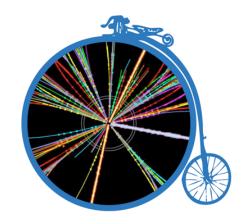
Normalizing VV at the LHC

J.Campbell, E.Castaneda, Y.Fang, N.Kauer B.Mellado and Sau Lan Wu (Not an ATLAS talk)



Special thanks S.Dawson, J.Qian and D.Rebuzzi LMET Workshop, UC Davis, 04/01/09

Overview

□ Introduction ☐ Relevance of VV at LHC □ Normalizing VV with Z^(*) ☐ Tools ☐ Inclusive rates **□ZZ** production **□ZW, WW production** ☐ Dependence on pp center of mass energy ☐ Jet veto survival probability **■ WW production □ ZW** production □ Outlook and conclusions

VV at the LHC

- □ VV (V=Z,W) is a major contributor to NL+MET and a background for a number of searches ranging from Higgs searches to SUSY and models beyond the SM
 - ☐ Attractive NL+MET signatures with first data
 - **☐ WW** gives a large rate of 2L+MET
 - With a jet veto it becomes the leading background for H->WW searches
 - **□ ZW** gives 3L+MET
 - With a jet veto it becomes the leading background to gaugino production
 - □ ZZ is the leading 4L production mechanism
 - Gives MET when T present or instrumental MET

Normalizing VV with Z^(*)

- □ Strong similarities of diagrams since dominant cross-section comes from qq->V(V) via EW couplings
- □ Ratios VV/V expected to reduce pdf and a significant portion of the scale uncertainty
 - □ This is an asset especially at the very beginning of data taking when global pdf fits will not be available

Prediction Theory Experimental efficiencies Observed
$$N(VV) = \left(\frac{\sigma(pp \to VV)}{\sigma(pp \to Z^{(*)})}\right)_{Th} \cdot \epsilon(ll \to Nl) \cdot N_{Obs}(Z^{(*)})$$

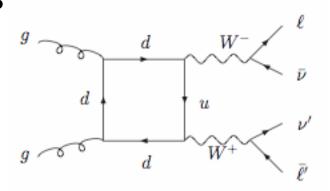
Abdullin et al. in hep-ph/0604120 computed the ratio ZZ/Z to NLO

Tools

- □ For qq->Z^(*) and qq->ZZ use MCFM v5.3 with bug fixes provided by John Campbell
 - ☐ Two independent analyses with independent MC samples. Cross-checks with Pythia
 - Foreseeing additional check with Sherpa
- □ For gg->ZZ use gg2ZZ. The numbers of the nominal cross-sections and the scale error uncertainties made by Nikolas Kauer
 - □ Evaluating with Pavel Nadolski (Resbos) and the K factor for gg->\gamma\gamma for M_{gg}~M_ZZ
 - Nikolas evaluating potential differences
- ☐ Studies with ALPGEN and MC@NLO

Ratio ZZ(WW)/Z*

- □ The production of ZZ and WW is enhanced by large contributions from gg->VV with gluons in the initial state
 - ☐ Formally a part of the NNLO contribution, but enhanced due to the large gluon flux



$$R = \frac{\sigma_{q\overline{q}\to ZZ}^{NLO} + \sigma_{gg\to ZZ}^{LO}}{\sigma_{q\overline{q}\to Z}^{NLO}}$$

Event selection & Settings

- □ The analysis is done at the "parton level". The theoretical errors are decoupled from the experimental errors
- \Box These studies are only for $M_{zz}>2M_z$
 - \Box Four (two) leptons with P_T>20 GeV, |eta|<2.5
 - □ Requirement of 71<M_{II}<111 on lepton pairs
 - □ \DeltaR_{II}>0.2 and \DeltaR_{Ij}>0.7
 - ☐ EW settings as default in MCFM taken by gg2ZZ
 - □ Set scales to M_V
 - We also have results with dynamic scales \mu=M_{Z^(*),ZZ}.

