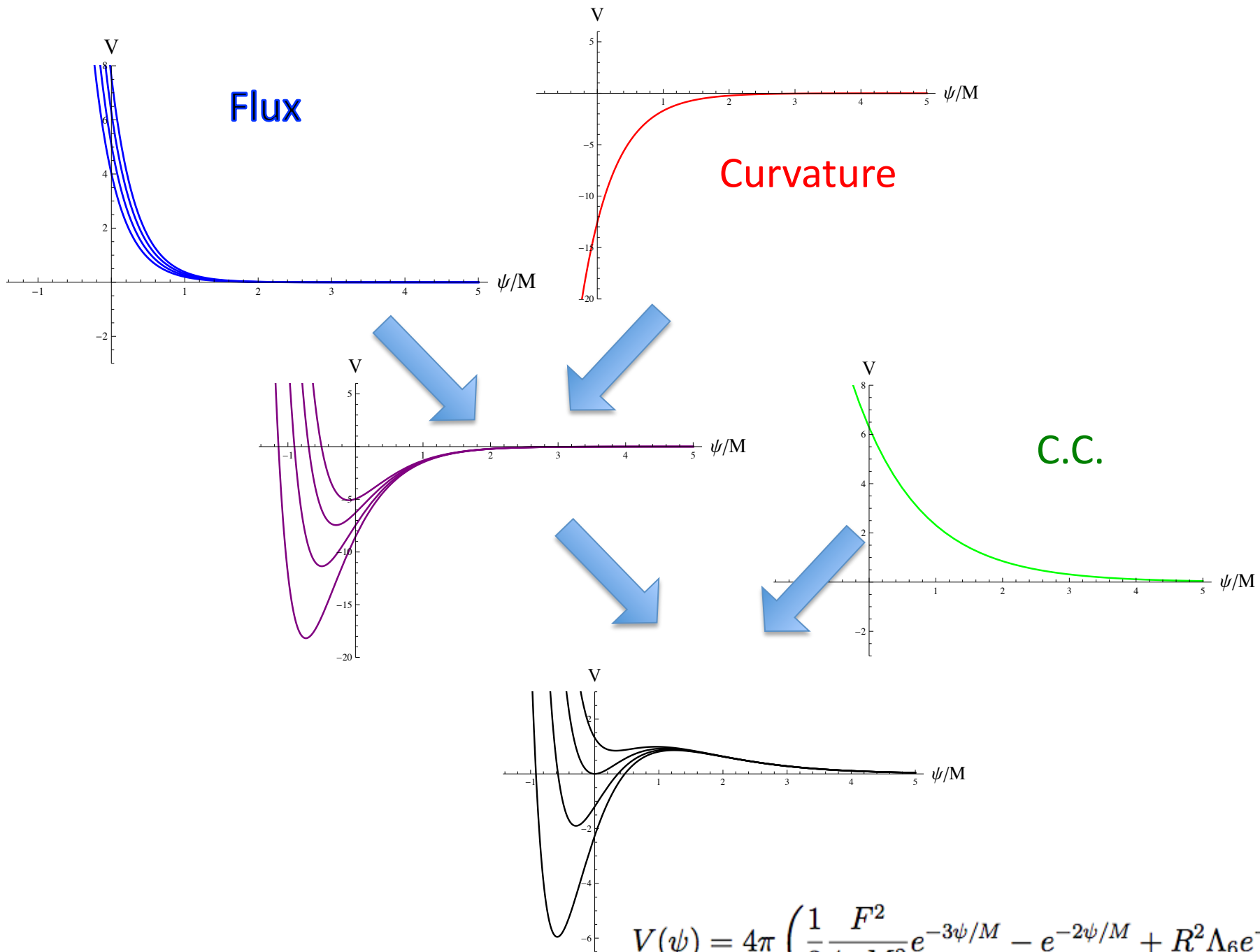
$$V(\psi) = 4\pi \left(\frac{1}{2} \frac{F^2}{4\pi M^2} e^{-3\psi/M} - e^{-2\psi/M} + R^2 \Lambda_6 e^{-\psi/M} \right)$$

canonically
normalized
RADION

FLUX
'repulsive'
short-range

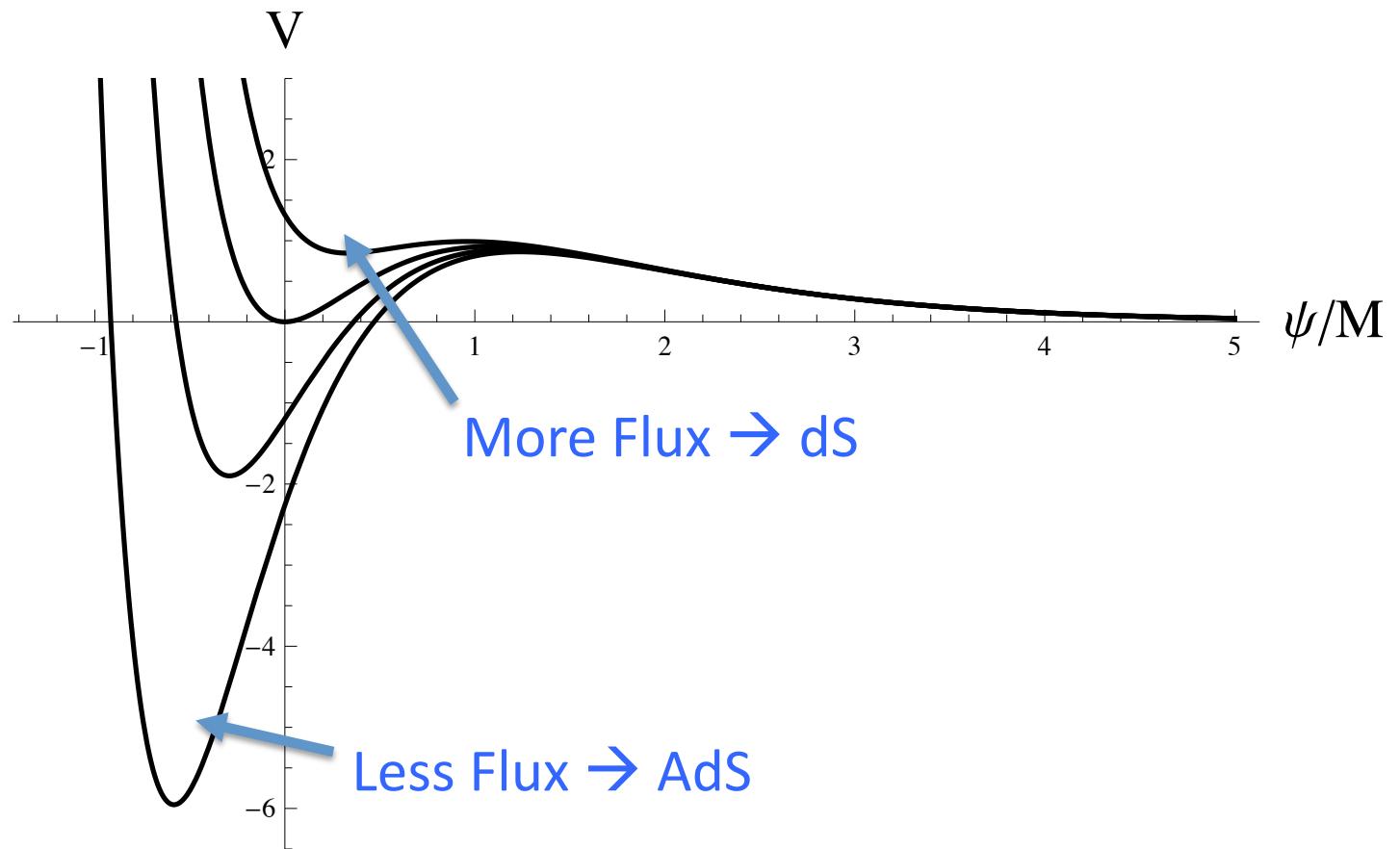
CURVATURE
'attractive'
medium-range

C.C.
'reulsive'
long-range

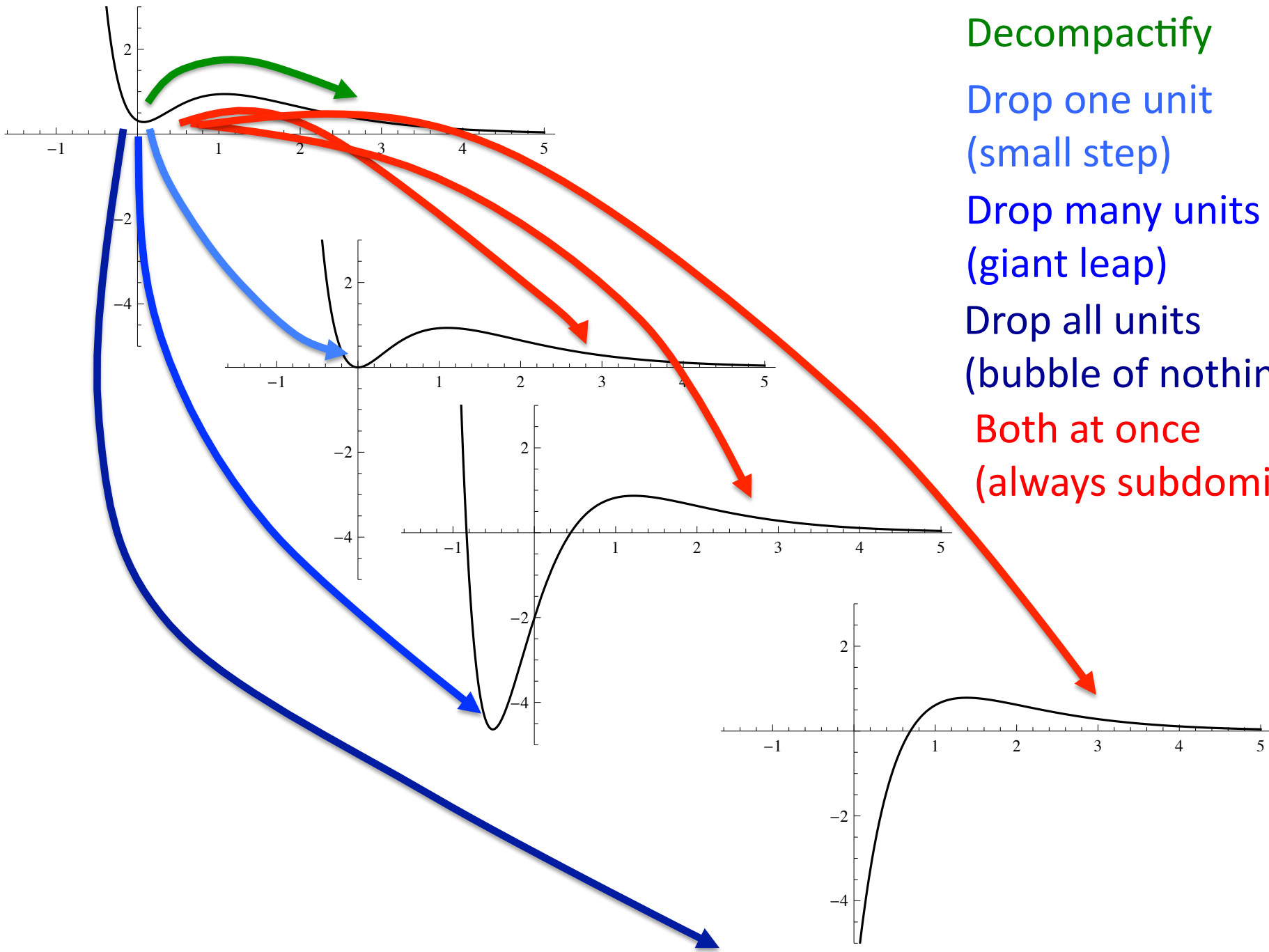


$$V(\psi) = 4\pi \left(\frac{1}{2} \frac{F^2}{4\pi M^2} e^{-3\psi/M} - e^{-2\psi/M} + R^2 \Lambda_6 e^{-\psi/M} \right)$$

A landscape of (discrete) vacua



But how to transition?



Decompactify

Drop one unit
(small step)

Drop many units
(giant leap)

Drop all units
(bubble of nothing)

Both at once
(always subdominant)

2 reasons back-reaction helps giant leaps

Radion mediates attractive force

$$S_{\text{brane tension}} = -T \int_{\Sigma} \sqrt{-\gamma} d^3\xi \quad ds^2 = e^{-\psi(x)/M_4} g_{\mu\nu} dx^\mu dx^\nu + e^{\psi(x)/M_4} R^2 d\Omega_2^2$$



$$S_{\text{brane tension}} = -T e^{-3\psi(\xi)/2M_4} \times (\text{surface area})$$

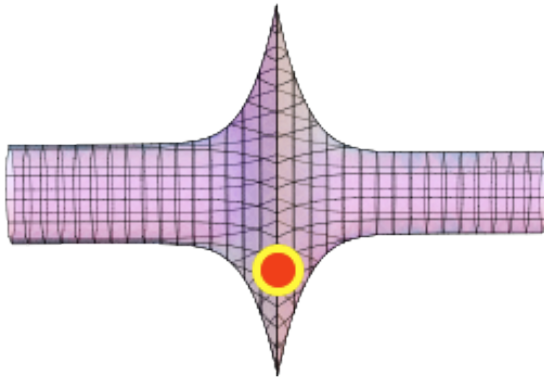
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2 reasons back-reaction helps giant leaps

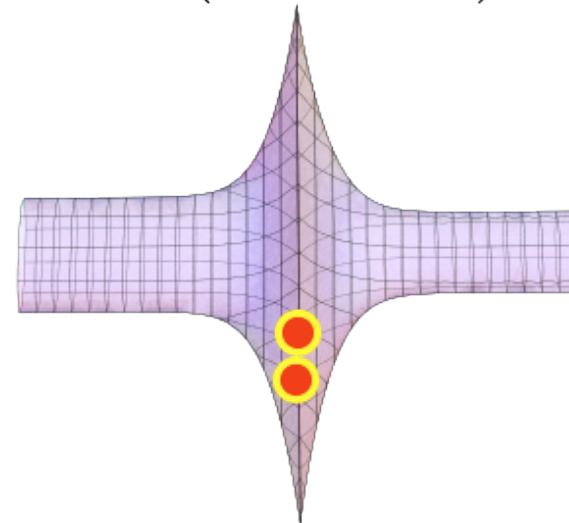
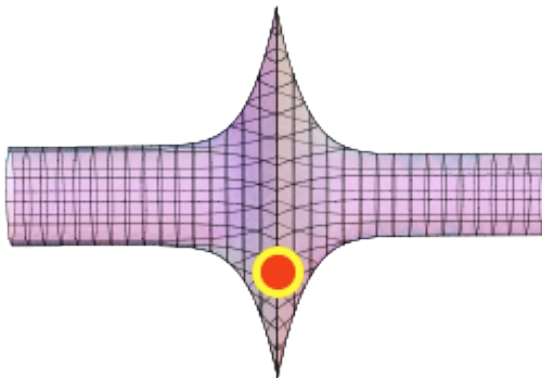
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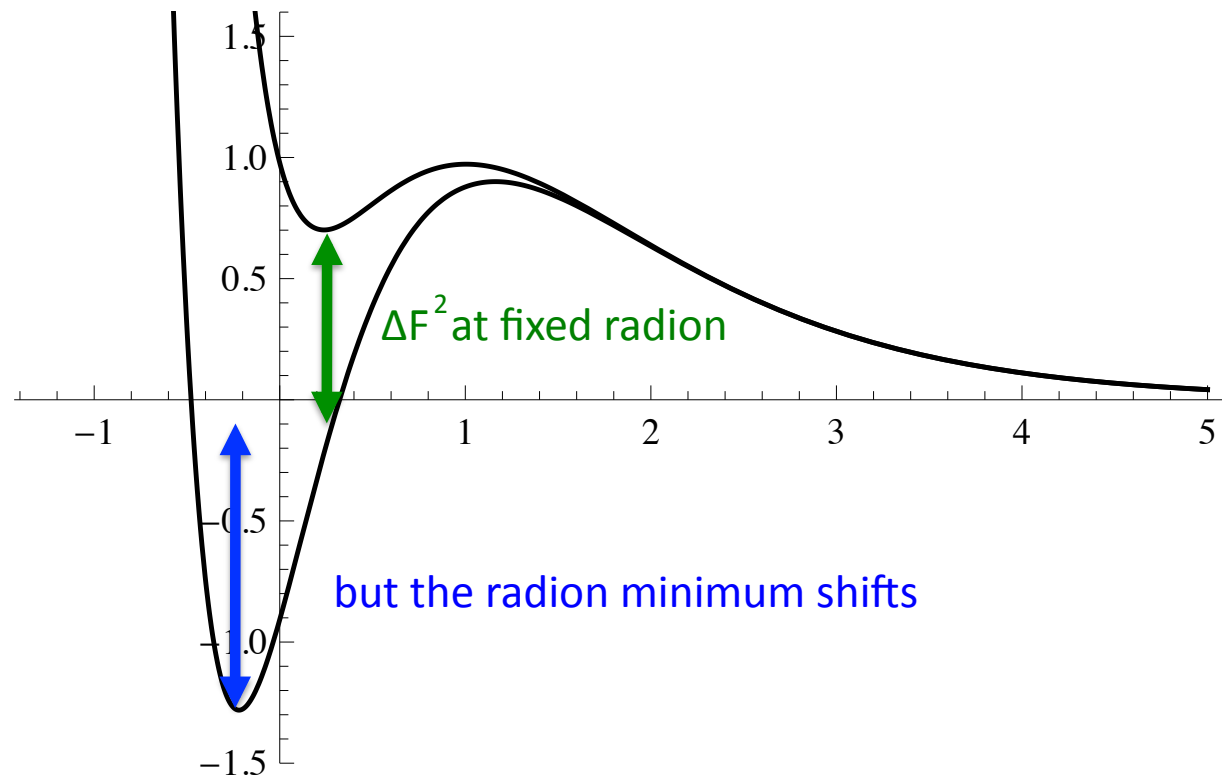


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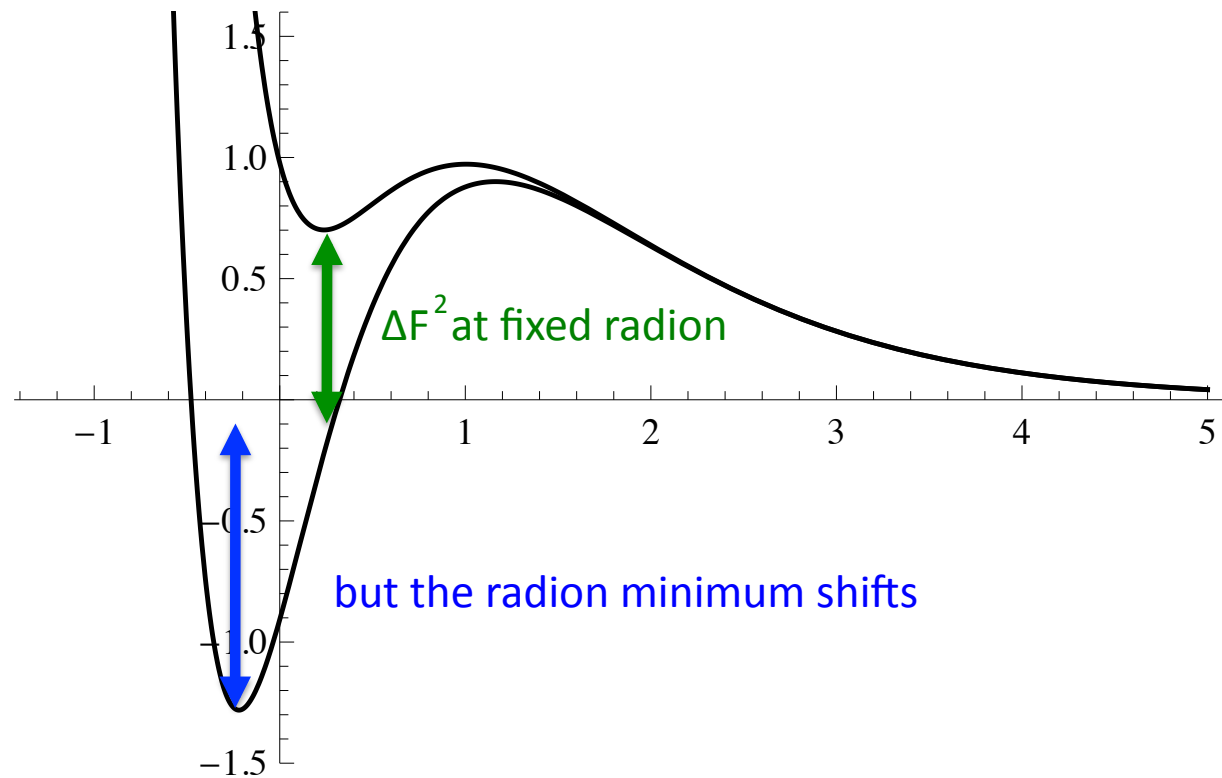
2 reasons back-reaction helps giant leaps

Radion can readjust for extra energy



2 reasons back-reaction helps giant leaps

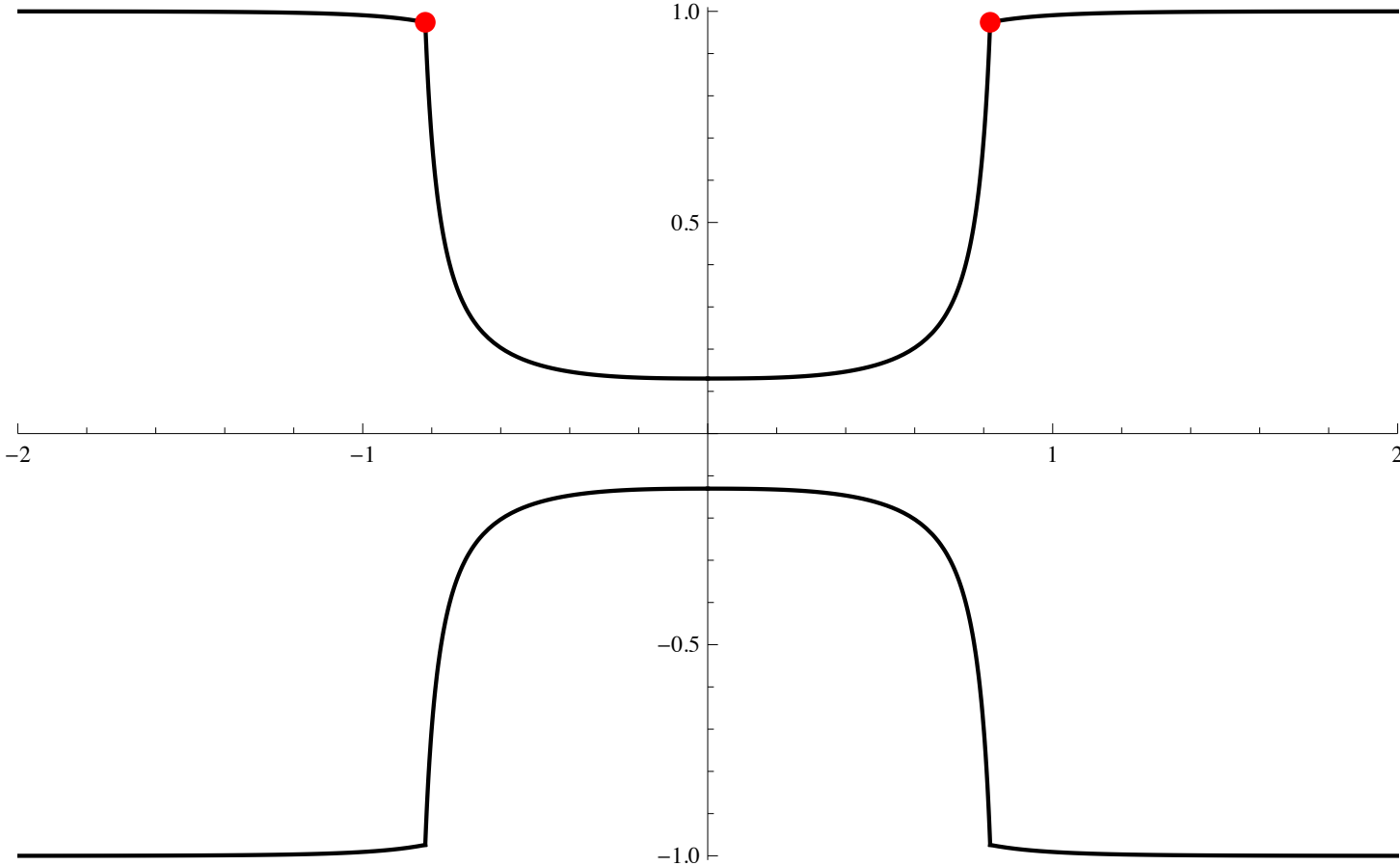
Radion can readjust for extra energy



1 reason it hurts giant leaps:

Thick wall means less region in the true vacuum

Sample instanton profile



Let's look at two extreme examples

$$F^2 = \sum_{i=1}^{\mathfrak{N}} g_i^2 N_i^2$$
$$T = \frac{2}{\sqrt{3}} \left(\sum_{i=1}^{\mathfrak{N}} g_i^2 n_i^2 \right)^{\frac{1}{2}}$$

MONOFLUX

A single type of flux and many units of it

$$\mathfrak{N} = 1 \quad N \gg 1 \quad F^2 = g^2 N^2$$

MULTIFLUX

Many different types of flux, each with a single unit

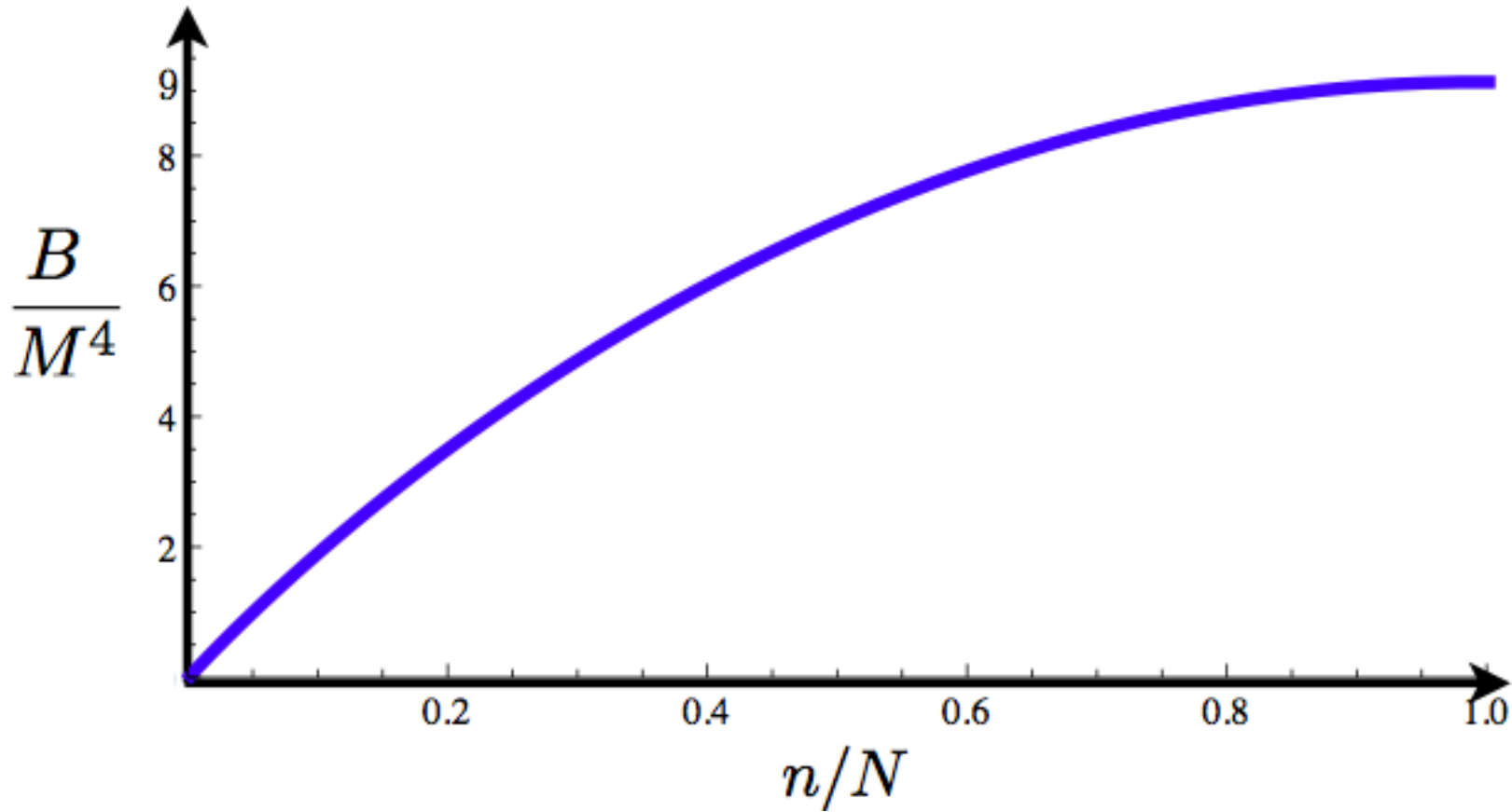
$$\mathfrak{N} \gg 1 \quad N_i = 1 \quad F^2 = g^2 \mathfrak{N}$$

Decay Rates

$$\Gamma \sim e^{-B/\hbar}, \quad B = S_E(\text{instanton}) - S_E(\text{false vacuum})$$

MONOFLUX

From $V=0$

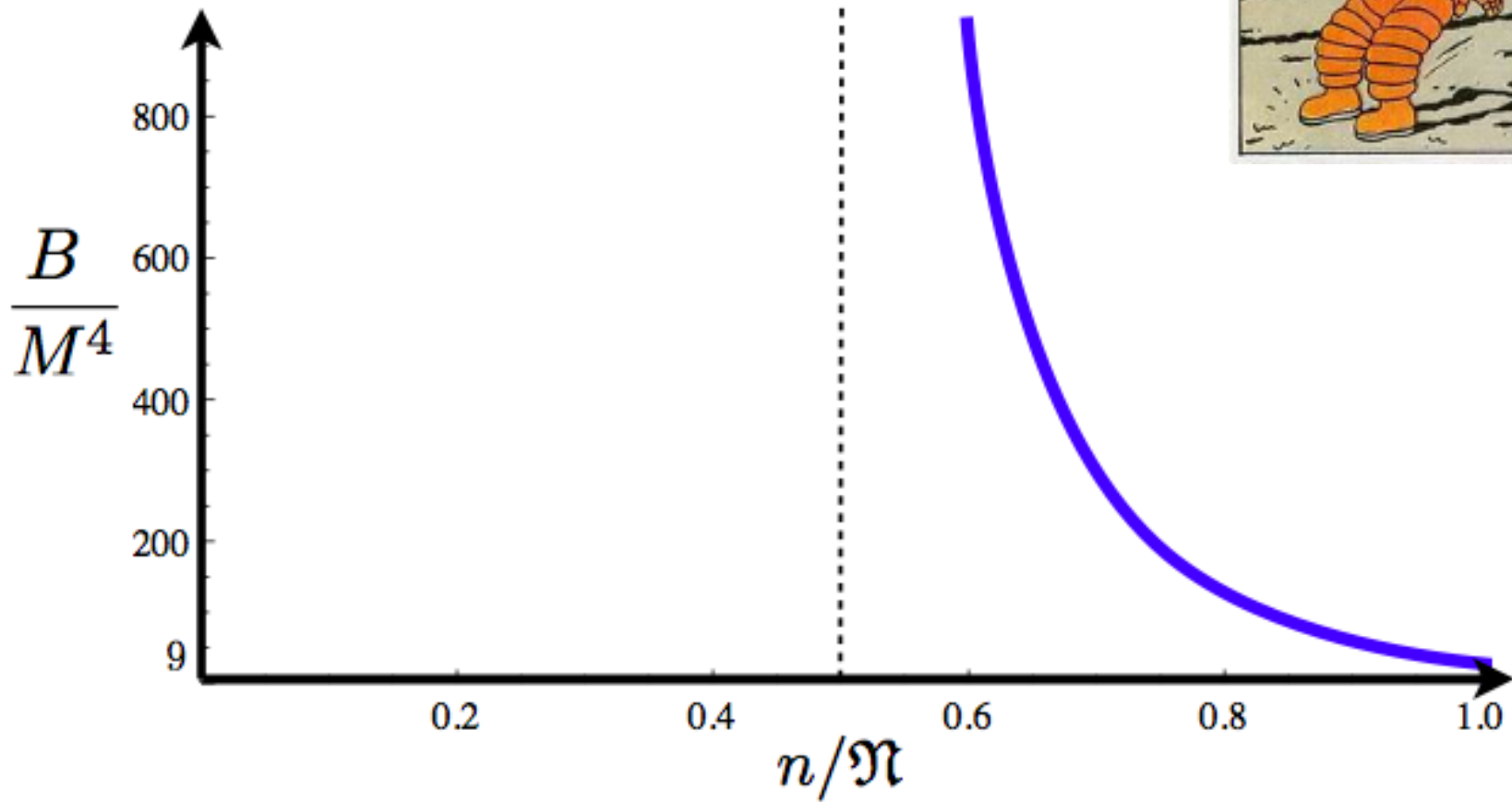


SMALL STEPS WIN

Decay Rates

MULTIFLUX

From $V=0$

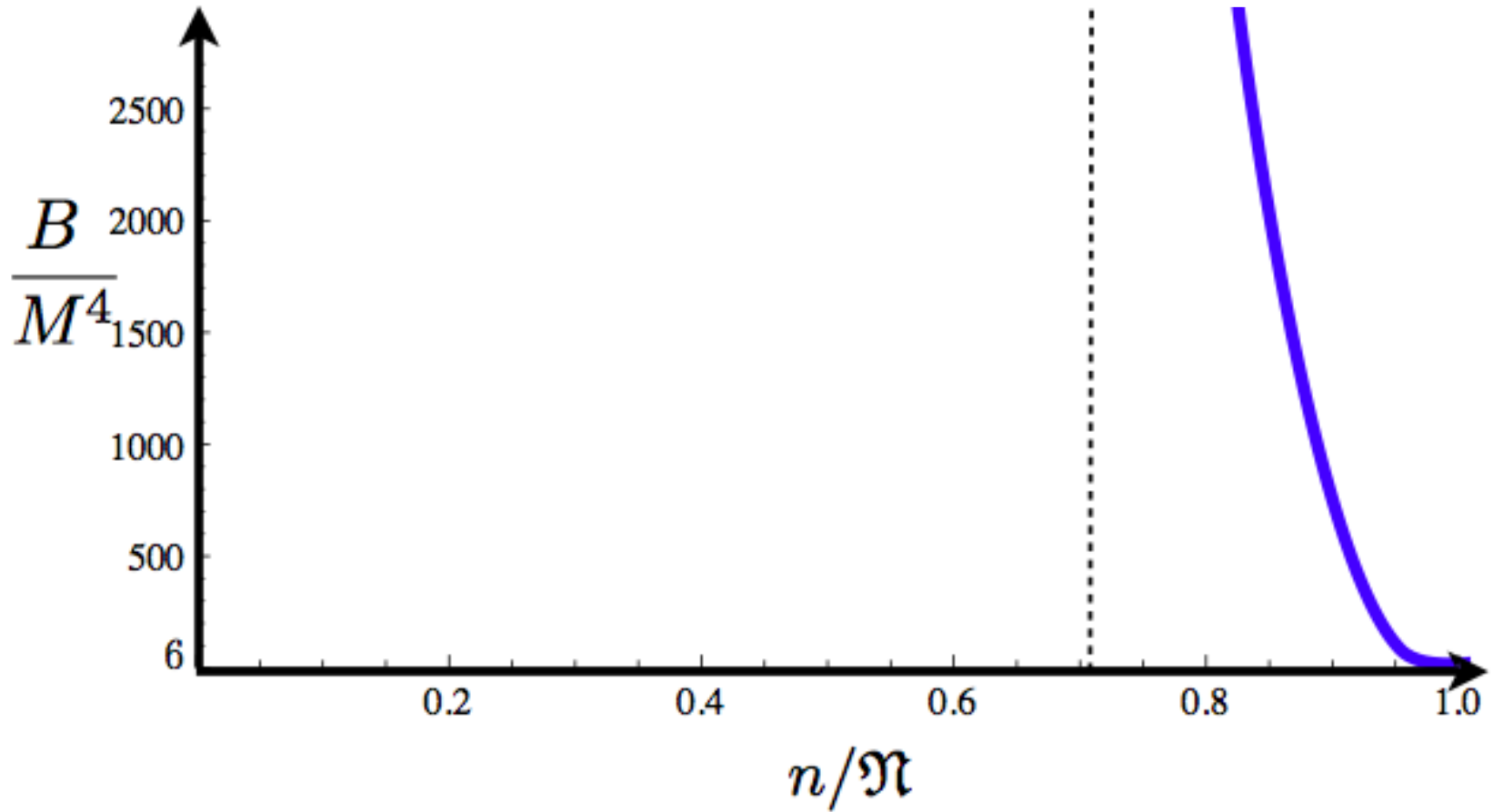


GIANT LEAPS WIN

Decay Rates

MULTIFLUX

From $V < 0$

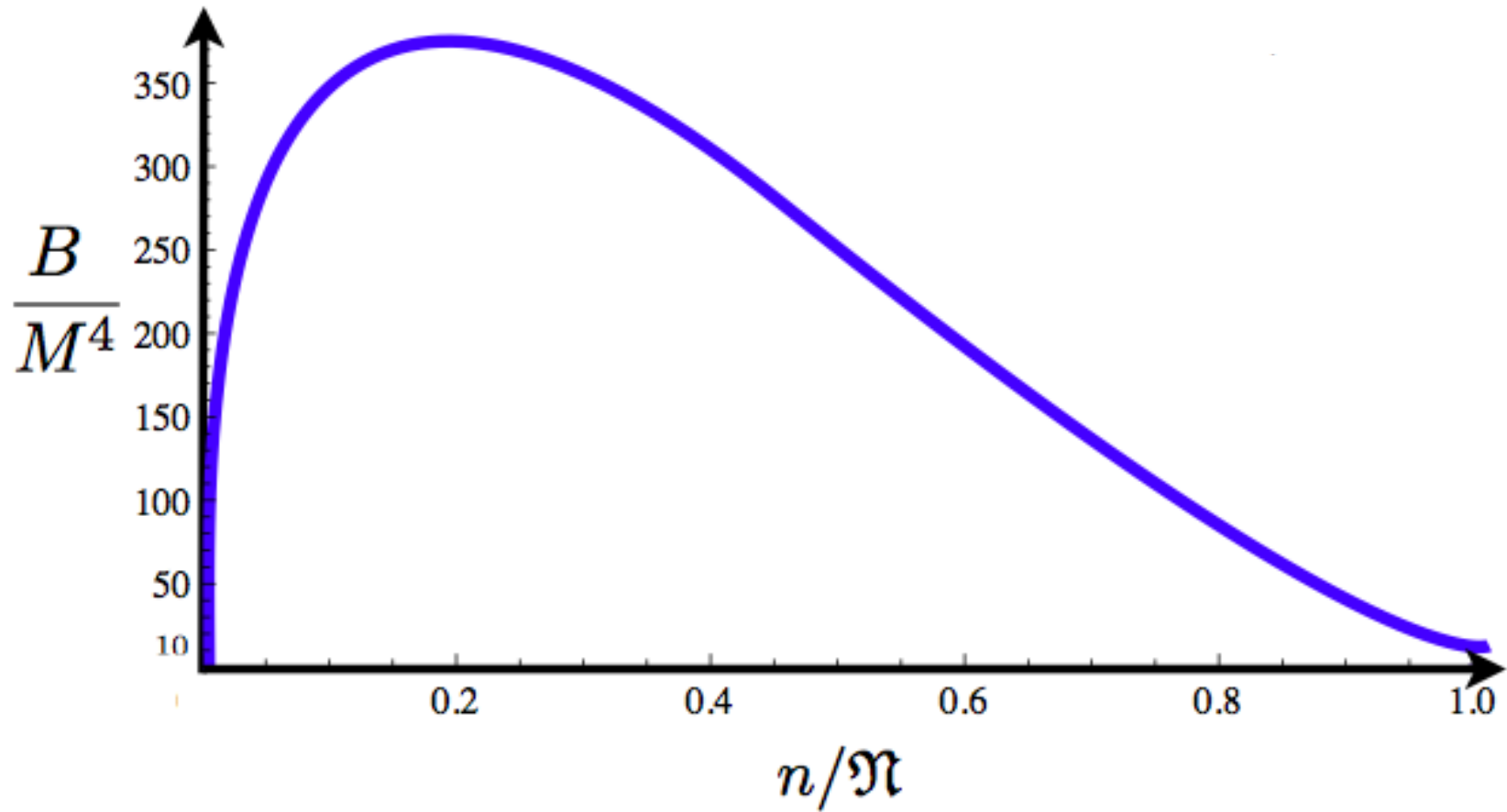


GIANT LEAPS WIN

Decay Rates

MULTIFLUX

From $V > 0$

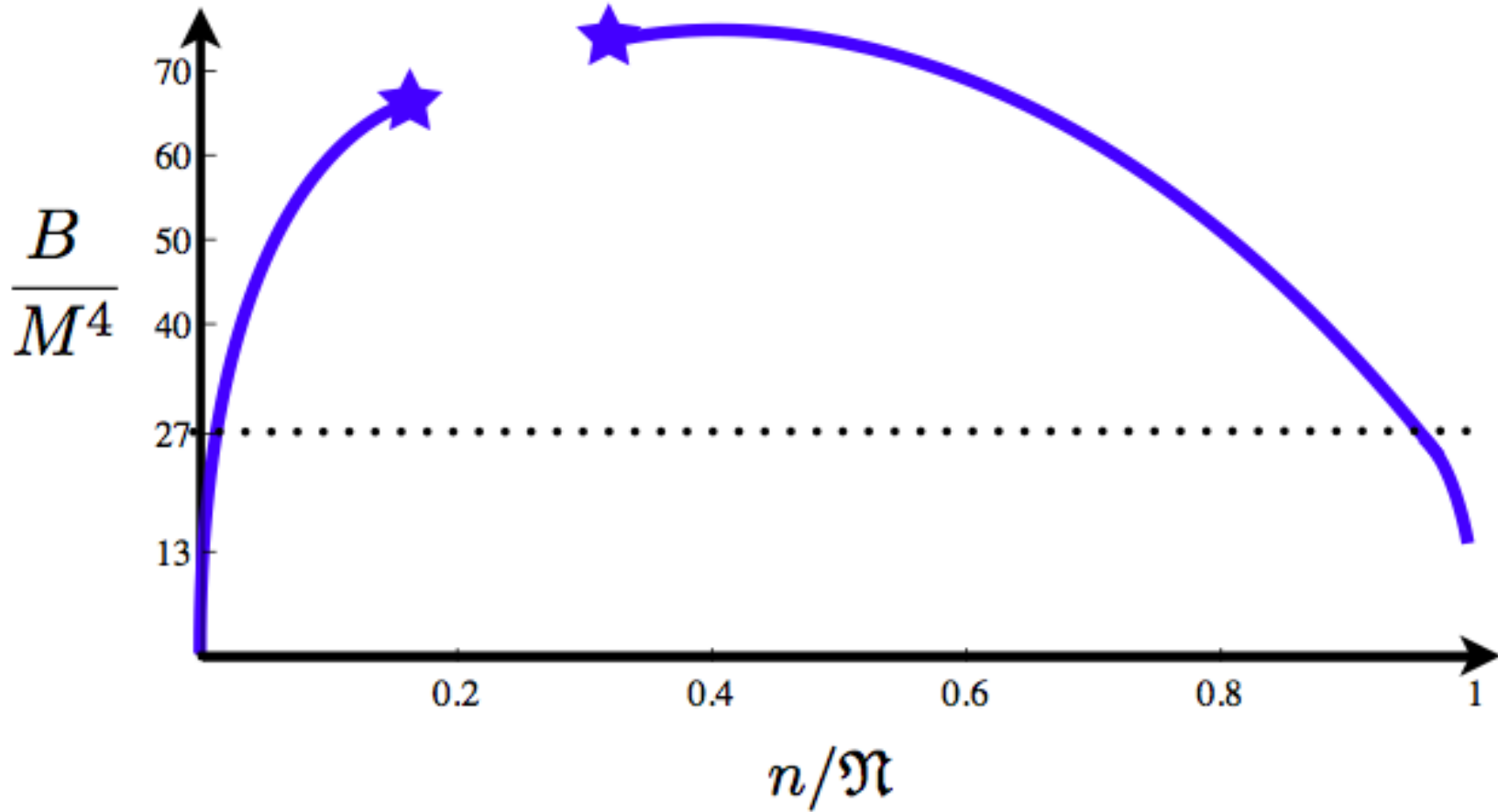


GIANT LEAPS MAY WIN

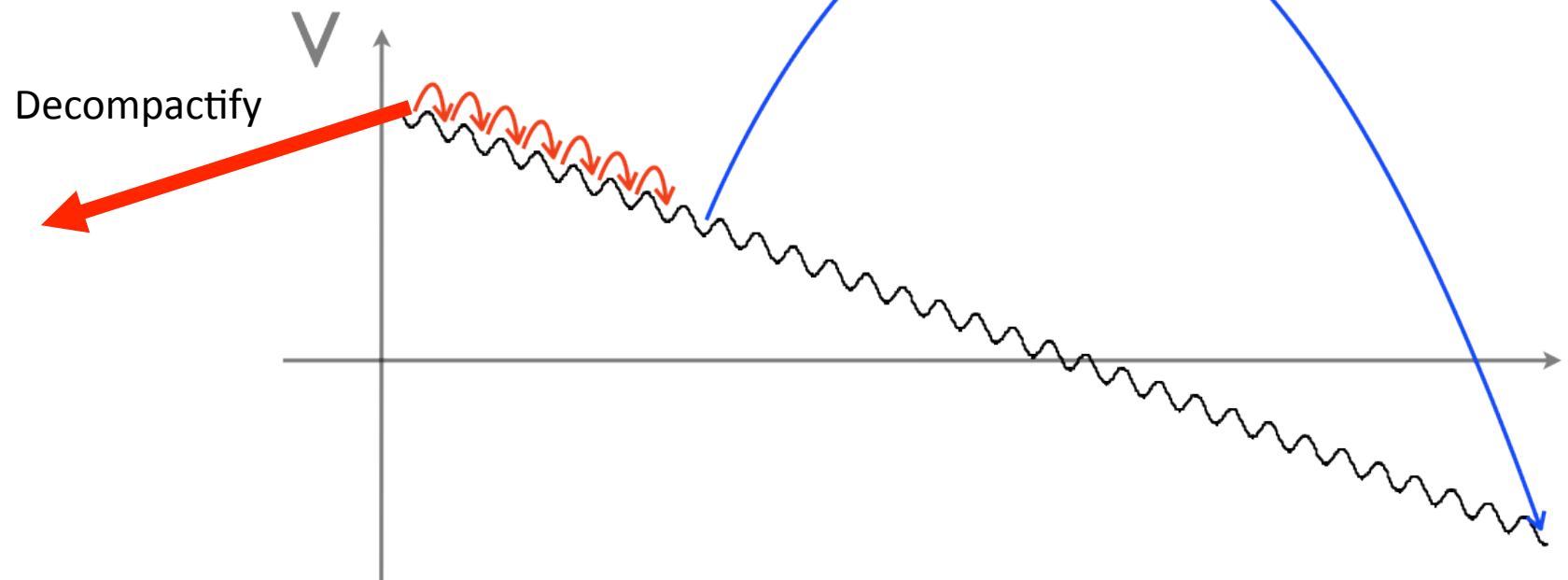
Decay Rates

MULTIFLUX

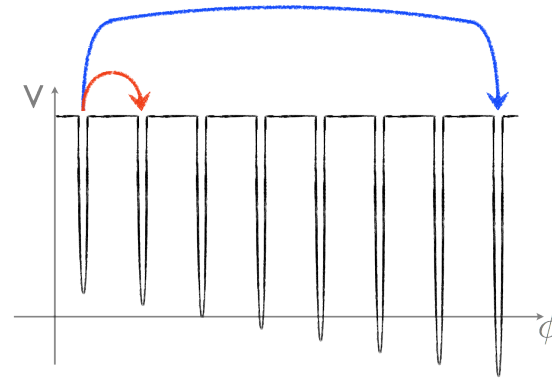
From $V \gg 0$



DECOMPACTIFICATION WINS



1. This potential gives small steps



2. Flux compactifications with many fluxes give giant leaps

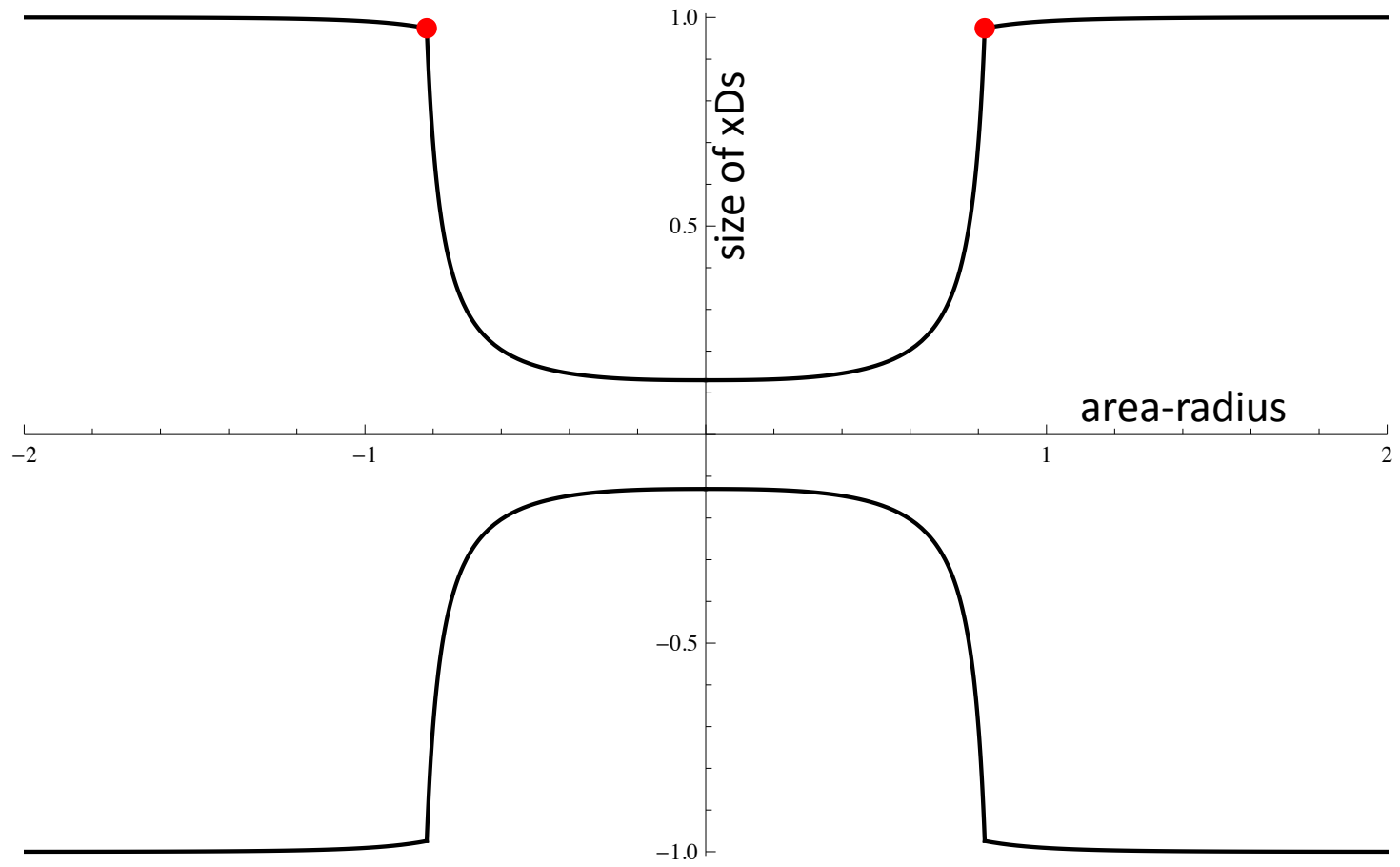
3. Monkey branes

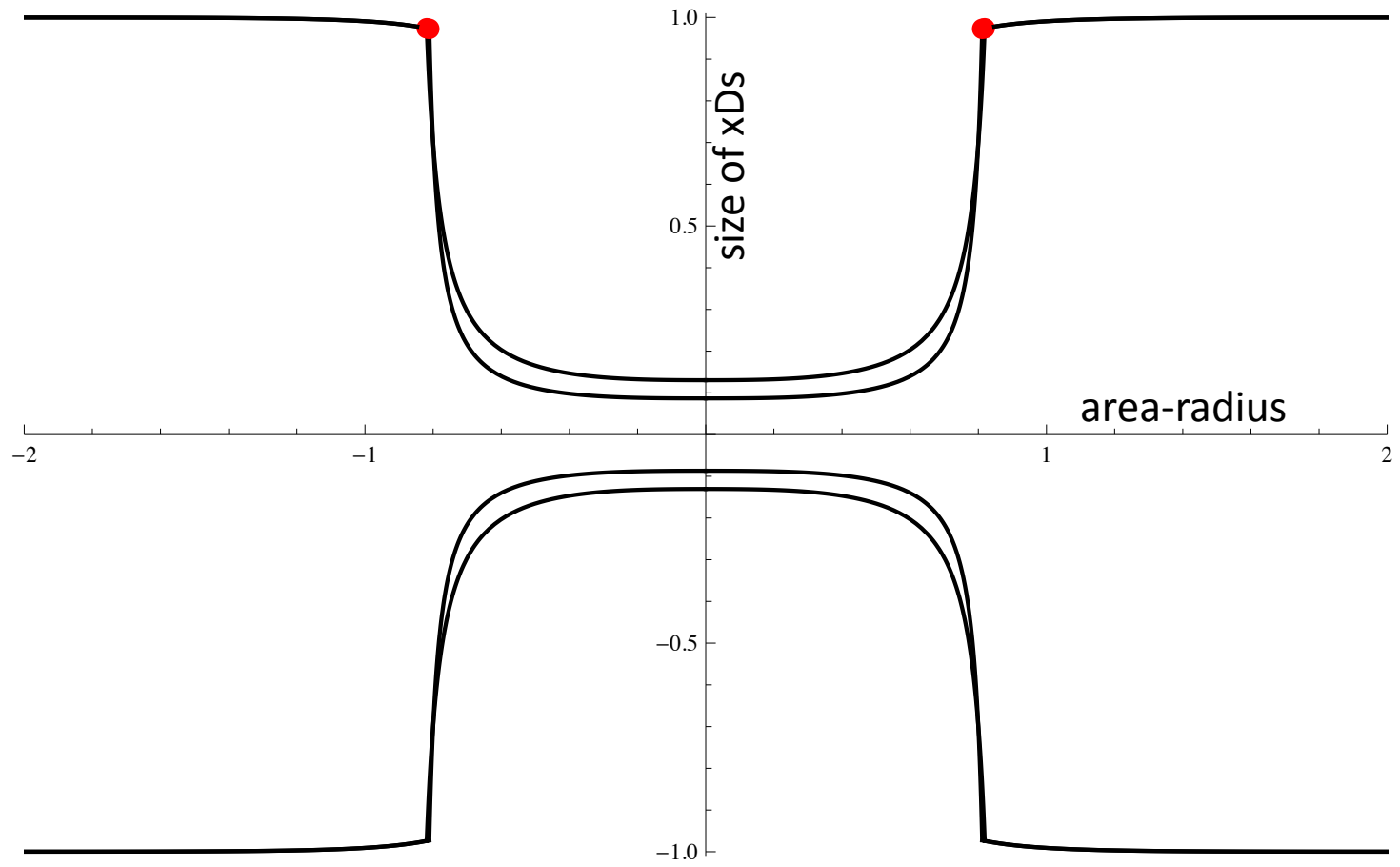


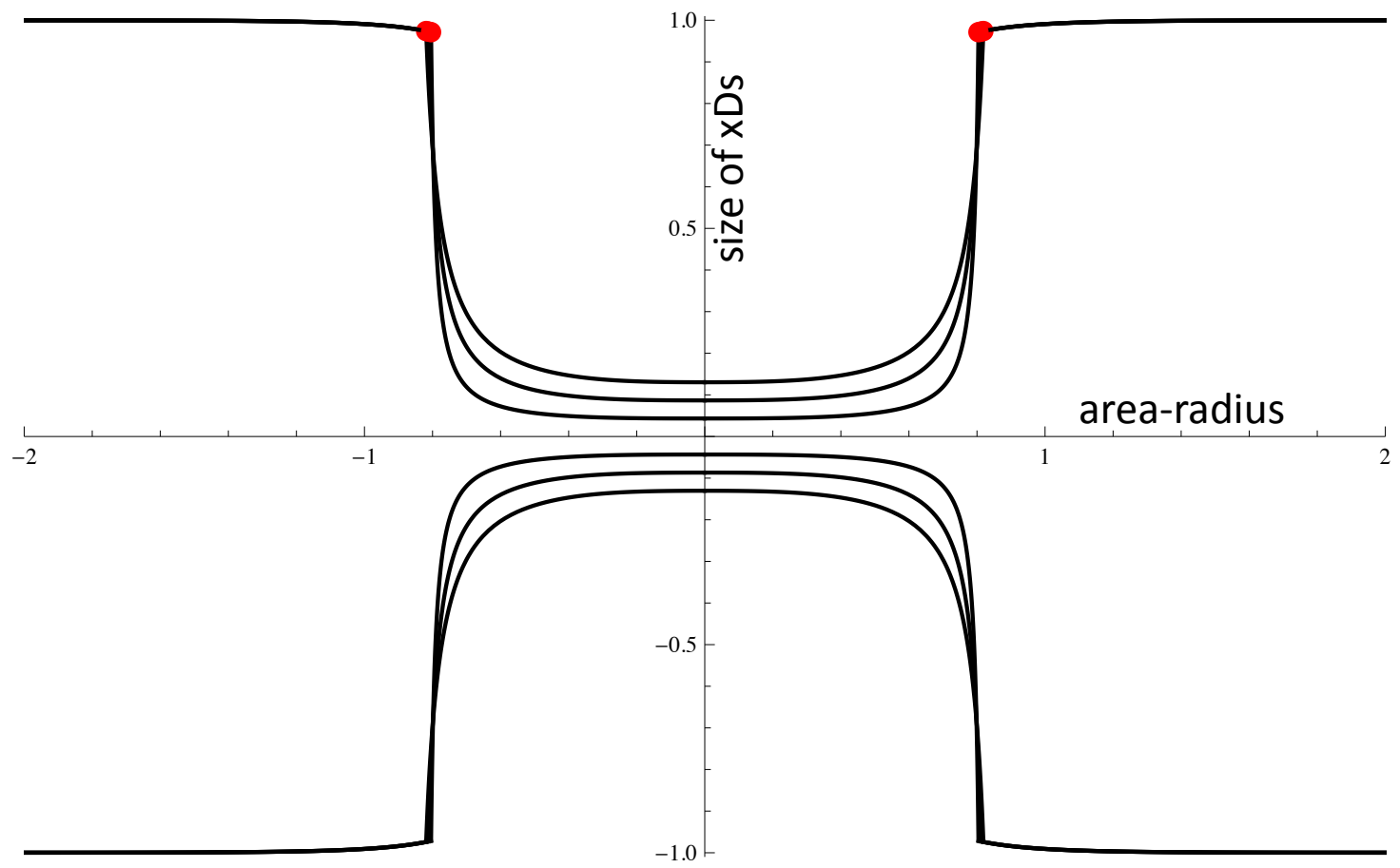
4. Adding back-reaction

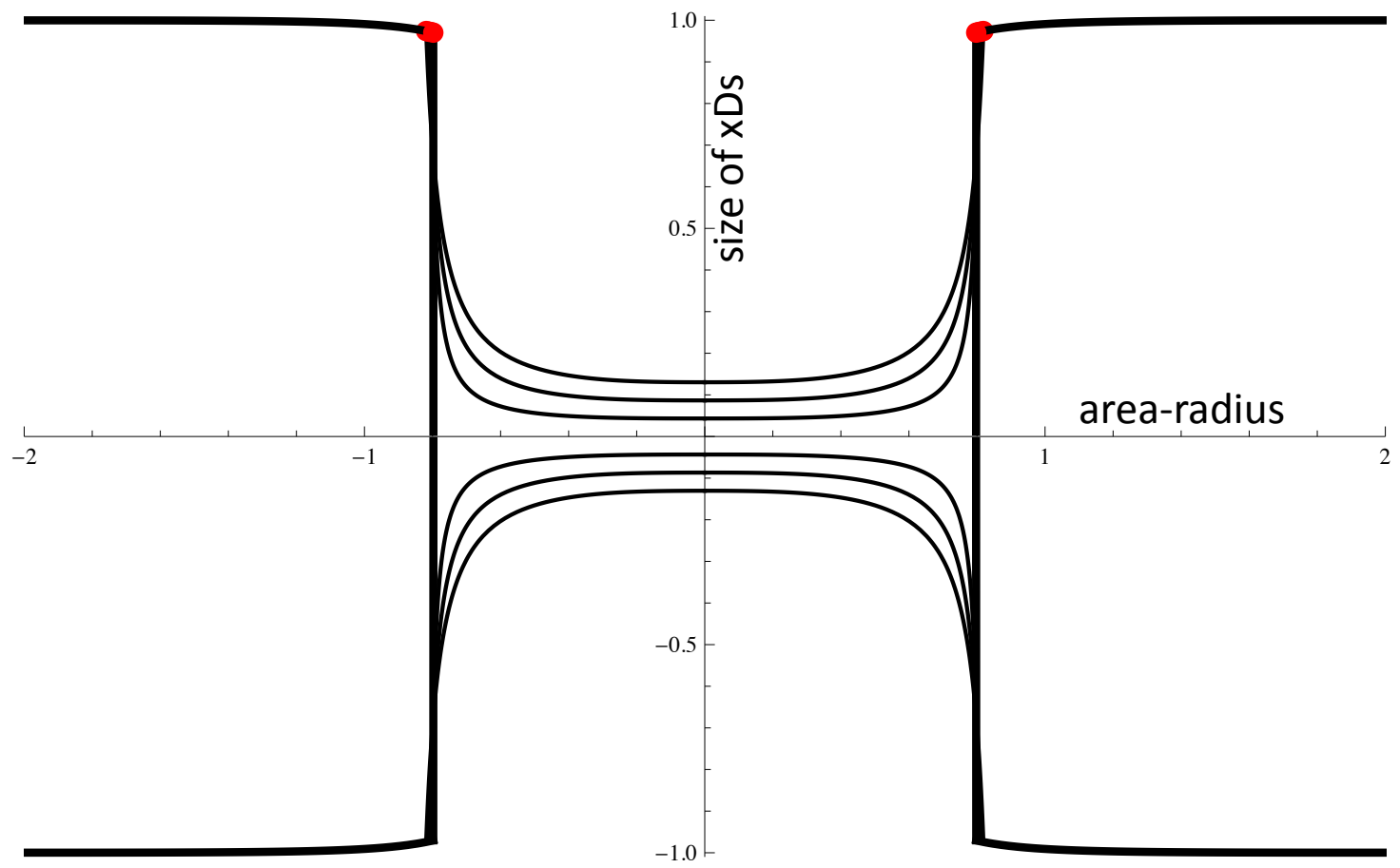
5. The giantest leap of all is a bubble of nothing





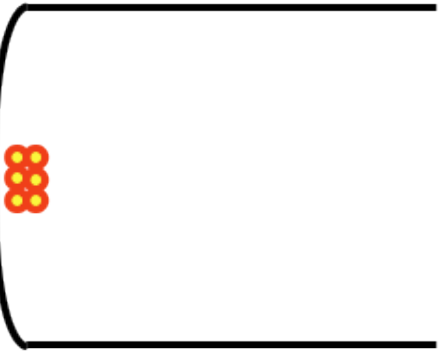




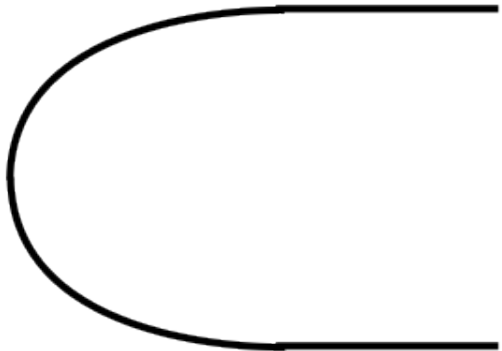
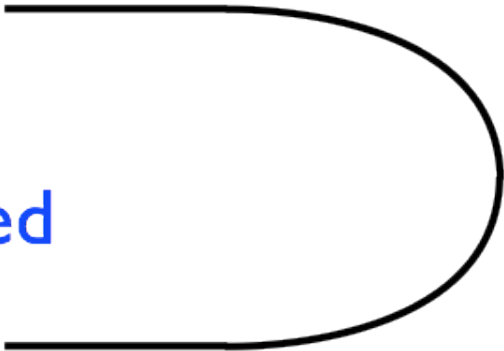


Bubble of Nothing

6d
stabilized



5d
unstabilized

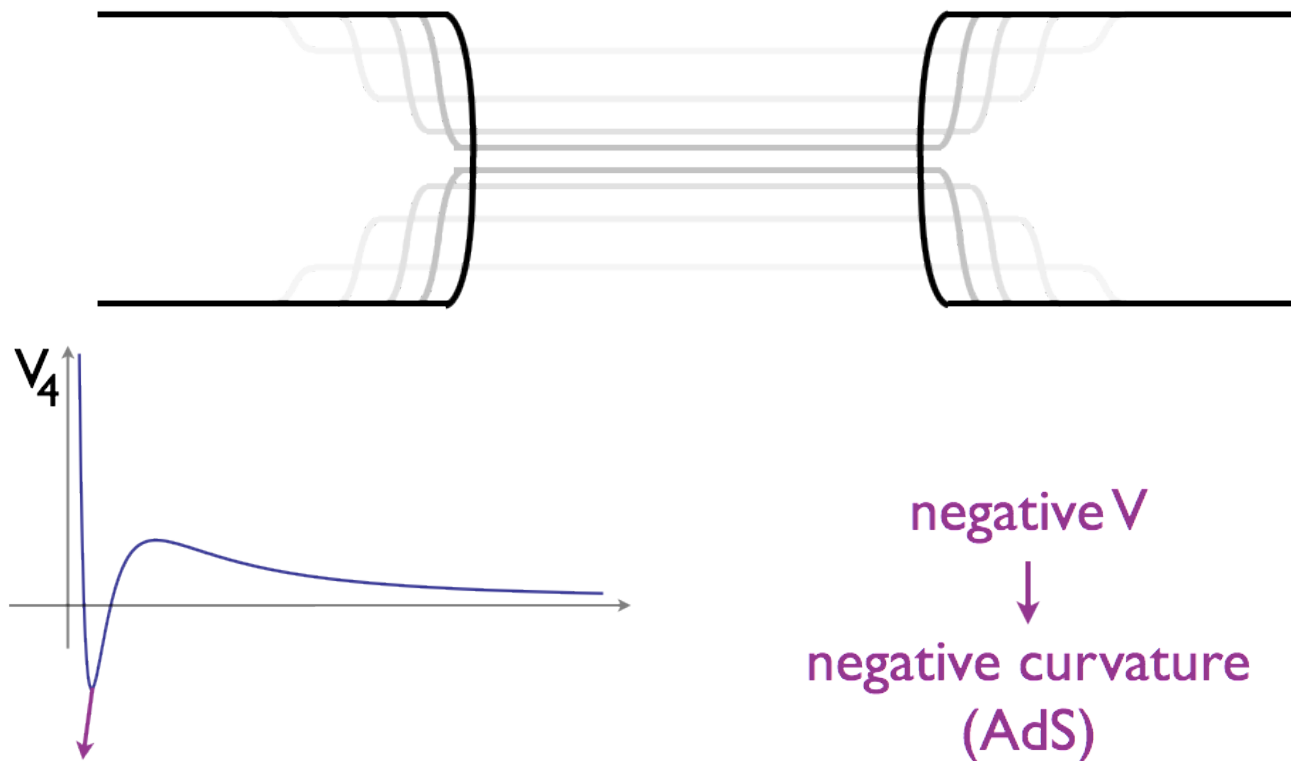


Witten, 1981

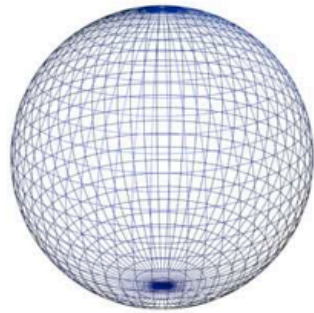


In the limit that ALL flux discharged:

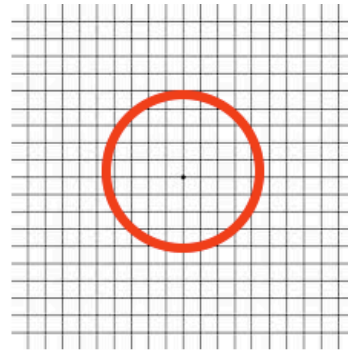
- I. Extra dimensions shrink to zero size
- II. What about the **3-volume** of a slice through the bubble?
like an infinitely thin pancake?



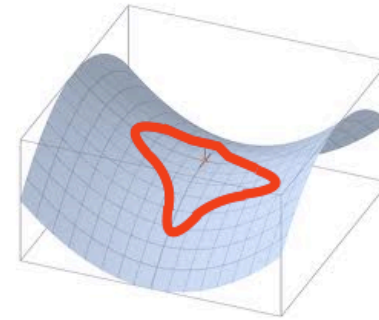
II. What about the **3-volume** of a slice through the bubble?
like an infinitely thin pancake?



$$\text{Area} > \pi r^2$$

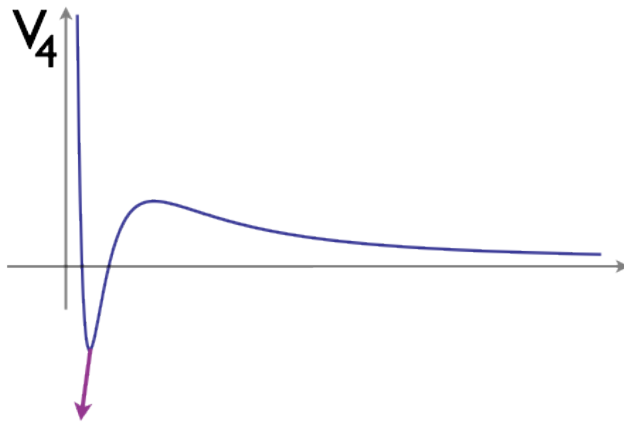


$$\text{Area} = \pi r^2$$



$$\text{Area} < \pi r^2$$

$$\text{Area} \sim r l_{\text{curv}}$$



negative V
↓
negative curvature
(AdS)

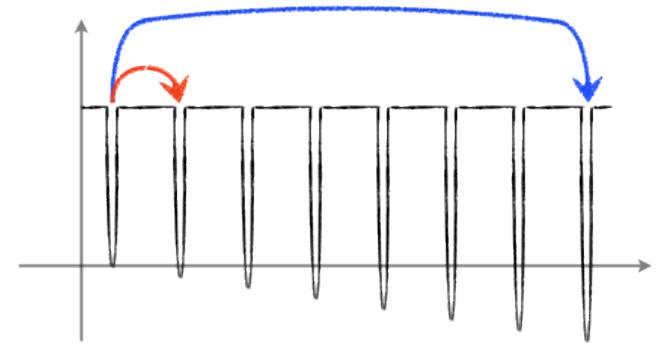
II. What about the **3-volume** of a slice through the bubble?
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$$\text{Volume} \sim \text{Area} \times l_{\text{curv}}$$
$$l_{\text{curv}} \rightarrow 0$$

- II. What about the **3-volume** of a slice through the bubble?
like an infinitely thin pancake?