Nominal Values of ZZ/Z*

- □ Ratios are constructed such that the invariant mass of Z(*) and ZZ are in the same bin
 - □ Contribution from gg->ZZ increases sigma by ~13%
 - □ Ratio depends weakly with Mass (nice surprise!)
 - Need to understand better behavior at very large masses

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Mass Range	$\sigma^{NLO}_{q\overline{q} ightarrow Z^*}$	$\sigma^{NLO}_{q\overline{q} ightarrow ZZ}$	$\sigma^{LO}_{gg o ZZ}$	$\frac{\sigma_{ZZ}}{\sigma_{Z^*}} \times 10^3$
200 - 250	1773.7	7.99	1.182	5.17
250 - 300	753.2	3.65	0.530	5.54
300 - 350	372.4	1.86	0.246	5.66
350 - 400	205.7	1.07	0.131	5.83
400 - 450	121.0	0.64	0.082	5.94
450 - 500	76.0	0.40	0.055	6.01
500 - 750	143.9	0.74	0.114	5.92
750 - 1000	27.4	0.16	0.033	6.88

Scale Errors of ZZ/Z*

- ☐ Treat qq->ZZ and gg->ZZ independently
 - ☐ This somewhat overestimates error on factorization scale due to expected anti-correlation for qq and gg
 - ☐ Get maximum deviation by changing renormalization and factorization scales in opposite directions

Cross-sections in fb

Change scale by *4, /4

Mass Range	ass Range $\sigma^{NLO}_{qar q o Z^*}$ $\sigma^{NLO}_{qar q o ZZ}$		σ_{gg}^{Lo}	$\sigma^{LO}_{gg o ZZ}$		< 10 ³		
200 - 250	1858.8	4.8	8.34	4.3	1.92	62.0	5.52	6.6
	1586.8	-10.5	7.14	-10.6	0.75	-36.4	4.98	-3.8
250 - 300	792.0	5.2	3.86	5.9	0.83	57.3	5.93	6.9
	683.8	-9.2	3.32	-9.0	0.35	-33.9	5.36	-3.3
300 - 350	390.5	4.9	1.94	4.2	0.38	53.6	5.94	4.9
	340.7	-8.5	1.70	-8.5	0.17	-31.5	5.50	-2.9
350 - 400	214.7	4.4	1.10	3.3	0.20	49.3	6.05	3.8
	195.3	-5.0	0.96	-10.0	0.09	-29.8	5.40	-7.5
400 - 450	125.8	4.0	0.67	5.8	0.12	46.0	6.31	6.2
	114.8	-5.1	0.60	-6.4	0.06	-28.5	5.70	-4.1
450 - 500	79.5	4.5	0.43	6.5	0.08	44.3	6.38	6.3
	72.4	-4.8	0.38	-6.0	0.04	-26.7	5.78	-3.8
500 - 750	147.6	2.6	0.78	5.9	0.16	40.9	6.39	7.8
	140.4	-2.5	0.70	-4.8	0.09	-22.0	5.64	-4.7
750 - 1000	28.1	2.6	0.16	2.0	0.04	30.1	7.17	4.2
	28.2	2.9	0.15	-4.9	0.03	-17.8	6.21	-9.8

Scale Errors of ZZ/Z*

□ Multiply the contribution of gg->ZZ by a factor of 2 (potential QCD NLO K factor) but keep the relative errors at the LO level

Mass Range	$igg \sigma^{NL}_{qar q-}$	$O \to Z^*$	$\sigma_{\!q{\overline q}}^{N}$	LO i→ZZ	σ_{gg}^{Lo}	$\stackrel{O}{\rightarrow}ZZ$	$\frac{\sigma_{ZZ}}{\sigma_{Z^*}}$	$\times 10^3$
200 - 250	1858.8	4.8	8.34	4.3	3.83	62.0	6.55	12.1
	1586.8	-10.5	7.14	-10.6	1.50	-36.4	5.45	-6.7
250 - 300	792.0	5.2	3.86	5.9	1.67	57.3	6.98	11.7
	683.8	-9.2	3.32	-9.0	0.70	-33.9	5.88	-6.0
300 - 350	390.5	4.9	1.94	4.2	0.76	53.6	6.91	9.2
	340.7	-8.5	1.70	-8.5	0.34	-31.5	5.99	-5.3
350 - 400	214.7	4.4	1.10	3.3	0.39	49.3	6.97	7.7
	195.3	-5.0	0.96	-10.0	0.18	-29.8	5.87	-9.3
400 - 450	125.8	4.0	0.67	5.8	0.24	46.0	7.26	9.7
	114.8	-5.1	0.60	-6.4	0.12	-28.5	6.22	-6.2
450 - 500	79.5	4.5	0.43	6.5	0.16	44.3	7.37	9.7
	72.4	-4.8	0.38	-6.0	0.08	-26.7	6.33	-5.9
500 - 750	147.6	2.6	0.78	5.9	0.32	40.9	7.47	11.3
	140.4	-2.5	0.70	-4.8	0.18	-22.0	6.27	-6.5
750 - 1000	28.1	2.6	0.16	2.0	0.08	30.1	8.68	7.5
	28.2	2.9	0.15	-4.9	0.05	-17.8	7.16	-11.3