1. This potential gives small steps



2. Flux compactifications with many fluxes give giant leaps

3. Monkey branes



4. Back-reaction on the compactification

5. The giantest leap of all is a bubble of nothing



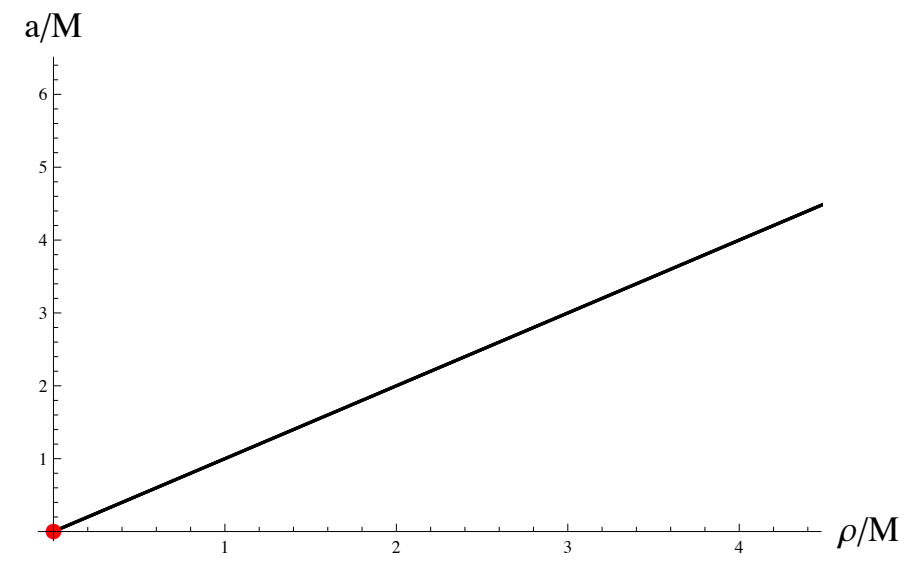
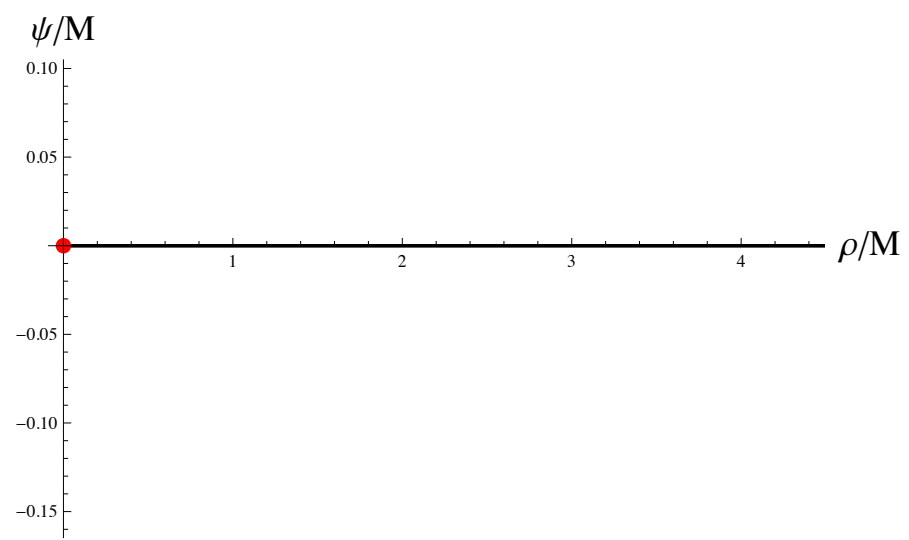
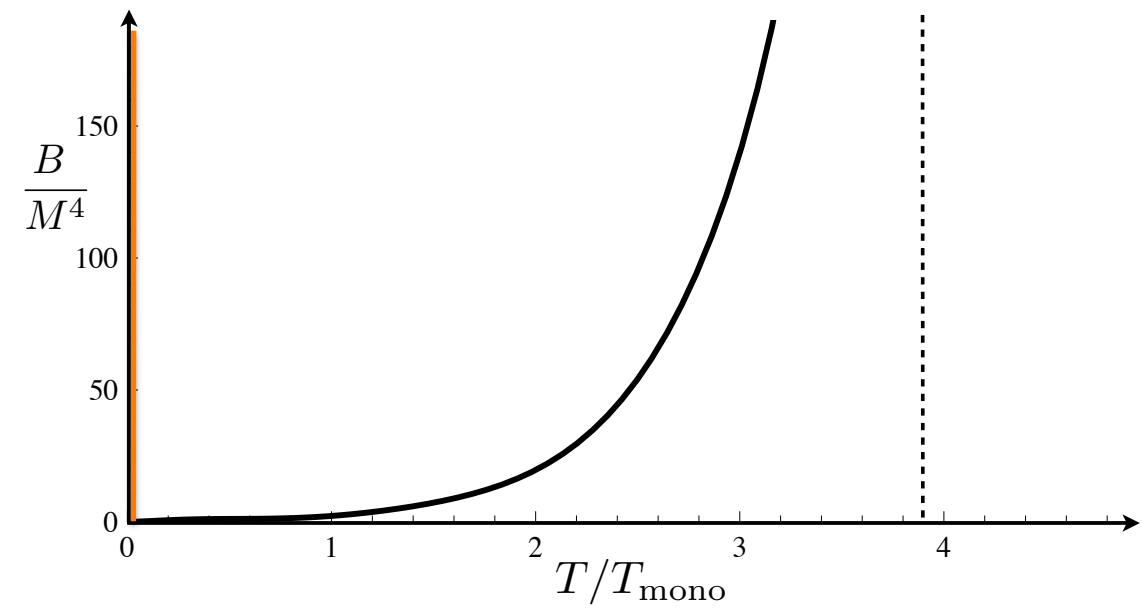
Summary

- Giant leaps beat small steps in landscapes
- The bubble of nothing can be the fastest decay
- It is realized smoothly as the limit of flux tunneling
- When the many fluxes come from higher dimensional model, still expect enhancement. Richer behavior.
- If we live in a multiflux landscape, then we draw two conclusions. We got here by an exponentially subdominant decay, and we will leave here by a bubble of nothing.

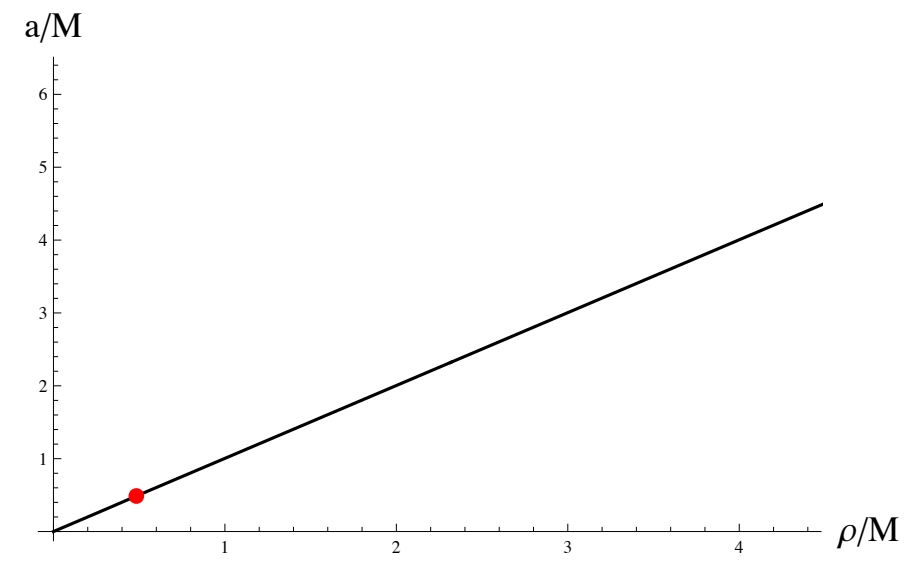
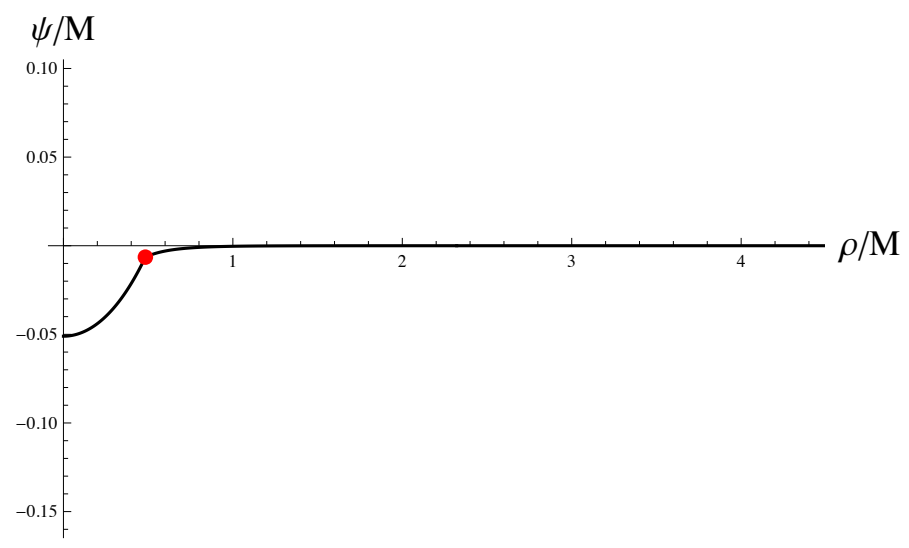
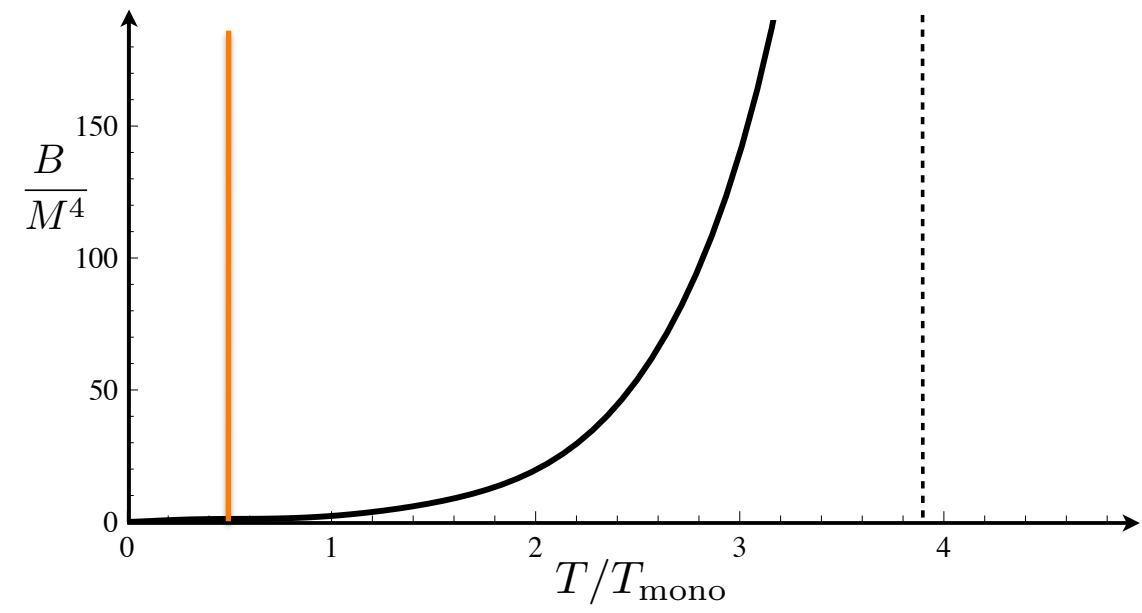
the end



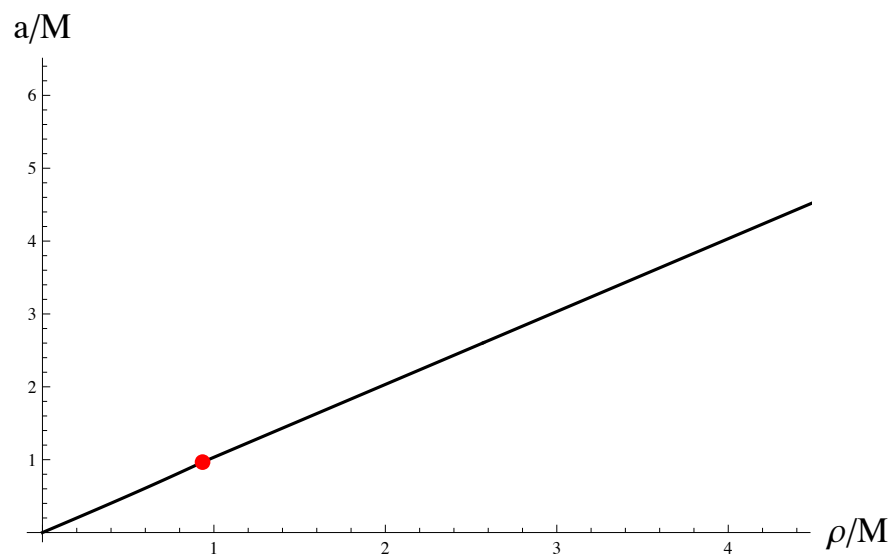
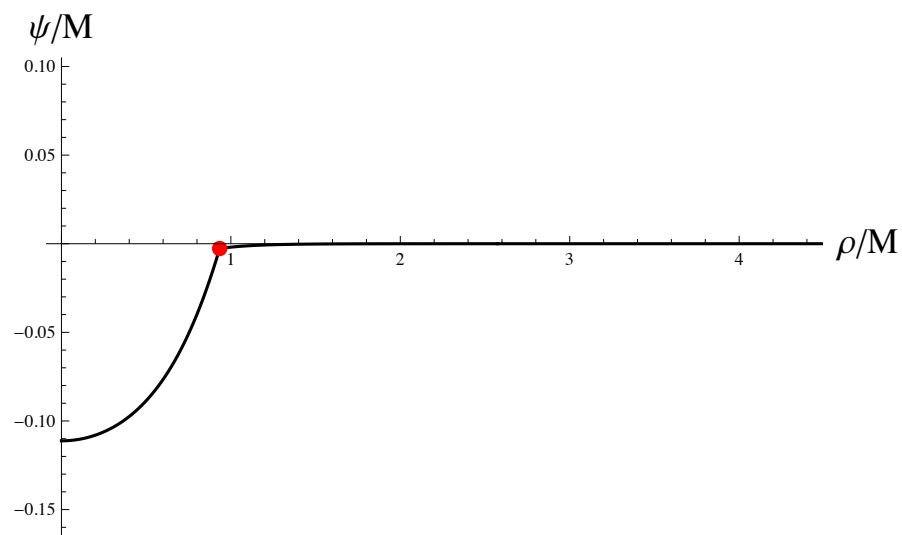
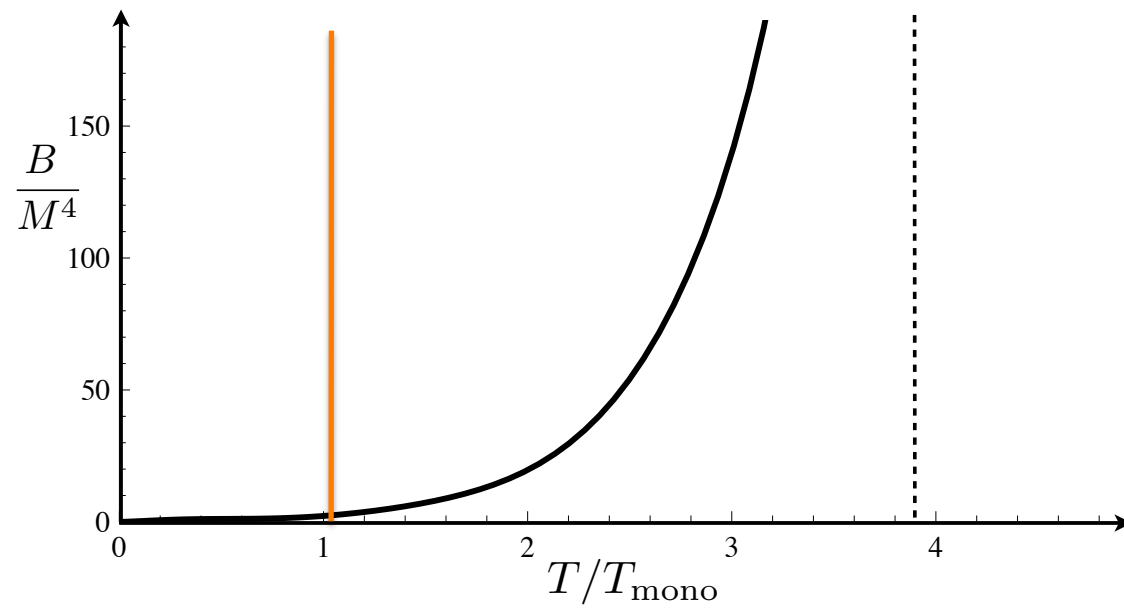
Case 1:
Decay from
Minkowski



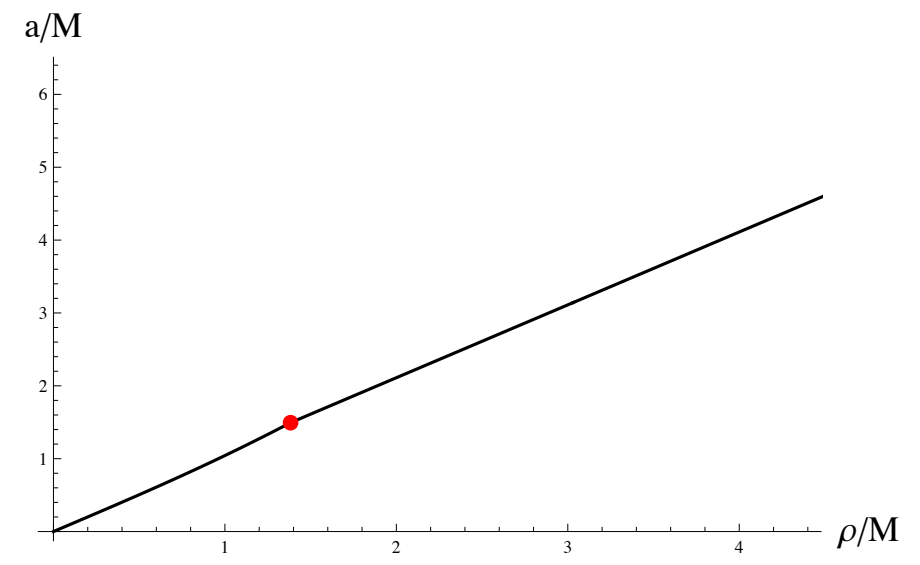
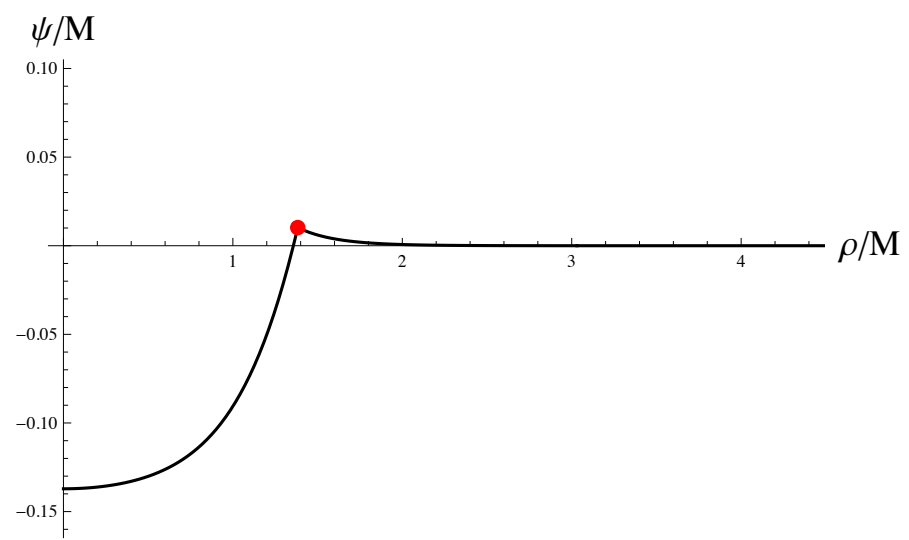
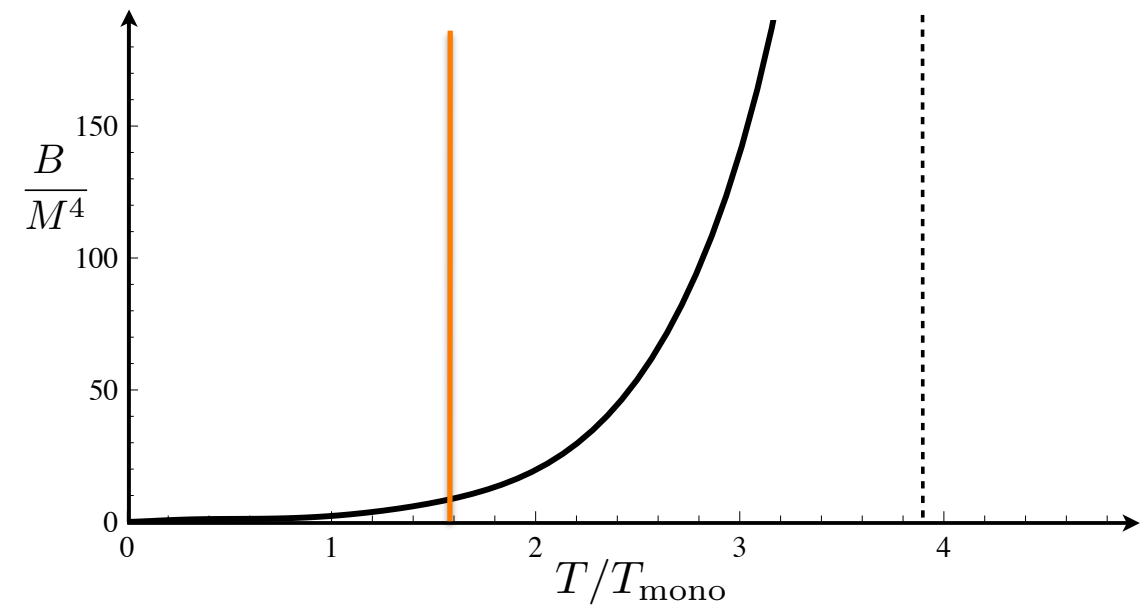
Case 1:
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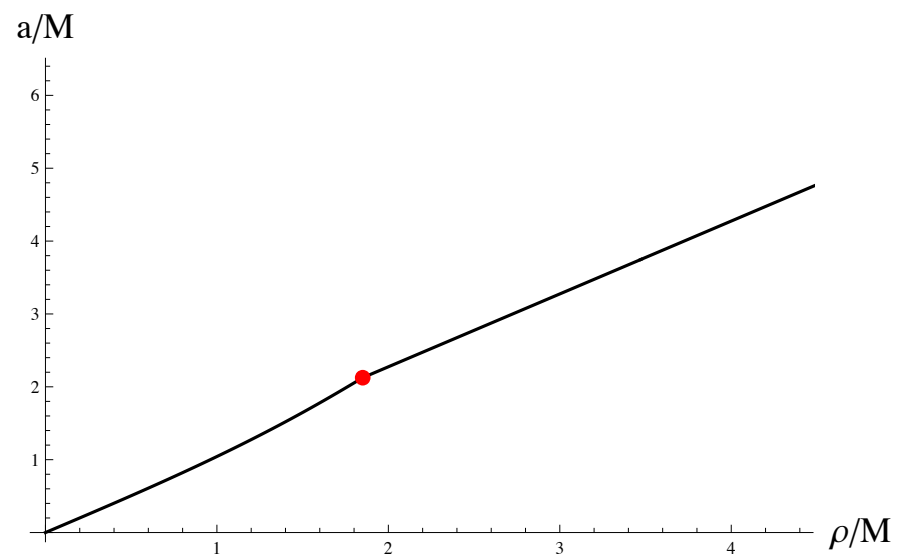
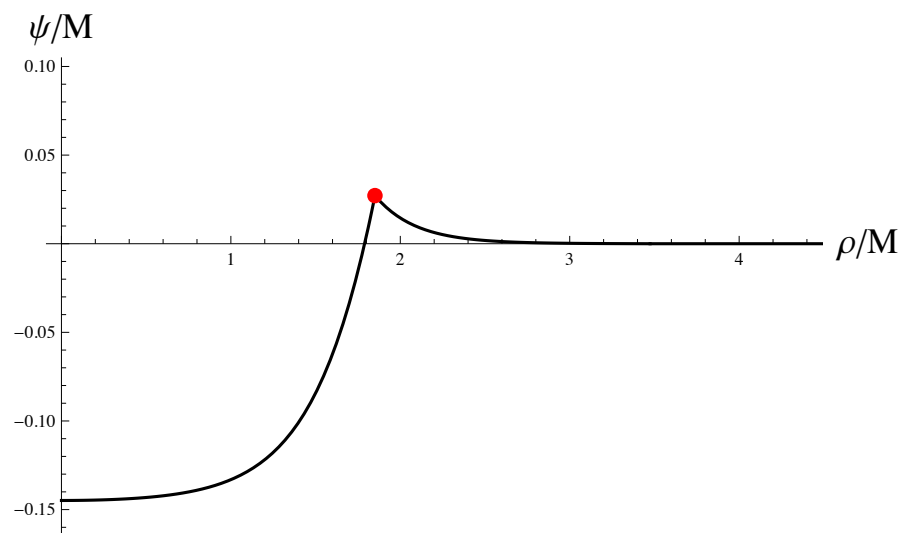
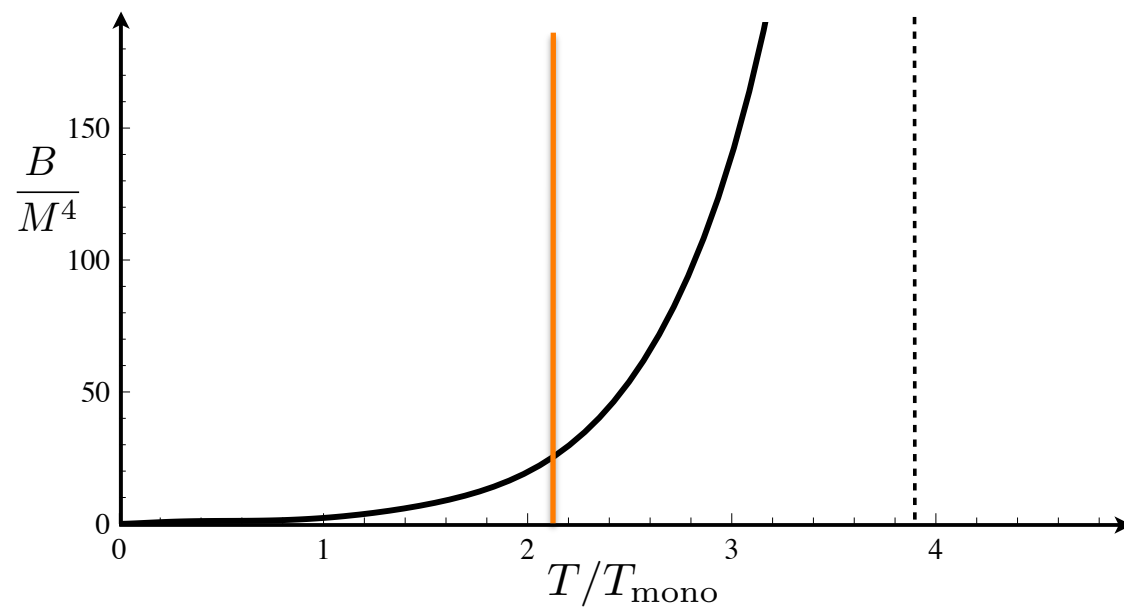
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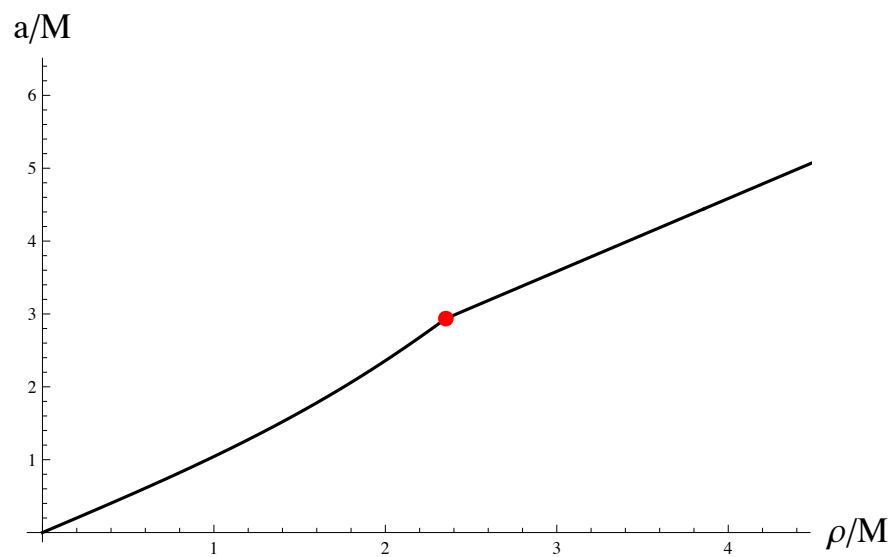
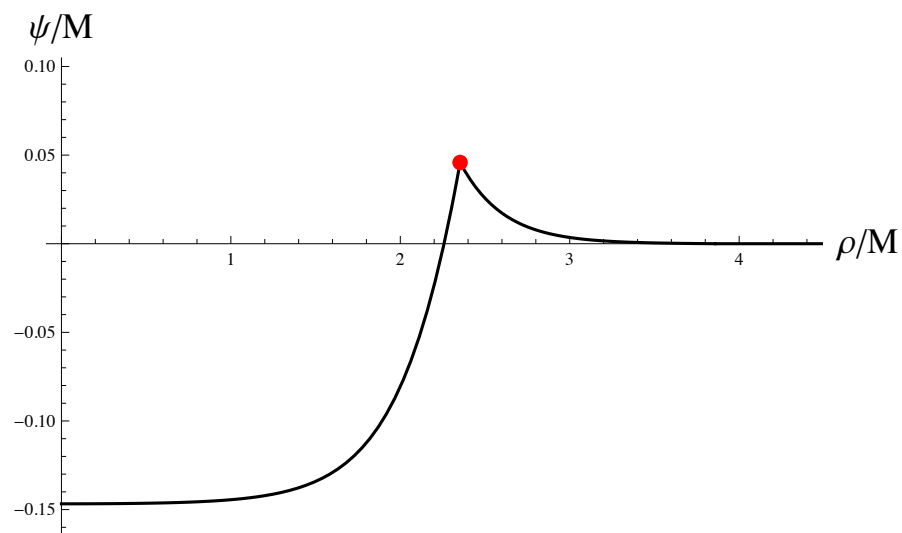
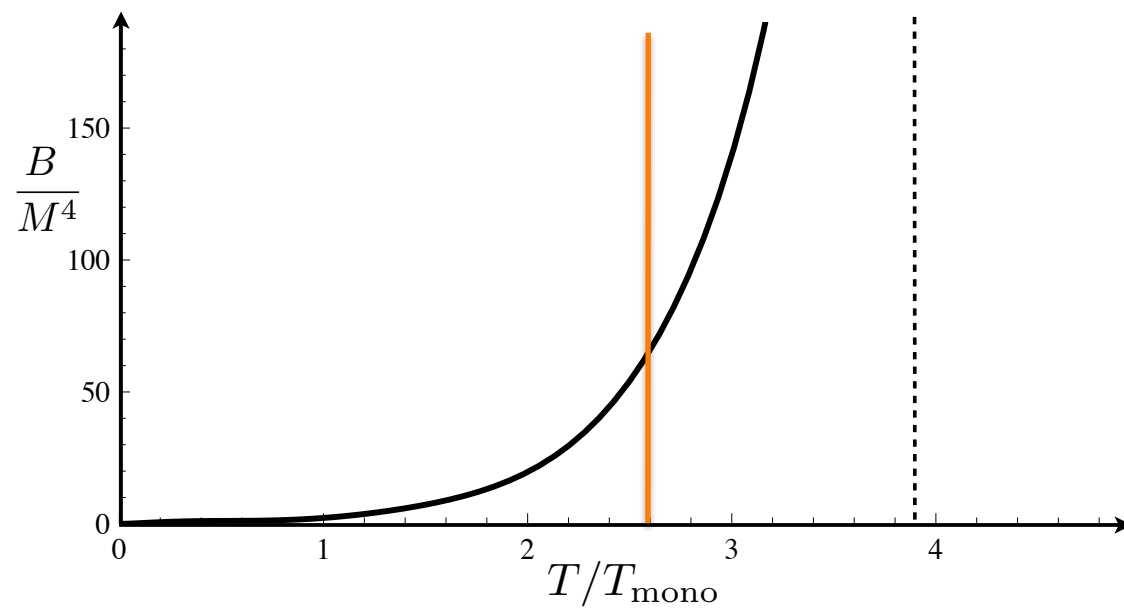
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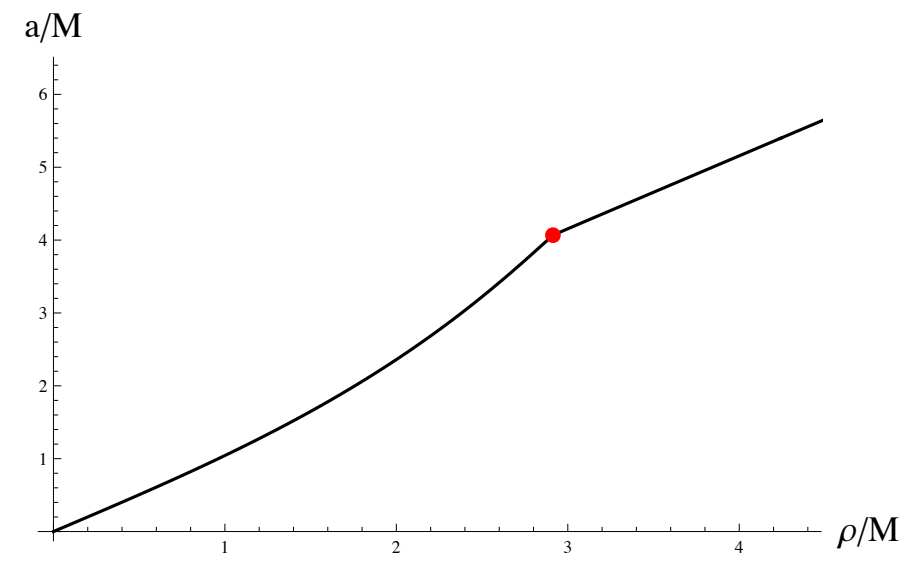
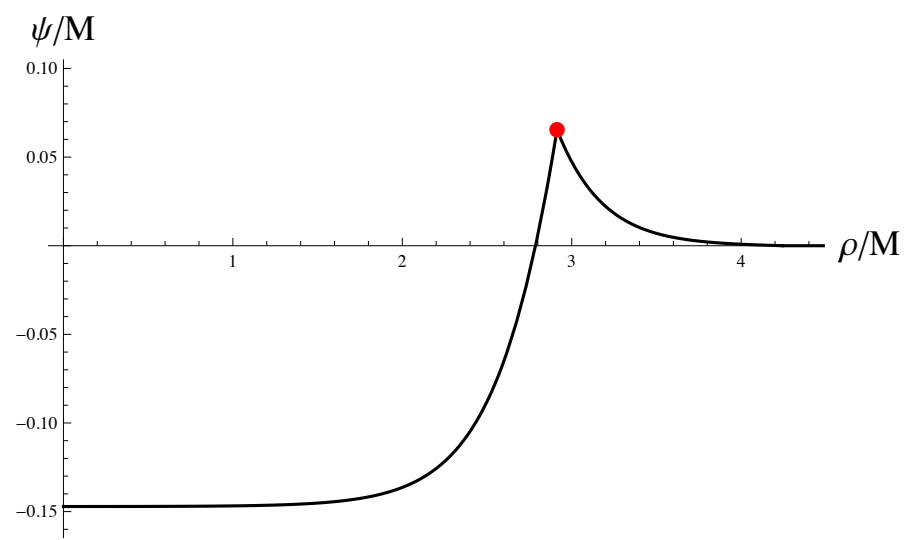
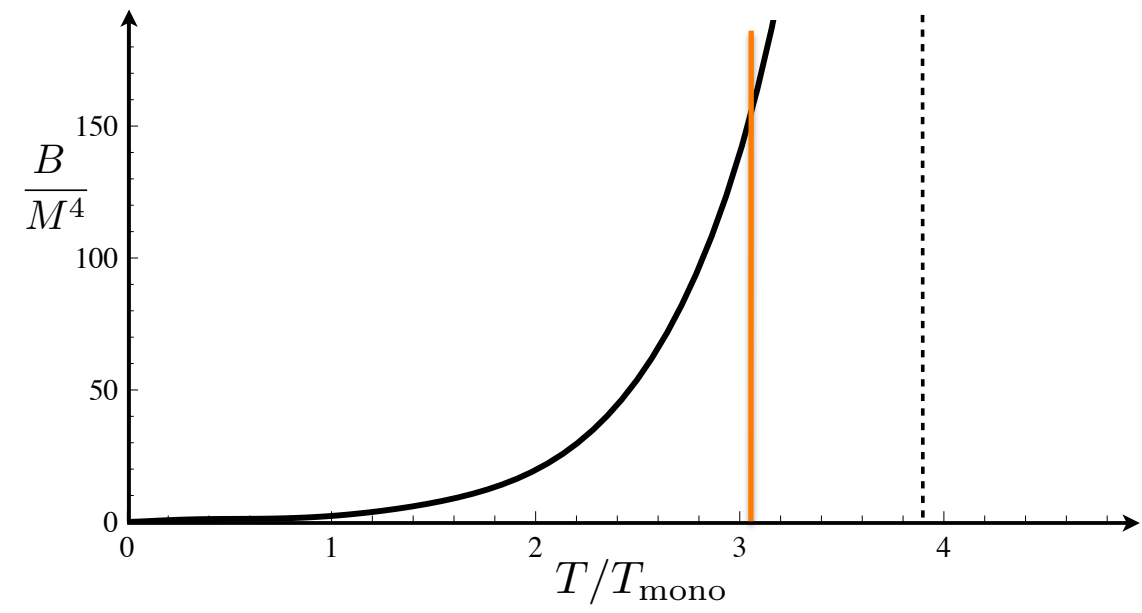
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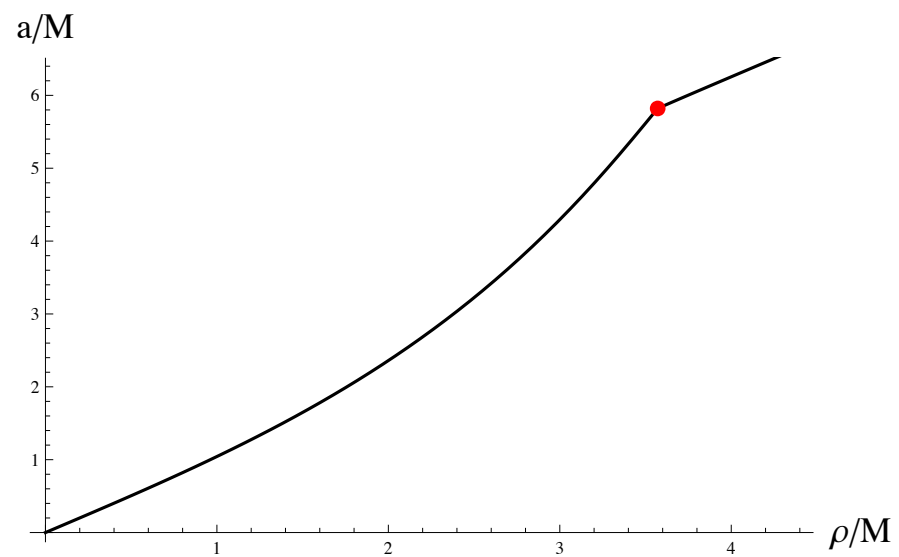
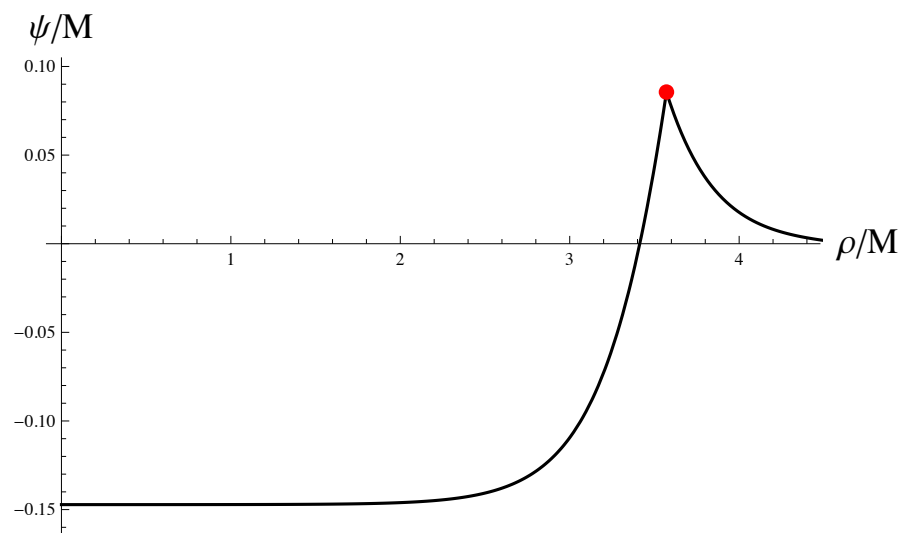
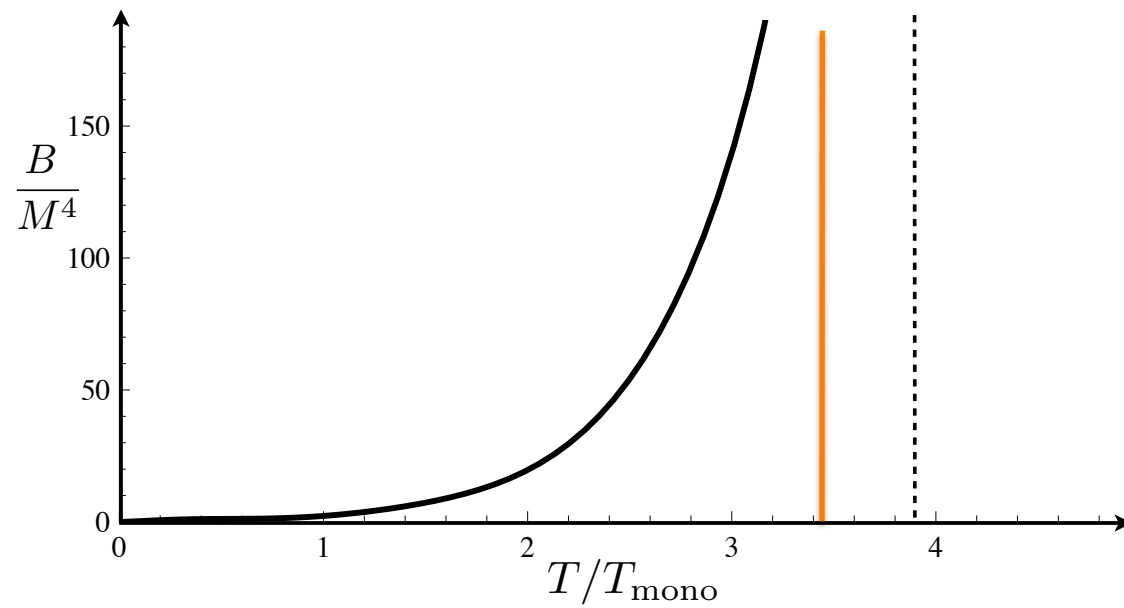
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Decay from
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Case 1:
Decay from
Minkowski