Cross-sections in fb

Comments on gg->ZZ to NLO

- □ The QCD NLO corrections gg->ZZ are expected to be significant
 - We can get a feeling of the size by looking into high mass gg->\gamma\gamma. Work in progress to compare with the numbers of 2002

PHYSICAL REVIEW D 66, 074018 (2002)

TABLE I. NLO QCD K factors for $\gamma\gamma$ Higgs signal and gluon fusion background. Both LO and NLO cross sections are computed using NLO parton distributions.

$M_{\gamma\gamma}$ (GeV)	$K_{ m Higgs}$	$K_{gg o \gamma\gamma}$
98	2.92	1.82
118	2.54	1.61
138	2.39	1.55

\sqrt{s} Dependence

- □ The contribution off the gg->ZZ to the table is not added yet
- □ The Ratio ZZ/Z* seems to be flat as a function of \sqrt{s} and different mass ranges
 - ☐ The ratio ZZ/Z is less flat

Table 4: Stability of the ratio $\frac{\sigma_{q\bar{q}\to ZZ}^{NLO}}{\sigma_{q\bar{q}\to Z^*}^{NLO}} \times 10^3$ for different ranges of the invariant mass of the leptonic system (in gev) as a function of the p-p collision center of mass energy (in TeV).

\sqrt{s}	200 - 250	250 - 300	300 - 500	> 500
14	4.51	4.87	5.18	5.06
12	4.52	4.88	5.07	4.98
10	4.45	4.88	4.96	4.98
8	4.47	4.82	4.97	4.98
6	4.44	4.73	4.93	5.04

Flatness of ratio indicates reduction of pdf uncertainties

Ratio WW/Z*

- ☐ Computation similar to ratio ZZ/Z*
 - □ Require two leptons with same cuts as ZZ events selection + MET>20 GeV
 - □ Contribution from gg->WW is about 10% of the total WW cross-section

$$R = \frac{\sigma_{q\overline{q} \to WW}^{NLO} + \sigma_{gg \to WW}^{LO}}{\sigma_{q\overline{q} \to Z^{(*)}}^{NLO}}$$

	$\sigma^{NLO}_{qar{q} o Z^\star}$	$\sigma^{NLO}_{q\overline{q} ightarrow WW}$	$\sigma^{LO}_{gg o WW}$	$rac{\sigma_{\!WW}}{\sigma_{\!Z^*}}$
Nominal	4513	1272	62.08	0.296
Max	+4.6	+11.5	+62.1	+8.9
Min	-9.9	-13.4	-35.9	-5.0

Results to NLO

\sqrt{s} Dependence of VV/Z*

□ Ratios are relatively stable w.r.t. \sqrt{s}

$ \sqrt{s} [\text{TeV}] $	$\frac{\sigma(WW)}{\sigma(Z^*)}$	$\frac{\sigma(ZW)}{\sigma(Z^*)}$	$\left \begin{array}{c} \sigma(ZZ) \\ \sigma(Z^*) \end{array} \right $
14	0.280	.0481	.0063
12	0.294	.0473	.0062
10	0.271	.0462	.0062
8	0.265	.0452	.0062
6	0.256	.0435	.0062

Ratio WW/ZZ

- ☐ The ratio WW/ZZ will diminish the error due to the gg->VV contribution.
 - ☐ The errors will probably be dominated by the experimental uncertainties

\sqrt{s} [TeV]	$\sigma(WW)$	$\delta\sigma(WW)$	$\sigma(ZZ)$	$\delta\sigma(ZZ)$	$\frac{\sigma(ZZ)}{\sigma(WW)} \times 10^2$	$\delta rac{\sigma(ZZ)}{\sigma(WW)}$
14	1272	11.5	20.09	4.0	1.58	3.9
		-13.4		-10.0		-6.7
10	881.9	7.8	14.13	4.5	1.60	3.9
		-10.1		-7.4		-3.0
8	675.6	8.7	11.15	4.2	1.65	2.3
		-7.1		-5.7		-4.2

- ☐ Results above are computed to NLO
- □ The scale errors due to gg->VV cancel almost completely

Jet Veto Survival Prob.