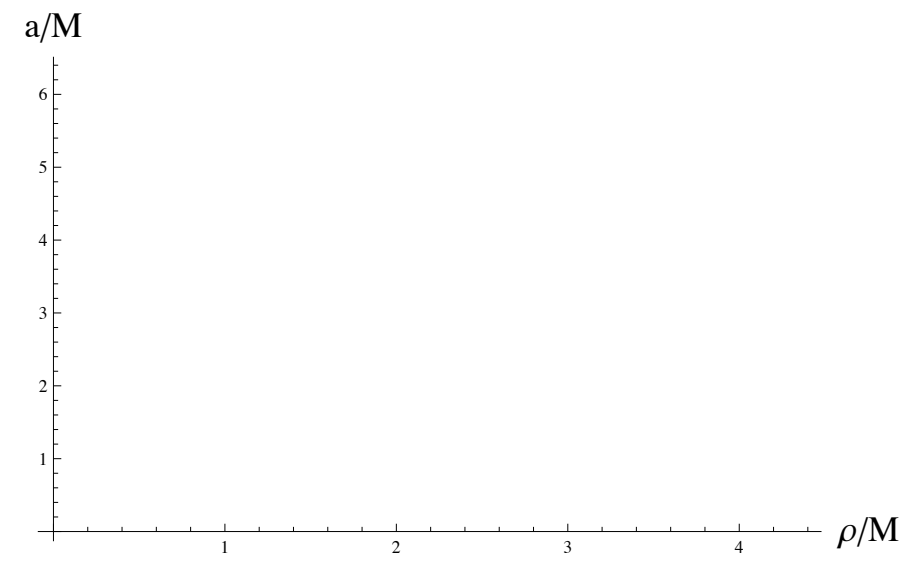
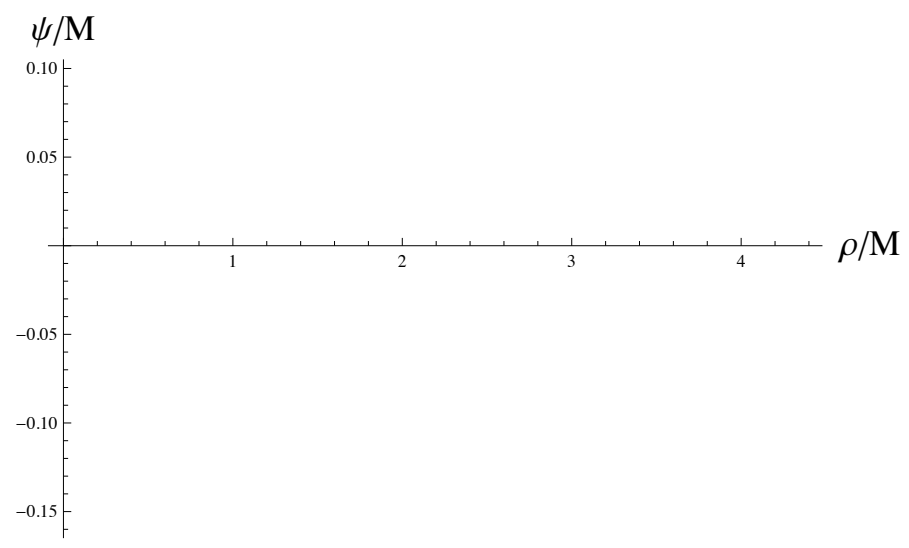
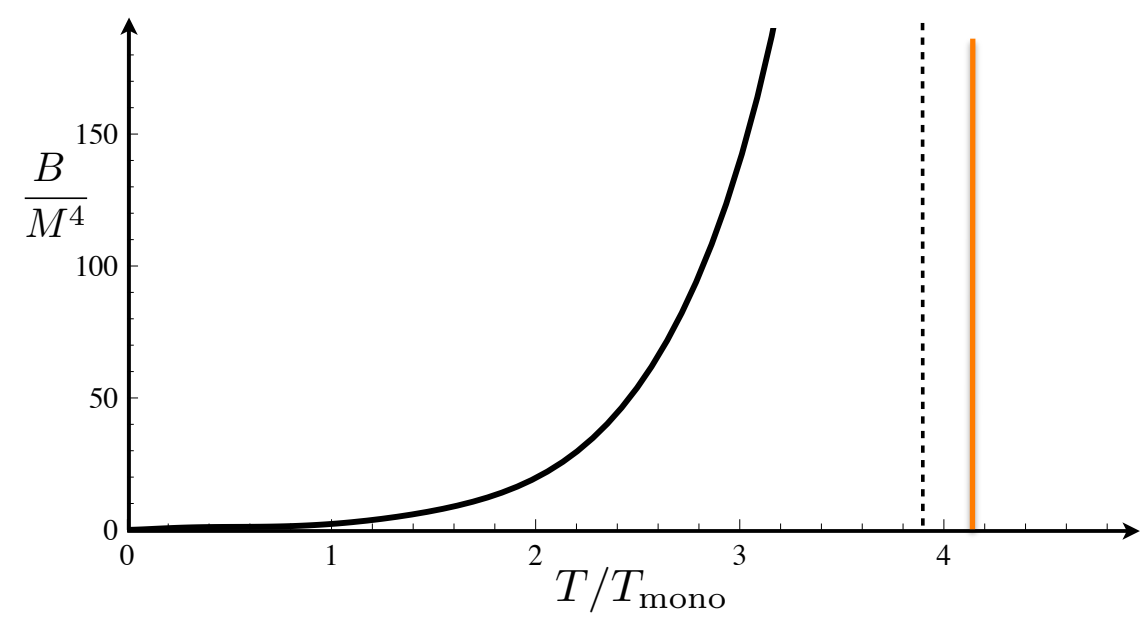


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Decay from
Minkowski

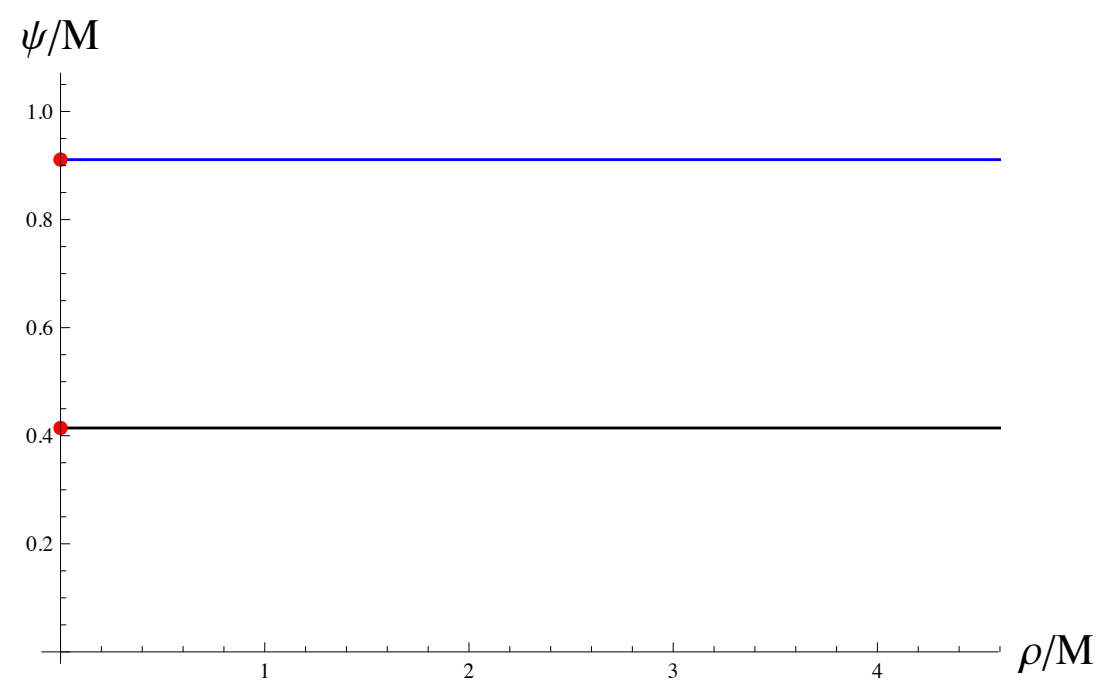
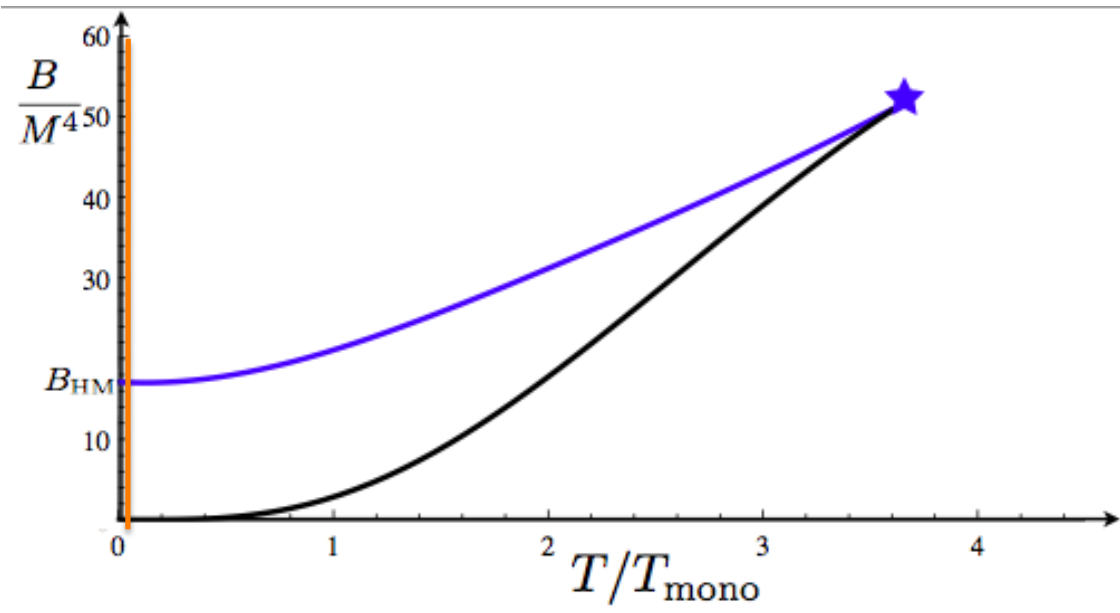


Case 1:
Decay from
Minkowski

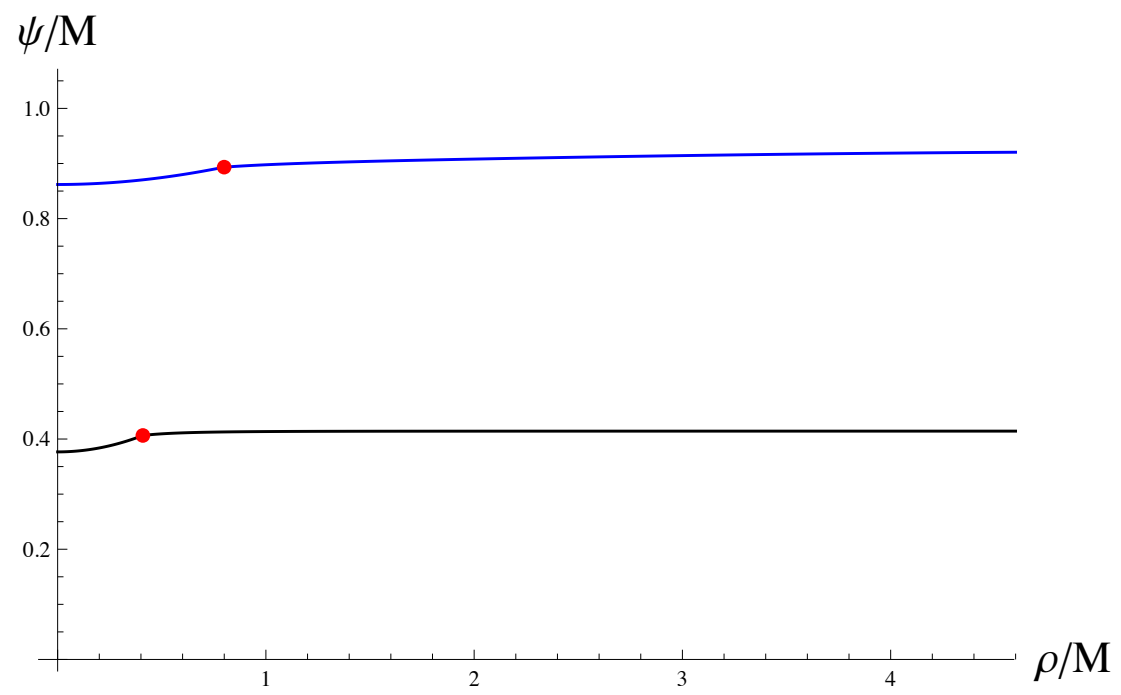
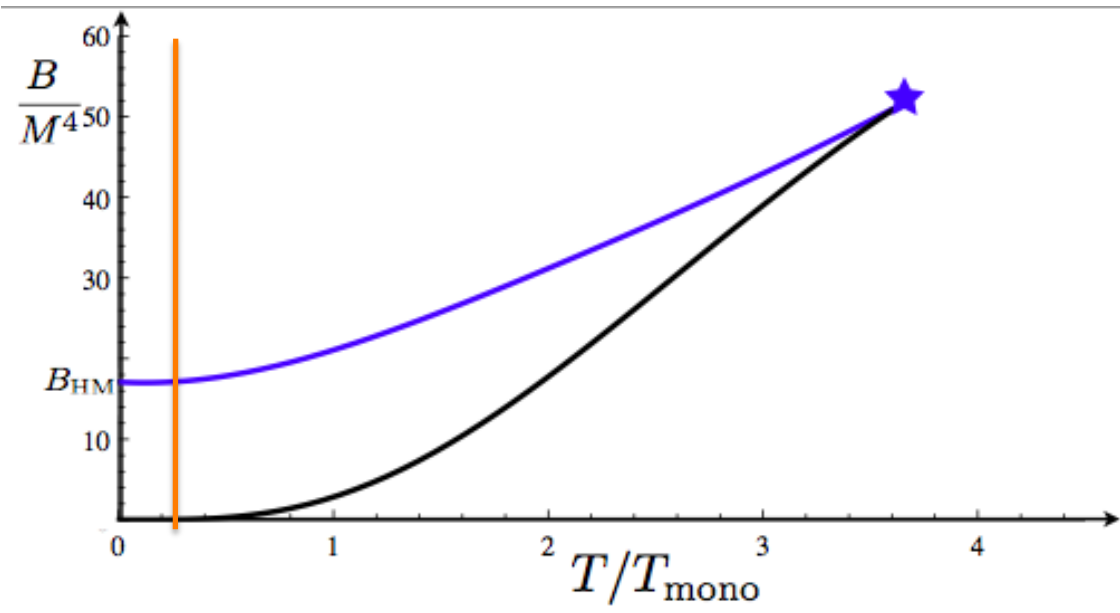
Gravitational
Blocking



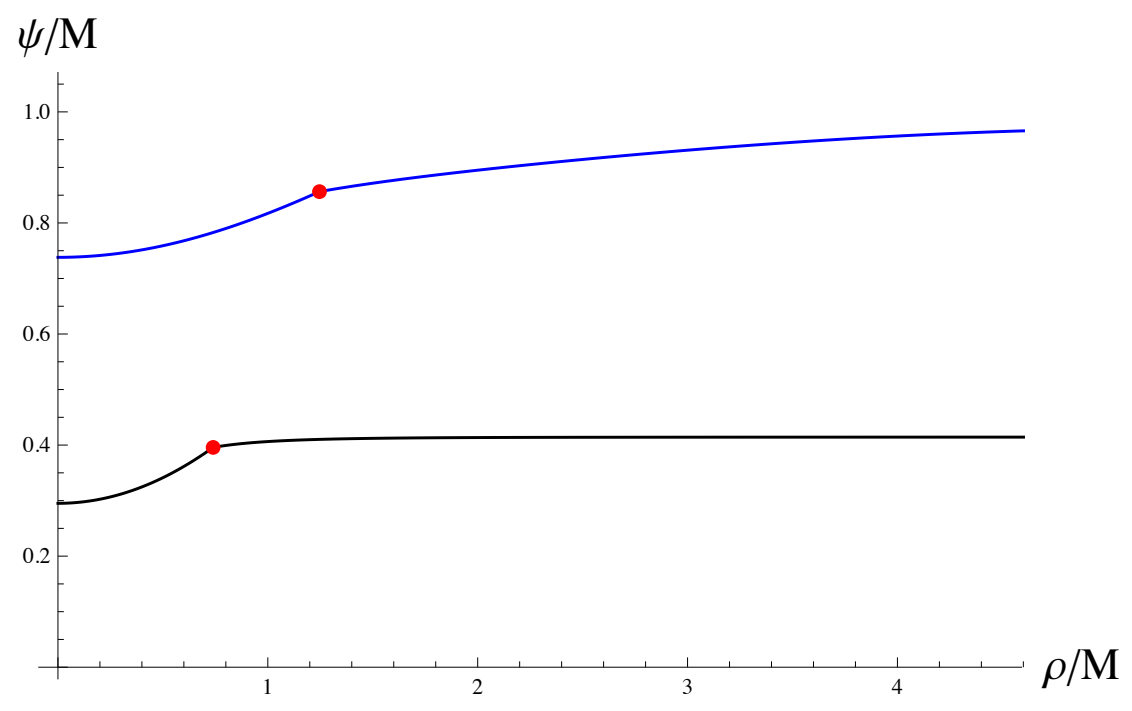
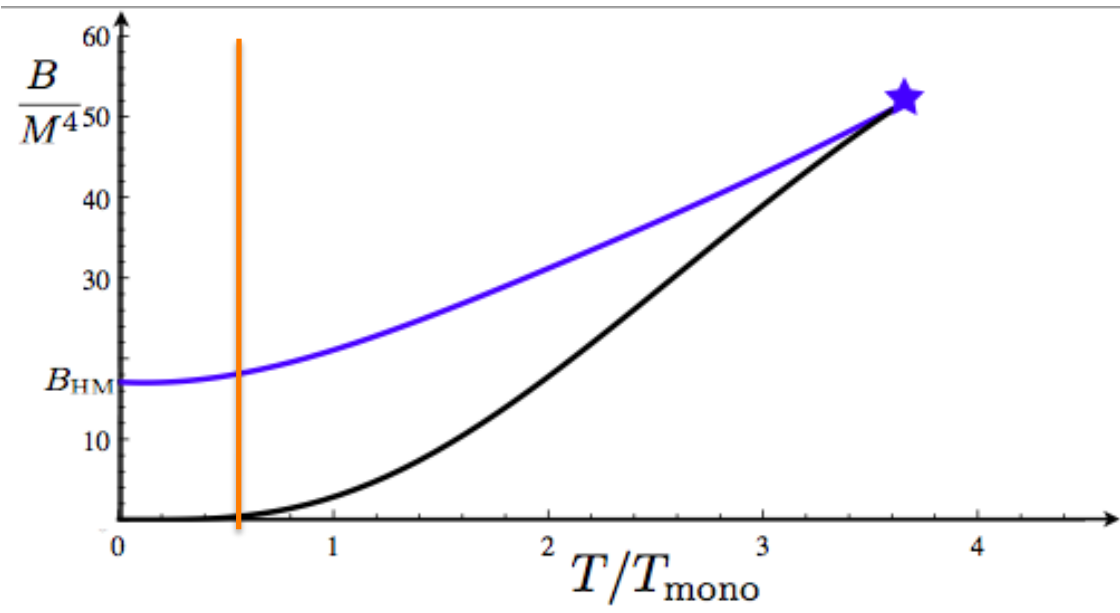
Case 2a:
Decay from
high dS



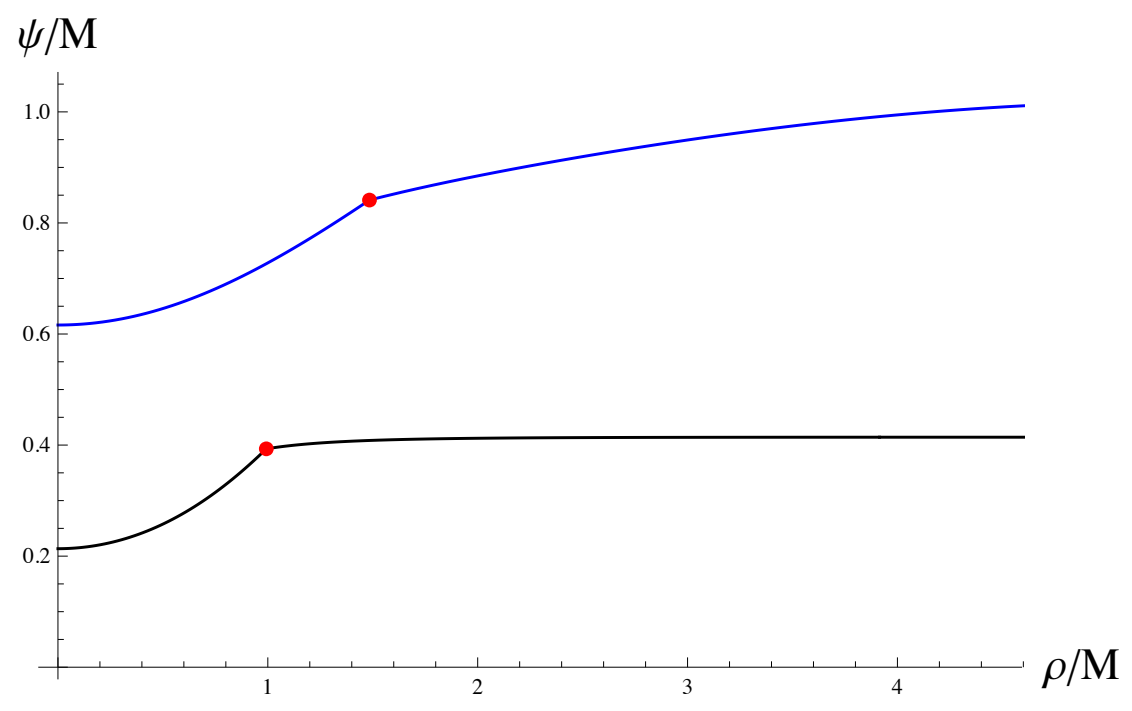
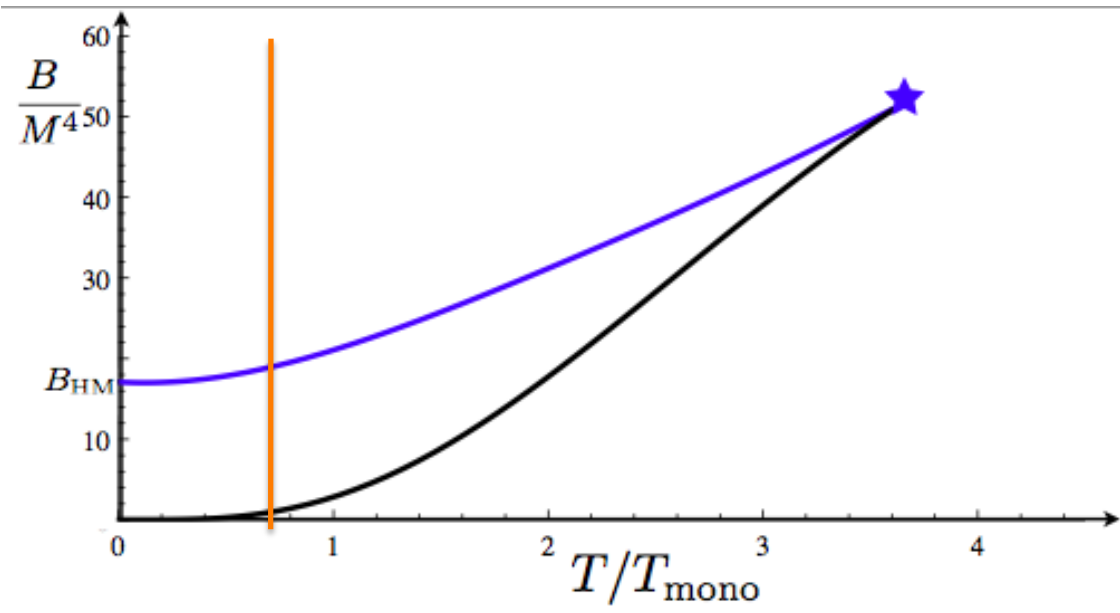
Case 2a:
Decay from
high dS



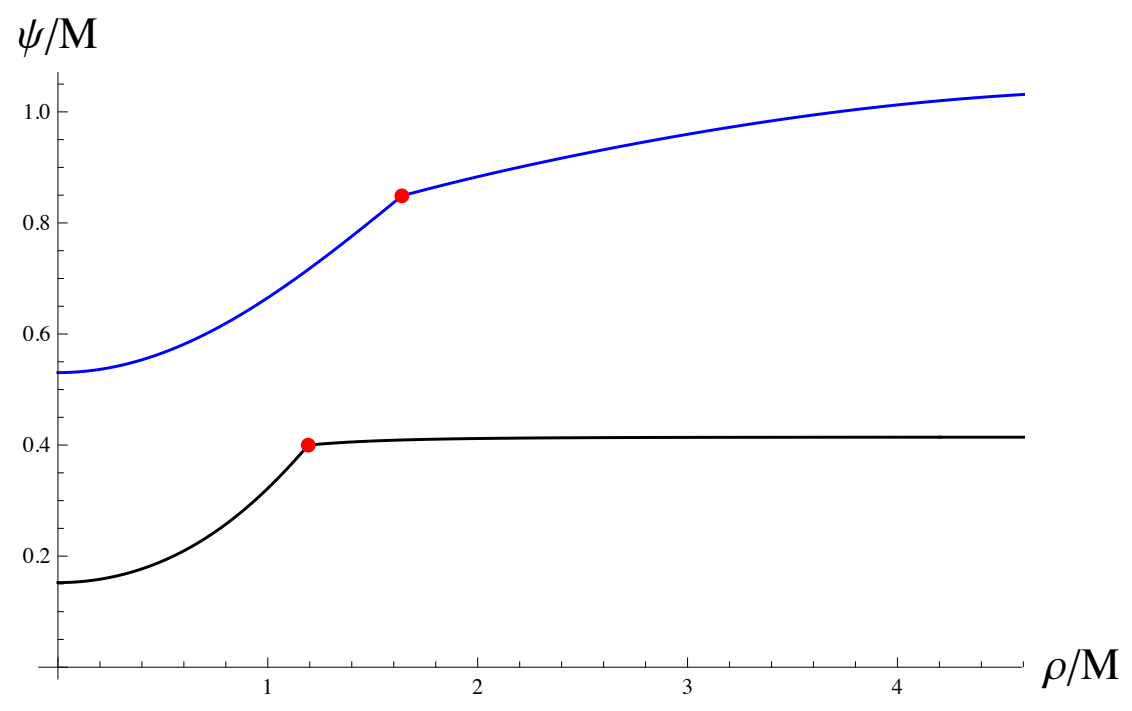
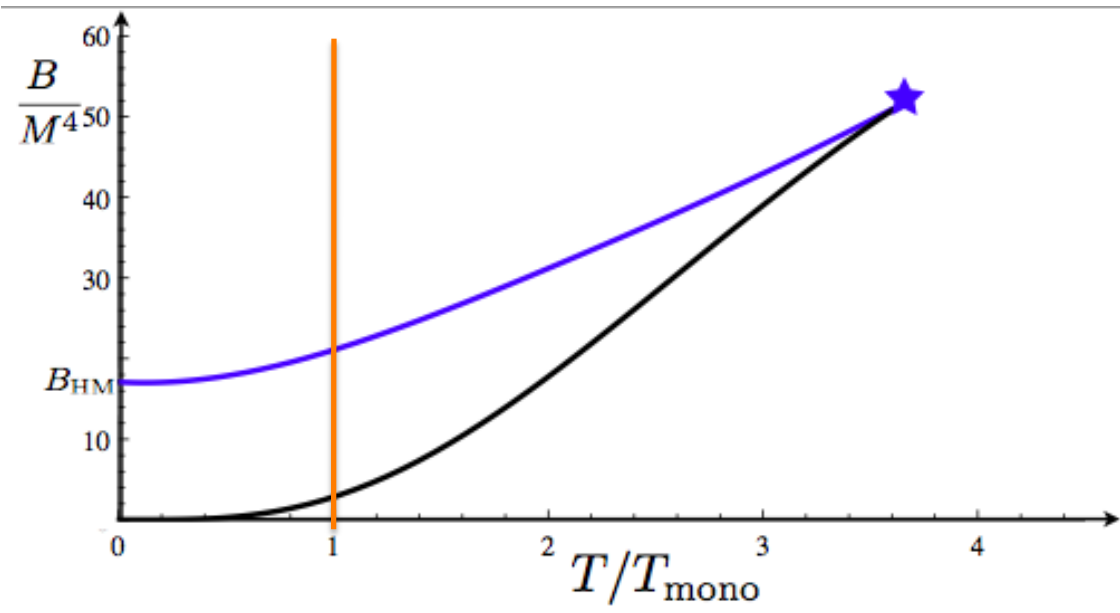
Case 2a:
Decay from
high dS