- □ In order to suppress top backgrounds a full (\eta|<5) jet veto is usually applied</p>
 - □ Define ε(P_T) as a fraction of the events that survive a cut on events with a jet of above a certain P_T threshold in (\eta|<5)</p>
- ☐ Here we address the possibility of predicting the jet veto survival probability of WW an ZW from Z*
- □ The P_T spectra of the leading jet in VV and V are different, but the ratio of the jet veto survival probability should have small uncertainties

$\sqrt{s} = 14 \text{ TeV}$

$p_T [\text{GeV}]$	$\varepsilon(Z^*)$	$\delta \varepsilon(Z^*)$	$\varepsilon(WW)$	$\delta \varepsilon(WW)$	$\frac{\varepsilon(WW)}{\varepsilon(Z^*)}$	$\delta \frac{\epsilon(WW)}{\epsilon(Z^*)}$
20	0.67	8.5	0.52	11.9	0.78	5.1
		-13.2		-15.2		-3.2
25	0.72	6.4	0.58	9.6	0.81	4.0
		-9.9		-11.8		-2.9
30	0.76	5.1	0.63	8.3	0.82	3.6
		-7.8		-9.1		-2.1
35	0.79	4.1	0.67	7.4	0.84	3.3
		-6.3		-7.3		-2.1
40	0.82	3.5	0.70	6.6	0.85	3.0
		-5.3		-5.9		-1.9
45	0.84	3.2	0.72	6.0	0.86	2.8
		-4.4		-5.4		-1.8
50	0.86	2.9	0.75	5.5	0.87	2.6
		-3.8		-5.0		-1.8
55	0.87	2.6	0.77	5.1	0.88	2.5
		-3.3		-4.7		-1.7
60	0.88	2.4	0.79	4.7	0.89	2.3
		-2.9		-4.3		-1.6
65	0.89	2.1	0.80	4.3	0.90	2.1
		-2.6		-4.0		-1.4
70	0.90	1.9	0.81	4.1	0.90	2.1
		-2.4		-3.7		-1.3
75	0.91	1.8	0.83	3.7	0.91	1.9
		-2.2		-3.3		-1.1
80	0.92	1.6	0.84	3.4	0.91	1.7
		-2.0		-3.2		-1.1
85	0.93	1.5	0.85	3.1	0.92	1.6
		-1.9		-3.1		-1.2
90	0.93	1.4	0.86	3.0	0.92	1.5
		-1.8		-2.9		-1.1
95	0.94	1.3	0.87	2.7	0.93	1.4
		-1.7		-2.7		-1.1
100	0.94	1.2	0.88	2.6	0.93	1.3
		-1.6		-2.6		-1.1