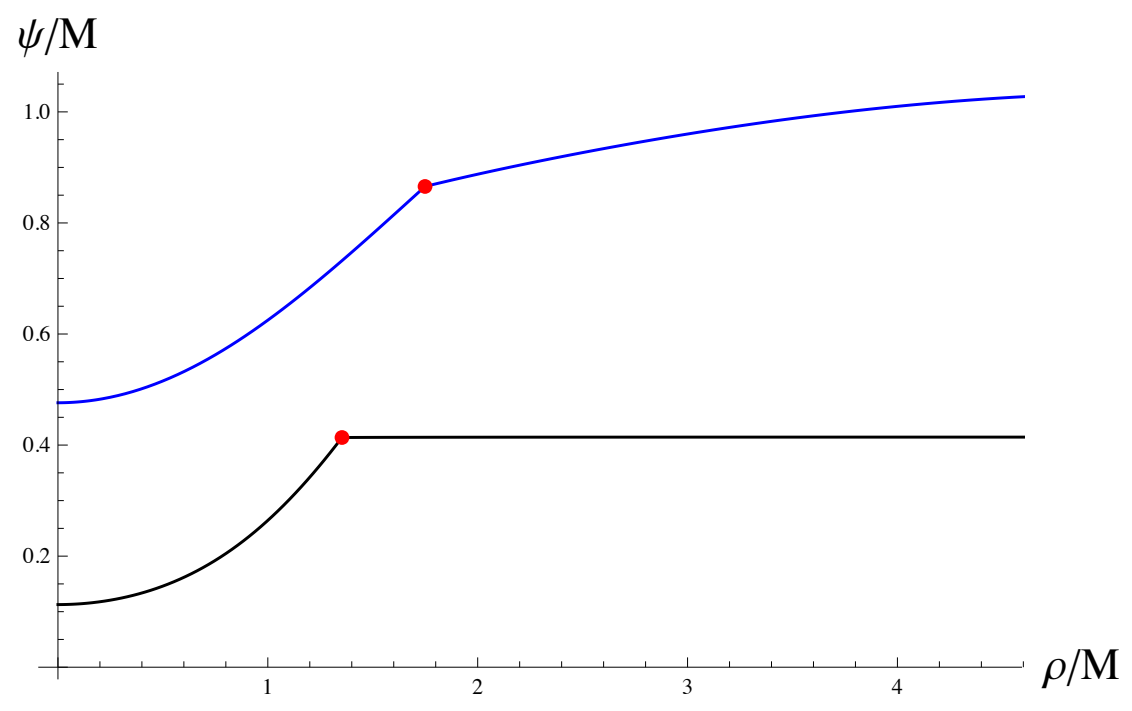
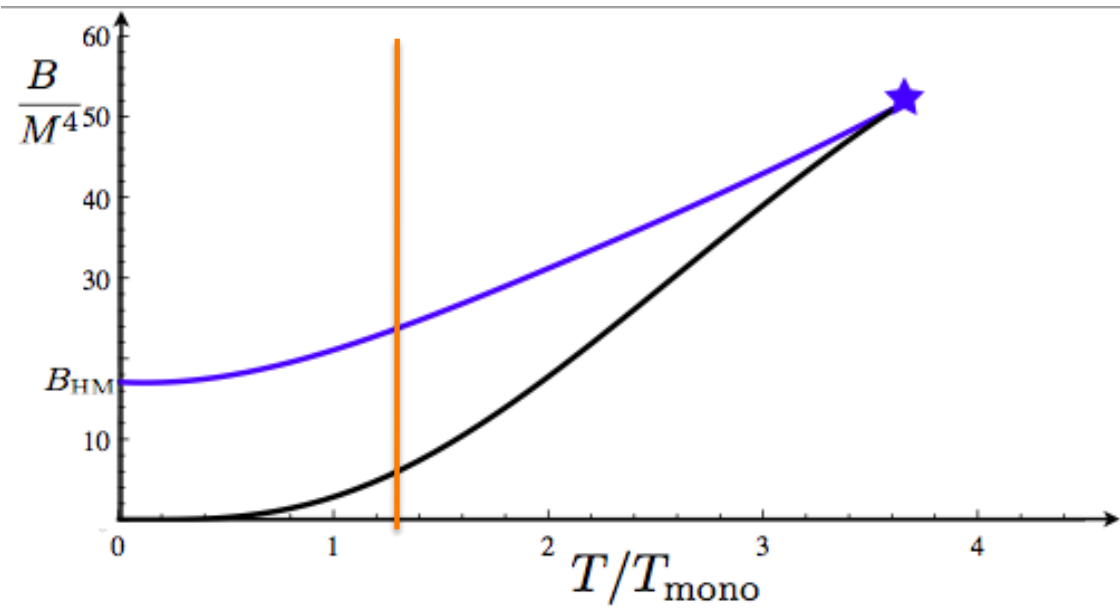
Case 2a:
Decay from
high dS



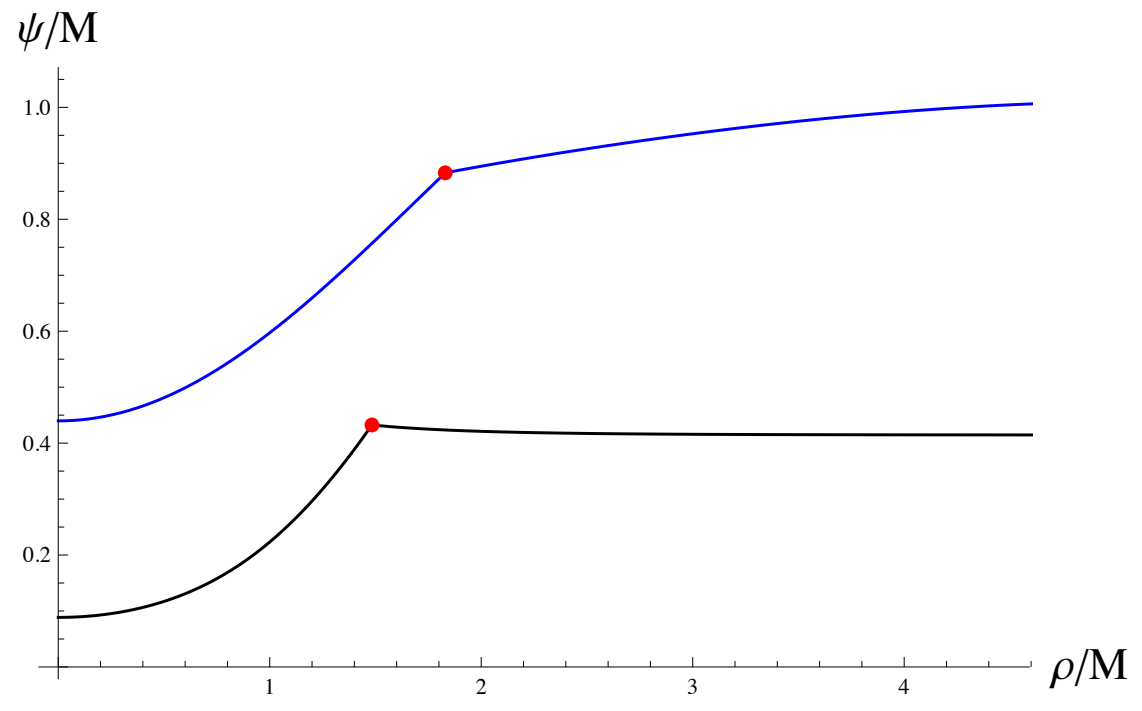
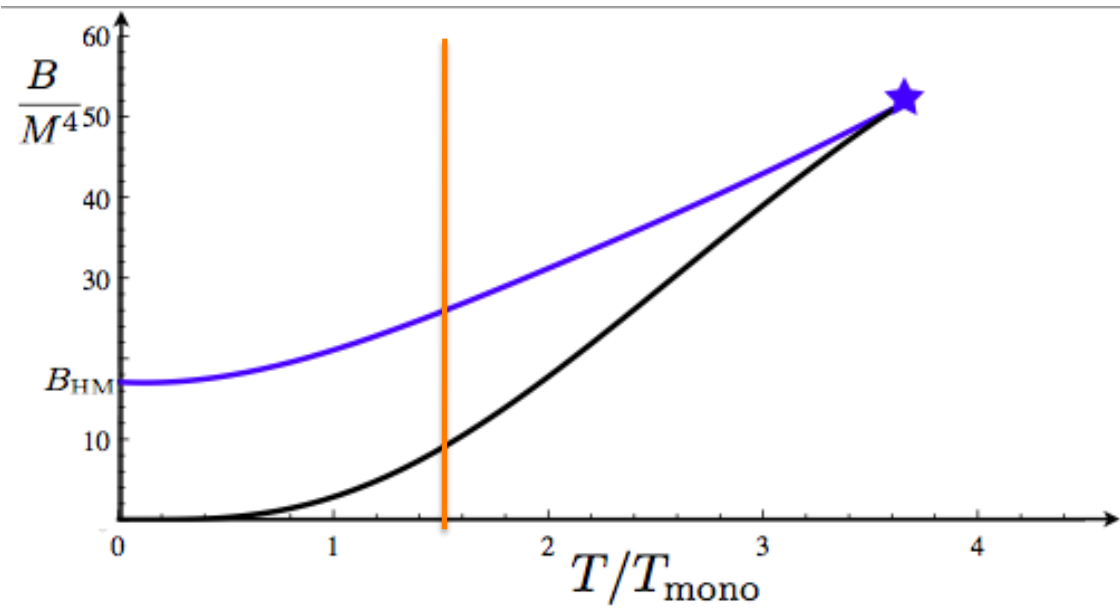
Case 2a:
Decay from
high dS



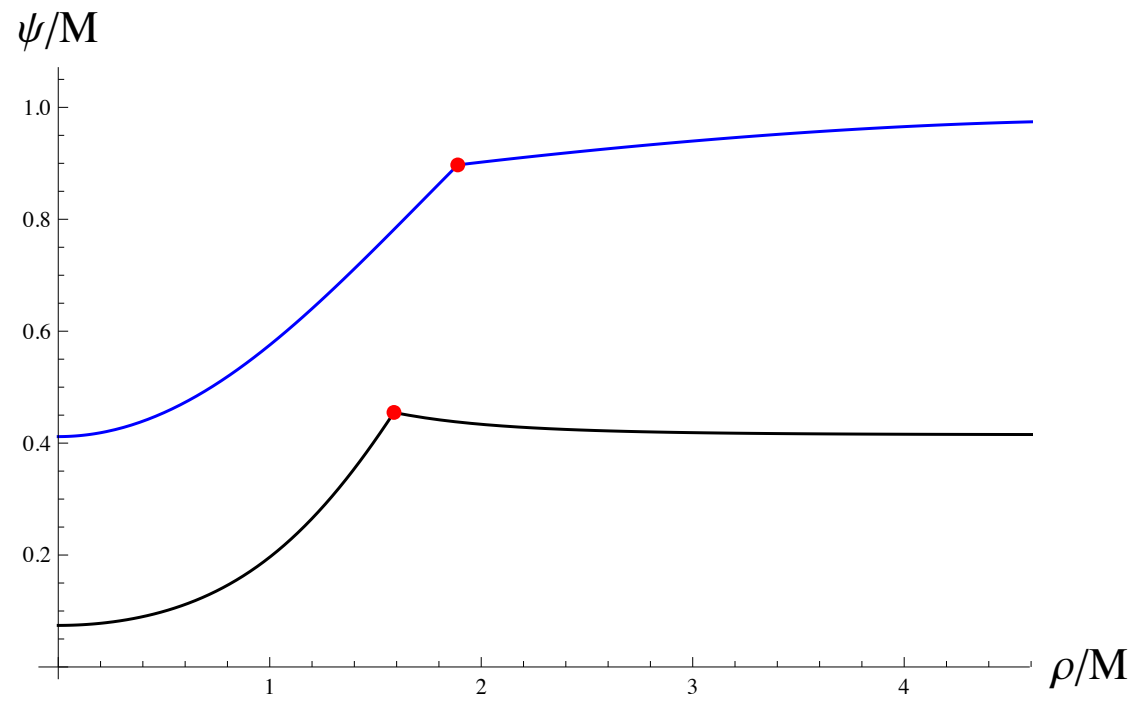
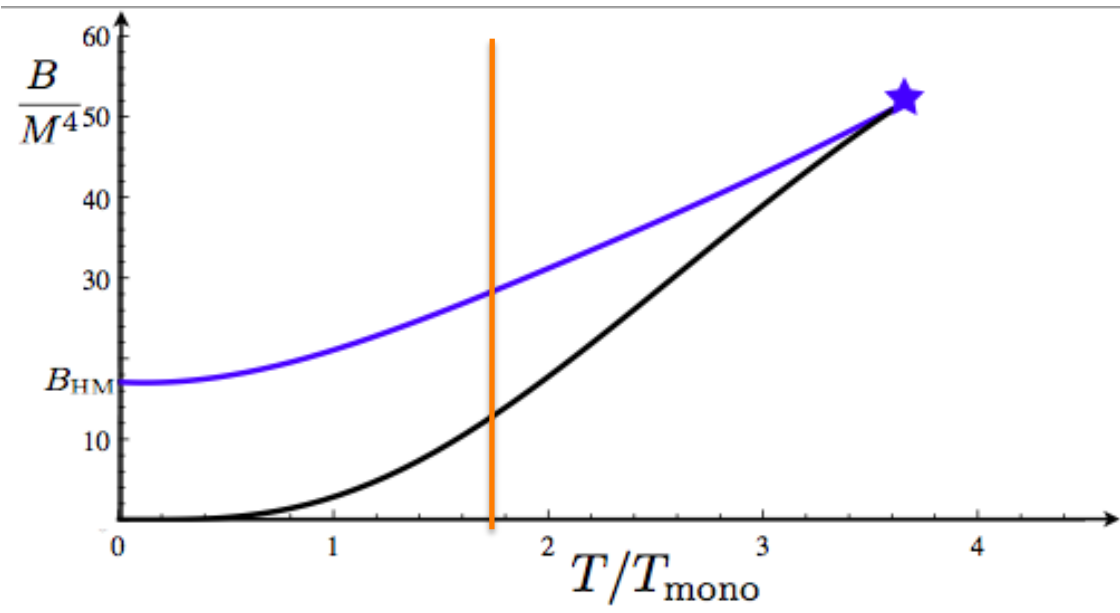
Case 2a:
Decay from
high dS



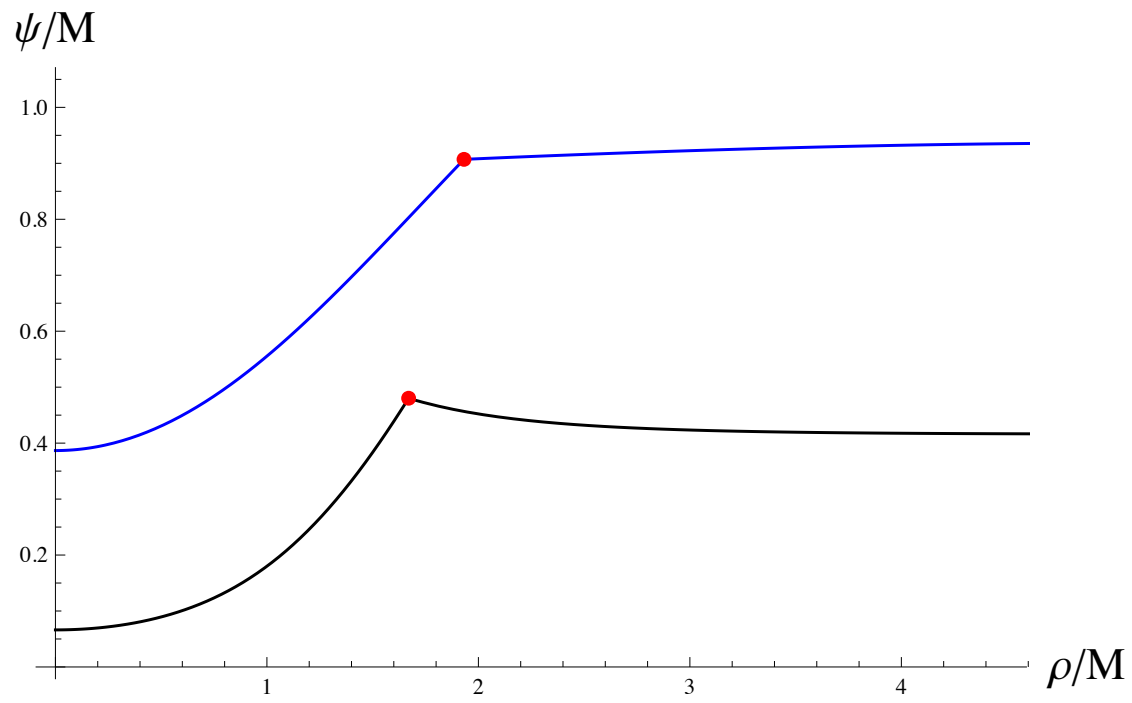
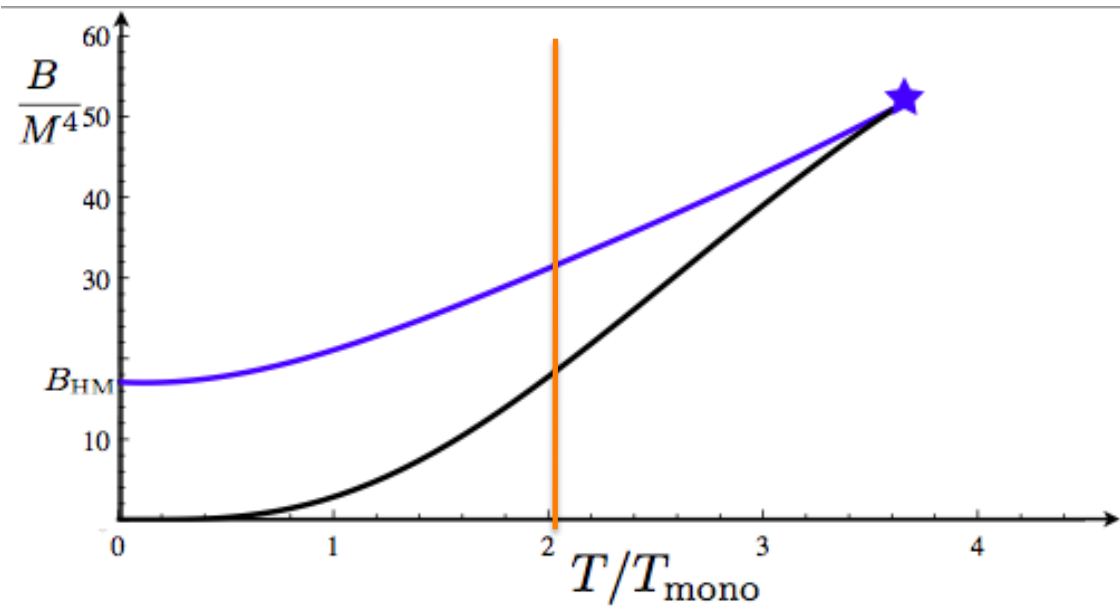
Case 2a:
Decay from
high dS



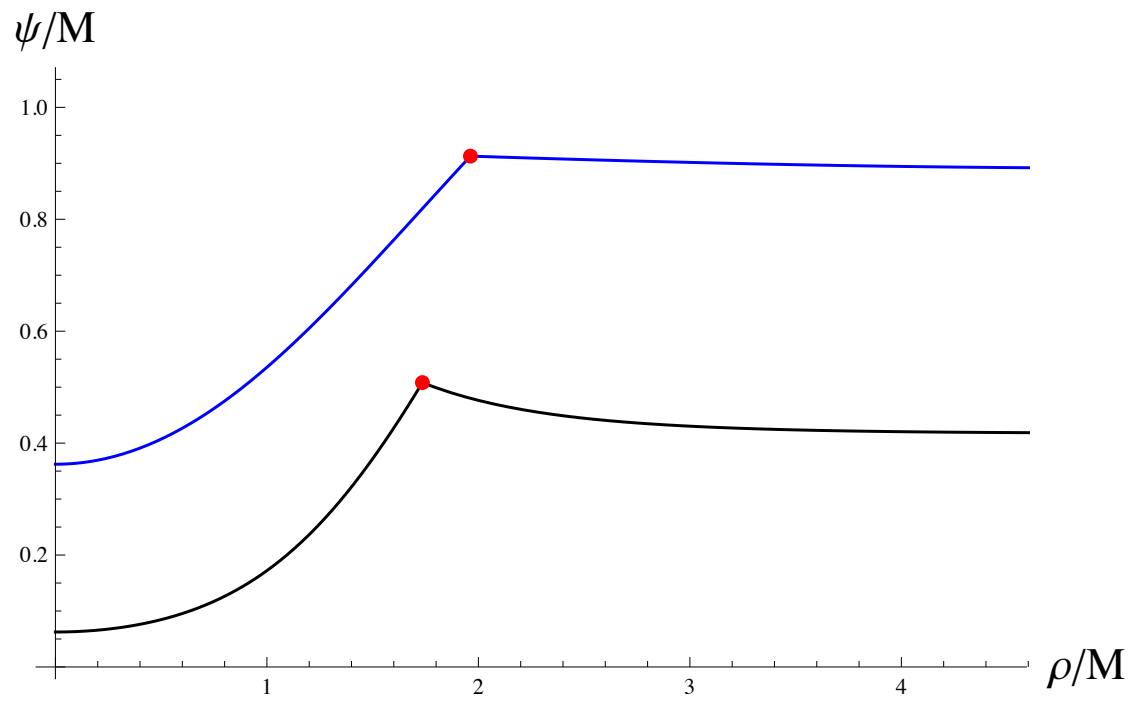
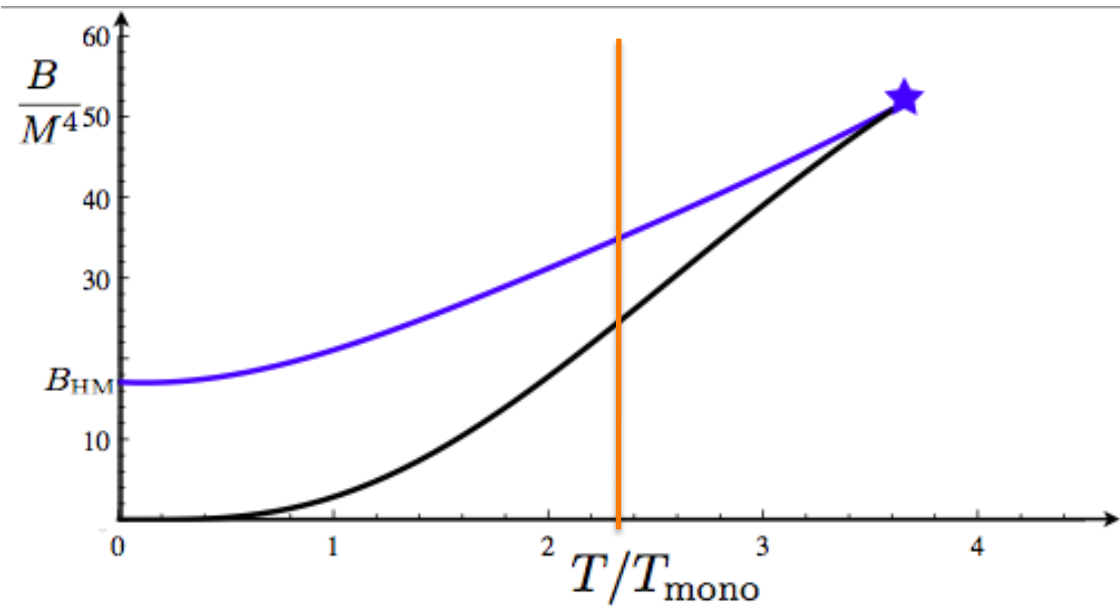
Case 2a:
Decay from
high dS



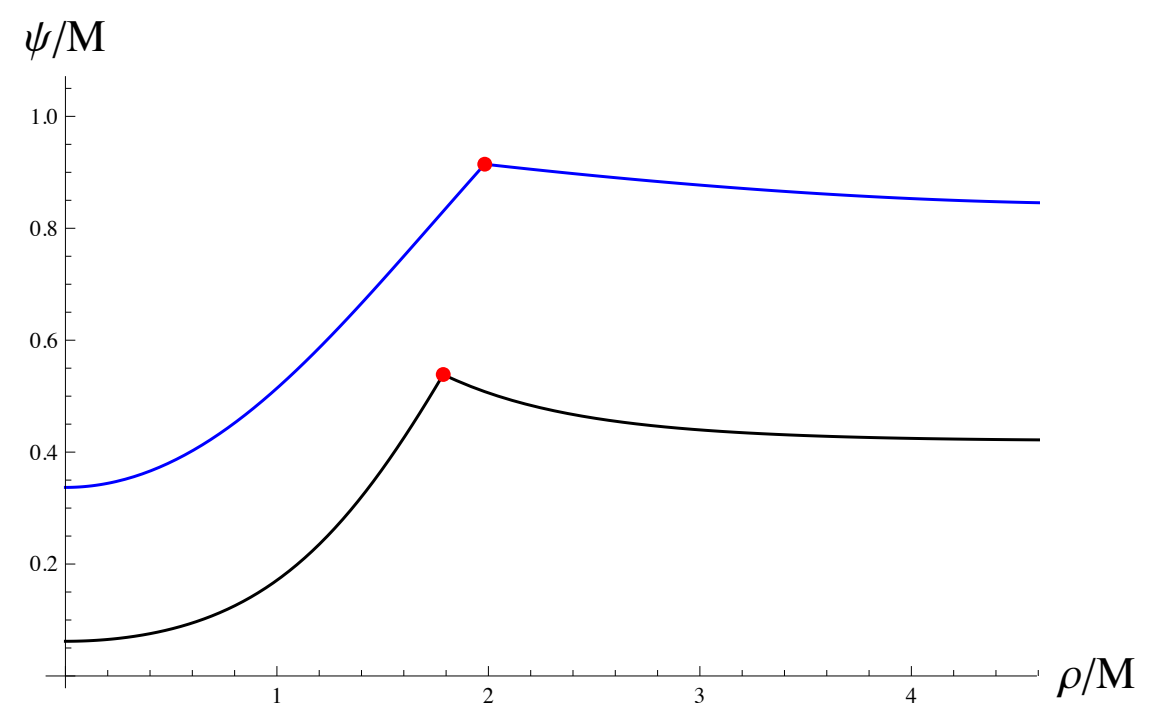
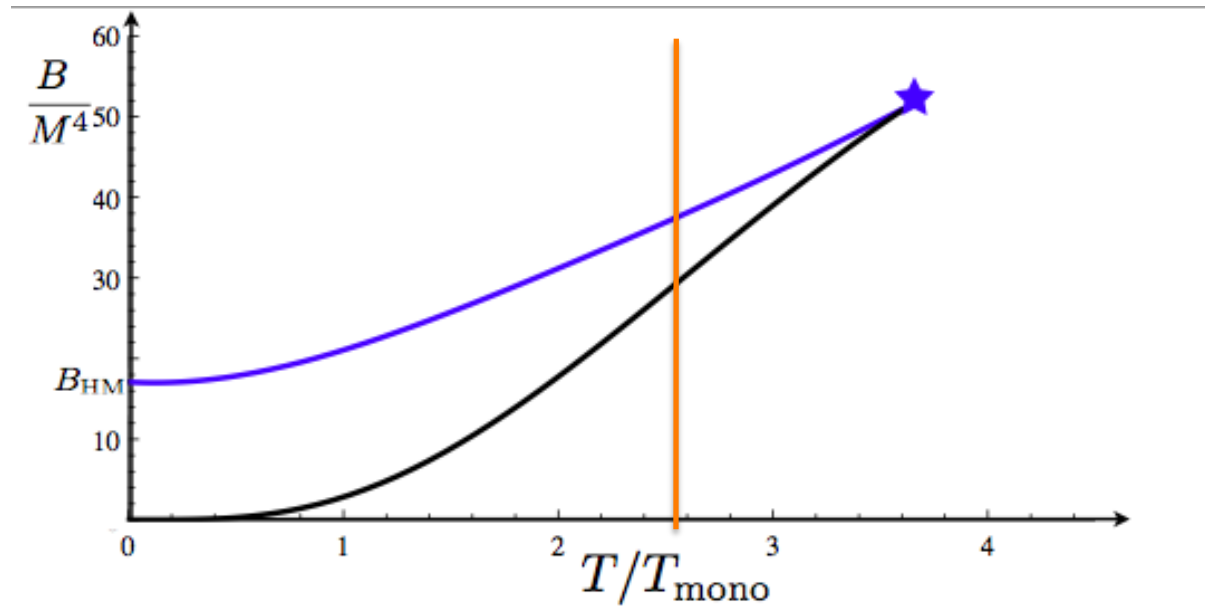
Case 2a:
Decay from
high dS



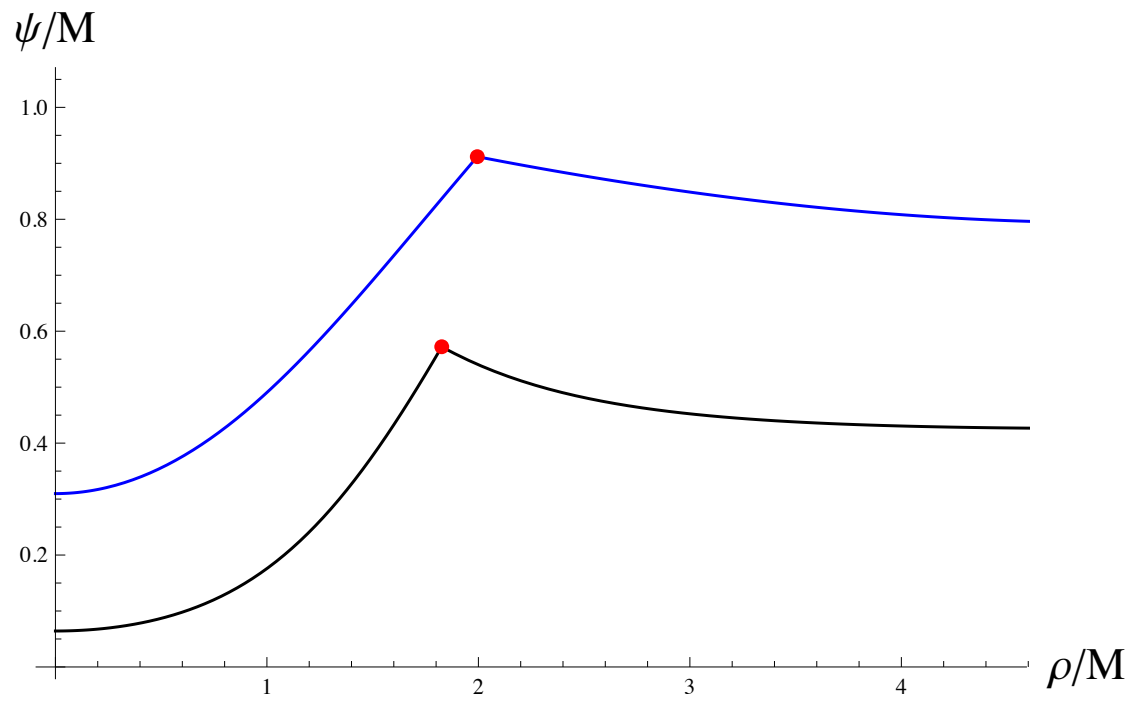
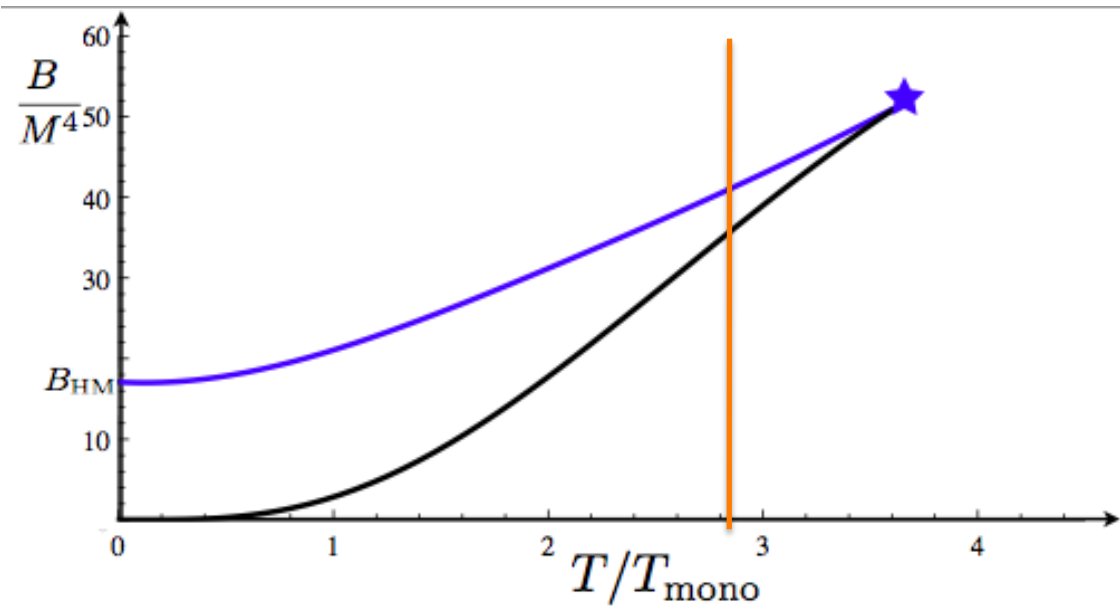
Case 2a:
Decay from
high dS