$\sqrt{s} = 10 \text{ TeV}$

p _T [GeV]	$\varepsilon(Z^*)$	$\delta \varepsilon(Z^*)$	$\varepsilon(WW)$	$\delta \varepsilon(WW)$	$\frac{\varepsilon(WW)}{\varepsilon(Z^*)}$	$\delta \frac{\varepsilon(WW)}{\varepsilon(Z^*)}$
20	0.69	8.6	0.58	10.9	0.83	2.1
		-7.2		-13.5		-7.8
25	0.75	6.8	0.63	9.2	0.85	2.2
		-5.5		-10.9		-6.3
30	0.78	5.6	0.68	8.0	0.86	2.2
		-4.3		-9.8		-5.8
35	0.81	4.8	0.71	7.1	0.87	2.1
		-3.7		-9.0		-5.4
40	0.84	4.1	0.74	6.2	0.88	2.0
		-3.3		-8.1		-4.9
45	0.86	3.6	0.76	5.4	0.89	1.7
		-3.0		-7.3		-4.4
50	0.87	3.2	0.79	4.8	0.90	1.6
		-2.7		-6.8		-4.2
55	0.89	2.9	0.80	4.4	0.91	1.5
		-2.4		-6.3		-4.0
60	0.90	2.6	0.82	4.0	0.91	1.4
		-2.3		-6.0		-3.8
65	0.91	2.3	0.83	3.8	0.92	1.4
		-2.1		-5.5		-3.4
70	0.92	2.1	0.85	3.4	0.93	1.2
		-1.9		-5.4		-3.5
75	0.92	2.0	0.86	3.2	0.93	1.2
		-1.8		-5.1		-3.4
80	0.93	1.8	0.87	2.9	0.94	1.1
		-1.7		-4.8		-3.2
85	0.94	1.7	0.88	2.8	0.94	1.1
		-1.6		-4.5		-2.9
90	0.94	1.5	0.89	2.6	0.94	1.0
		-1.5		-4.2		-2.8
95	0.95	1.4	0.90	2.4	0.95	0.9
		-1.3		-4.0		-2.7
100	0.95	1.3	0.90	2.2	0.95	0.9
		-1.3		-3.9		-2.7

\sqrt{s} Dependence of Jet Veto

Results of ε shown for P_T =30 GeV

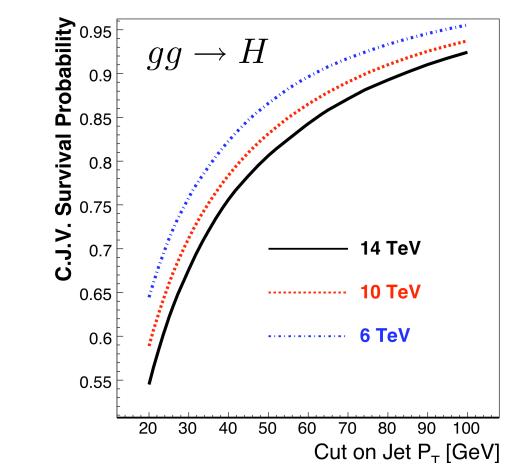
		WW	W Z^*]
\sqrt{s}	$ < p_{Tj} > $	$ < \eta >$	$arepsilon_{jv}$	$ < p_{Tj} >$	$ < \eta >$	$oldsymbol{arepsilon}_{jv}$	$rac{arepsilon_{j_{\mathcal{V}}}^{WW}}{arepsilon_{j_{\mathcal{V}}}^{Z^*}}$
14	38.6	0.76	0.64	22.3	0.58	0.77	0.83
12	34.8	0.68	0.67	20.7	0.54	0.78	0.86
10	32.1	0.66	0.69	18.9	0.50	0.80	0.86
8	27.7	0.59	0.72	17.2	0.47	0.81	0.89
6	22.7	0.51	0.76	14.3	0.40	0.84	0.90

\sqrt{s} Dependence of Jet Veto

	12		10		8		6	
$p_T[\text{ GeV}]$	$\varepsilon^R(WW)$	$\frac{\varepsilon^R(WW)}{\varepsilon^R(Z^*)}$	$\varepsilon^R(WW)$	$\frac{\varepsilon^R(WW)}{\varepsilon^R(Z^*)}$	$\varepsilon^R(WW)$	$\frac{\varepsilon^R(WW)}{\varepsilon^R(Z^*)}$	$\varepsilon^R(WW)$	$\frac{\varepsilon^R(WW)}{\varepsilon^R(Z^*)}$
20	1.07	1.04	1.10	1.04	1.16	1.08	1.25	1.11
25	1.06	1.03	1.08	1.04	1.14	1.07	1.21	1.09
30	1.05	1.03	1.08	1.04	1.12	1.06	1.18	1.08
35	1.04	1.03	1.06	1.03	1.11	1.06	1.16	1.08
40	1.04	1.03	1.06	1.03	1.10	1.05	1.15	1.07
45	1.04	1.02	1.05	1.03	1.09	1.05	1.13	1.07
50	1.03	1.02	1.05	1.03	1.08	1.04	1.12	1.06
55	1.03	1.02	1.05	1.02	1.08	1.04	1.11	1.06
60	1.02	1.01	1.04	1.02	1.07	1.04	1.10	1.05
65	1.02	1.01	1.04	1.02	1.07	1.04	1.10	1.05
70	1.02	1.01	1.04	1.02	1.06	1.03	1.09	1.05
75	1.01	1.01	1.03	1.02	1.06	1.03	1.09	1.05
80	1.01	1.01	1.03	1.02	1.05	1.03	1.08	1.04
85	1.01	1.01	1.03	1.02	1.05	1.03	1.07	1.04
90	1.01	1.01	1.03	1.02	1.05	1.03	1.07	1.04
95	1.01	1.00	1.03	1.02	1.04	1.03	1.07	1.04
100	1.01	1.00	1.03	1.01	1.04	1.02	1.06	1.04