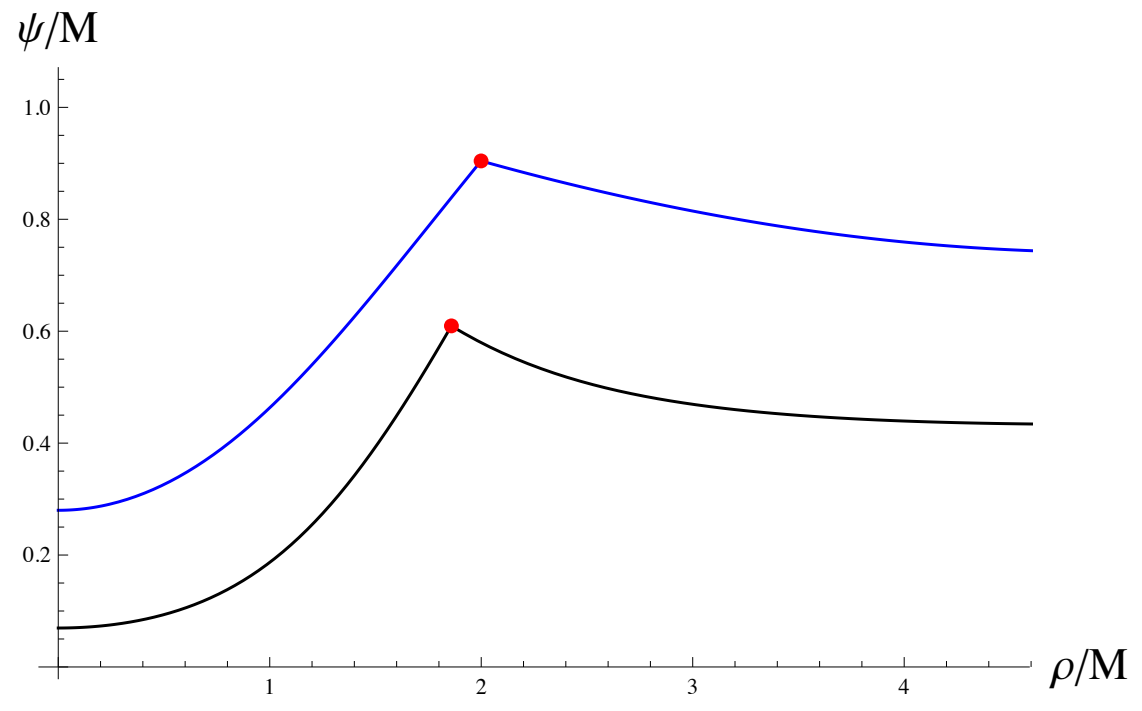
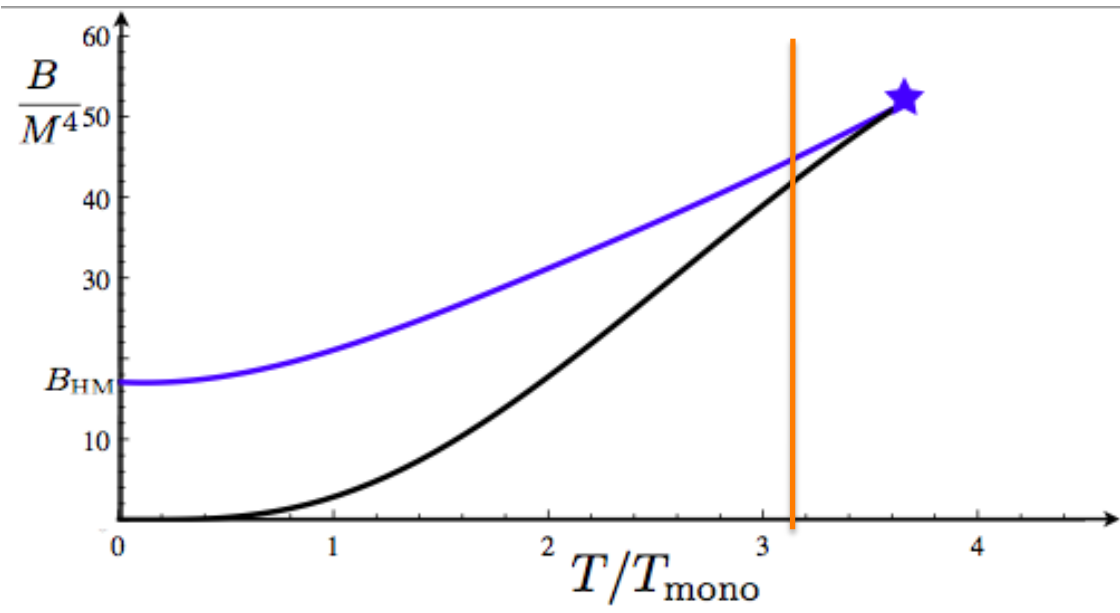
Case 2a:
Decay from
high dS



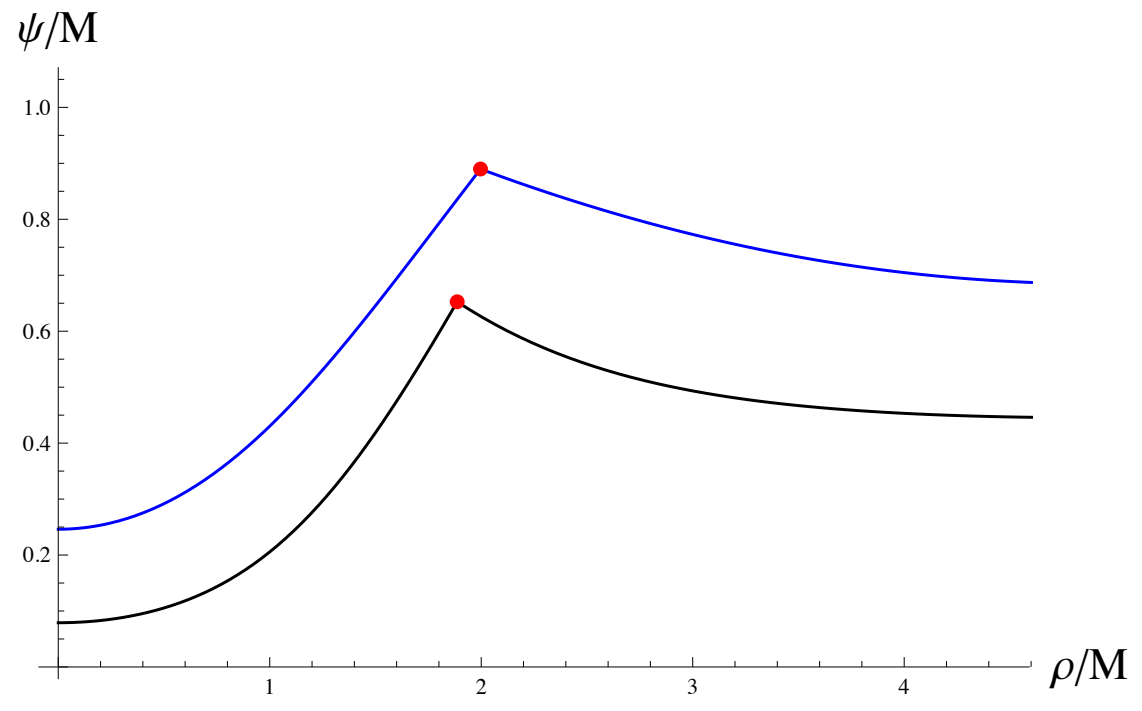
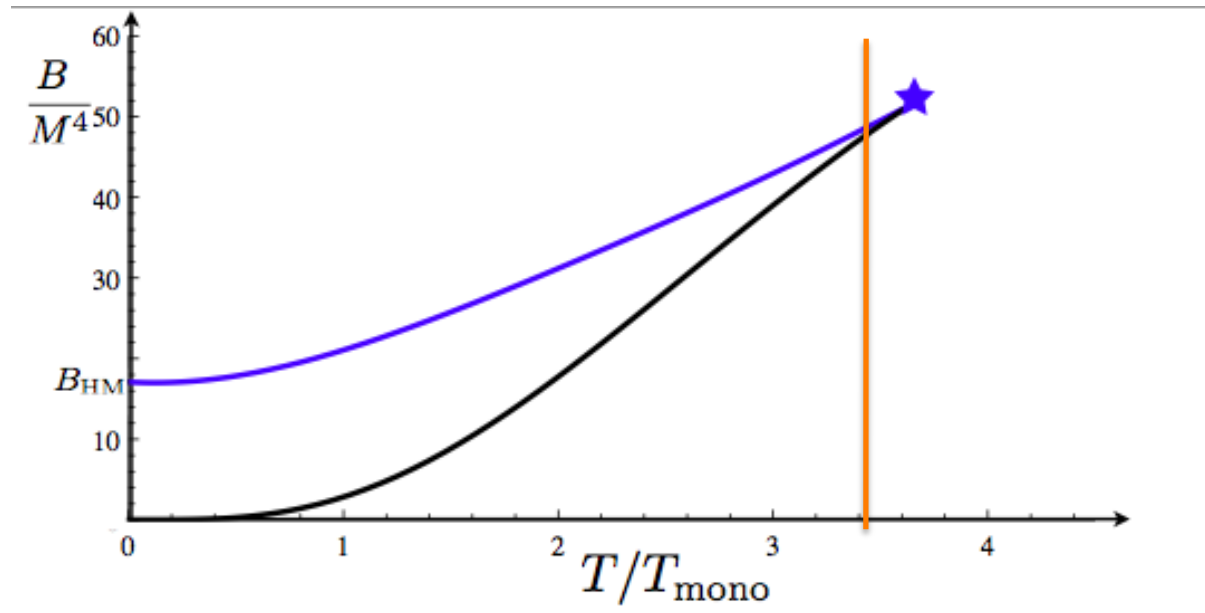
Case 2a:
Decay from
high dS



Case 2a:
Decay from
high dS

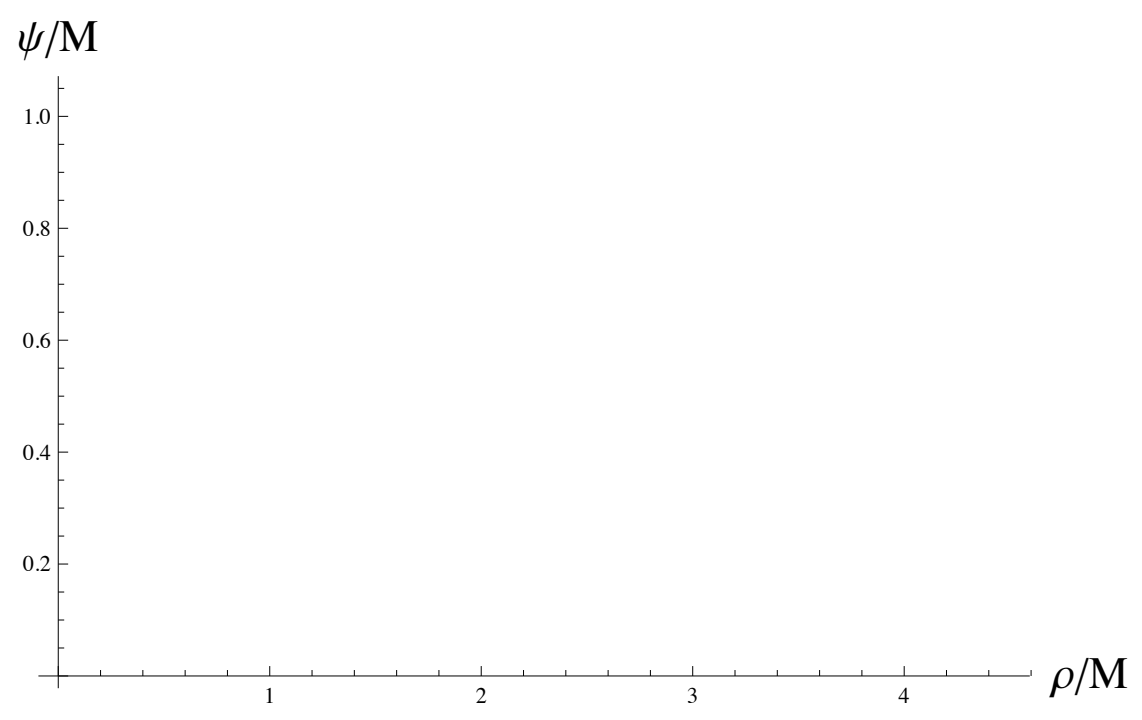
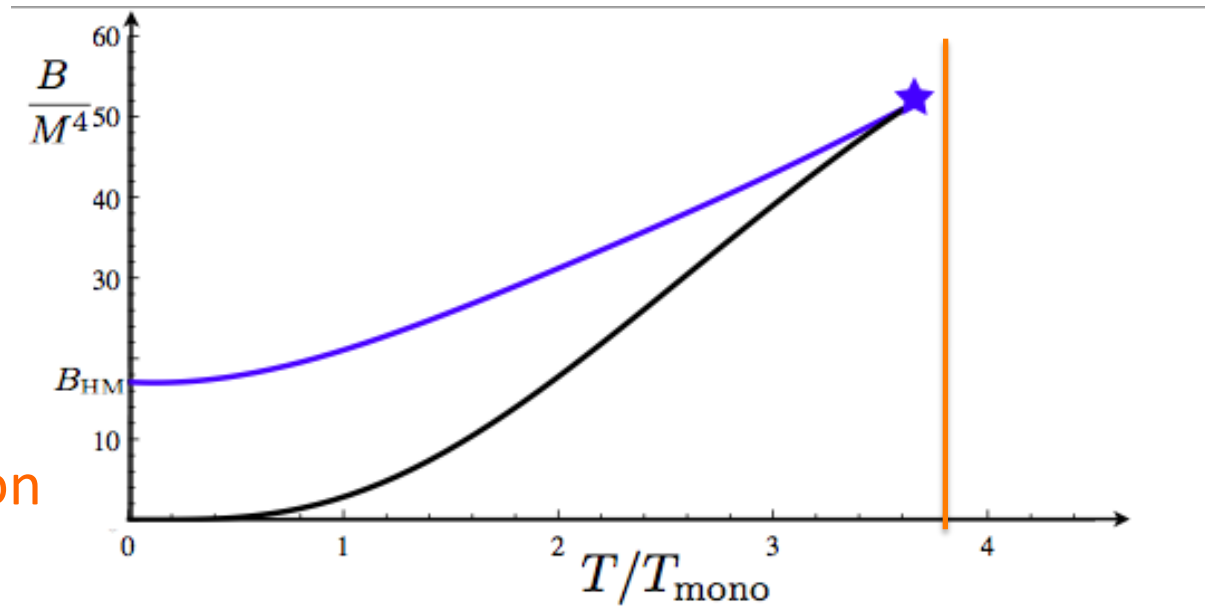


Case 2a:
Decay from
high dS

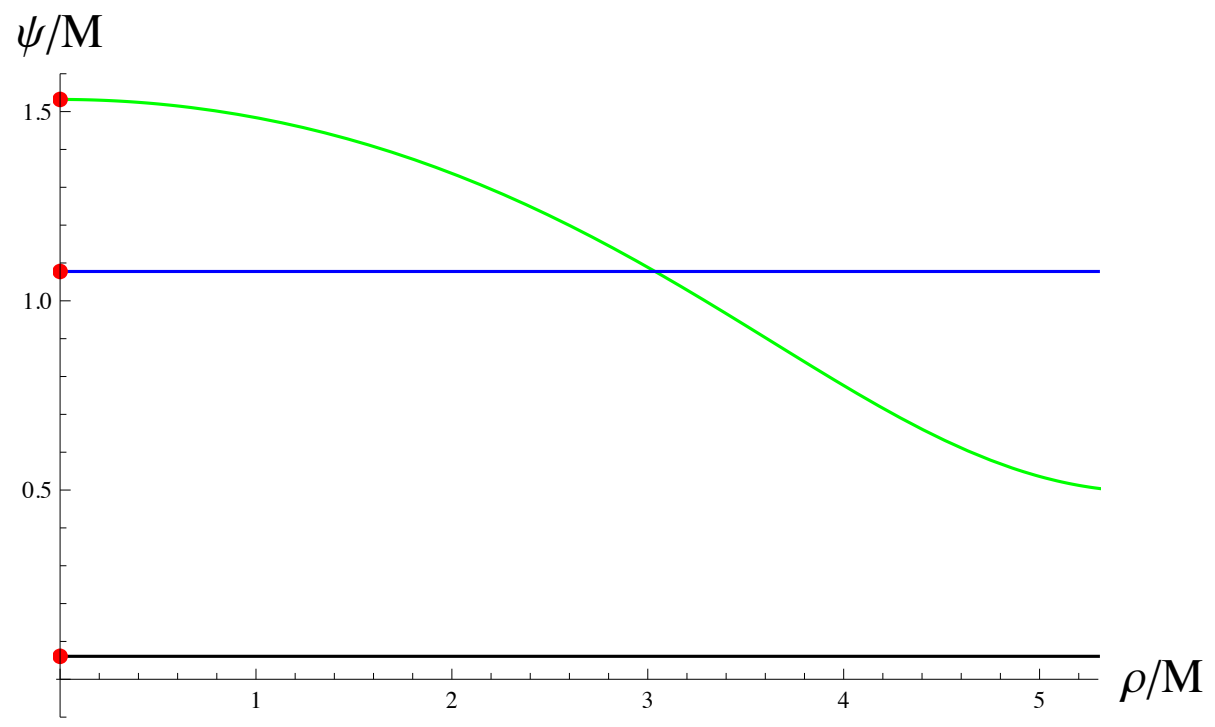
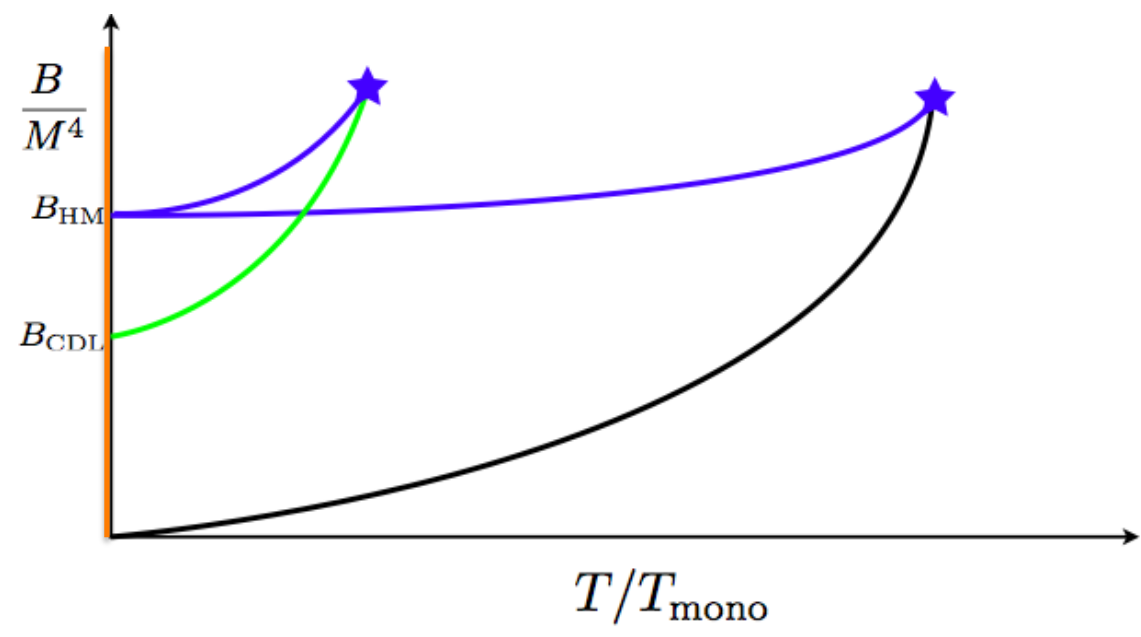


Case 2a:
Decay from
high dS

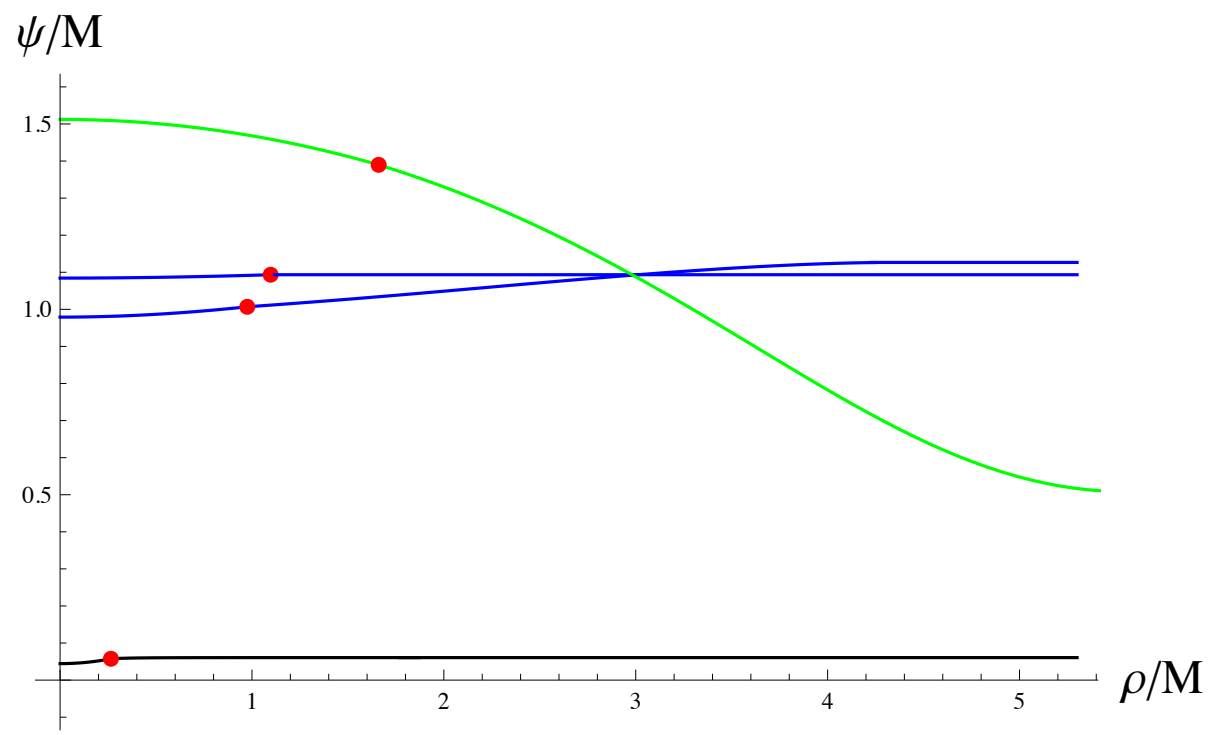
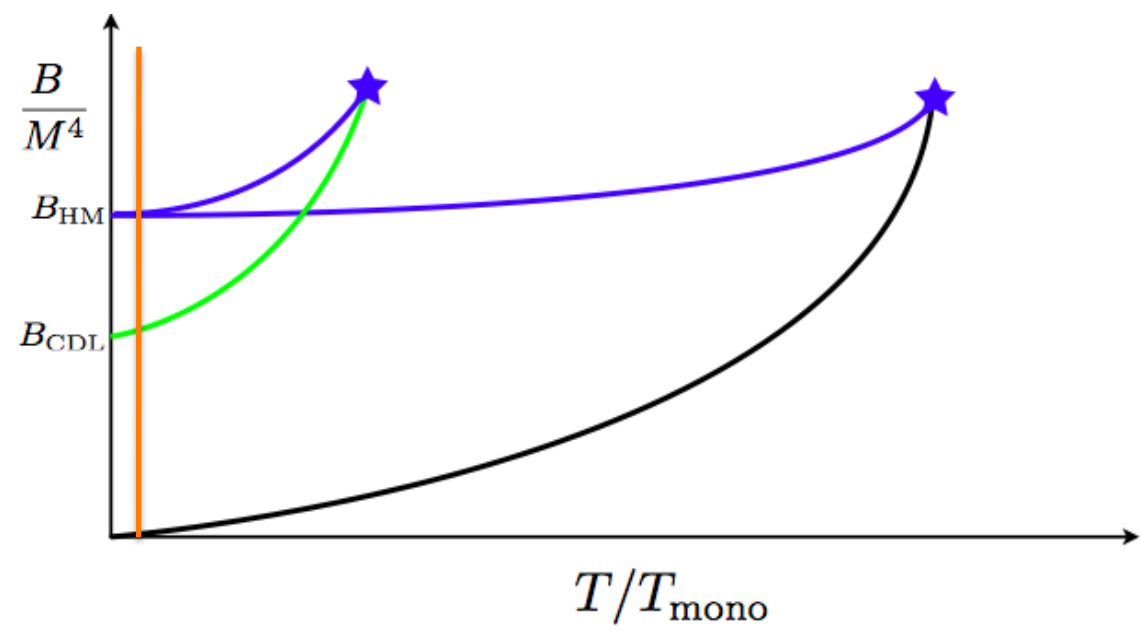
Instanton
Annihilation



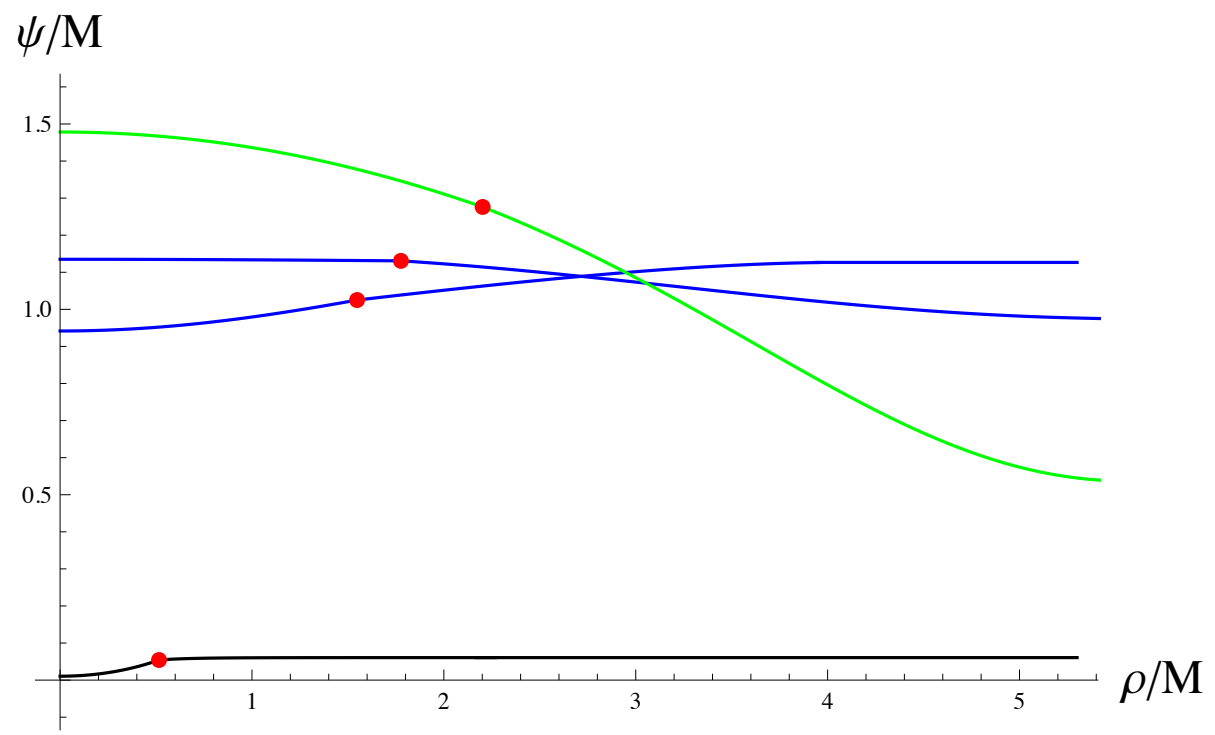
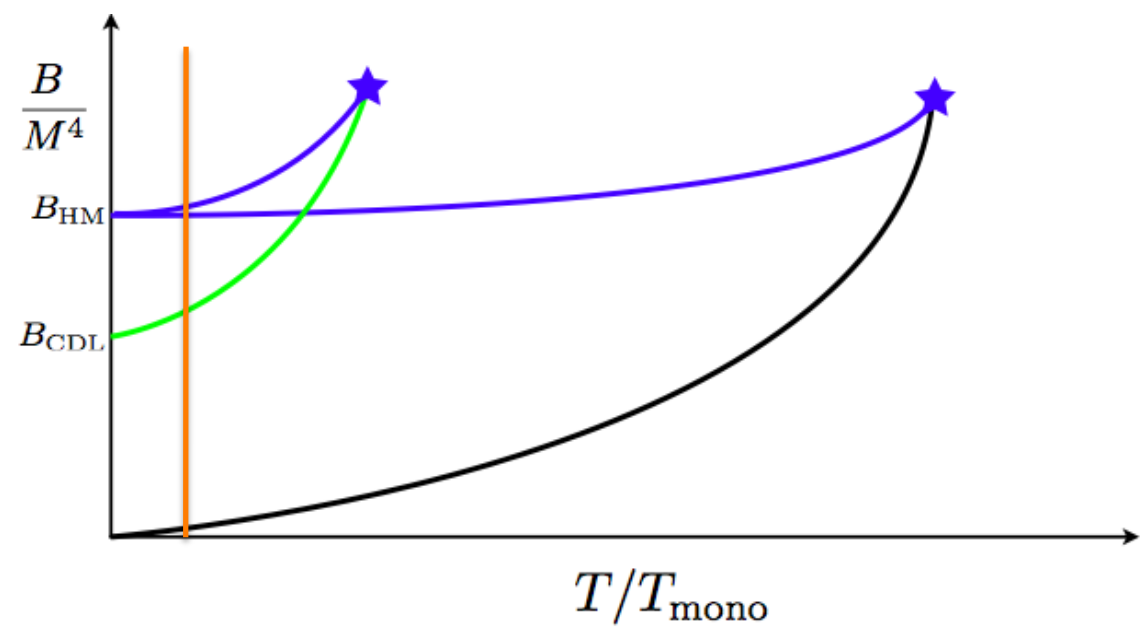
Case 2b:
Decay from
low dS



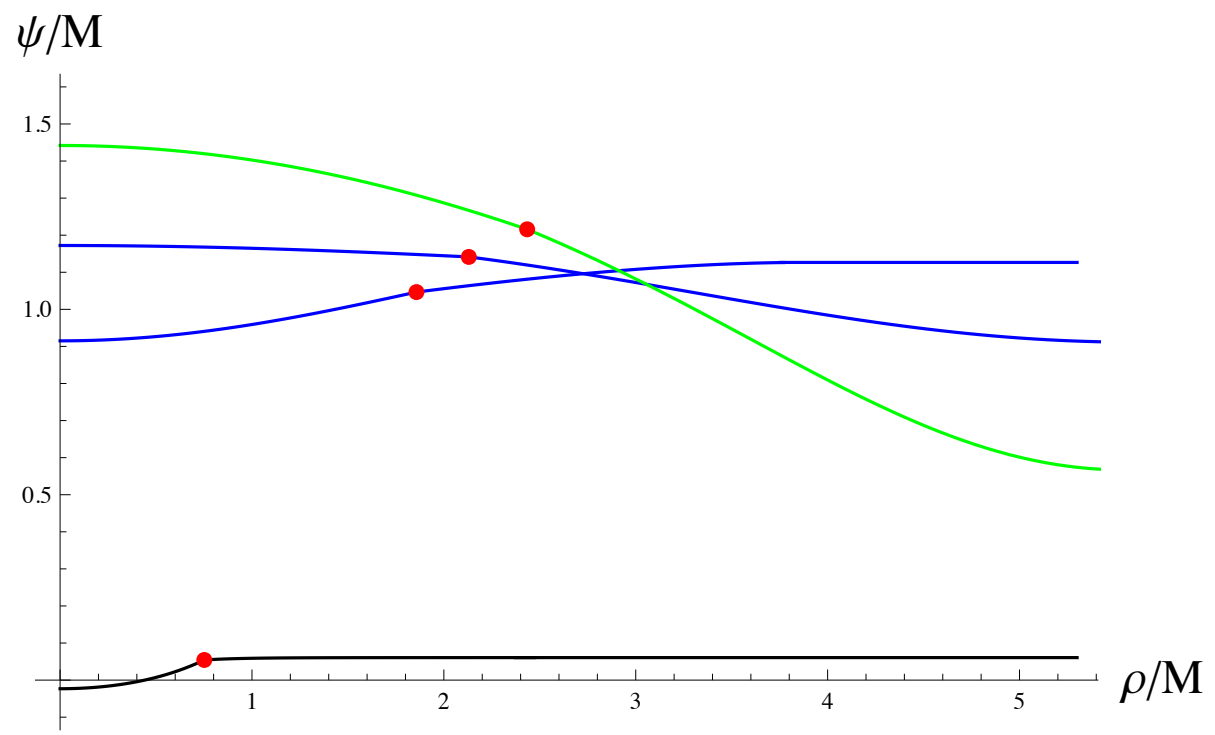
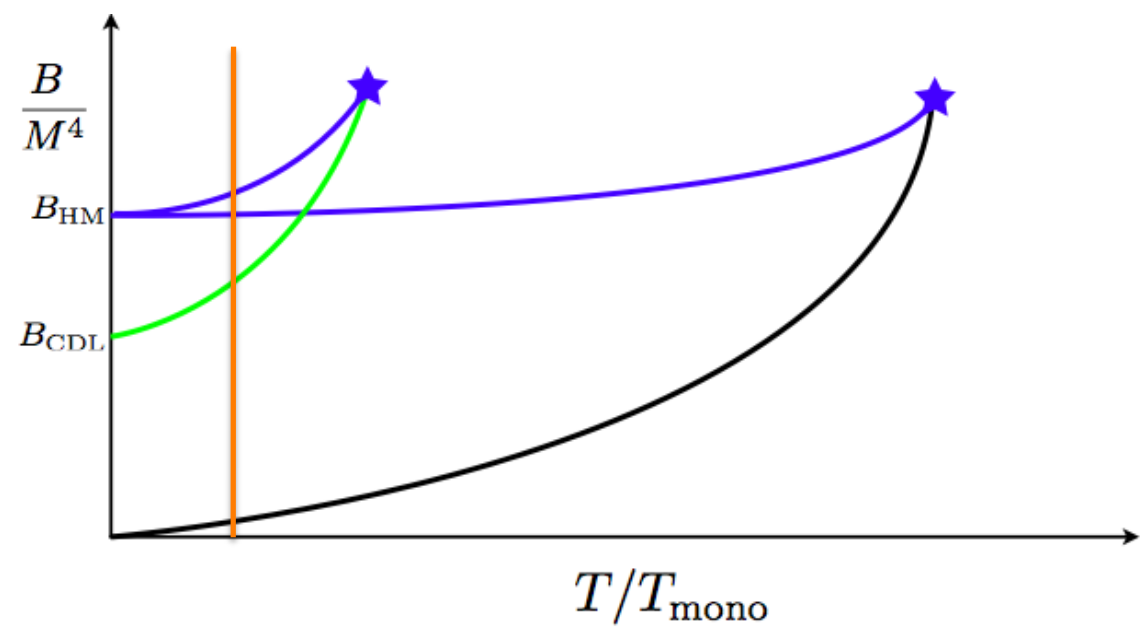
Case 2b:
Decay from
low dS