Jet Veto Survival Prob for gg->VV

- ☐ This quantity has not been calculated.
- □ We can get a feeling of how different it is from the qq->VV process by looking into gg->H



Jet veto survival probability for Higgs is very close to that of qq->WW

	WW				
\sqrt{s}	$< p_{Tj} >$	$< \eta >$	$arepsilon_{jv}$		
14	38.6	0.76	0.64		
12	34.8	0.68	0.67		
10	32.1	0.66	0.69		
8	27.7	0.59	0.72		
6	22.7	0.51	0.76		

Results of ε shown for P_T =30 GeV

$\sqrt{s} = 14 \text{ TeV}$

p_T [GeV]	$\varepsilon(Z^*)$	$\delta \varepsilon(Z^*)$	$\varepsilon(ZW)$	$\delta \varepsilon(ZW)$	$\frac{\varepsilon(ZW)}{\varepsilon(Z^*)}$	$\delta \frac{\varepsilon(ZW)}{\varepsilon(Z^*)}$
20	0.67	8.5	0.48	13.2	0.71	6.3
		-13.2		-15.3		-7.3
25	0.72	6.4	0.53	11.1	0.73	5.5
		-9.9		-12.2		-6.4
30	0.76	5.1	0.57	9.7	0.75	5.0
		-7.8		-10.8		-5.9
35	0.79	4.1	0.61	8.7	0.76	4.6
		-6.3		-9.9		-5.5
40	0.82	3.5	0.64	7.7	0.78	4.0
		-5.3		-9.2		-5.3
45	0.84	3.2	0.66	7.1	0.79	3.8
		-4.4		-8.5		-5.0
50	0.86	2.9	0.68	6.5	0.80	3.5
		-3.8		-7.8		-4.7
55	0.87	2.6	0.70	6.1	0.81	3.4
		-3.3		-7.5		-4.6
60	0.88	2.4	0.72	5.6	0.82	3.2
		-2.9		-7.0		-4.4
65	0.89	2.1	0.74	5.4	0.83	3.2
		-2.6		-6.6		-4.1
70	0.90	1.9	0.75	5.0	0.83	3.0
		-2.4		-6.3		-4.0
75	0.91	1.8	0.77	4.7	0.84	2.9
		-2.2		-6.1		-4.0
80	0.92	1.6	0.78	4.4	0.85	2.7
		-2.0		-5.8		-3.8
85	0.93	1.5	0.79	4.2	0.86	2.6
		-1.9		-5.6		-3.7
90	0.93	1.4	0.80	4.0	0.86	2.5
		-1.8		-5.2		-3.5
95	0.94	1.3	0.81	3.8	0.87	2.4
		-1.7		-4.9		-3.2
100	0.94	1.2	0.82	3.6	0.87	2.3
		-1.6		-4.6		-3.0

$\sqrt{s} = 10 \text{ TeV}$

$p_T [\text{GeV}]$	$\varepsilon(Z^*)$	$\delta \varepsilon(Z^*)$	$\varepsilon(ZW)$	$\delta \varepsilon(ZW)$	$\frac{\varepsilon(ZW)}{\varepsilon(Z^*)}$	$\delta \frac{\varepsilon(ZW)}{\varepsilon(Z^*)}$
20	0.69	8.6	0.52	16.2	0.75	7.0
		-7.2		-13.9		-8.4
25	0.75	6.8	0.58	13.2	0.77	5.9
		-5.5		-12.1		-7.5
30	0.78	5.6	0.62	11.7	0.79	5.7
		-4.3		-10.5		-6.5
35	0.81	4.8	0.65	10.1	0.80	5.1
		-3.7		-9.4		-5.9
40	0.84	4.1	0.68	8.9	0.81	4.5
		-3.3		-8.9		-5.7
45	0.86	3.6	0.70	8.1	0.82	4.4
		-3.0		-8.2		-5.4
50	0.87	3.2	0.72	7.5	0.83	4.1
		-2.7		-7.6		-5.0
55	0.89	2.9	0.74	6.9	0.84	3.8
		-2.4		-7.1		-4.8
60	0.90	2.6	0.76	6.3	0.85	3.6
		-2.3		-6.7		-4.5
65	0.91	2.3	0.78	5.9	0.86	3.5
		-2.1		-6.3		-4.2
70	0.92	2.1	0.79	5.5	0.86	3.3
		-1.9		-5.9		-4.0
75	0.92	2.0	0.80	5.2	0.87	3.1
		-1.8		-5.5		-3.8
80	0.93	1.8	0.82	4.8	0.88	3.0
		-1.7		-5.2		-3.6
85	0.94	1.7	0.83	4.5	0.88	2.8
		-1.6		-5.0		-3.5
90	0.94	1.5	0.84	4.2	0.89	2.6
		-1.5		-4.8		-3.4
95	0.95	1.4	0.85	4.0	0.89	2.5
		-1.3		-4.5		-3.2
100	0.95	1.3	0.86	3.8	0.90	2.4
		-1.3		-4.2		-3.0

Outlook and Conclusions

- ☐ The use of Z^(*) events is a powerful sample to normalize VV production
- \Box Consider ratios of $\sigma(VV)/\sigma(Z^{(*)})$
 - **□** Considered inclusive rates
 - The theoretical error of $\sigma(ZZ,WW)/\sigma(Z^{(*)})$ is dominated by the LO uncertainties of gg->VV
 - Errors remain at the level of 10%
 - Ratios depend weakly on \sqrt{s} and the mass
 - Also considered $\sigma(ZZ)/\sigma(WW)$ for which the uncertainty due to gg->VV cancels out
 - ☐ Consider Jet veto survival probability
 - Errors of ε(WW,ZW)/ε(Z*) stay below 10%

EXTRA SLIDES

MCFM Settings

Paramter	Name	Input Value	Output Value determined by ewscheme			
	$(_inp)$		-1	0	1	2
G_F	Gf	1.16639×10^{-5}	input	calculated	mput	input
$\alpha(M_Z)$	aemmz	1/128.89	input	input	calculated	input
$\sin^2 \theta_w$	XW	0.2312	calculated	input	calculated	input
M_W	wmass	$80.419~\mathrm{GeV}$	input	calculated	input	calculated
M_Z	zmass	91.188 GeV	input	input	input	calculated
m_t	mt	172.5 GeV	calculated	$_{ m input}$	input	input

Parameter	Fortran name	Default value	
$m_{ au}$	mtau	$1.777 \mathrm{GeV}$	
$m_{ au}^2$	mtausq	$3.1577 \; \mathrm{GeV^2}$	
m_c^2	mcsq	$2.25 \mathrm{GeV^2}$	
m_b^2	mbsq	$17.64 \mathrm{GeV^2}$	
$\Gamma_{ au}$	tauwidth	$2.269 \times 10^{-12} \text{ GeV}$	
Γ_W	wwidth	2.06 GeV	
Γ_Z	zwidth	2.49 GeV	
V_{ud}	Vud	0.975	
V_{us}	Vus	0.222	
V_{ub}	Vub	0.	
V_{cd}	Vcd	0.222	
V_{cs}	Vcs	0.975	
V_{cb}	Vcb	0.	

Petrielo et al $pp \rightarrow W+X$ $pp \rightarrow (Z, \gamma^*) + X$ W^{-} W^+ **NNLO** 400 $d^2\sigma/dM/dY~[pb/GeV]$ $\mathrm{d}^2\sigma/\mathrm{dM}/\mathrm{dY}\;[\mathrm{pb/GeV}]$ 300 200 $\sqrt{s} = 14 \text{ TeV}$ $M = M_Z$ 100 $\sqrt{s} = 14 \text{ TeV}$ $M/2 \le \mu \le 2M$ $M = M_{W}$ $M/2 \le \mu \le 2M$ Y -2 0 Y

The NLO band does contain the NNLO result for Z,W⁺,W⁻ production Same applies for the gg->H production