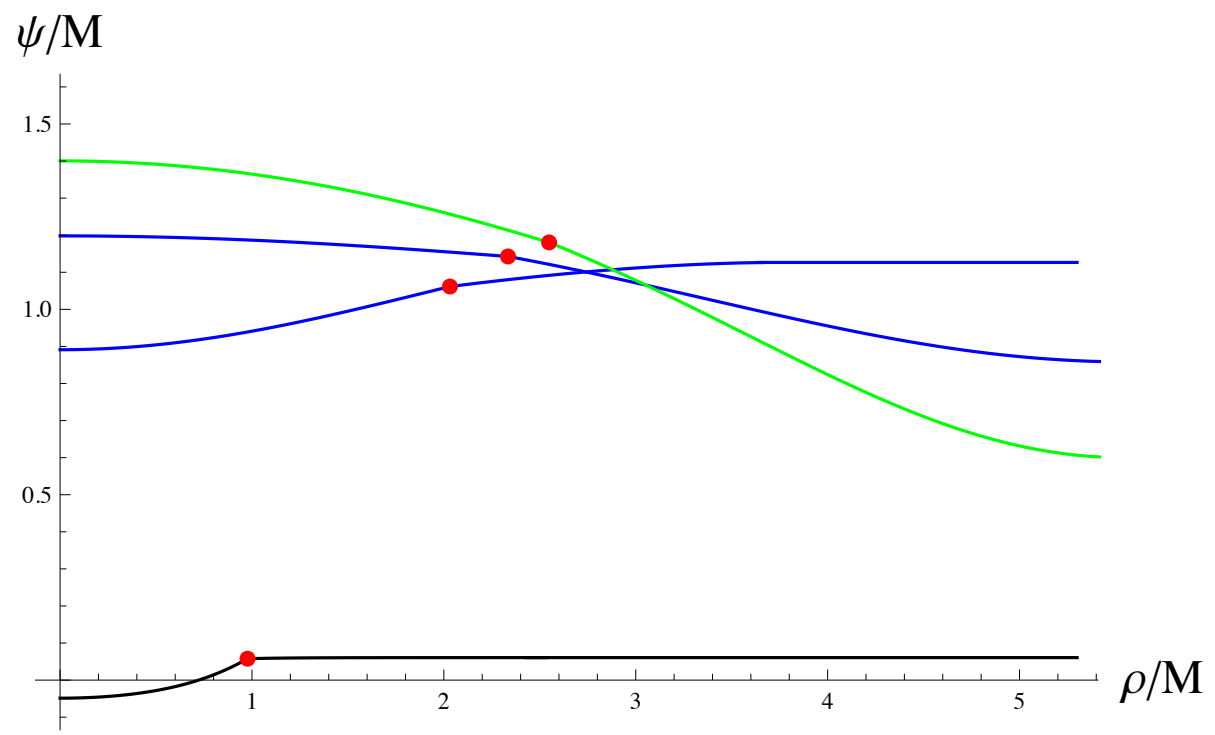
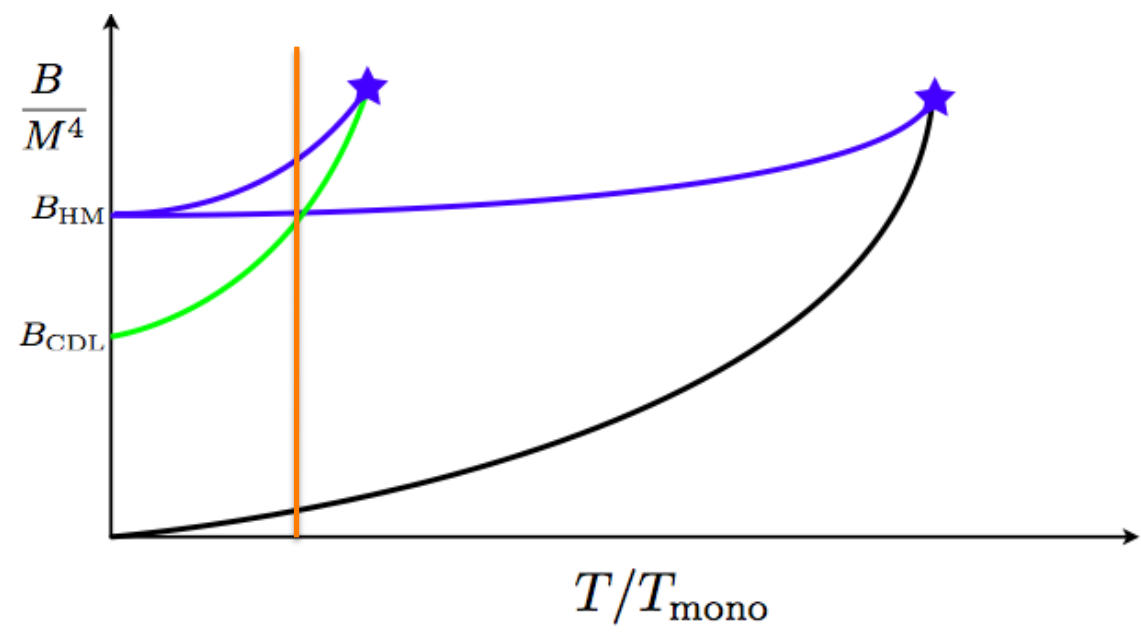
Case 2b:
Decay from
low dS



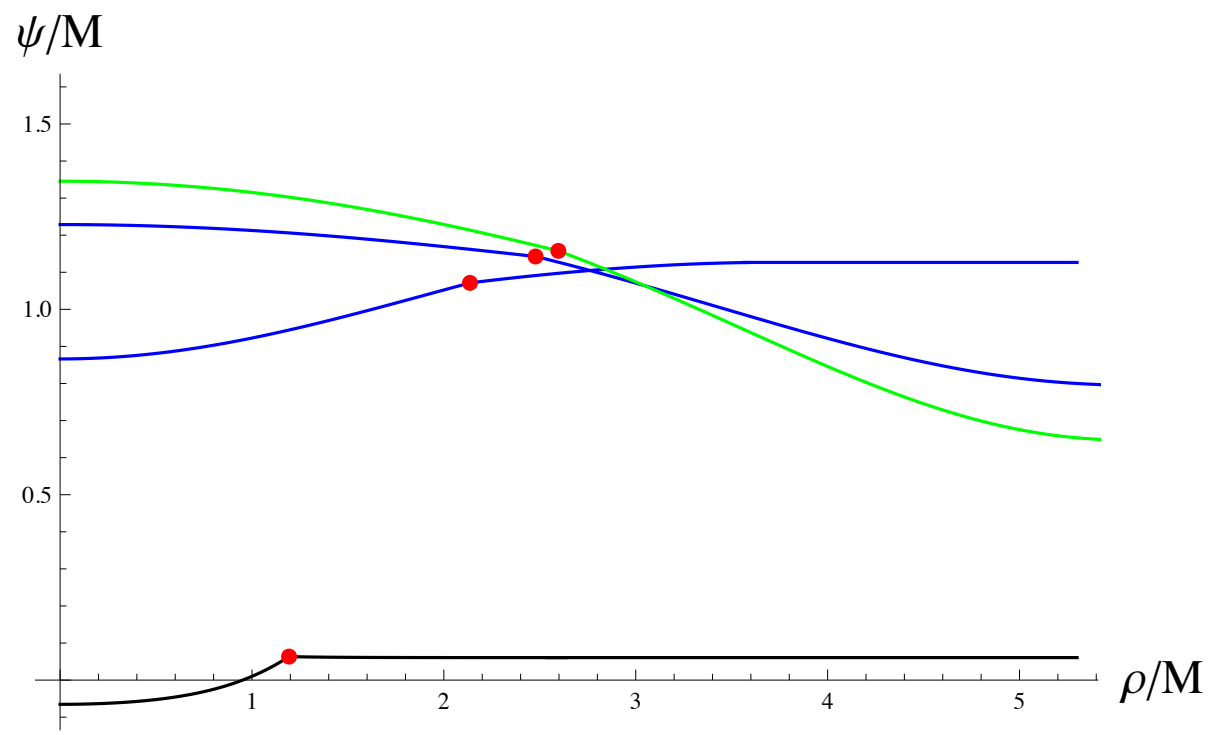
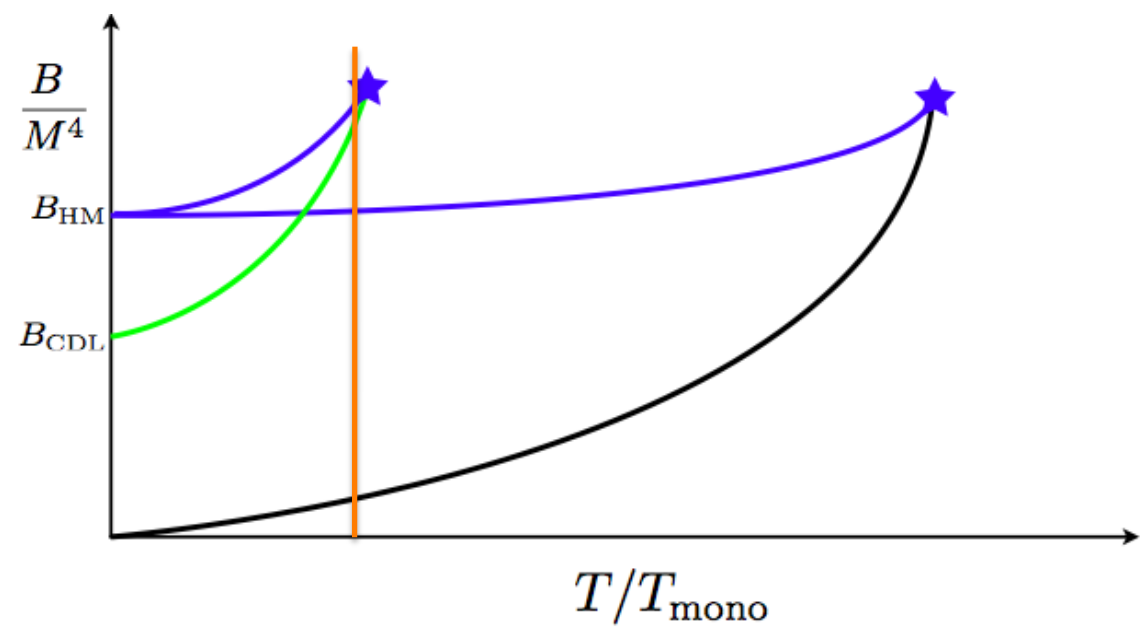
Case 2b:
Decay from
low dS



Case 2b:
Decay from
low dS

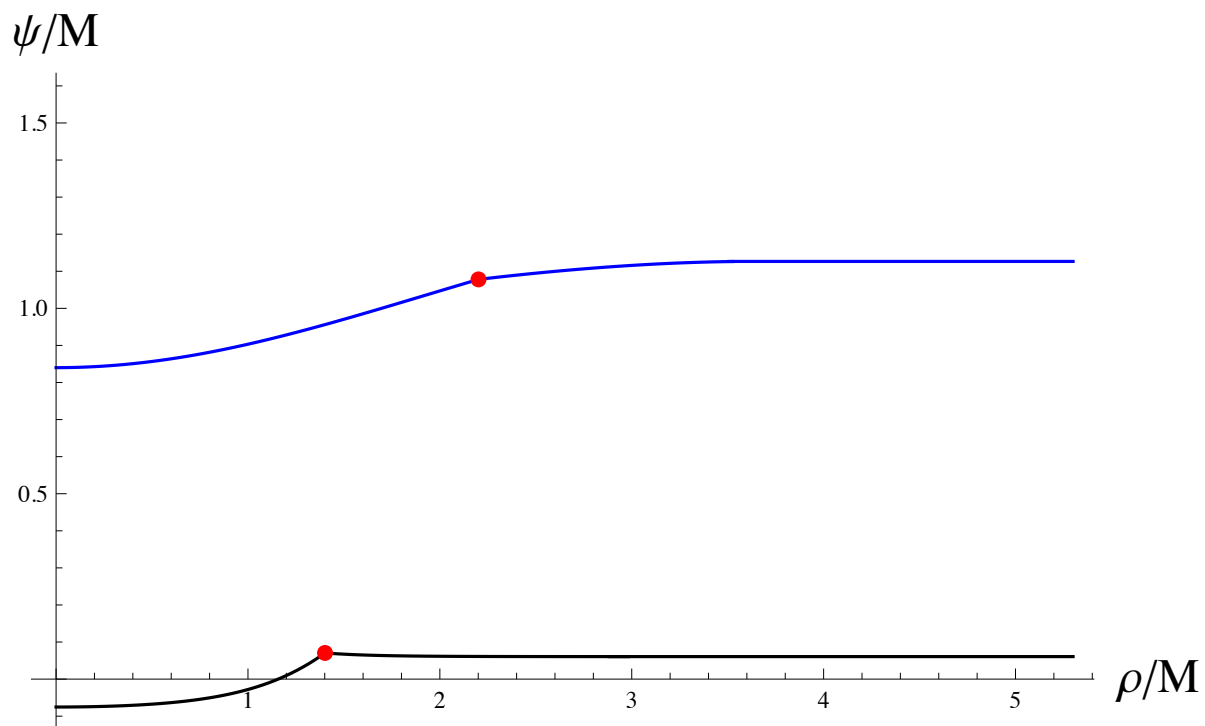
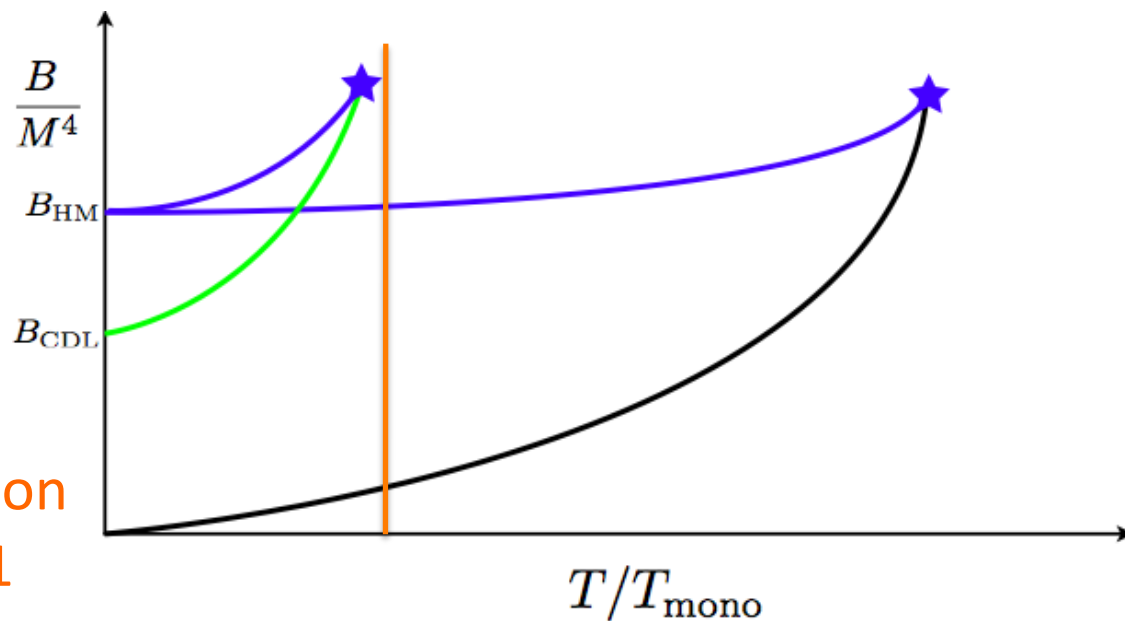


Case 2b:
Decay from
low dS

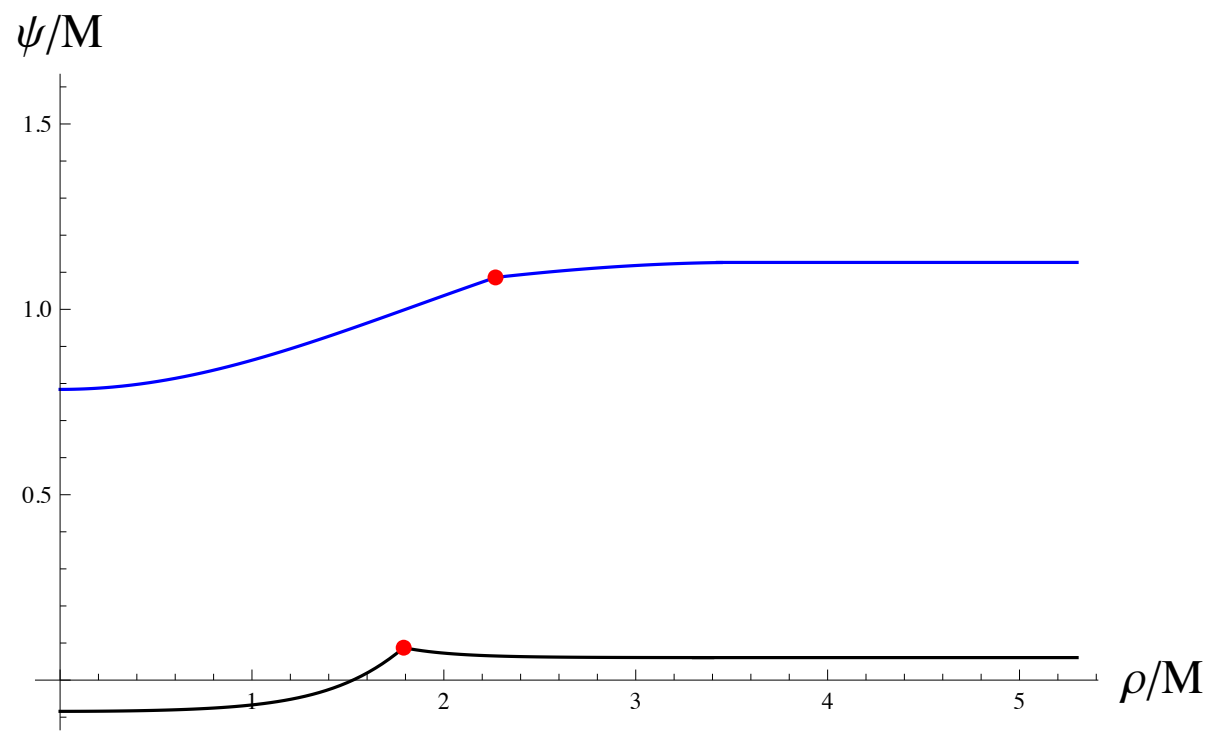
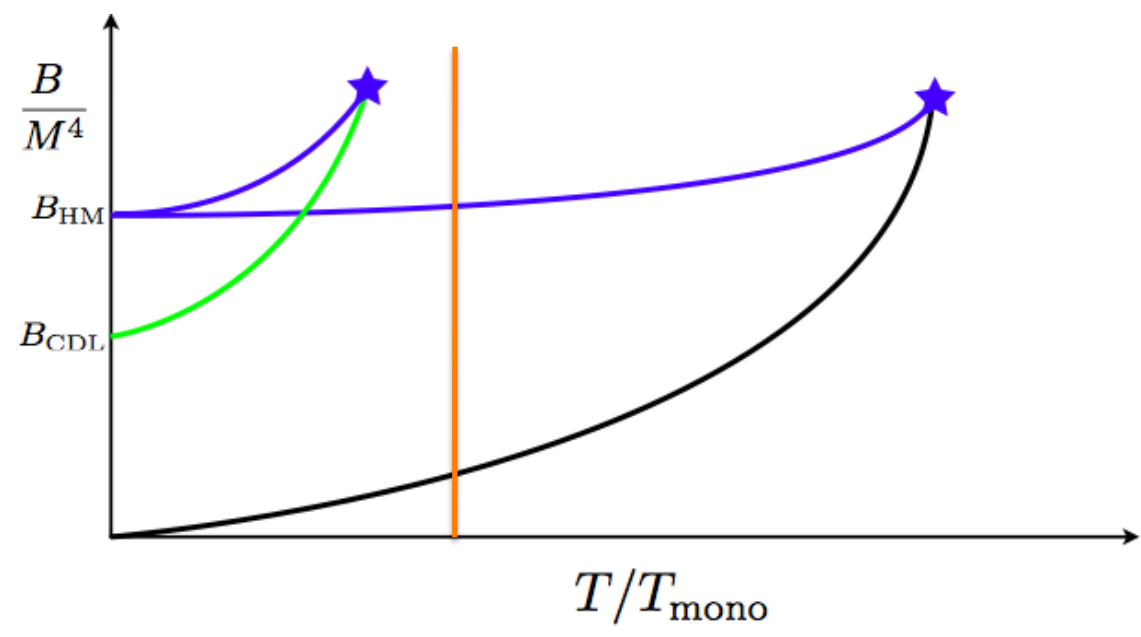


Case 2b:
Decay from
low dS

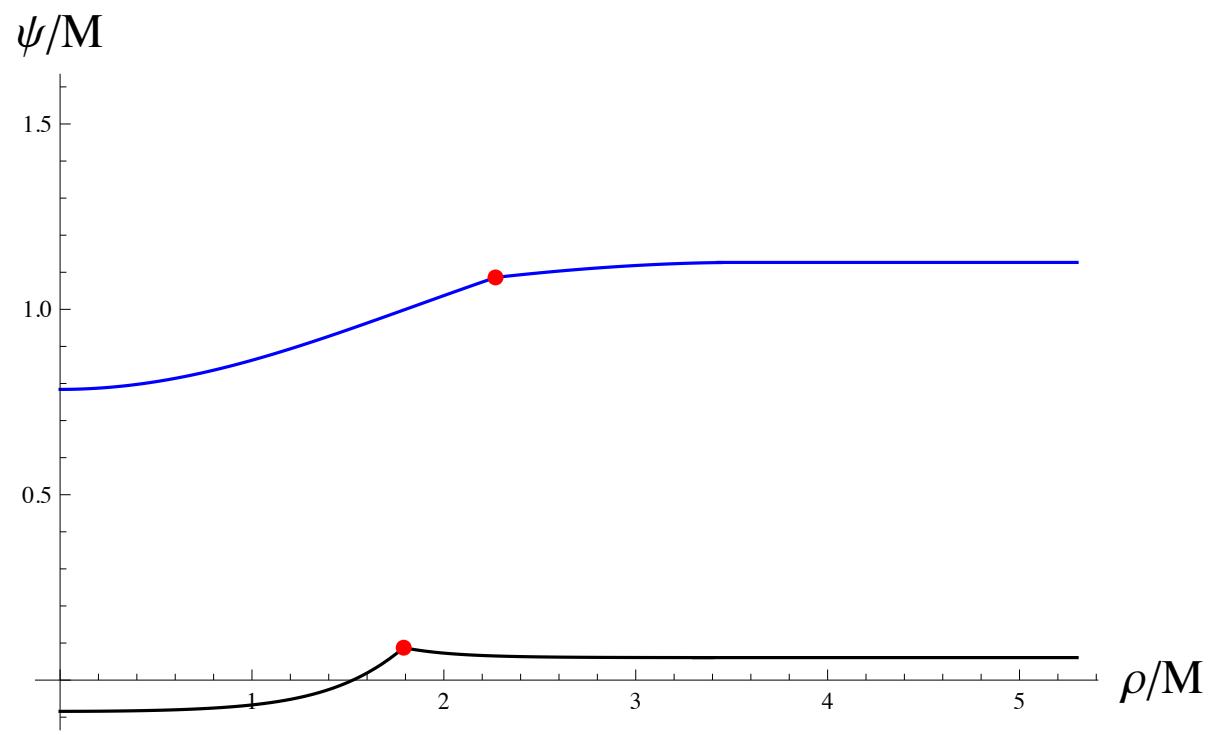
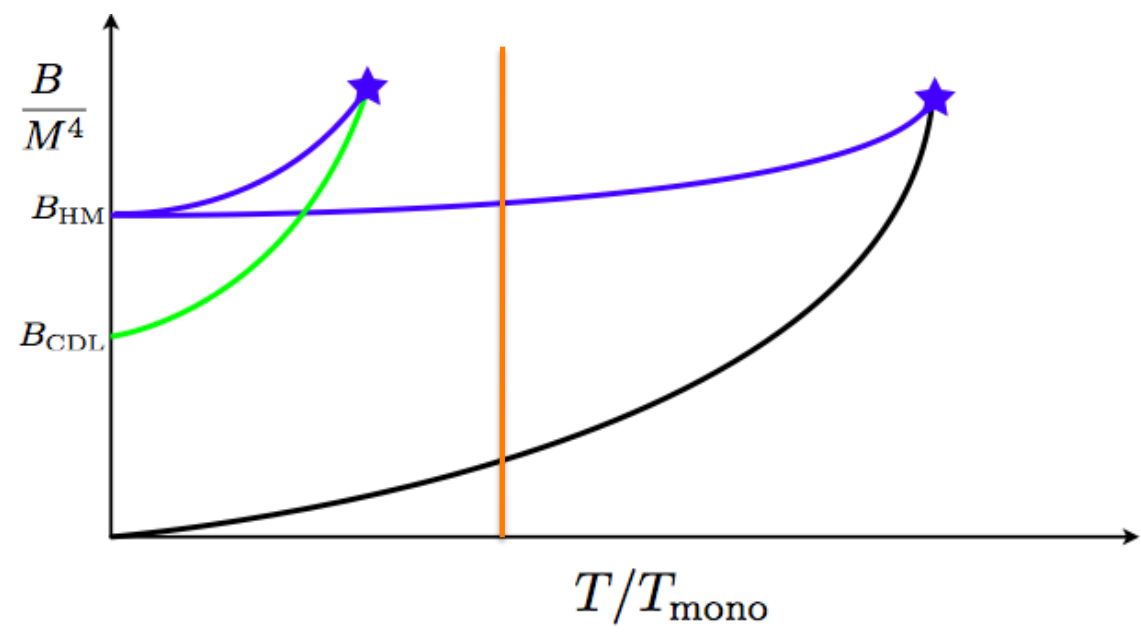
Instanton
Annihilation
Number 1



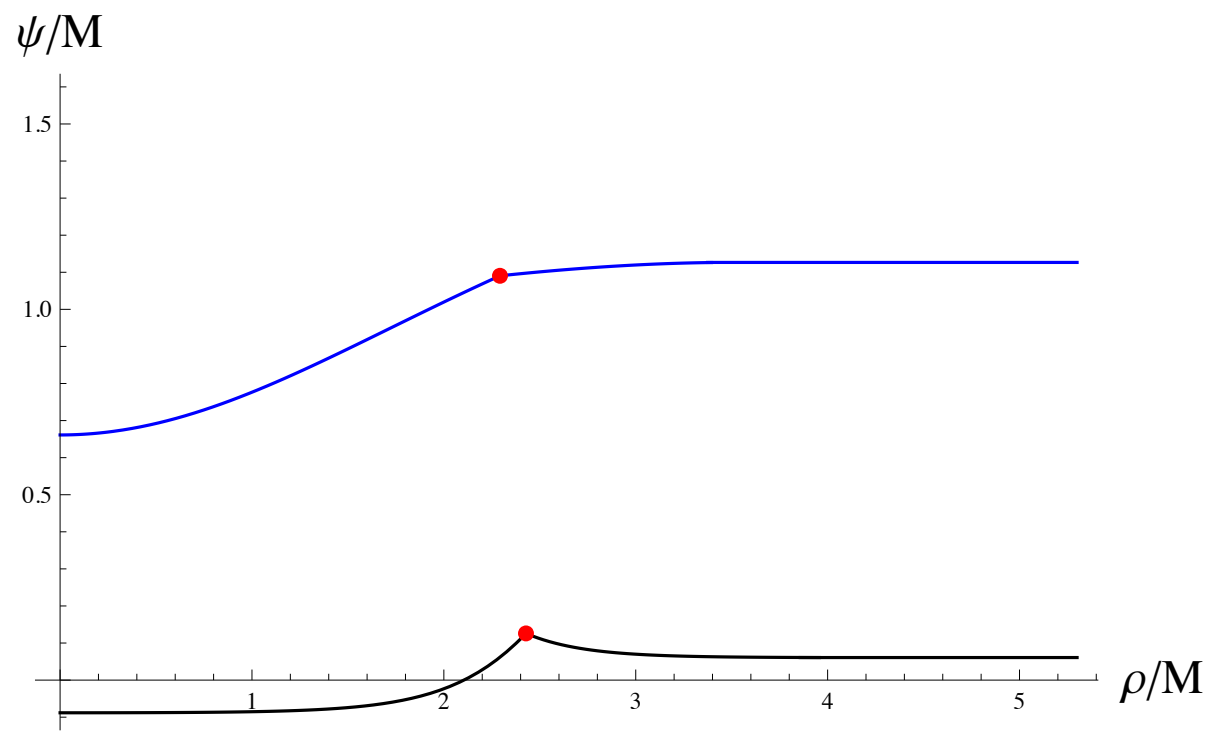
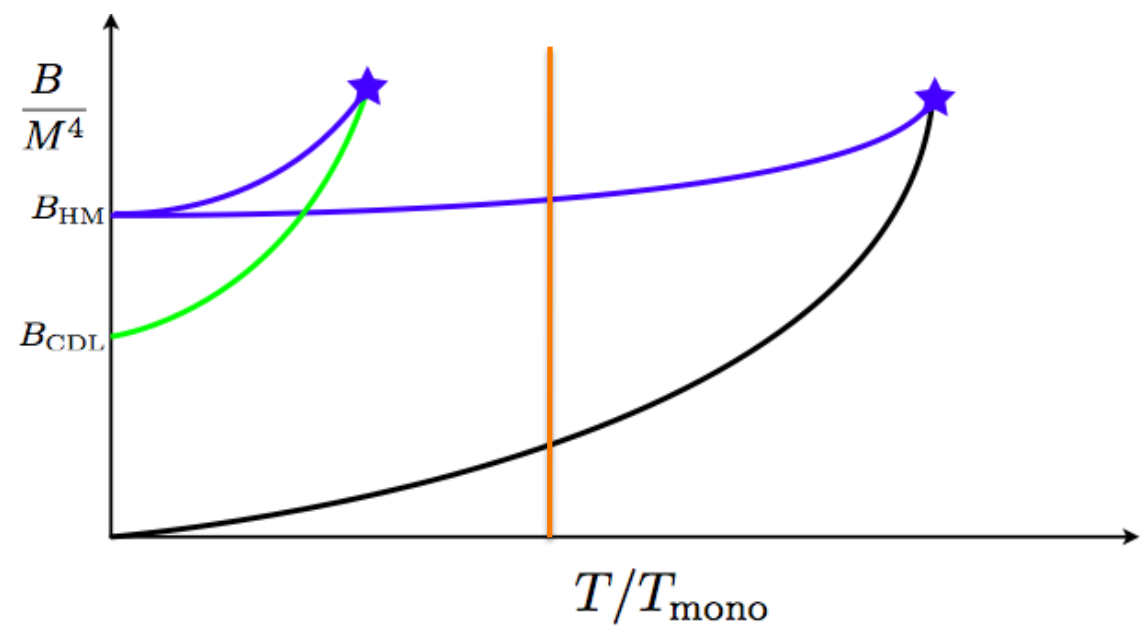
Case 2b:
Decay from
low dS



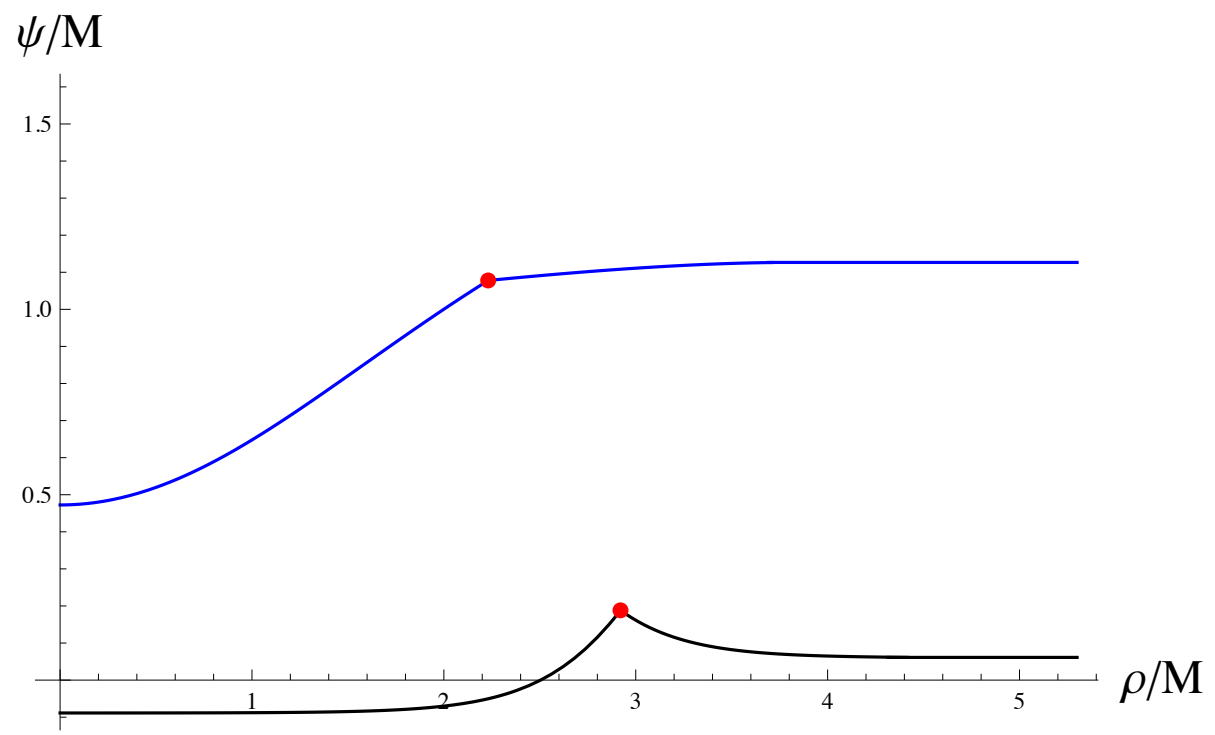
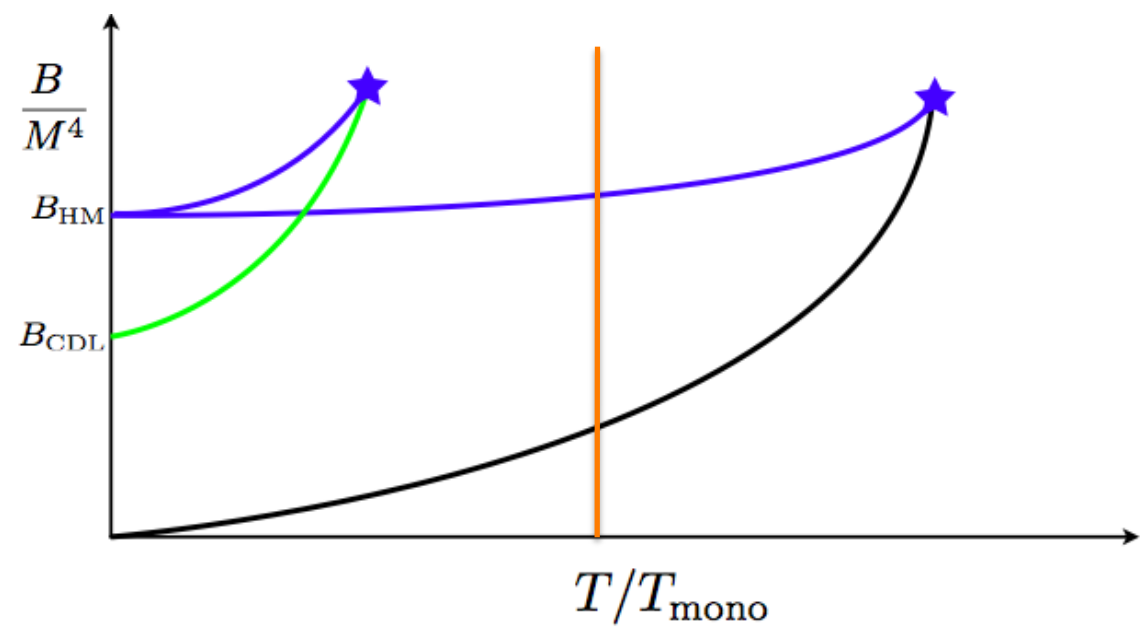
Case 2b:
Decay from
low dS



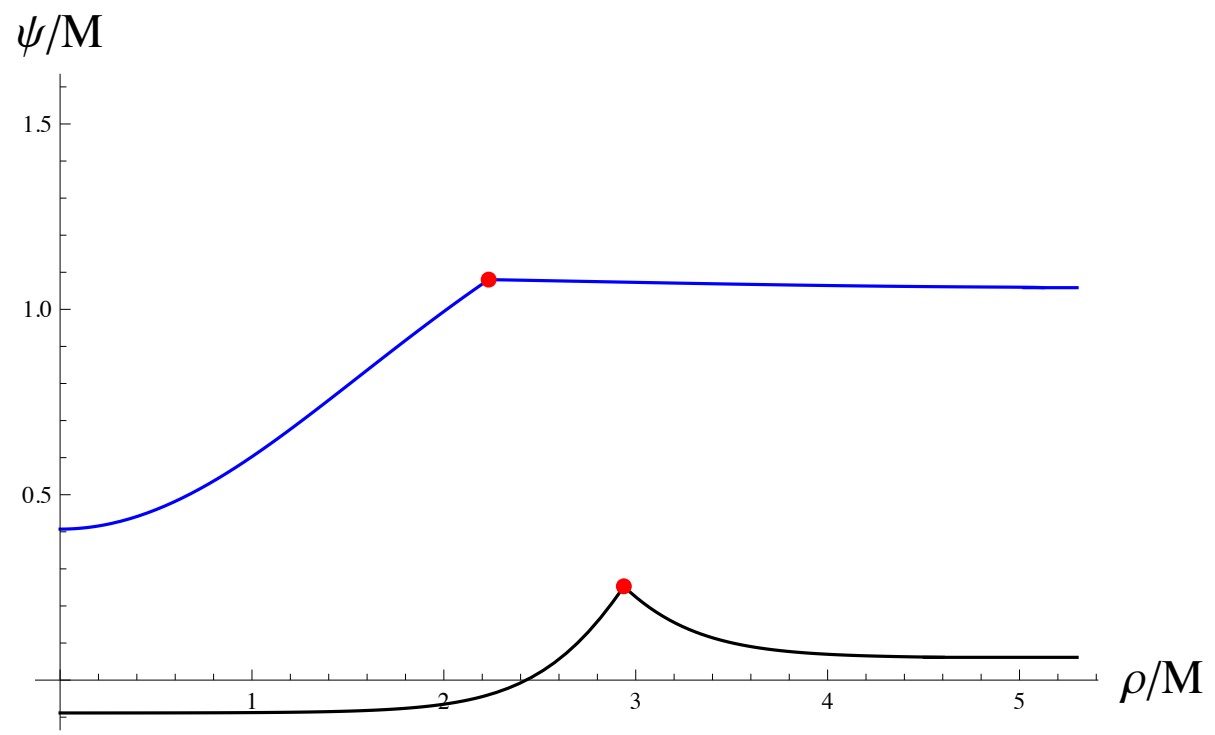
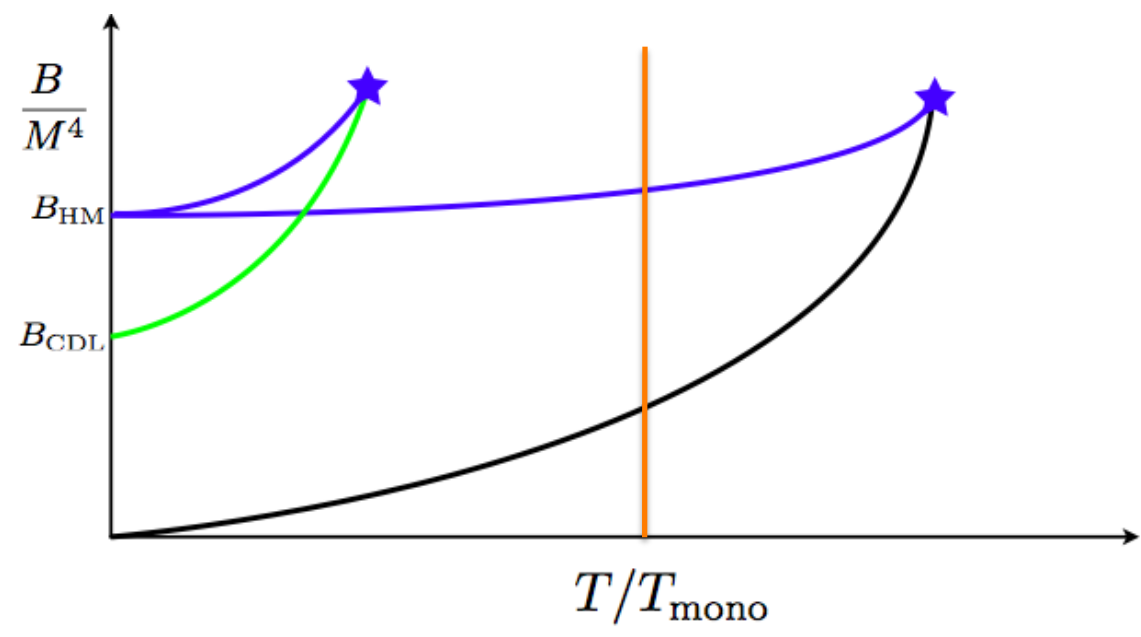
Case 2b:
Decay from
low dS



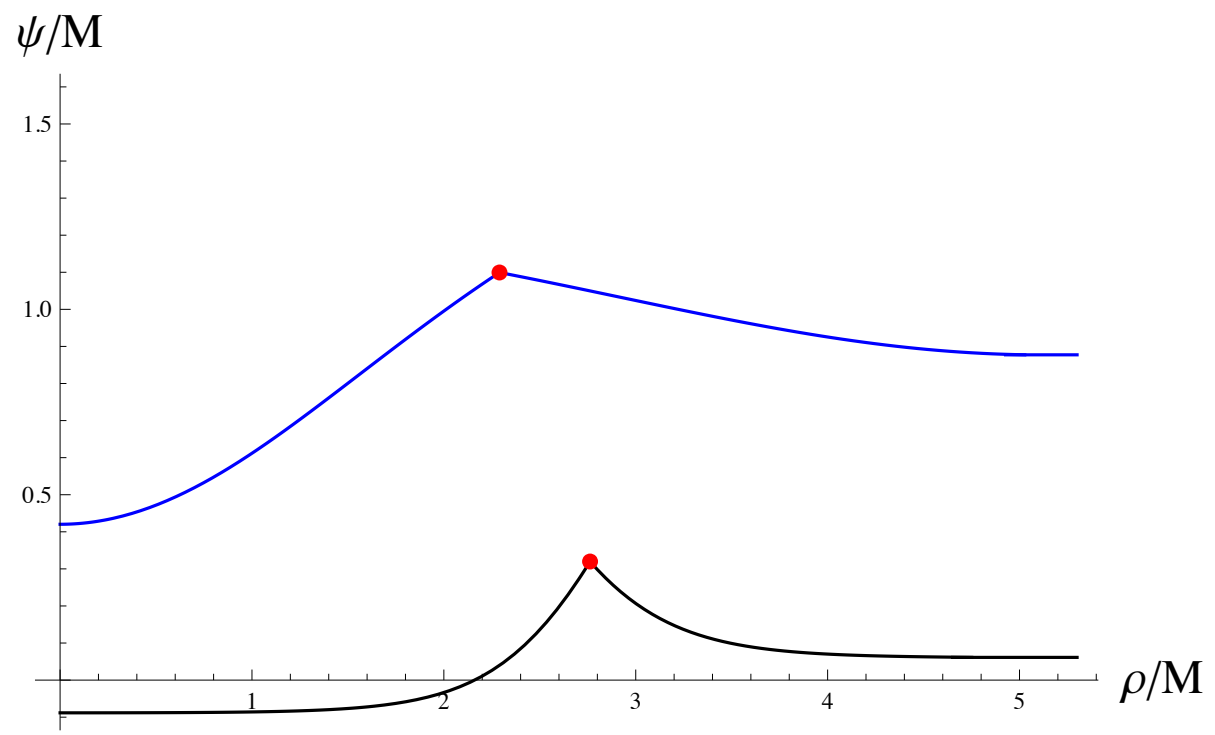
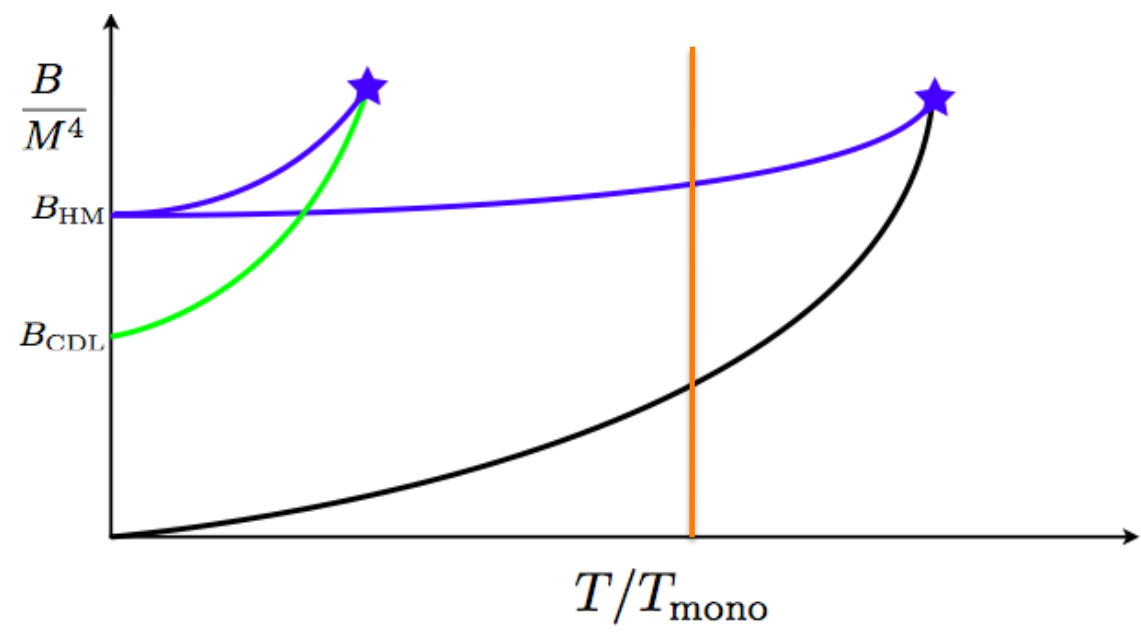
Case 2b:
Decay from
low dS



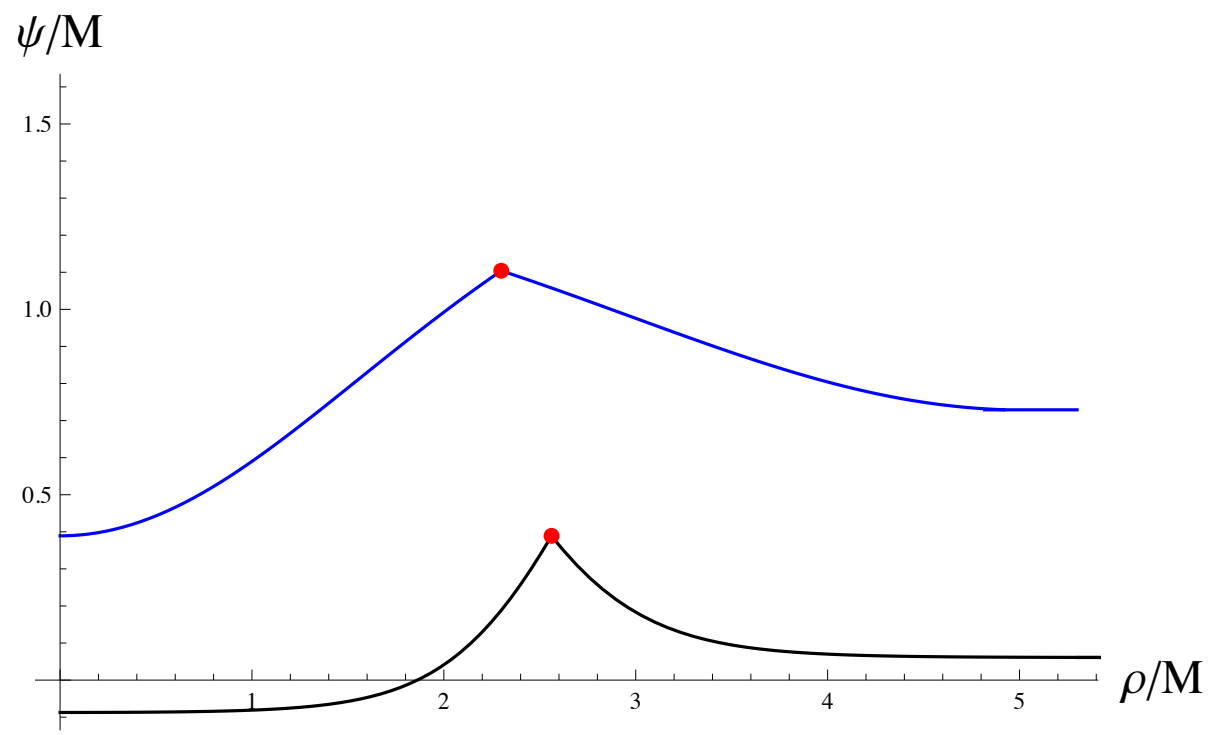
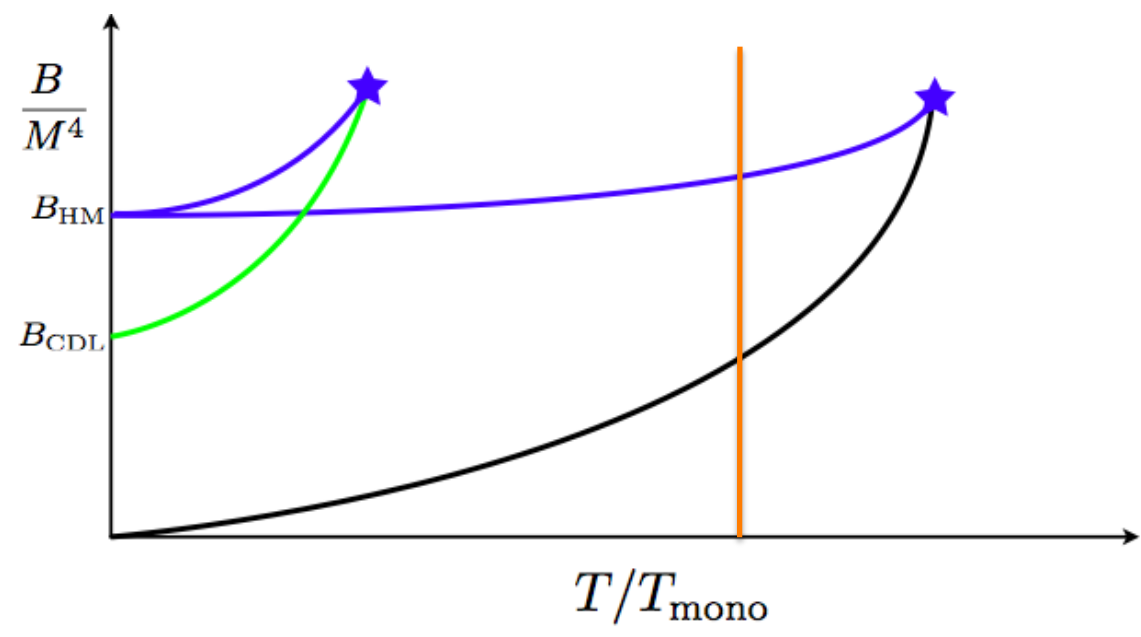
Case 2b:
Decay from
low dS



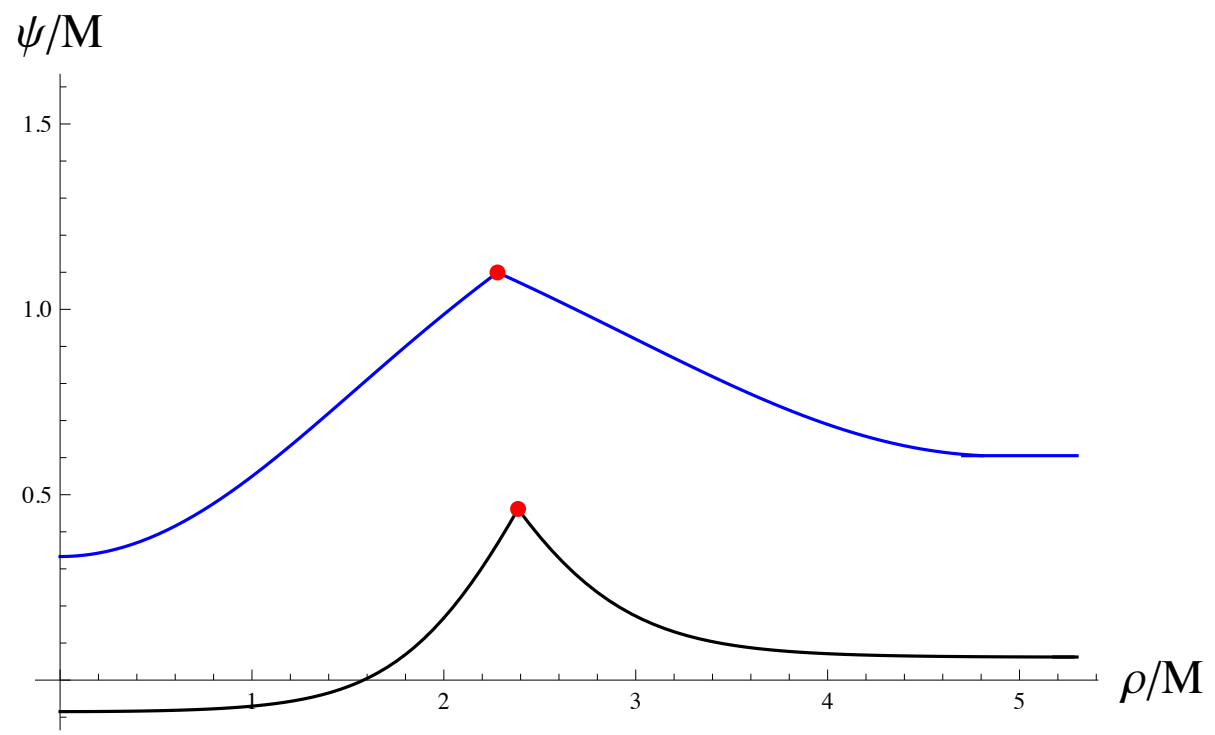
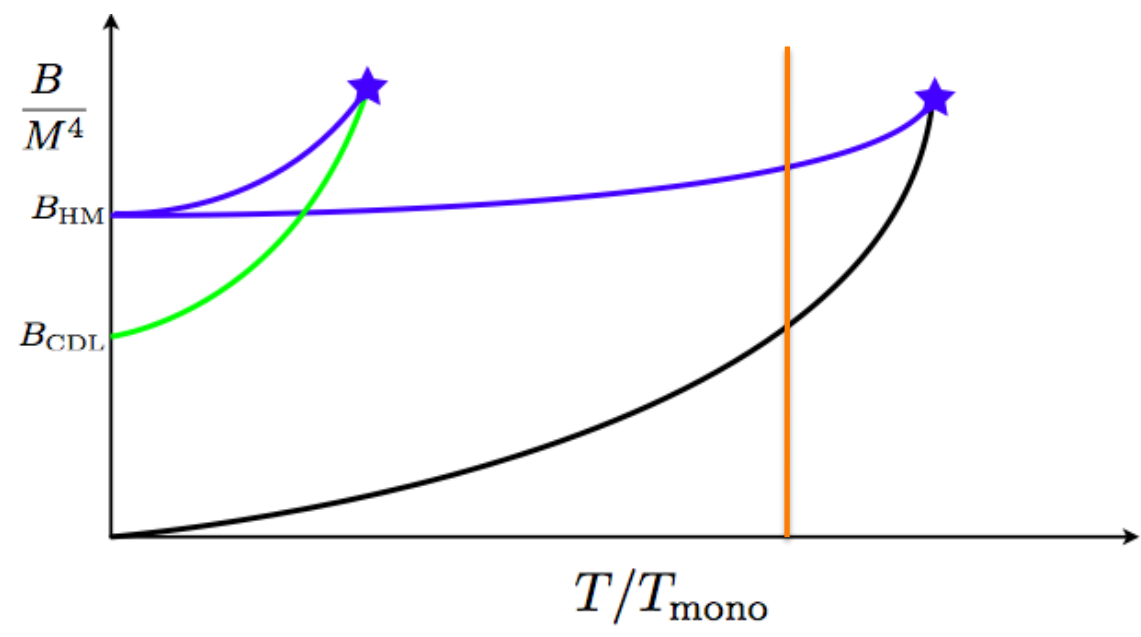
Case 2b:
Decay from
low dS



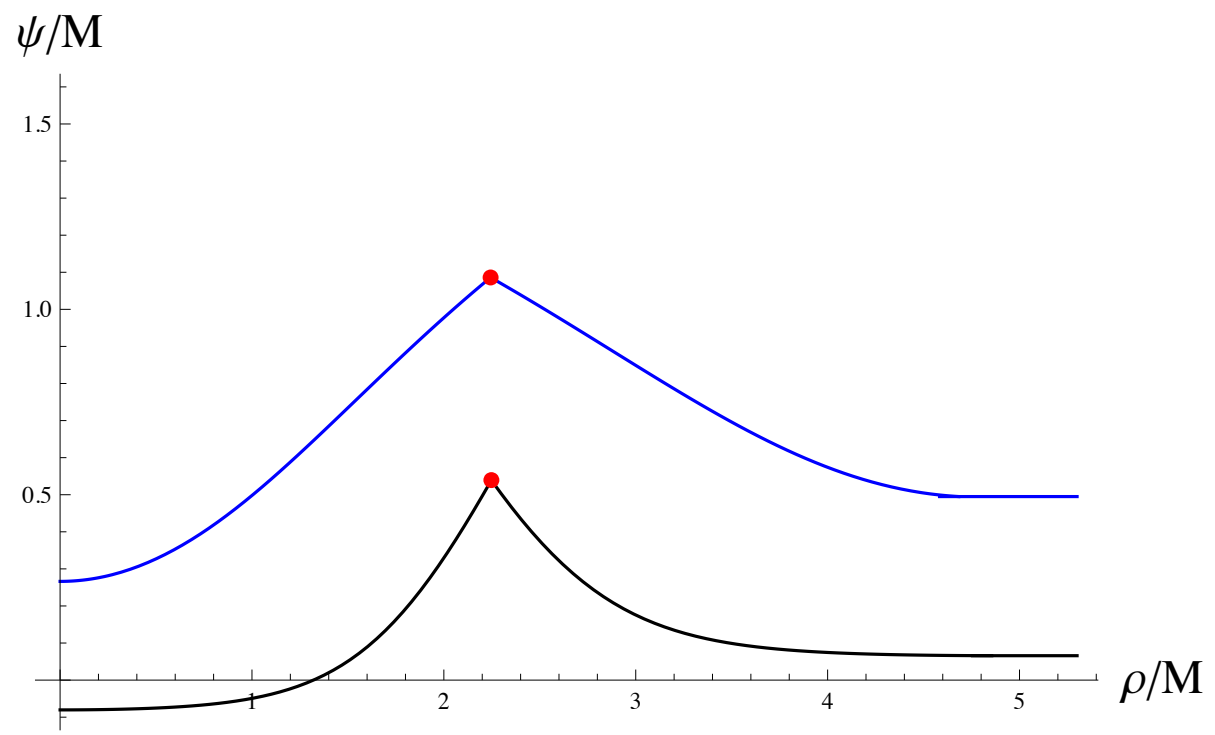
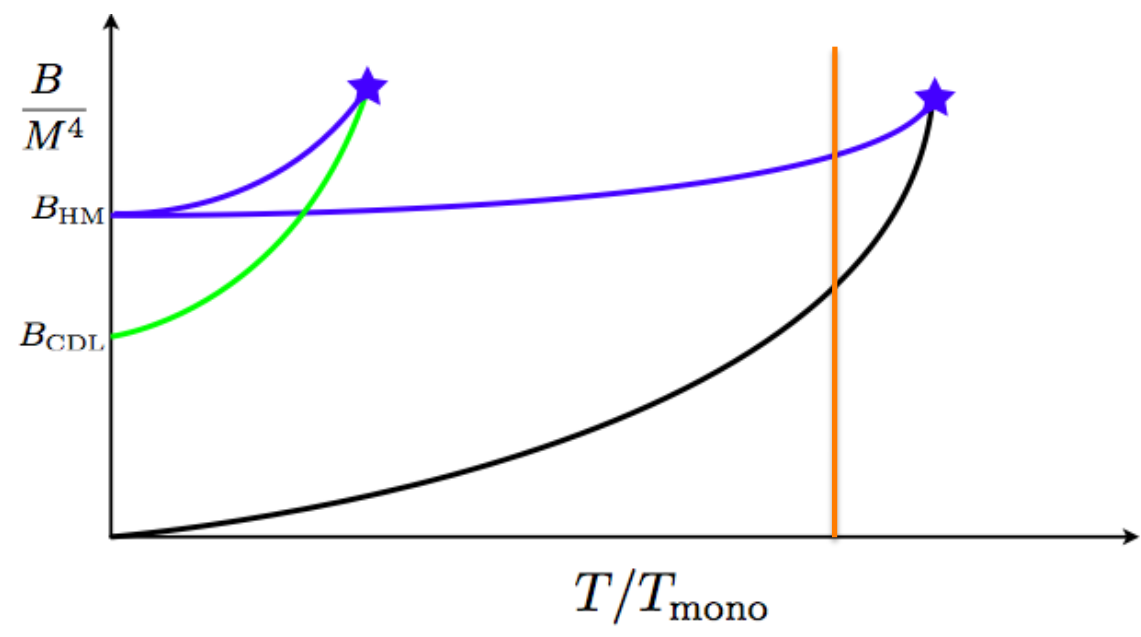
Case 2b:
Decay from
low dS



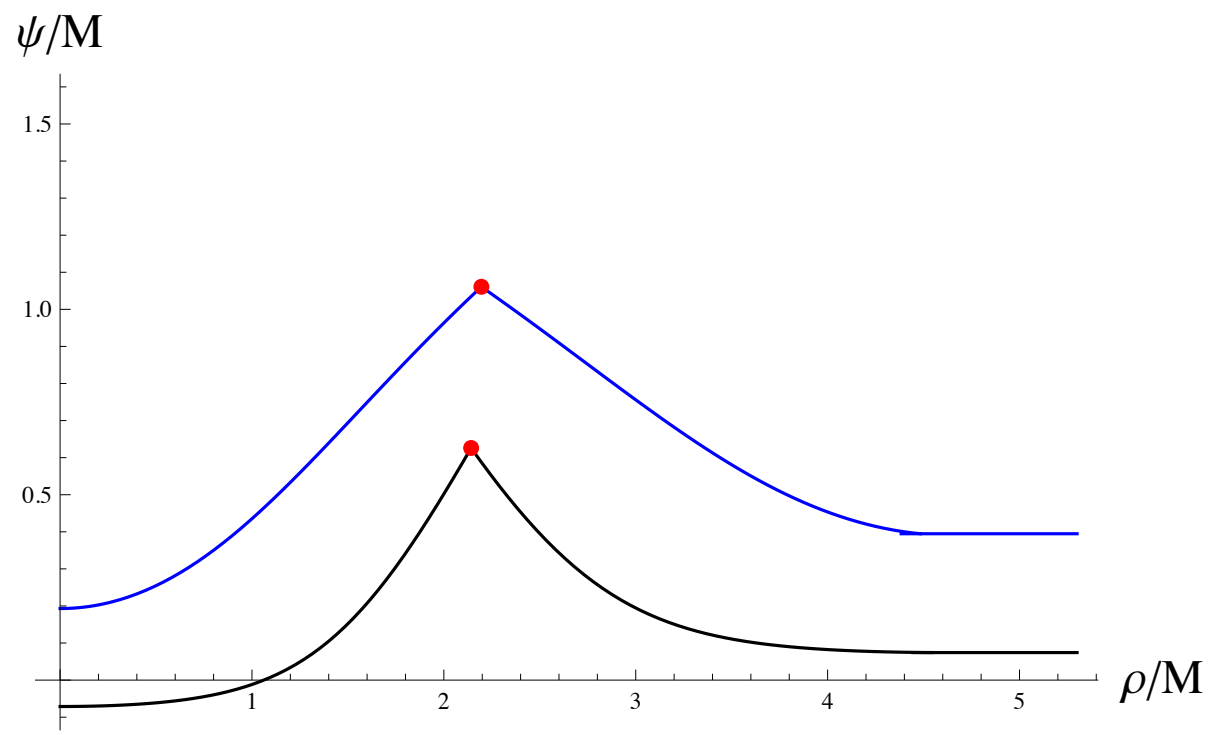
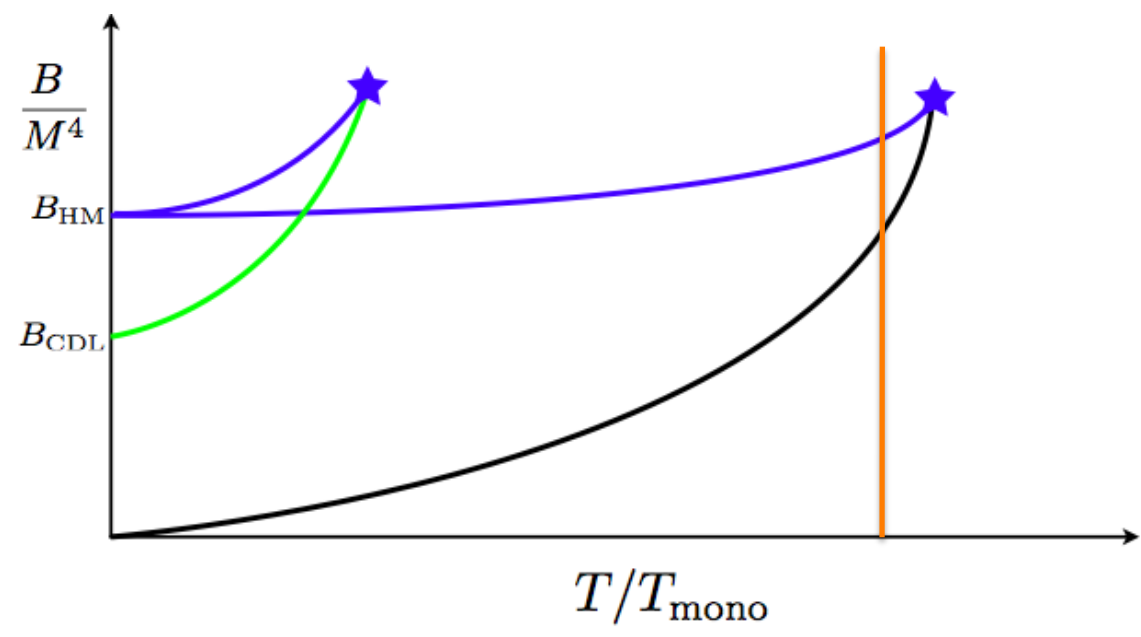
Case 2b:
Decay from
low dS



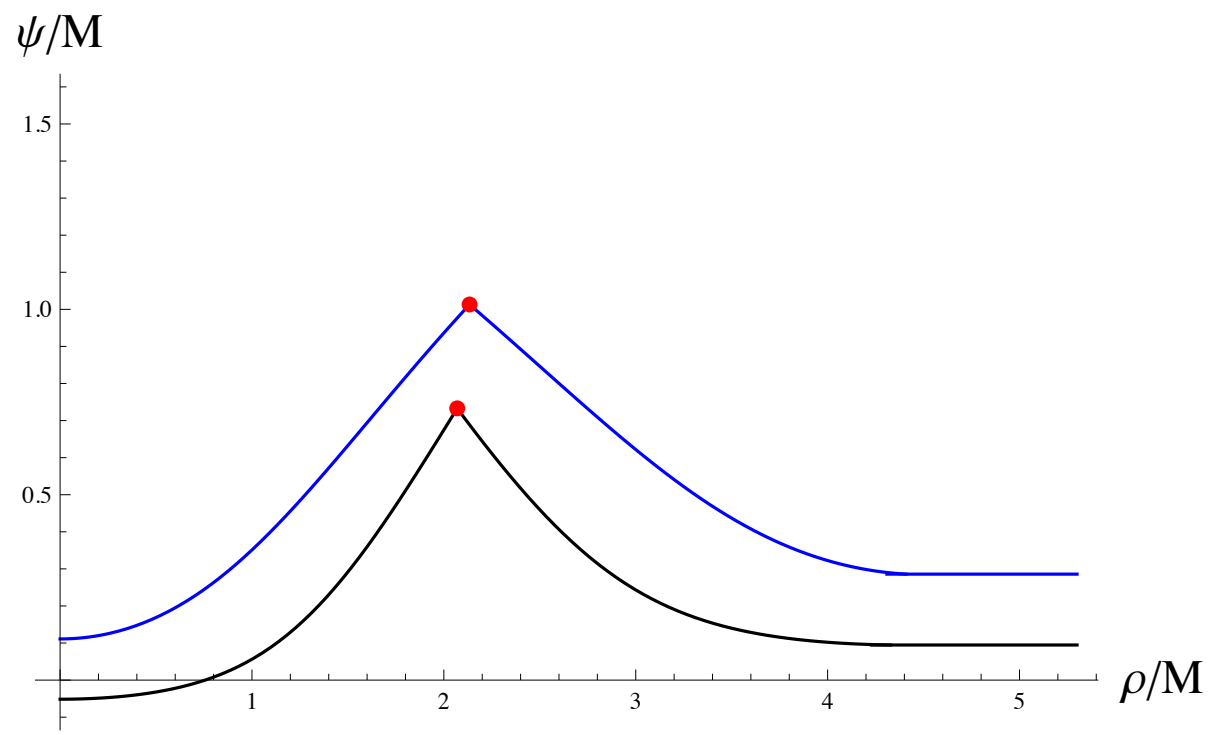
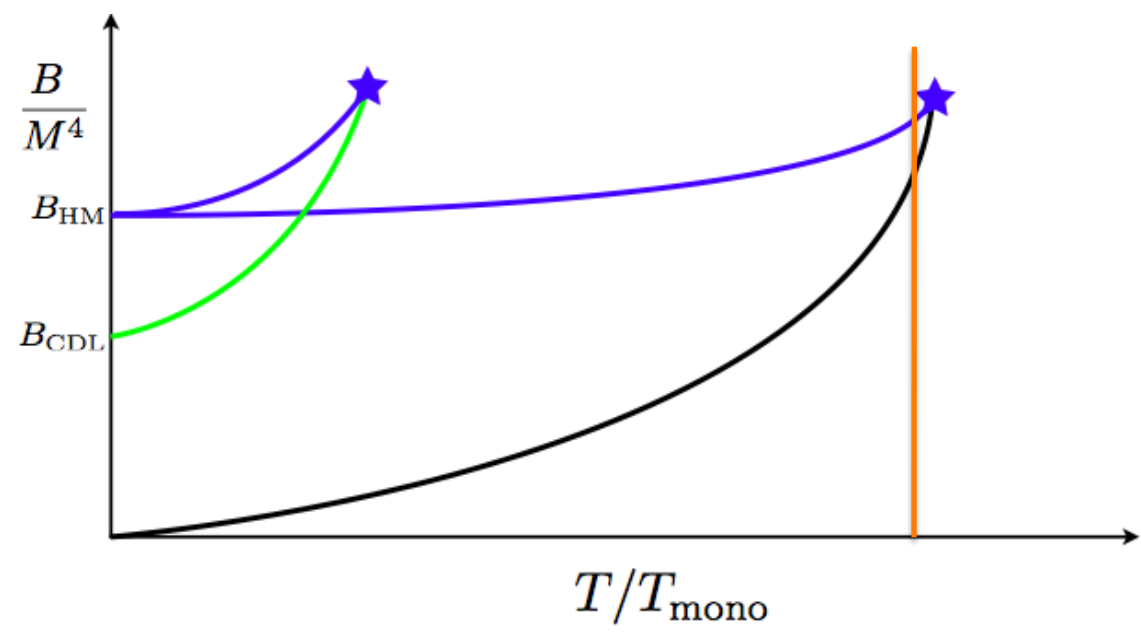
Case 2b:
Decay from
low dS



Case 2b:
Decay from
low dS



Case 2b:
Decay from
low dS



Case 2b:
Decay from
low dS

Instanton
Annihilation
Number 2

